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Citation for final published version:

Ma, Lin, Wang, Qiang, Kerr, Andrew , Li, Zheng-Xiang, Dan, Wei, Yang, Ya-Nan, Zhou, Jin-Sheng and Li, Cheng 2023. Eocene magmatism in the Himalaya: response to lithospheric flexure during early Indian collision? *Geology* 51 (1) , pp. 96-100. 10.1130/G50438.1

Publishers page: <https://doi.org/10.1130/G50438.1>

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1 **Eocene magmatism in the Himalaya: A response to**
2 **lithospheric flexure during early Indian collision?**

3

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20 Eocene mafic magmatism in the Himalaya provides a crucial window for probing the
21 evolution of crustal anatexis processes within the lower-plate in a collisional orogen. Here we
22 report geochemical data from the earliest post-collision ocean island basalt-like mafic dikes
23 intruding the Tethyan Himalaya near the northern edge of the colliding Indian plate. These
24 dikes occurred coevally, and spatially overlap with, Eocene granitoids in the cores of gneiss
25 domes and are likely derived from interaction of melts from the lithosphere-asthenosphere
26 boundary with the Indian continental lithosphere. We propose that these mafic magmas were
27 emplaced along lithospheric fractures in response to lithospheric flexure during initial
28 subduction of the Indian continent and that the underplating of such mafic magmas resulted in
29 orogen-parallel crustal anatexis within the Indian continent. This mechanism can explain the
30 formation of coeval magmatism and the geological evolution of collisional orogen on both
31 sides of the suture zone.

32

33 INTRODUCTION

34 The Himalaya belt is the most active collisional orogen in the world. It exposes the former
35 passive margins of the Indian continent, and it is characterized by widespread Cenozoic
36 crustal anatexis, high-grade metamorphism and some orogen-scale normal and strike-slip
37 faulting ([Harrison et al., 1997](#); [Yin, 2006](#)) ([Fig. 1a](#)). These magmatic and metamorphic units
38 distributed parallel to the suture within the lower plate/previously passive side of the
39 collisional orogen, can yield significant information on the collision and related crustal
40 reworking processes ([Hou et al., 2012](#); [Vanderhaeghe and Teyssier, 2001](#); [Wang et al., 2021](#);
41 [Weller et al., 2021](#); [Zeng et al., 2011](#)).

42 Two critical episodes of Cenozoic metamorphism and magmatism have been identified
43 from the Himalaya: 1) 48–35 Ma Barrovian-type prograde metamorphism with Eocene mafic
44 and Na-rich adakitic melts (e.g., [Hou et al., 2012](#); [Ji et al., 2016](#)), and 2) Late Oligocene–
45 Early Miocene retrograde metamorphism associated with leucogranite formation ([Harrison et](#)
46 [al., 1997](#); [Vanderhaeghe and Teyssier, 2001](#); [Weller et al., 2021](#)) (Fig. 1).

47 Himalayan uplift and associated crustal anatexis ([Hou et al., 2012](#)) has been linked to a
48 range of possible processes including thin-skinned thrusting ([DeCelles et al., 2002](#); [Yin, 2006](#)),
49 middle-crustal melting and ductile flow ([Nelson et al., 1996](#); [Vanderhaeghe and Teyssier,](#)
50 [2001](#)), or extrusion of an Indian crustal wedge ([Chemenda et al., 2000](#)). A variety of anatexis
51 mechanisms beneath the Himalayas have been proposed, including: shear heating ([Harrison et](#)
52 [al., 1998](#)), decompression melting ([Davidson et al., 2008](#)), radiogenic heating ([Searle et al.,](#)
53 [2003](#)) and heat transferred from mantle-derived melts ([Zheng et al., 2016](#)). However, the links
54 between the crustal anatexis event(s) and coeval tectonic developments remain unclear ([Guo](#)
55 [and Wilson, 2012](#); [Hou et al., 2012](#) and references therein).

56 In this paper we report geochemical data from Eocene mafic dikes found in the Tethyan
57 Himalaya that have ocean island basalt (OIB)-like compositions. This mafic magmatism
58 occurred coevally, and spatially overlaps with, the well-developed Tethyan Himalayan
59 granitoids and associated metamorphic event ([Hou et al., 2012](#)) (Fig. 1a-b). These dikes
60 provide a rare opportunity to examine the origin of such enigmatic intraplate magmatism and
61 related geodynamic evolution along the margin of the lower plate in a continental collision
62 zone. Similar magmatism is also reported in other collisional orogens ([Vanderhaeghe et al.,](#)
63 [2020](#); [Weller et al., 2021](#)) and in this study we present a new geodynamic model for such

64 orogen-parallel lower-plate magmatic and metamorphic belts.

65

66 **BACKGROUND AND SAMPLES**

67 The Himalaya-Tibet orogen was formed by the collision of the Indian continent with
68 Eurasia that started along the Yarlung Zangbo Suture (YZS) (Fig. 1). The Lhasa block,
69 immediately north of the YZS, represents the southern edge of the Asian upper plate, and
70 experienced long-lived subduction of Neo-Tethyan ocean crust with extensive late Triassic to
71 Eocene calc-alkaline plutonism and volcanism before the terminal collision (Zhu et al., 2013,
72 2019). Following the collision and a magmatic flare-up with a peak of 51 ± 3 Ma, arc
73 magmatism in southern Tibet waned, as cold Indian lithosphere underthrust Tibet (Chung et
74 al., 2005; Zhu et al., 2019).

75 On the opposite side of the suture is the lower Indian plate representing a pre-collisional
76 stable craton, with slightly younger (ca. 48–35 Ma) high Sr/Y granitoids and limited OIB-type
77 gabbros and medium-temperature eclogite–high pressure (EHP) granulite metamorphism (Fig.
78 1; Hou et al., 2012; Ji et al., 2016; Weller et al., 2021; Zeng et al., 2011). The high Sr/Y
79 granitoids are believed to have derived from partial melting of amphibolite at $\sim 880^\circ\text{C}$ and ~ 10
80 kbar (Hou et al., 2012; Zeng et al., 2011). The 45 Ma Langshan gabbro, on the other hand, has
81 HIMU (high μ , $\mu = {}^{238}\text{U}/{}^{204}\text{Pb}$)-type OIB signatures with depleted Sr-Nd isotopes. These
82 gabbros are thought to have been generated by partial melting of the asthenosphere during
83 detachment of the subducted Neo-Tethyan slab (Ji et al., 2016). Subsequent to this magmatic
84 episode, kilometer-scale Himalayan leucogranite bodies (~ 25 – 15 Ma and ~ 8 Ma) were
85 emplaced in the northern Himalaya (Fig. 1; Vanderhaeghe and Teyssier, 2001; Weller et al.,

2021). These leucogranites were produced by muscovite-dehydration melting of meta-sediments (Weinberg, 2016) at ultrahigh temperature (UHT) conditions (900–970°C and 6–11 kbar [$\sim 40^\circ\text{C}/\text{km}$]), mostly between 25 and 15 Ma (Wang et al., 2021).

Several ENE-trending, broadly orogen-parallel, 5–8 m wide diabase dikes that intrude the Early Jurassic Ridang Formation limestone, marl limestone and shale, have recently been discovered near Gyangze (Figs. 1 and DR2). These dikes are coarse-grained and consist of clinopyroxene, plagioclase and amphibolite with secondary chlorite and sericite.

GEOCHRONOLOGY AND GEOCHEMISTRY

Zircon grains from the Gyantse diabase sample JZ18-2-1 yielded a U-Pb age of 48.6 ± 0.5 Ma (Fig. 2d, methodology and detailed analytical results in the Data Repository), that is slightly older than the Langshan OIB-type gabbro (45 ± 1.4 Ma, Ji et al., 2016) but overlaps with the earliest Cenozoic high Sr/Y granitoids (ca. 48–45 Ma; e.g., Hou et al., 2012). Zircon $\delta^{18}\text{O}$ values ranging from 5.1‰ to 6.4‰ with a mean of $5.9 \pm 0.6\text{‰}$, are consistent (within error) with mantle zircon values ($5.3 \pm 0.6\text{‰}$).

The Gyantse diabase dikes have relatively broad ranges of SiO_2 (46.0 to 54.5 wt.%) and MgO (5.0 to 12.4 wt.%) contents, and plot in the field of alkali basaltic rocks (Fig. 2a). They have OIB-like element patterns with enriched Nb (11.4–21.5 ppm) and TiO_2 (1.9–2.9 wt.%) and a slight Eu anomaly (chondrite normalized $\text{Eu}/\sqrt{\text{Sm} \times \text{Gd}} = 0.85 - 1.10$). Initial Sr-isotope ratios (0.7076–0.7115) and $\varepsilon_{\text{Nd}}(t)$ (–2.7 to –2.0), are more enriched than the Langshan gabbro but less so than the coeval granitoids (Fig. 2c).

108 **PETROGENESIS AND GEOTECTONIC IMPLICATIONS**

109 The Gyantse mafic dikes have relatively low Nb (11.7–21.5 ppm) and more enriched Sr-Nd
110 isotope compositions than the Langshan gabbro which was likely formed by partial melting of
111 asthenosphere (Ji et al., 2016), that contained more-enriched components. Given the
112 insignificant crustal contamination of the Gyantse mafic magmas (see details in Data
113 Repository), the Indian lithospheric mantle (Shellnutt et al., 2014) is the most likely source of
114 this enriched component. Our modeling indicates that the Gyantse mafic dikes were most
115 likely derived from the lithosphere-asthenosphere boundary (LAB) melts at the top of the
116 asthenosphere with a contribution from the Indian lithospheric mantle (Figs. 2c and DR3).
117 Seismic profiles show that the present depth of the LAB below eastern Himalaya ranges from
118 140 km to 100 km (Zhao et al. 2010), which is likely to be similar to, or greater than, the
119 lithosphere thickness at the early stage of the collision (~50 Ma). Peridotite with 1.0–2.5 wt.%
120 CO₂ or 200–300 ppm H₂O can produce stable partial melts at 3 Gpa (Hirschmann et al., 2009).
121 The LAB has recently been documented to be volatile enriched (Blatter et al. 2022), and thus
122 provides the most likely source for mafic magmas in Tethyan Himalaya during early collision.

123 The Early Eocene (51–45 Ma) magmatism, dominated by crustal anatexis, occurred on
124 both sides of the YZS during early collision. The southern margin of Asian plate is
125 characterized by thickened juvenile crust and a relatively high crustal thermal state (Ma et al.,
126 2017). In contrast, the northern edge of the Indian plate is marked by a thickened ancient crust
127 with a moderate geothermal gradient during the early stage of collision (Hou et al., 2012;
128 Weller et al., 2021 and references therein). The 48–35 Ma Tethyan Himalaya magmatism,
129 consisting of high Sr/Y granitoids within gneiss dome and coeval gabbros and dikes, is slightly

130 younger than the magmatic peak in the Lhasa block (51 ± 3 Ma) (Fig. 1). After that, another
131 episode (25–8 Ma) of crustal anatexis of accreted sediments dominated the melting in the
132 Himalaya (Guo and Wilson, 2012; Weller et al., 2021).

133 The sources of the two episodes of crustal anatexis along the northern Indian continental
134 margin (one at around 48–35 Ma, and the other around 25–8 Ma) may be very different. The
135 Miocene crustal anatexis accompanied by ultrahigh temperature (900–970 °C) metamorphism
136 indicates a hotter crustal thermal state than the Eocene crustal anatexis events with moderate
137 temperature (~600–750 °C) metamorphism (e.g., Wang et al., 2021; Weller et al., 2021).
138 Numerous studies over the past few decades have been carried out on the Miocene crustal
139 anatexis and related process, which led to the Cenozoic rise of the Himalaya orogen
140 (Chemenda et al., 2000; DeCelles et al., 2002; Harrison et al., 1998; Nelson et al., 1996; Yin,
141 2006; Wang et al., 2021). However, the Eocene crustal anatexis event is poorly understood
142 due to the lack of critical evidence. A previous model for the generation of mafic magmas in
143 Tethyan Himalaya involves decompression melting of the asthenosphere triggered by the
144 break-off of the Neo-Tethyan lithosphere (e.g., Ji et al., 2016). However, such a model has
145 difficulties in accounting for the following geological observations. 1) Similar OIB-like
146 magma has not been found in the Lhasa terrane below which the break-off of the
147 Neo-Tethyan lithosphere is proposed to have occurred. 2) It would have been extremely
148 difficult for OIB-like magma to migrate southwards from beneath the Lhasa terrane and
149 emplace as the Indian plate continued to push northward against Eurasia. An alternative
150 model is therefore required to reconcile coeval metamorphic and magmatic records and
151 geological observations along both sides of the suture zone in this collisional orogen.

152 It has been noted that a sudden increase in the convergence rate of the Indian continent
153 toward Eurasia occurred prior to its initial collision with Eurasia was likely a response to
154 enhanced pull caused by the steepening subduction of the Neo-Tethyan plate (Fig. 3; Chung et
155 al., 2005). Such a steepening of subduction may also have resulted in the likely steep
156 geometry of the early subduction of the Indian continental margin (Qi et al., 2020). This
157 steepening Neo-Tethyan and Indian lithospheric subduction may also have triggered
158 subduction channel widening and asthenospheric upwelling under the southern Lhasa terrane
159 (Kelly et al., 2019). This would have eventually resulted in break-off of the Neo-Tethyan
160 oceanic slab (Hou et al., 2012), and/or lithospheric delamination of southern Lhasa terrane
161 (Qi et al., 2020), causing melting to produce the ca. 51 Ma magmatism in southern Lhasa
162 block along the suture zone (Fig. 3).

163 We propose here that after the 51 Ma magmatic event, a lithospheric flexure formed along the
164 northern Indian continent parallel to the suture, causing brittle cracking (i.e., bending-induced
165 faults; Romeo and Álvarez-Gómez, 2018) and the 48–45 Ma melting in northern Himalaya
166 (Fig. 3a). The lithospheric flexure could have resulted from either: 1) the break-off of the
167 subducted Neo-Tethyan lithosphere at ca. 50 Ma (Fig. 3a), and the resultant
168 buoyancy-induced upward bending of the leading edge of the Indian continental lithosphere;
169 or 2) the slowdown of subduction along the leading-edge of the subducting Indian lithosphere
170 at ca. 50 Ma due to the buoyancy of the Indian continental crust while the Indian plate was
171 still continuously pushing northward. This is consistent with the rapid slowdown of the Indian
172 continent's northward movement since 50 Ma (Fig. 1c, Cande et al., 2010). In addition, the
173 loading caused by crustal thickening after the collision may also have contributed to

174 lithospheric bending (Fig. 3a).

175 In our model, the melts derived by decompression melting of the LAB intruded to the
176 shallow levels of the lithosphere along extensional fractures below neutral plane of the
177 downwarped lithosphere (Fig. 3). Given the rapid thickening of the Indian continental crust
178 during the early collision, the neutral plane of the bended lithosphere could have been close to
179 or above the Moho, with shortening above the neutral plane mostly absorbed by the series of
180 crust thrusting (Fig. 3a). The high buoyancy of volatile-rich LAB melts and potential
181 thermal–mechanical–chemical erosion could have driven the migration of the LAB melts
182 through the lithosphere (Spence and Turcotte, 1985). Emplacement of such mafic magmas
183 may not only form the reported gabbros and mafic dikes which are now exhumed to the
184 surface by kilometers of erosion, associated mafic underplating in the crust likely also
185 provided the heat source for the formation of coeval (51–40 Ma) orogen-parallel crustal
186 anatexis, and thus the orogen-parallel gneiss domes (Hou et al., 2012; Weller et al., 2021).
187 After this 48–35 Ma magmatic event, the ongoing over-thrusting and crustal compression
188 enabled the accumulation of radiogenic heat in the lower crust. This radiogenic decay resulted
189 in an elevated geotherm and causing the subsequent larger-scale crustal anatexis during the
190 Miocene (Fig. 1). However, discussion and modelling of this process is beyond the scope of
191 the present paper.

192 Overall, our model provides new insights into the mechanisms of magma generation and
193 orogenic evolution within the lower plate (a previous passive margin)-side of convergent
194 orogens. Such mechanisms may also be applicable to converging oceanic lower plates, where
195 explanations for the mechanism of orogen-parallel magmatism range from a mantle transition

zone origin (Yang and Faccenda, 2020) to melts formed at the LAB due to either lithospheric flexure-related extension (e.g., Hirano et al., 2006; Pilet et al., 2016) or enhanced pull of the subducting plate (Dan et al., 2021). Elements of our model also share similarities to that of Yuan et al. (2010) proposed for the formation of Triassic granitoids in the eastern Songpan-Ganzi Fold Belt. Our model may also be applicable to magmatism along the passive side of other collisional orogens.

ACKNOWLEDGMENTS

We appreciate valuable suggestions of Prof. U. Schaltegger, Prof. D. Brown, Prof. B. White, Prof. A. Yin, Prof. O. Vanderhaeghe and an anonymous reviewer. We also thank Prof. Yin-De Jiang and Dr. Xiu-Zheng Zhang for beneficial discussions. Financial support was provided by the Second Tibetan Plateau Scientific Expedition and Research (2019QZKK0702), the National Natural Science Foundation of China (Nos. 42122022, 41872062 and 42021002). ZXL acknowledge support from the Australian Research Council (Laureate Fellowship grant FL150100133). This is contribution XXXXXX from GIGCAS.

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314 Figure 1. (a) Simplified geologic map of the Himalayan orogenic belt, southern Tibet
 315 (after [Zeng et al., 2011](#)) showing locations of the magmatism and metamorphism in
 316 the northern Himalaya, as well as the location of the study area. Locations and ages of
 317 Eocene magmatism in Tethyan Himalaya are listed in [Table DR1](#). YZS: Yarlung
 318 Zangbo suture; STDS: Southern Tibet Detachment System; MCT: Main Central
 319 Thrust; MBT: Main Boundary Thrust; LH: Lower Himalaya. (b) Geological map of
 320 the study area. (c) Histogram of ages for Eocene magmatic rocks in the Tethyan
 321 Himalaya. The convergence rate of Indian continent is shown as green dotted line.
 322 The dark blue line shows the kernel density estimate ([Chapman and Kapp, 2017](#)) for
 323 the age of the southern Lhasa magmatism. (d) Plots of representative
 324 pressure-temperature (P-T) and thermal gradients for Cenozoic metamorphism in the
 325 Himalaya. The data and references are listed in [Table DR5](#).

326

327 Figure 2. Geochemistry and Tera-Wasserburg diagram of the Gyantse dikes. The
 328 Langshan gabbro data are from [Ji et al. \(2016\)](#). The data for Tethyan MORB, Panjal
 329 basalt, Eocene granite and Oligo-Miocene leucogranite shown in [Table DR6](#).
 330 Langshan gabbro (12FW63, $[\text{}^{87}\text{Sr}/\text{}^{86}\text{Sr}]_i = 0.706571$, $\epsilon_{\text{Nd}}(t) = 5.8$) and Panjal low-Ti
 331 basalt (PJ2-014, $[\text{}^{87}\text{Sr}/\text{}^{86}\text{Sr}]_i = 0.712667$, $\epsilon_{\text{Nd}}(t) = -6.4$) represent the asthenospheric
 332 and lithospheric mantle end-members for modeling, respectively. The numbers along
 333 the blue tick-line indicate percentage contribution of asthenosphere material. MSWD
 334 = mean square of weighted deviation.

335

336 Figure 3. Schematic diagram illustrating the formation of the Eocene (ca. 50-35 Ma)
337 orogen-parallel magmatic and metamorphic zone in the Tethyan Himalaya due to
338 lithospheric flexure. Mafic melts from the lithosphere-asthenosphere boundary
339 percolate into the Indian continental lithosphere and underplate the continental crust
340 along fractures, causing coeval orogen-parallel thickened crustal anatexis. The
341 steepening subduction of Neo-Tethyan and Indian lithosphere resulted in the
342 subduction channel widening and asthenospheric upwelling and/or a slab break-off,
343 causing melting to produce coeval magmatism in the Lhasa block. (b) Elevation and
344 S-wave receiver function profiles along the dark blue dotted line in [Fig.1a](#), are
345 adapted from [Zhao et al. \(2010\)](#). The blue low velocity zone indicates a possible
346 partial melting zone.





