

This is an Open Access document downloaded from ORCA, Cardiff University's institutional repository: <https://orca.cardiff.ac.uk/id/eprint/74367/>

This is the author's version of a work that was submitted to / accepted for publication.

Citation for final published version:

Raschke, Ulli, Schmitt, Ralf Thomas, McDonald, Iain , Reimold, Wolf Uwe, Mader, Dieter and Koeberl, Christian 2015. Geochemical studies of impact breccias and country rocks from the El'gygytgyn impact structure, Russia. *Meteoritics and Planetary Science* 50 (6) , pp. 1071-1088. 10.1111/maps.12455

Publishers page: <http://dx.doi.org/10.1111/maps.12455>

Please note:

Changes made as a result of publishing processes such as copy-editing, formatting and page numbers may not be reflected in this version. For the definitive version of this publication, please refer to the published source. You are advised to consult the publisher's version if you wish to cite this paper.

This version is being made available in accordance with publisher policies. See <http://orca.cf.ac.uk/policies.html> for usage policies. Copyright and moral rights for publications made available in ORCA are retained by the copyright holders.



METEORITICS & PLANETARY SCIENCE

**GEOCHEMICAL STUDIES OF IMPACT BRECCIAS AND
COUNTRY ROCKS FROM EL'GYGYTGYN IMPACT STRUCTURE,
RUSSIA**

Journal:	<i>Meteoritics & Planetary Science</i>
Manuscript ID:	MAPS-2222.R1
Manuscript Type:	Article
Date Submitted by the Author:	02-Feb-2015
Complete List of Authors:	Raschke, Ulli; Museum für Naturkunde Berlin, Leibniz-Institut für Evolution und Biodiversitätsforschung, Forschung Schmitt, Ralf-Thomas; Institut für Mineralogie, McDonald, Iain; Cardiff University, School of Earth, Ocean and Planetary Sciences ; Reimold, Wolf; Museum für Naturkunde, ; Mader, Dieter; Universität Wien, Lithospheric Research Koeberl, Christian; University of Vienna, Department of Geological Sciences
Keywords:	Geochemistry, El-gygytgyn, Russia < Impact crater, Inductively coupled plasma-mass spectrometry, Instrumental neutron activation analysis

SCHOLARONE™
Manuscripts

GEOCHEMICAL STUDIES OF IMPACT BRECCIAS AND COUNTRY ROCKS FROM THE EL'GYGYTGYN IMPACT STRUCTURE, RUSSIA

Ulli RASCHKE^{1,2*}, Ralf Thomas SCHMITT¹, Iain MCDONALD³, Wolf Uwe REIMOLD^{1,4}, Dieter MADER⁵, and Christian KOEBERL^{5,6}

¹Museum für Naturkunde Berlin, Leibniz Institute for Evolution and Biodiversity Science, Invalidenstraße 43, 10115 Berlin, Germany

²Freie Universität Berlin, Institut für Geologische Wissenschaften, Malteser Str. 74-100, 12249 Berlin

³School of Earth & Ocean Sciences, Cardiff University, Cardiff CF10 3YE, UK

⁴Humboldt-Universität zu Berlin, Unter den Linden 6, 10099 Berlin, Germany

⁵Department of Lithospheric Research, Center for Earth Sciences, University of Vienna, Althanstrasse 14, 1090 Vienna, Austria

⁶Natural History Museum, Burgring 7, 1010 Vienna, Austria

*Corresponding author. E-mail: ulli.raschke@outlook.com

Abstract

The complex impact structure El'gygytgyn in northeastern Russia (age 3.6 Ma, diameter 18 km) was formed in ~88 Ma old volcanic target rocks of the Ochotsk-Chukotsky Volcanic Belt (OCVB). In 2009, El'gygytgyn was the target of a drilling project of the International Continental Scientific Drilling Program (ICDP), and in summer 2011 it was investigated further by a Russian-German expedition. Drill core material and surface samples, including volcanic target rocks and impactites, have been investigated by various geochemical techniques in order to improve the record of trace element characteristics for these lithologies and to attempt to detect and constrain a possible meteoritic component. The bedrock units of the ICDP drill core reflect the felsic volcanics that are predominant in the crater vicinity. The overlying suevites comprise a mixture of all currently known target lithologies, dominated by felsic rocks but lacking a discernable meteoritic component based on platinum group element (PGE) abundances. The reworked suevite, directly overlain by lake sediments, is not only comparatively enriched in shocked minerals and impact glass spherules, but also contains the highest concentrations of Os, Ir, Ru, and Rh compared to other El'gygytgyn impactites. This is - to a lesser extent - the result of admixture of a mafic component, but more likely the signature of a chondritic meteoritic component. However, the highly siderophile element contribution from target material akin to the mafic blocks of the ICDP drill core to the impactites remains poorly constrained.

INTRODUCTION AND GEOLOGICAL BACKGROUND

The El'gygytgyn impact structure is located on the Chukotka Peninsula of far northeast Russia; it is centered at 67°30'N and 172°34'E (Fig. 1). The 18 km diameter, near-circular depression is largely filled by the 12 km wide Lake El'gygytgyn. The impact age was determined at 3.58 ± 0.04 Ma (Layer 2000). The volcanic target rocks belong to the Late Cretaceous Ochotsk-Chukotsky Volcanic Belt (OCVB) that is of Albian to Campanian/Maastrichtian (86-106 Ma) age (Belyi and Belaya 1998; Raschke et al. 2014 and references therein). The target lithologies are generally known from the work of Belyi (1994), Belyi and Belaya (1998), and from Gurov and co-workers (Gurov et al. 1978, 2005, 2007; Gurov and Gurova 1983). These authors described the OCVB rocks as a suite comprising (from top to bottom): ignimbrites (mainly felsic, 250 m); tuffs and rhyolitic lavas (200 m); tuffs and andesitic lava (70 m, occurring especially to the southwest of the crater); and finally ash and welded tuffs of rhyolitic and dacitic compositions (100 m). Above this sequence a ca. 110 m thick

1
2
3 basalt sill occurs as a plateau at the northeastern crater rim (Gurov et al. 2004). Additionally, there are
4 previously unknown lithologies at the southeastern crater rim that were defined for the first time by
5 Raschke et al. (2014). Mount Otvevergin, on the northeastern lakeshore, is composed of reddish and
6 greenish ignimbrites. In the southeastern sector of the lake several mini-plateaus occur that are made
7 up of (sub)horizontal basalt or andesite layers; they are, on aggregate, $\sim 2 \text{ km}^2$ in area extent. To the
8 south of the lake, a suite of gray to reddish, basaltic-andesitic tuffs is present (see Fig.1).
9
10

11 **Figure 1**

12
13
14 The crater rim is well preserved, except for the southeastern part that has been eroded by the
15 Enmyvaam River, a periodic outflow from the lake. Previous studies have shown that rocks of the
16 crater rim did not reveal any characteristic shock metamorphic effects (Gurov et al. 2007; Raschke et
17 al. 2014). The originally in situ ejecta deposits (comprising a mélange of unshocked and shocked
18 rocks, and fragments of impact melt breccia) around the impact crater have been nearly completely
19 eroded by arctic weathering. Only a few allochthonous remnants have been found, embedded in the
20 lacustrine and fluvial terraces inside and outside of the crater rim. These include rounded cobbles (2-
21 15 cm in size), and larger, meter-sized blocks of dark impact melt breccia (Raschke et al. 2014;
22 Pittarello et al., 2013; and references therein). Aerodynamically shaped glass bombs occur together
23 with shock metamorphosed rocks in the lacustrine terraces inside the crater and also in terraces along
24 some streams (e.g., along the Enmyvaam river) in the environs of the crater. All recorded types of
25 impactites from the wider crater area are generally fresh and most of the samples described do not
26 display significant post-impact hydrothermal alteration and weathering (Gurov and Koeberl 2004;
27 Raschke et al. 2014). The impact origin was confirmed by Gurov and co-workers, who found evidence
28 for shock metamorphism in some samples from the crater region (Gurov et al. 1978, 1979, 2005). That
29 includes planar deformation features in quartz, diaplectic quartz glass, coesite and stishovite, and
30 planar fractures in quartz (which by themselves are not shock diagnostic).
31
32
33
34
35

36 In spring 2009 an International Continental Scientific Drilling Program (ICDP) drilling
37 campaign (summarized in Koeberl et al. 2013) recovered a $\sim 520 \text{ m}$ long drill core, comprising $\sim 318 \text{ m}$
38 of lacustrine sediments and $\sim 200 \text{ m}$ of impactites (drilling location shown in the cross-section of Fig.
39 2). The drilled impactites can be stratigraphically divided (from top to bottom, see Fig. 3) into $\sim 12 \text{ m}$
40 of reworked suevite (316.77 – 328.00 m below lake floor [mblf]), $\sim 63 \text{ m}$ of suevite (328.00 – 390.74
41 mblf), and $\sim 30 \text{ m}$ of upper (390.74 – 420.89 mblf) and $\sim 96 \text{ m}$ of lower bedrock (420.89 – 517.00
42 mblf) (Raschke et al. 2013a). The lower bedrock is interpreted as (parautochthonous) crater basement.
43 It is crosscut by a single, thin polymict impact breccia dike at 471.42 – 471.96 mblf depth. The upper
44 bedrock unit contains different ignimbrites, and three meter-sized mafic blocks (at ~ 391 , 420, and 422
45 mblf depth). The bedrock units are mainly unshocked but intensely fractured. The suevitic units
46 contain shocked minerals and relatively rare impact melt particles. Only in the reworked suevite, at the
47 top of the drilled sequence, stronger shocked lithic clasts, melt particles and impact-produced glass
48 spherules are abundant (cf. also Wittmann et al. 2013). All drilled rocks are moderately to strongly
49 weathered (for detailed petrographic information, see Raschke et al. 2013b, Pittarello et al. 2013).
50
51
52
53

54 **Figures 2, 3**

55
56 In addition, one of us (UR) participated in a 2011 Russian-German expedition to El'gygytyn
57 to supplement the existing surface geological data base with new mapping results and to obtain surface
58 samples of country rocks and impactites for comparison with drill core lithologies. Based on the 2011
59 surface exploration, an upgraded geological map of the El'gygytyn area was compiled (Raschke et al.
60 2014). The Zr/TiO₂ vs. Nb/Y diagram of Fig. 4 (data from Raschke et al. 2013b, 2014) illustrates the
variability of the compositions of the drill core and surface samples. Both sample sets cover the same

1
2
3 range of compositions. Obviously, the predominance of target rocks in the basaltic or andesitic-
4 basaltic field of Fig. 4 is based on the proportionally higher number of samples analyzed from these
5 lithologies.
6
7

8 **Figure 4**

9 **Impact and volcanic melt rocks in the crater area and in the drill core**

10
11 The distinction between the volcanic and impact melt rocks has proven to be a complex task in
12 the study of the El'gygytgyn crater (cf. Pittarello and Koeberl 2013a). In contrast to the majority of
13 other impact craters on Earth, the classification of melt particles is a basic requirement for the
14 distinction between impact-generated and volcanic melt particles. Furthermore, the determination of a
15 meteoritic component in impact produced melt particles can help to confirm the type of projectile and
16 its role as well as its dissipation in the impact process.
17
18
19

20
21 Volcanic melt particles occur in the ignimbritic rocks of the upper and lower bedrock. They
22 are generally recrystallized and similar in their composition to the rhyolitic or rhyodacitic host rocks.
23 Alkali feldspar and mafic minerals (biotite and amphibole) occur as phenocrysts in the fine-grained
24 melt. Altered glassy fragments are found inside the pumice fragments of the rhyolitic or rhyodacitic
25 ignimbrite. A detailed description of these volcanic melt particles was given by Raschke et al. (2013b,
26 2014).
27

28
29 Impact melt occurs in four different settings: i) blocks of impact melt breccia and glass bombs
30 in the lake terraces; ii) tiny (0.5 - 1.5 mm) glass spherules on the lake terrace and along the Enmyvaam
31 River (Glushkova and Smirnov 2007); iii) similar spherules in the reworked suevite section of the
32 ICDP drill core (Wittmann et al. 2013; Goderis et al. 2013); and iv) small (altered) melt particles in the
33 drilled suevite section (Pittarello et al. 2013; Raschke et al. 2013b).
34

35
36 i) Impact melt breccia sampled on the surface (Gurov and Koeberl 2004) outside the crater
37 structure occurs as a fresh, heterogeneous mélange of glassy, mostly blackish but also translucent
38 "schlieren", which may be rich in vesicles, but relatively poor in mineral or lithic inclusions. Other
39 melt breccia resembles a volcanic scoria with larger clasts of unmelted or only partially molten rock
40 fragments. The composition of such breccia depends on the host rock material and can include pieces
41 of, e.g., pumice, ignimbrite, andesite, or basalt. The minerals in these clasts often show shock features,
42 for example planar fractures, planar deformation features, and diaplectic glass (see Raschke et al.
43 2013b; Pittarello and Koeberl 2013b).
44
45

46
47 ii) Up to 1.5 mm size glass spherules found in lacustrine sediments to the south of the crater
48 (during the Quaternary, lake El'gygytgyn covered a larger surface area and had a higher lake level)
49 and in fluvial terraces along the Enmyvaam River (Gurov 1979; Glushkova and Smirnov 2007) were
50 analyzed by Adolph and Deutsch (2010), Smirnov et al. (2011), and Wittmann et al. (2013). All these
51 authors concluded, on the basis of geochemical data, that the spherules were impact-produced melt
52 droplets that had been deposited from the collapsing ejecta plume (with lithic debris) in a thin layer on
53 the juvenile post-impact surface. Overall, the spherules are strongly heterogeneous, ranging in
54 composition from basaltic to rhyolitic, and are probably derived from the different volcanic lithologies
55 in the target area, which requires, in turn, that the spherules did not undergo homogenization in the
56 ejecta plume (see Wittmann et al. 2013).
57
58

59
60 iii) An accumulation of spherules occurs on top of the reworked suevite section between 317
and 322 mblf. The spherules are very heterogeneous and occur in different types. First, there are
hollow spherules with a glassy margin and may contain a few crystal inclusions or microfragments of

1
2
3 different minerals (e.g. feldspar, quartz and zeolite). Another type of spherule is filled by
4 aluminosilicate glassy melt, which contains microlites of feldspar or of mafic composition (Raschke et
5 al. 2013b and references therein).
6

7
8 iv) Impact melt was identified in the matrix of the suevite section of the drill core between 328
9 and 391 mblf (Raschke et al. 2013b). This comprises very small melt particles, ~1 mm in size, which
10 are generally altered to secondary phyllosilicates (e.g., smectites and chlorites). These particles
11 amount to much less than 1 vol% of the whole suevite package.
12

13 **Previous studies of siderophile elements, Platinum Group Elements, and Rare Earth Elements**

14
15
16 Pittarello et al. (2013) analyzed rare earth element (REE) concentrations of drill core rocks and
17 compared these with volcanic rocks from the regional geological setting. With the exception of data
18 for the mafic blocks from the drill core, all other impactite samples, including the suevites, plot in the
19 same space as the volcanic target rocks. Raschke et al.'s (2014) chemical comparison between
20 impactites of the drill core and regionally occurring lithologies revealed very similar chemical
21 compositions of upper and lower bedrock and the suevitic units, as well as the surface rocks from the
22 crater rim that are dominated by the rhyolitic or rhyodacitic ignimbrites.
23
24

25
26 The enrichment of siderophile elements in microtektites (or microkrystites) is generally a very
27 useful tool for the determination of a projectile signature (Koeberl 2014; Koeberl et al. 2012).
28 According to Wittmann et al. (2013), the siderophile element contents in the spherules of the reworked
29 suevite are highly variable. The El'gygytyn glass spherules show a wide range of compositions,
30 reflecting the geochemical signature of the target lithology assemblage composed of both mafic and
31 felsic rocks (Raschke et al. 2013b; Wittmann et al., 2013). The siderophile element contents of the
32 spherules in the reworked suevite are highly variable (Ni ~30 to 1400 ppm), similar to the spherules
33 from outside of the crater (Ni ~300 to 1100 ppm), and are probably related to projectile contamination
34 (see also Wittmann et al. 2013).
35
36

37
38 Foriel et al. (2013) found that some impact glass samples from the surface of the El'gygytyn
39 area have a chromium isotopic anomaly that agrees best with an ureilite source. They suggested that
40 the impactor could have had a composition similar to that of the Almahata Sitta meteorite from Sudan,
41 which is an ureilite with clasts of ordinary chondrite (Jenniskens et al. 2009).
42

43
44 Platinum group element (PGE) analyses were undertaken by Goderis et al. (2013) on the
45 spherule-bearing deposits, as well as on a few hand specimens of impact melt recovered from the
46 crater rim. Together with their Os isotope and Ir concentration analysis, these authors concluded that
47 rather than an achondritic (ureilitic) impactor composition, an ordinary chondrite type was probable.
48

49
50 Based on these previous studies, especially the instrumental neutron activation analysis
51 (INAA) data of Pittarello et al. (2013), as well as work done on drill core and country rock samples by
52 Raschke et al. (2013b, 2014), we decided to try to derive more information about the geochemical
53 character of the impactites and their target rocks, including the comparison with impact melt breccia
54 that was collected on the lake terraces within the crater. Another goal has been the identification of a
55 meteoritic component using siderophile element abundances in impactites from the El'gygytyn
56 crater.
57
58
59
60

SAMPLES AND ANALYTICAL METHODS

1
2
3 A suite of 17 samples from the ICDP drill core (impactites, including suevite and bedrock
4 lithologies) was selected for INAA. A second suite of samples (7 ICDP drill core and 10 surface
5 specimens) was used for PGE analysis. Some petrographic and chemical details about the surface
6 samples have previously been presented in Raschke et al. (2014). Sampled drill core depths (this work
7 and from Pittarello et al. 2013) are given in Table 1.
8
9

10 **Table 1**

11
12 The measurements by INAA were carried out at the Department of Lithospheric Research,
13 University of Vienna. The contents of some major (Na, K, and Fe) and many trace elements (including
14 the REE) were determined using this method. In general, about 130 mg of powdered sample was
15 sealed in a polyethylene capsule and irradiated in the 250 kW Triga Mark-II reactor of the Atomic
16 Institute in Vienna. For calibration three international rock standards were used: (i) Allende
17 carbonaceous chondrite (Smithsonian Institution, Washington DC, see Jarosewich et al. 1987); (ii)
18 Ailsa Craig Granite AC-E (Centre de Recherche Petrographique et Geochimique, Nancy, France, see
19 Govindaraju 1989); and (iii) Devonian Ohio Shale SDO-1 (USGS, see Govindaraju 1994). Further
20 details about the method, technique, and accuracy on results is given by Koeberl (1993) and Mader
21 and Koeberl (2009). The INAA data for the various lithologies of the ICDP drill core are reported in
22 Table 2.
23
24
25
26

27 **Table 2**

28
29 The contents of the PGE and Au were determined in Cardiff by inductively coupled plasma-
30 mass spectrometry (ICP-MS) after pre-concentration by Ni-sulfide fire assay with co-precipitation,
31 using external calibration. For each sample 15 grams of material was used. Two low level
32 concentrations of powdered reference material were used for the validation of PGE analysis: i) WITS-
33 1 (a silicified komatiite and ultramafic rock from the Barberton area, South Africa), and ii) TDB-1, a
34 basaltic (diabase) rock sample from Canada (Tredoux and McDonald 1996). More details regarding
35 the analytical technique and the related precision and accuracy values have been published in Huber et
36 al. (2001) and McDonald and Viljoen (2006). For drill core and surface samples the PGE and Au
37 abundance data are reported in Table 3.
38
39
40

41 **Table 3**

42
43 In addition, we used the datasets of siderophile elements from petrographic and geochemical
44 studies, which we have already published for the drill core material (Raschke et al. 2013b) and for the
45 surface samples of the wider crater region (Raschke et al. 2014). Additional trace element data for the
46 ICDP drill core from Pittarello et al. (2013) measured by INAA in the same laboratory as our samples
47 were used to extend the data set, especially for scarce lithologies such as the mafic blocks. All samples
48 are listed in Table 1 and Fig. 3. Using this large dataset we tried to discriminate special characteristics
49 of the reworked suevite (including layers of impact produced glass spherules) and the other impactites
50 from the drill core in contrast to the target rocks from the crater vicinity, inclusive of impact melt
51 breccia from the lake terrace. Furthermore, we compared our results with respect to the data of
52 Goderis et al. (2013), Wittmann et al. (2013), Foriel et al. (2013), and Pittarello et al. (2013).
53
54
55
56
57

58 **RESULTS**

59 **Composition of the drill core material and target rocks**

1
2
3 The El'gygytgyn drill core material and the surface samples mainly comprise felsic volcanic
4 rocks. Rhyolitic or rhyodacitic ignimbrites are the predominant rock types in the drill core (lower
5 bedrock unit, ~50 % of the impactite section) as well as regarding the country rocks. In the vicinity of
6 the crater more than 90 % of the country rocks are SiO₂-rich volcanics (Raschke et al. 2014). The
7 mafic rocks, i.e., basalts, andesitic basalts, and their eruptive equivalents (phreatomagmatic tuffs),
8 form a minor contribution in the area and are only found in the southeastern sector of the crater
9 environs.
10
11

12
13 In this work, we focus on four types of lithologies for chemical discrimination and
14 interpretation: i) the reworked suevite with accumulated impact glass spherules in the groundmass (see
15 Raschke et al. 2013b and Wittmann et al. 2013, as well as references therein); ii) the impact melt
16 breccia from the lake terrace that might carry a possible meteoritic component; iii) the suevite, a
17 mélange of all possible target lithologies and impact melt particles; and iv) the mafic blocks from the
18 drill core between upper and lower bedrock unit. These blocks are possibly derived from basaltic
19 intrusions (sills) and are highly altered and fractured. These altered samples are characterized by a
20 high loss on ignition (LOI) as well as an extraordinary chemical signature in comparison to all other
21 target rocks; they are enriched in a wide range of metal oxides and easily recognizable in the
22 compositional discrimination diagrams.
23
24

25 26 Rare Earth Elements

27
28 The average REE contents of the different lithologies of the ICDP drill core from this and
29 previous studies are summarized in Table 4. The CI chondrite normalized REE patterns for sampled
30 lithologies are shown in Figs. 5a-c. The patterns of the average upper and lower bedrock of the ICDP
31 drill core (Fig. 5a) are very similar. They indicate enrichments for the average upper and lower
32 bedrock by factors of 75 to 89 for La, and 10 to 8 for Yb, respectively, compared to CI chondrite
33 composition. The light REE (LREE) are enriched compared to the heavy REE (HREE) (average
34 La_N/Yb_N 8-10), and a negative Eu anomaly (average Eu/Eu* ~ 0.6 to 0.7; Eu/Eu* = Eu_N/(Sm_N x
35 Gd_N)^{0.5}) is characteristic for these rocks. Another prominent feature of the upper and lower bedrock is
36 a flat pattern of HREE. In comparison to the rocks of the Ochotsk-Chukotsky Volcanic Belt (OCVB),
37 the upper and lower bedrock show less fractionation and slightly lower REE ratios, namely La_N/Yb_N
38 ratios of 7.9 and 10.8 for the upper and lower bedrock, respectively, compared to ~ 8 to 18 for the
39 OCVB, and La/Sm ratios of 3.7 and 4.9, respectively, compared to 5 to 8 for the OCVB (Tikhomirov
40 et al. 2008).
41
42
43
44

45 In contrast to the felsic target rocks, the mafic blocks of the ICDP drill core display different
46 REE patterns (**Fig. 5b**). The CI chondrite-normalized REE patterns of the mafic block samples show
47 comparable signatures characterized by an enrichment of the LREE compared to the HREE and a
48 slightly fractionated profile for the HREE. The REE patterns show different enrichments for the mafic
49 blocks at 391, 420 and 422 mblf by factors of 134, 50, and 143 for La, and 9, 9, and 15 for Yb,
50 respectively, compared to CI chondrite composition. The enrichment of the LREE is more prominent
51 in the blocks at 391 and 422 mblf with La_N/Yb_N ratios of 14 and 9.5, respectively, compared to the
52 block at 420 mblf with a La_N/Yb_N ratio of 5.9. The REE patterns for the mafic blocks at 391 and 422
53 mblf do not show distinct Eu anomalies, whereas the block at 420 mblf displays - in contrast to all
54 other lithologies - a slightly positive Eu anomaly with a Eu/Eu* ratio of 1.15. However, these blocks
55 are very heterogeneous, and it is difficult to compare these with each other or with other lithologies
56 from the drill core, crater, and the OCVB.
57
58
59
60

The average signatures for suevite, the polymict impact breccia dike, and the reworked suevite
of the ICDP drill core display similar REE patterns (Fig. 5c). All lithologies show an enrichment of

1
2
3 the REE compared to the CI chondrite composition by factors of 90, 79, and 90 for La, and 8, 8, and
4 10 for Yb for the suevite, polymict impact breccia dike, and reworked suevite, respectively. The LREE
5 are enriched compared to the HREE in these lithologies with La_N/Yb_N ratios of 10.6, 9.4, and 9.4,
6 respectively, and negative Eu anomalies are present, with Eu/Eu^* ratios of 0.60, 0.69, and 0.58 for the
7 suevite, polymict impact breccia dike, and reworked suevite, respectively. The REE patterns of the
8 suevite and polymict impact breccia dike show strong similarities to those of the upper and lower
9 bedrock, and indicate that the suevite mainly formed from these target lithologies. This is also visible
10 in the Yb vs. Gd diagram (Fig. 5d). The reworked suevite indicates some slight differences in the REE
11 patterns from those for the suevite. The absolute concentrations of the REE and the enrichments of the
12 REE compared to CI chondrite composition are slightly higher, and the negative Eu anomaly is lower
13 in the reworked suevite in comparison to the suevite and the lower and upper bedrock. This behavior
14 could be explained by an additional admixture of mafic material in the reworked suevite compared to
15 the suevite, as suggested in the Yb vs. Gd diagram (Fig. 5d).
16
17
18
19
20
21

22 **Figure 5, Table 4**

23 **Siderophile elements**

24
25
26
27
28 The concentrations of the siderophile elements Co, Ni, and Cr, and the Ni/Cr, Ni/Co, and
29 Cr/Co ratios are summarized for the different lithologies of the ICDP drill core in Table 5. Our results
30 show that, in general, the siderophile element concentrations are low in the felsic (lower and upper
31 bedrock) and distinctly higher in the mafic target lithologies (mafic blocks), with the highest
32 concentrations of siderophile elements having been measured for the mafic block at ~420 mblf. The
33 concentrations of the siderophile elements and their ratios within the suevite are quite similar to the
34 respective concentrations and ratios in the lower and upper bedrock. The concentrations of siderophile
35 elements reported for impact melt rocks and glass bombs collected at the surface around the crater are
36 also in this range, with concentrations of <50 ppm Cr, <7 ppm Co, and <21 ppm Ni (Gurov and
37 Koeberl 2004; Gurov et al. 2005). Therefore, a contamination of the suevite and the impact melt rocks
38 by a meteoritic component is not obvious in these siderophile element abundances.
39
40
41
42

43 Slightly higher concentrations of siderophile elements together with lower Ni/Cr and higher
44 Ni/Co and Cr/Co ratios in comparison to the suevite unit are observed in the reworked suevite and
45 within a polymict impact breccia dike occurring in the lower bedrock at ~471 mblf. For the impact
46 spherules (Wittmann et al. 2013) the contents of siderophile elements (measured by LA-ICP-MS) are
47 much higher in comparison to all other target lithologies (Table 5), e.g. the Ni data for some samples
48 (sph6 at 317.60 mblf) show high values up to 1400 ppm (Wittmann et al. 2013). Regarding to the
49 moderately siderophile element budget of the reworked suevite (Table 5), these spherules are
50 negligible. These observations agree with the results of Pittarello et al. (2013) and Goderis et al.
51 (2013). Therefore, the higher concentrations of siderophile elements in the reworked suevite and
52 polymict impact breccia dike, and their different ratios in comparison to the suevite, are most likely
53 the result of a higher amount of mafic material within these impactites. Overall, the observed
54 siderophile element ratios for the suevite, reworked suevite, and polymict impact breccia dike do not
55 match meteoritic ratios (e.g., Tagle and Berlin 2008; Koeberl 2014).
56
57
58
59
60

Table 5

Platinum Element Group analysis – the presence of a meteoritic component

Results of the PGE and Au analysis are given in Table 3 and plotted in Figs. 6 and 7. The Ir contents of the target rocks vary between < 0.03 and 0.52 ppb (Table 3). The Ir concentrations of the felsic lithologies are generally low (< 0.10 ppb), whereas higher Ir contents (0.52 ppb) were measured for the basaltic target lithologies, especially for the highly altered and metal oxide enriched mafic blocks at ~ 420 and 422 mblf in the drill core. The high Ir concentrations in the mafic blocks are associated with high Os concentrations, but also with elevated concentrations of Pt, Pd, and Au that are typical of many mafic lavas (e.g., Barnes et al. 1985; Tredoux et al. 1995; McDonald 1998; Crocket 2002).

The Ir contents of the suevite, impact melt breccia and polymict impact breccia dike samples are in the range of 0.04 to 0.09 ppb, and in good agreement with data previously presented by Goderis et al. (2013), who determined a range from 0.05 to 0.20 ppb for similar samples. Gurov and Koeberl (2004) reported Ir concentrations of 0.02 to 0.11 ppb for impact melt rocks and glass bombs from El'gygytgyn, which also corresponds well with our new measurements.

Notably part of the reworked suevite has a significantly higher PGE concentration in comparison to the suevite, impact melt breccia, and polymict impact breccia dike, as well as most of the felsic and mafic target lithologies (Table 3), in terms of Os (0.40 ppb), Ir (0.42 ppb), Ru (0.64 ppb), and Rh (0.19 ppb) (Fig. 6c). Additionally, these values are very similar to those for the mafic block at ~ 420 mblf, but also considerably increased in comparison with the mafic blocks at ~ 391 and 422 mblf. The Os/Ir ratio of the reworked suevite is higher (~ 1) compared to the values for the mafic blocks at ~ 420 and 422 mblf (~ 0.8 ; an Os/Ir-ratio < 1 is typical for mafic magmas (Barnes et al. 1985).

Figure 6, 7

DISCUSSION

Goderis et al. (2013) analysed a wide range of siderophile element contents in the mafic block at ~ 391 mblf, in the dike of polymict impact breccia (471 mblf), and in the reworked suevite at 318.9 mblf (named by these authors as “bottom of reworked fallout deposit”) of the ICDP drill core. Raschke et al. (2013b) also reported high concentrations of Ni, Cr, and Co for the mafic blocks from the drill core (423 to 391 mblf). Goderis et al. (2013) reported that the $^{187}\text{Os}/^{188}\text{Os}$ isotopic signal of the mafic block at 391.6 mblf is much more radiogenic (2.8 ± 0.1) than the reworked suevite ($0.148 \pm 0.001 - 0.239 \pm 0.006$). This suggests the Os in the reworked suevite cannot be derived from the mafic component. Consequently, the mafic blocks and similar lithologies cannot be the only contributors to the moderate siderophile elements budget of the drilled impactites. The Ni/Cr and Cr/Co abundance for some samples are between the values of chondritic and primitive achondritic (ureilitic) meteoritic components, especially for impact glass spherules from outside of the crater. The Ni/Co ratios fall between values for ureilites, branchinites, and chondrites (Warren et al. 2006).

The distribution of spherules in the reworked suevite section is reminiscent of similar impact spherules found in the ICDP drill core LB-5 from the Bosumtwi crater in Ghana (Koeberl et al. 2007). Bosumtwi is a 10.5 -km diameter complex impact structure in the same size range as El'gygytgyn. These spherules were preserved in what has been interpreted as the youngest fallback deposit (Koeberl et al. 2007). At Bosumtwi, despite the presence of a high indigenous component linked to ultramafic target rocks, the spherule-bearing deposit shows a slightly elevated and distinct (i.e., unfractionated) PGE signature (Goderis et al. 2007).

1
2
3 Quantitative chemical analysis by EMPA-EDX has indicated that the glasses in these
4 spherules are compositionally heterogeneous (Koeberl et al. 2007a). The detection of the projectile
5 component is a difficult and complicated task, because some of the target lithologies with high PGE
6 contents mask the presence of an extraterrestrial component. For the El'gygytgyn impact crater,
7 Goderis et al. (2013) determined generally very low PGE contents in the impactites (> 50 % under
8 quantification limit) with the result that Ir, Ru, Pt, and Rh are slightly enriched in the reworked suevite
9 and the impact melt breccia, while Pd and Au are not equally elevated. In general, the PGE and Au
10 plots show that the El'gygytgyn samples are generally comparable to chondritic patterns. Based on the
11 slight Ir enrichment with flat, nonfractionated CI-normalized PGE patterns for the reworked suevite,
12 Os isotope ratios for the spherule-bearing deposit that are inconsistent with the target rock
13 composition, and mixing models for the major and Cr, Co, and Ni composition of the spherules
14 characterized by LA-ICP-MS, Goderis et al. (2013) favored an ordinary chondrite (possible LL-type)
15 as the most likely type of projectile for El'gygytgyn.
16
17
18
19

20
21 Foriel et al. (2013) compiled analytical data from Pittarello et al. (2013) of the ICDP drill core
22 and a glass bomb, which was collected at the crater surface. Additionally, these authors used data by
23 Val'ter et al. (1982) and Gurov and Koeberl (2004). Similar to Goderis et al. (2013), Foriel et al.
24 (2013) found an enrichment of siderophile elements (Cr, Co, and Ni) for the suevite of the drill core,
25 but could not substantiate a meteoritic component, because it was not possible to constrain the
26 influence of mafic target rocks (indigenous component). Nonetheless, they found in one of their
27 impact glass samples non-terrestrial Cr isotopic values. Such values are close to those of ureilitic
28 meteorites, but also within analytical error of the range determined for eucrites and ordinary
29 chondrites. These authors concluded that the ratios for siderophile elements did match neither
30 chondritic nor achondritic meteorite compositions. Based on the Cr isotope data, Foriel et al. (2013)
31 favored a ureilite type impactor, although an ordinary chondrite could not be excluded. Other types of
32 meteorites were considered unlikely though.
33
34
35

36 Here, we present new results on trace element compositions, including siderophile elements,
37 especially the PGE, of the impactites and target rocks from the El'gygytgyn impact crater (Tables 2-5).
38 The concentrations of the siderophile elements (Cr, Co, and Ni) are typically very low in the felsic
39 volcanics/ignimbrites, but slightly enriched in the mafic target lithologies and extraordinarily high in
40 the three mafic blocks of the drill core (Raschke et al. 2013b, 2014; Pittarello et al. 2013). The
41 siderophile element, as well as the REE abundances and patterns, for the upper and lower bedrock of
42 the drill core correspond to those for suevite samples (Figs. 5a-c, Tables 4, 5). These observations are
43 in agreement with those of Goderis et al. (2013). Therefore, the suevite represents mixtures of all
44 target lithologies in accordance with their regional proportions. The contribution of the mafic target
45 lithologies (~ 7 % based on surface geology, Raschke et al. 2014) to the trace element budget of the
46 suevite is negligible.
47
48
49

50 Generally, the PGE concentrations (Table 3), their ratios (Fig. 6), and the CI-normalized PGE
51 patterns (Fig. 7) for the suevite are also in the same range as the data for the felsic to intermediate
52 target lithologies. The PGE data confirm the observations based on siderophile element abundances,
53 and, therefore, a meteoritic component could not be detected in the suevite based on trace element data
54 alone. The parautochthonous origin of the lower bedrock drilled in the crater basement, as discussed in
55 Raschke et al. (2013b), could be confirmed by these trace element data. The chemical characteristics
56 of the felsic surface rocks and the lower bedrock are similar and represent the same lithology, namely
57 rhyodacitic ignimbrite.
58
59
60

The reworked suevite at the top of the impactite section of the drill core contains a larger amount of strongly shocked lithoclasts, impact melt particles, and impact glass spherules, and is

1
2
3 chemically characterized by an enrichment of Fe-, Al-, and Mg-oxides compared with all other
4 impactites (Raschke et al. 2013b). Also, the REE concentrations and patterns (Fig. 5, Table 4) display
5 a slight difference to the suevites and the felsic target lithologies. A comparatively higher proportion
6 of a mafic component in the reworked suevite could provide an explanation for these differences. For
7 this process two different scenarios, or a combination of these, can be imagined: (i) First, suevite is
8 formed as a ground surge inside the inner crater. This is followed by addition of highly shocked clasts
9 from all target rock types, and intercalation of mafic and intermediate rocks especially at the top of the
10 suevite sequence due to debris coming off the collapsing crater rim - besides mixing in of some
11 material from the ejecta plume. (ii) Second, the pre-impact geology of the target volume could have
12 contained a higher proportion of mafic and intermediate rocks than indicated by the crater environs
13 today. This could be supported by the actual stratigraphy of the crater rim (Raschke et al. 2014). The
14 older rocks (felsic ignimbrites of the Pykarvaam Formation) are partly covered in the SE and E of the
15 crater by sub-horizontal layers of younger (Voron'in and Koekvun' formations) basalts and andesites.
16 In addition, phreatomagmatic tuffs of basaltic-andesitic composition occur to the south of the crater
17 (Raschke et al. 2014).

18
19
20
21
22
23 However, the siderophile elements and PGE are significantly enriched in the reworked suevite
24 in comparison to all other impactites and most of the target lithologies (Figs. 6, 7, Tables 3, 5). The
25 idea of admixture of a mafic component to form the package of reworked suevite, as mentioned
26 before, cannot explain the high values of Os, Ir, Ru, and Rh found for this unit, in comparison to the
27 composition of the mafic target lithologies (Table 3). Only the mafic blocks drilled in the ICDP core,
28 especially the mafic block at ~420 mblf, have significantly enriched PGE values, which are in the
29 range of the PGE values of the reworked suevite. Nevertheless, it is not plausible that a very strong
30 mafic contamination similar to the composition of the mafic blocks would alone be responsible for the
31 high PGE concentrations in the reworked suevite based on mass balance for the major and other trace
32 elements, including the REE and iron (see Figs. 5-7). However, a hitherto undiscovered, additional
33 ultramafic lithology is possible but so far remains hypothetical. Therefore, a contamination by a
34 meteoritic component in this uppermost reworked suevite seems plausible. A combination of the two
35 scenarios described above, a mixing during the crater collapse with an additional input from meteoritic
36 components and a proportion of basaltic target rocks, would probably be the best-fit hypothesis. This
37 is similar to the findings of Goderis et al. (2013), who also suggested the likely admixture of a
38 meteoritic component to the reworked suevite.

39
40
41
42
43 The average PGE concentrations of the El'gygytgyn target (Table 6) were calculated using the
44 surface area proportions of the target lithologies from Raschke et al. (2014), and the PGE
45 concentrations of these lithologies from Table 3. Based on these data, we attempt to reproduce the
46 PGE content of the reworked suevite, especially the Os, Ir, and Ru concentrations, by mixing the
47 average El'gygytgyn target with different proportions of average ureilite (Warren et al. 2006), LL and
48 CI chondrite (Tagle and Berlin 2008). The best fits for these mixtures, based on a fixed Os
49 concentration according to the content of the reworked suevite, were achieved with an admixture of
50 0.12 % ureilite, 0.10 % LL chondrite, and 0.07 % CI chondrite component, respectively (see Table 6).
51 A comparison between these three meteoritic components shows that the best match could be achieved
52 with admixture of both chondritic components. A better calculation including major and siderophile
53 elements is currently not possible, because the majority of data were measured by XRF and not by
54 INAA or LA-ICP-MS.

55
56
57
58
59
60 A similar finding is revealed by comparison of the Os/Ir and Os/Ru ratios, which are 0.95 and
0.63, 1.23 and 0.82, 1.08 and 0.70, and 1.06 and 0.70, for the reworked suevite, average ureilite, LL,
and CI chondrite, respectively (data for ureilites from Warren et al. 2006, and for chondrites from
Tagle and Berlin 2008). These results suggest the possible admixture of a chondritic component to the

1
2
3 reworked suevite similar to the findings of Goderis et al. (2013). Taking into account the moderately
4 siderophile element ratios reported by these authors for the spherules in the reworked suevite section,
5 an ordinary chondrite component seems to provide the best option as a possible impactor for the
6 El'gygytgyn impact, based on the PGE data.
7
8

9 The method used by Foriel et al. (2013) to determine the nature of projectile component by Cr
10 isotopic measurements would be difficult to use on the reworked suevite samples, because the Cr
11 isotope method is generally capable of detecting only ≥ 1 % extraterrestrial component, whereas PGE
12 abundances allow to determine somewhat lower meteoritic admixtures (in rare cases to about 0.2 %)
13 (cf. Koeberl 2014; Koeberl et al. 2002). Nevertheless, the uncertainties about the role of the mafic
14 blocks with their relatively high PGE concentrations and their possible contribution to the reworked
15 suevite prevent the unambiguous detection of a meteoritic component. The nature of these impactites
16 requires further investigation.
17
18

19 20 21 22 CONCLUSIONS

- 23
24 1. Impact melt breccia found at the surface is obviously a mélange of mainly rhyo(dacitic)
25 ignimbrite and rare basaltic andesite, based on major and trace element compositions.
26 Compared with the drilled rocks, the composition of the suevite and the upper bedrock
27 unit closely matches the impact melt breccia. The PGE content of the impact melt breccia
28 is also similar to that of the suevite sequence between 328 and 391 mblf of the ICDP drill
29 core.
30
- 31 2. Based on PGE analyses, the suevite in the drill core does not show evidence of any
32 unambiguous meteoritic contamination.
33
- 34 3. The mafic blocks of the drill core (between suevite and lower bedrock) at ~420 and 422
35 mblf are very unusual in their composition, compared to all other drill core and surface
36 lithologies. Their siderophile and PGE concentrations are much higher than the respective
37 concentrations of investigated basaltic rocks at the surface. The probable enrichment with
38 metal oxides (TiO_2 , Al_2O_3 , Fe_2O_3 , MgO) and trace elements (Sc, V, Cr, Co, Ni, Cu, Zn), as
39 well as the PGE, during a hydrothermal alteration process seems plausible as indicated by
40 a high loss on ignition (LOI) and the strongly altered state of these blocks.
41
- 42 4. The concentrations of PGE in the reworked suevite are much higher compared to all other
43 impactites. These elevated PGE contents are most likely the result of an admixture of a
44 meteoritic component, probably of chondritic composition – in good agreement with the
45 previous work of Goderis et al. (2013) and Gurov and Koeberl (2004).
46
- 47 5. Nevertheless, the reworked suevite contains also a higher proportion of a mafic
48 component, as indicated by the REE content, in comparison to the suevite. The
49 composition of this mafic component and its PGE content cannot clearly be determined
50 because of the possible contribution of the chemically unusual mafic blocks to the element
51 budget. Therefore, it is not possible at this stage to unambiguously determine the nature of
52 the meteoritic projectile from the new results of this study either.
53
54
55
56
57

58 **Acknowledgments**– Funding for the El'gygytgyn drilling project was provided by the
59 International Continental Scientific Drilling Program (ICDP), the U.S. National Science Foundation
60 (NSF), the German Federal Ministry of Education and Research (BMBF), Alfred Wegener Institute
(AWI), Deutsches GeoForschungsZentrum Potsdam (GFZ), the Russian Academy of Sciences Far
East Branch (RAS FEB), the Russian Foundation for Basic Research (RFBR), the Arctic and Antarctic

1
2
3 Research Institute (AARI) in St. Petersburg, and the Austrian Federal Ministry of Science and
4 Research (BMWF). The Russian GLAD 800 drilling system was developed and operated by DOSECC
5 Inc., and the downhole logging was performed by the ICDP-OSG. The work in Vienna was supported
6 by the Austrian Science Foundation (FWF), project P21821-N19. This work is supported by Deutsche
7 Forschungsgemeinschaft (DFG) grant RE 528/10-2 to WUR and RTS. Special thanks go to the
8 reviewers B. Simonson and S. Goderis for their helpful comments and suggestions to improve the
9 quality of the original manuscript.
10
11

12 13 14 15 REFERENCES

16
17 Adolph L. and Deutsch A. 2010. Trace element analysis of impact glass spherules of the El'gygytyn
18 crater, Siberia (abstract #2421). *41th Lunar and Planetary Science Conference*. CD-ROM.

19
20 Barnes S.-J., Naldrett A. J., and Gorton M. P. (1985) The origin of platinum-group elements in
21 terrestrial magmas. *Chemical Geology* 53:303-323.

22
23 Belyi V. F., 1994. *The Geology of Okhotsk-Chukchi Volcanogenic Belt*. North East Interdisciplinary
24 Research Institute Magadan, Russian Academy of Sciences Far East Branch. 76 p. In Russian.

25
26 Belyi V. F. and Belaya B. V., 1998. *The Late Stage of the Okhotsk - Chukchi Volcanic Belt*
27 *Development (the Enmyvaam River Upperrun Area*. North East Interdisciplinary Research Institute,
28 Magadan, Russian Academy of Sciences Far East Branch. 108 p. In Russian.

29
30 Crocket J. H. (2002) Platinum-group element geochemistry of mafic and ultramafic rocks. In: *The*
31 *Geology, Geochemistry, Mineralogy and mineral beneficiation of Platinum-group elements*. Special
32 Volume 54 (ed. L. J. Cabri). Canadian Institute of Mining, Metallurgy and Petroleum, pp. 177-210.

33
34 Foriel J., Moynier F., Schulz T., and Koeberl C. 2013. Chromium isotope anomaly in an El'gygytyn
35 crater impactite: Evidence for a ureilite projectile. *Meteoritics & Planetary Science* 48:1339-1350.

36
37 Glushkova O. Y., and Smirnov V. N. 2007. Pliocene to Holocene geomorphic evolution and
38 paleogeography of the El'gygytyn Lake region, NE Russia. *Journal of Paleolimnology* 37:37-47.

39
40 Goderis S., Wittmann A., Zaiss J., Elburg M., Ravizza G., Vanhaecke F., Deutsch A., and Claeys P.
41 2013. Testing the ureilite projectile hypothesis for the El'gygytyn impact: Determination of
42 siderophile element abundances and Os isotope ratios in ICDP drill core samples and melt rocks.
43 *Meteoritics & Planetary Science* 48:1296-1324.

44
45 Govindaraju K. 1989. Compilation of working values and sample description for 272 geostandards.
46 *Geostandards Newsletter* 13:1-113.

47
48 Govindaraju K. 1994. 1994 compilation of working values and samples description for 383
49 geostandards. *Geostandards Newsletter* 18:1-158.

50
51 Gurov E. P., Valter A. A., Gurova E. P., and Serebrennikov A. I. 1978. Meteorite impact crater
52 El'gygytyn in Chukotka. *Doklady Akademii Nauk USSR* 240:1407-1410. In Russian.

53
54 Gurov E. P., Valter A., A. Gurova E. P., and Kotlovskaya F. I. 1979. El'gygytyn impact crater,
55 Chukotka: Shock metamorphism of volcanic rocks (abstract). *Lunar and Planetary Science*
56 *Conference* 10:479-481.
57
58
59
60

1
2
3 Gurov E. P., and Koeberl C. 2004. Shocked rocks and impact glasses from the El'gygytyn impact
4 structure (Russia). *Meteoritics and Planetary Science* 39:1495–1508.

6 Gurov E. P., Koeberl C., Reimold W. U., Brandstätter F., and Amare K. 2005. Shock metamorphism
7 of siliceous volcanic rocks of the El'gygytyn impact crater (Chukotka, Russia). In: Large Meteorite
8 Impacts III, edited by Kenkmann T., Hörz F. and Deutsch A. *Geological Society of America Special
9 Paper* 384:391-412.

12 Gurov E. P., Koeberl C., and Yamnichenkov A. 2007. El'gygytyn impact crater, Russia: Structure,
13 tectonics, and morphology. *Meteoritics and Planetary Science* 42:307–319

16 Gurov E. P., and Gurova E. P. 1983. Regularities of fault spreading around meteorite craters (on
17 example of the El'gygytyn crater). *Doklady Akademii Nauk USSR*. 275:958–961. In Russian.

19 Huber H., Koeberl C., McDonald I., and Reimold W.U. 2001. Geochemistry and petrology of
20 Witwatersrand and Dwyka diamictites from South Africa: search for an extraterrestrial component.
21 *Geochimica et Cosmochimica Acta* 65:2007-2016.

24 Jarosewich E., Clarke R.S., and Barrows J.N. 1987. The Allende Meteorite Reference Sample.
25 Smithsonian Contributions to the Earth Sciences 27: 49 pp.

27 Jenniskens P., Shaddad M.H., Numan D., Elsir S., Kudoda A.M., Zolensky M.E., Le L., Robinson
28 G.A., Friedrich J.M., Rumble D., Steele A., Chesley S.R., Fitzsimmons A., Duddy S., Hsieh H.H.,
29 Ramsay G., Brown P.G., Edwards W.N., Tagliaferri E., Boslough M.B., Spalding R.E., Dantowitz R.,
30 Kozubal M., Pravec P., Borovicka J., Charvat Z., Vaubaillon J., Kuiper J., Albers J., Bishop J.L.,
31 Mancinelli R.L., Sandford S.A., Milam S.N., Nuevo M., Worden S.P. 2009. The impact and recovery
32 of asteroid 2008 TC(3). *Nature* 458(7237):485-488.

35 Koeberl C. 1993. Instrumental neutron activation analysis of geochemical and cosmochemical
36 samples: A fast and proven method for small sample analysis. *Journal of Radioanalytical and Nuclear
37 Chemistry* I68:47–50.

40 Koeberl C. 2014. The Geochemistry and Cosmochemistry of Impacts. In: Holland H.D. and Turekian
41 K.K. (eds.) *Treatise on Geochemistry, Second Edition, vol. 2 (Planets, Asteroids, Comets and The
42 Solar System)*, pp. 73-118. Oxford. Elsevier.

44 Koeberl C., Peucker-Ehrenbrink B., Reimold W. U., Shukolyukov A., and Lugmair G.W. 2002. A
45 comparison of the osmium and chromium isotopic methods for the detection of meteoritic components
46 in impactites: Examples from the Morokweng and Vredefort impact structures, South Africa. In:
47 Catastrophic Events and Mass Extinctions: Impacts and Beyond (eds. C. Koeberl and K.G. MacLeod).
48 *Geological Society of America Special Paper* 356:607-617.

51 Koeberl C., Brandstätter F., Glass B. P., Hecht L., Mader D., and Reimold W. U. 2007. Uppermost
52 impact fallback layer in the Bosumtwi crater (Ghana): Mineralogy, geochemistry, and comparison
53 with Ivory Coast tektites. *Meteoritics and Planetary Science* 42:709-729.

56 Koeberl C., Milkereit B., Overpeck J. T., Scholz C. A., Amoako P. Y. O., Boamah D., Danuor S.,
57 Karp T., Kueck J., Hecky R. E., King J. W., and Peck J. A. 2007b. An international and
58 multidisciplinary drilling project into a young complex impact structure: The 2004 ICDP Bosumtwi
59 Crater Drilling Project-An overview. *Meteoritics & Planetary Science* 42:483-511.

- 1
2
3 Koeberl, C., Claeys, P., Hecht, L., and McDonald, I. 2012. Geochemistry of impactites. *Elements* 8:
4 37-42.
5
6 Koeberl C., Pittarello L., Reimold W. U., Raschke U., Brigham-Grette J., Melles M., and Minyuk P.
7 2013. El'gygytgyn impact crater, Chukotka, Arctic Russia: Impact cratering aspects of the 2009 ICDP
8 drilling project. *Meteoritics and Planetary Science* 48:1108-1129.
9
10 Layer P. W. 2000. ⁴⁰Argon/³⁹Argon-age of the El'gygytgyn event, Chukotka, Russia. *Meteoritics and*
11 *Planetary Science* 35:591–599.
12
13 Lodders K. 2003. Solar system abundances and condensation temperatures of elements. *The*
14 *Astrophysical Journal* 591:1220-1247.
15
16 Mader D. and Koeberl C. 2009. Using Instrumental Neutron Activation Analysis for geochemical
17 analyses of terrestrial impact structures: Current analytical procedures at the University of Vienna
18 Geochemistry Activation Analysis Laboratory. *Applied Radiation and Isotopes* 67:2100–2103.
19
20 McDonald I. 1998. The need for a common framework for collection and interpretation of data in
21 platinum-group element geochemistry. *Geostandards Newsletter* 22:85-91.
22
23 McDonald I. and Viljoen K. S. 2006. Platinum-group element geochemistry of mantle eclogites: A
24 reconnaissance study of xenoliths from the Orapa kimberlite, Botswana. *Transactions of the Institution*
25 *of Mining and Metallurgy, Section B* 115:81-93.
26
27 Melles M., Brigham-Grette J., Minyuk P., Koeberl C., Andreev A., Cook T., Fedorov G., Gebhardt C.,
28 Haltia-Hovi E., Kukkonen M., Nowaczyk N. R., Schwamborn G., Wennrich V. and the El'gygytgyn
29 Scientific Party 2011. The Lake El'gygytgyn Scientific Drilling Project – Conquering Arctic
30 Challenges through Continental. *Scientific Drilling* 11:29-40.
31
32 Pittarello L. and Koeberl C. 2013a. Clast size distribution (CSD) and other geometrical features in
33 shocked and unshocked rocks from the El'gygytgyn impact structure. *Meteoritics and Planetary*
34 *Science* 48:1325-1338.
35
36 Pittarello L. and Koeberl C. 2013b. Petrography of impact glasses and melt breccias from the
37 El'gygytgyn impact structure, Russia. *Meteoritics and Planetary Science* 48:1236-1250.
38
39 Pittarello L., Schulz T., Koeberl C., Hoffmann J. E., and Münker C. 2013. Petrography, geochemistry
40 and Hf-Nd isotope evolution of drill core samples and target rocks from the El'gygytgyn impact crater,
41 NE Chukotka, Arctic Russia. *Meteoritics and Planetary Science* 48:1160-1198.
42
43 Raschke U., Reimold W. U., Zaag P. T., Pittarello L., and Koeberl C. 2013a. Lithostratigraphy of the
44 impactite and bedrock section in ICDP drill core D1c from the El'gygytgyn impact crater, Russia.
45 *Meteoritics and Planetary Science* 48:1143-1159.
46
47 Raschke U., Schmitt R. T., and Reimold W. U. 2013b. Petrography and geochemistry of impactites
48 and volcanic bedrock in the ICDP drill core D1c from lake El'gygytgyn, NE Russia. *Meteoritics and*
49 *Planetary Science* 48:1251-1286.
50
51 Raschke U., Zaag P. T., Schmitt R. T., and Reimold W. U. 2014. The 2011 expedition to the
52 El'gygytgyn impact structure, Northeast Russia: Towards a new geological map for the crater area.
53 *Meteoritics and Planetary Science*, 49:978-1006.
54
55
56
57
58
59
60

1
2
3 Smirnov V. N., Savva N. E., and Glushkova O. Yu. 2011. New data on spherules from the region of
4 the El'gygytyn crater. *Geochemistry International* 49:314–318.

5
6 Tagle R. and Berlin J. 2008. A database of chondrite analyses including platinum group elements, Ni,
7 Co., Au, and Cr: Implications for the identification of chondritic projectiles. *Meteoritics and Planetary*
8 *Science* 43:541–559.

9
10 Taylor S. R. and McLennan S. M. 1985. *The Continental Crust: Its Composition and Evolution*.
11 Oxford: Blackwell Science Publishers. 312 p.

12
13 Tikhomirov P. L., Kalinina E. A., Kobayashi K. and Nakamura E. 2008. Late Mesozoic silicic
14 magmatism of the North Chukotka area (NE Russia): Age, magma sources, and geodynamic
15 implications. *Lithos* 105:329-346.

16
17 Tredoux M., Lindsay N. M., Davies G., and McDonald L. 1995. The fractionation of platinum-group
18 elements in magmatic systems with the suggestion of a novel causal mechanism. *South African*
19 *Journal of Geology* 98:157-167.

20
21 Tredoux M. and McDonald I. 1996. Komatiite WITS-1: a low concentration noble metal standard for
22 the analysis of non-mineralized samples. *Geostandards Newsletter* 20:267-276.

23
24 Val'ter A. A., Barchuk I. F., Bulkin V. S., Ogorodnik A. F., and Kotishevskaya E. Y. 1982. The
25 El'gygytyn meteorite—Probable composition. *Soviet Astronomy Letters* 8:59-62.

26
27 Warren P. H., Ulf-Møller F., Huber H., and Kallemayn G. W. 2006. Siderophile geochemistry of
28 ureilites: A record of early stages of planetesimal core formation. *Geochimica et Cosmochimica Acta*
29 *70*:2104-2126.

30
31 Winchester, J. A. and Floyd, P. A. 1977. Geochemical discrimination of different magma series and
32 their differentiation products using immobile elements. *Chemical Geology* 20:325-343.

33
34 Wittmann A., Goderis S., Claeys P., Vanhaecke F., Deutsch A., and Adolph F. 2013. Petrology of
35 impactites from El'gygytyn crater: Breccias in the ICDP-drill core 1C, glassy impact melt rocks and
36 spherules. *Meteoritics and Planetary Science* 48:1199-1235.

37 38 39 40 41 42 **Figure captions**

43
44 **Fig. 1:** Geological map of the El'gygytyn impact crater with drill core location. Small inset
45 indicating the geographic location of this impact structure in NE Siberia (Raschke et al. 2014)

46
47 **Fig. 2:** Simplified NW-SE cross-section through the El'gygytyn impact structure, showing the drill
48 core location and drilled lithologies. For more detail see Raschke et al. (2013a) and Koeberl et al.
49 (2013). Based on a diagram by Melles et al. 2011.

50
51 **Fig. 3:** Stratigraphic column of the ICDP drill core (modified after Raschke et al. 2013a). The
52 stratigraphic positions of samples used for INAA and PGE analyses are indicated, as well as those of
53 samples analyzed by INAA from Pittarello et al. (2013) used in this work.

54
55 **Fig. 4:** Zr/TiO₂ vs. Nb/Y diagram for classification of volcanic rocks after Winchester and Floyd
56 (1977). Note: The suevitic units (incl. reworked suevite) plot in the same field as the upper and lower
57 bedrock of the drill core as well as the rhyolitic and rhyodacitic ignimbrites from the crater rim. These
58 lithologies are illustrated by differently shaded fields that each include a larger number of data. The
59
60

1
2
3 symbols for mafic units represent a single analysis per sample. Data from Raschke et al. (2013b,
4 2014).

5
6
7 **Fig. 5:** CI chondrite normalized REE plots (normalization values from Taylor and McLennan 1985)
8 for samples of the ICDP drill core: **(a)** upper and lower bedrock; **(b)** three mafic blocks at depths of
9 391, 420, and 422 mblf; **(c)** reworked suevite, suevite, and polymict impact breccia dike. **(d)** Yb vs.
10 Gd-diagram displaying the distinctly increased concentrations of Gd and Yb in the mafic blocks at 391
11 and 422 mblf, and the admixture of such a mafic component to the reworked suevite. Note that surface
12 volcanic target lithologies and impact melt breccia are not plotted in this figure.

13
14
15 **Fig. 6:** **(a)** Os vs. Ir, **(b)** Rh vs. Ir, and **(c)** Ru vs. Ir abundance plots. Note the high concentrations of
16 these elements in the mafic block at 420 mblf and the reworked suevite.

17
18 **Fig. 7:** CI-normalized PGE plots (normalization values from Lodders 2003) of **(a)** surface volcanic
19 rocks including rhyolitic ignimbrite, rhyodacitic ignimbrite, andesite, andesitic-dacitic tuff, basalt,
20 and basaltic-andesitic tuff, **(b)** the three mafic blocks in the ICDP drill core at 391, 420, and 422 mblf
21 depths, and **(c)** reworked suevite, suevite, impact melt breccia, and polymict impact breccia dike. Note
22 the significantly higher concentrations of Os, Ir, Ru, and Rh in the reworked suevite.
23
24
25
26
27
28
29
30
31
32
33
34
35
36
37
38
39
40
41
42
43
44
45
46
47
48
49
50
51
52
53
54
55
56
57
58
59
60

Abstract

The complex impact structure El'gygytgyn in northeastern Russia (age 3.6 Ma, diameter 18 km) was formed in ~88 Ma old volcanic target rocks of the Ochotsk-Chukotsky Volcanic Belt (OCVB). In 2009, El'gygytgyn was the target of a drilling project of the International Continental Scientific Drilling Program (ICDP), and in summer 2011 it was investigated further by a Russian-German expedition. Drill core material and surface samples, including volcanic target rocks and impactites, have been investigated by various geochemical techniques in order to improve the record of trace element characteristics for these lithologies and to attempt to detect and constrain a possible meteoritic component. The bedrock units of the ICDP drill core reflect the felsic volcanics that are predominant in the crater vicinity. The overlying suevites comprise a mixture of all currently known target lithologies, dominated by felsic rocks but lacking a discernable meteoritic component based on platinum group element (PGE) abundances. The reworked suevite, directly overlain by lake sediments, is not only comparatively enriched in shocked minerals and impact glass spherules, but also contains the highest concentrations of Os, Ir, Ru, and Rh compared to other El'gygytgyn impactites. This is - to a lesser extent - the result of admixture of a mafic component, but more likely the signature of a chondritic meteoritic component. However, the highly siderophile element contribution from target material akin to the mafic blocks of the ICDP drill core to the impactites remains poorly constrained.

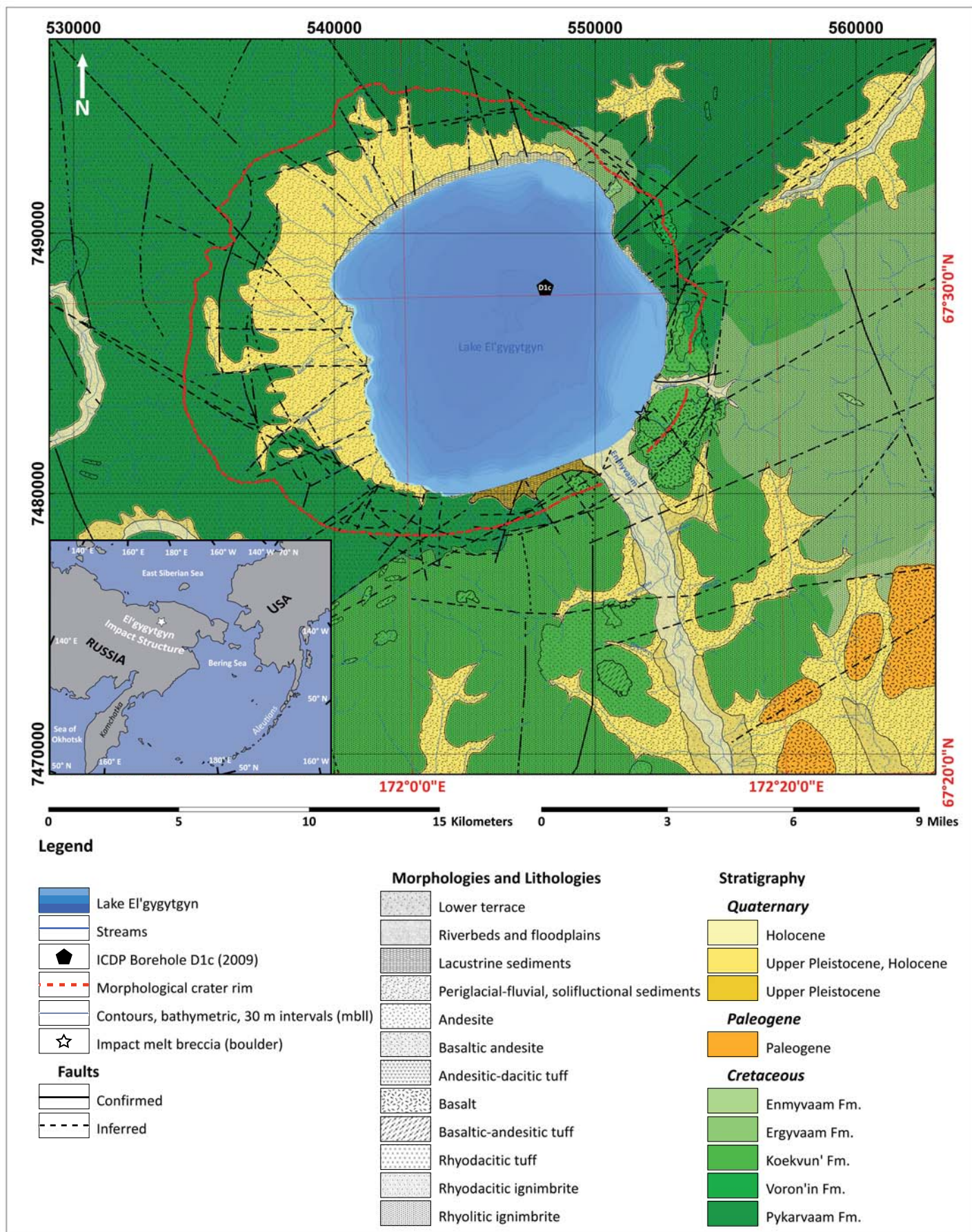


Fig. 1

1
2
3
4
5
6
7
8
9
10
11
12
13
14
15
16
17
18
19
20
21
22
23
24
25
26
27
28
29
30
31
32
33
34
35
36
37
38
39
40
41
42
43
44
45
46
47

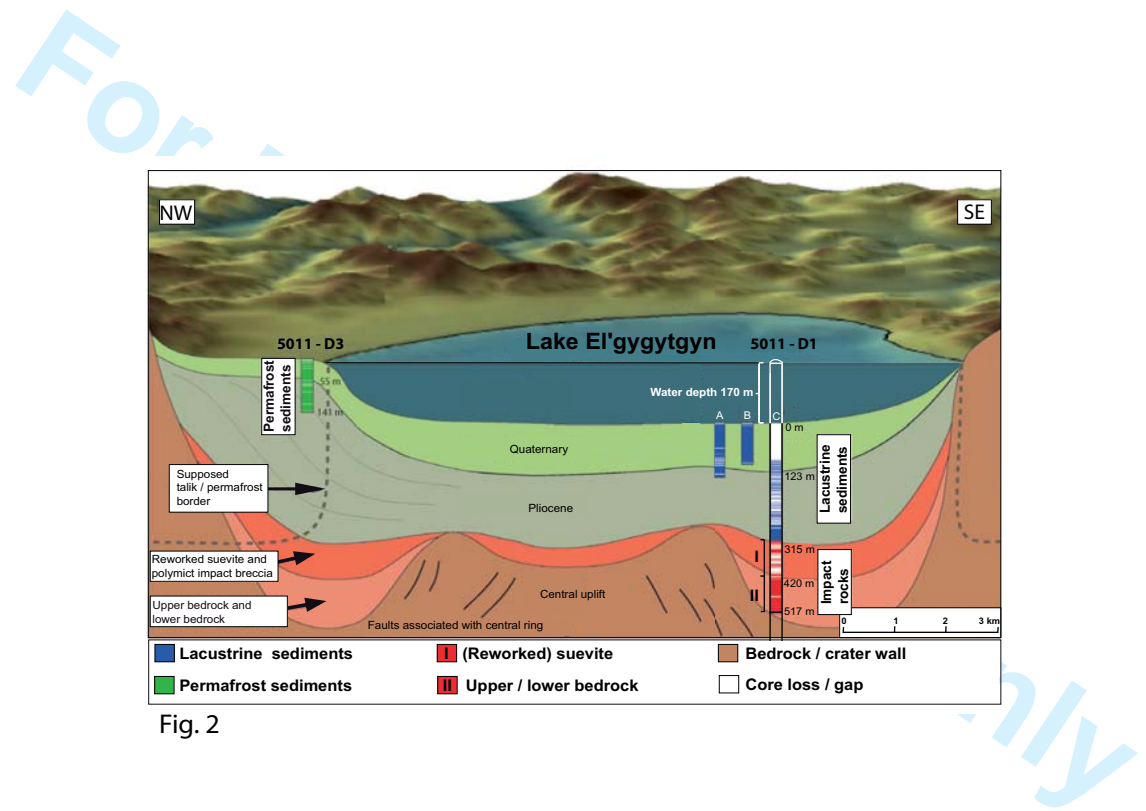


Fig. 2

Running Head

1
2
3
4
5
6
7
8
9
10
11
12
13
14
15
16
17
18
19
20
21
22
23
24
25
26
27
28
29
30
31
32
33
34
35
36
37
38
39
40
41
42
43
44
45
46
47
48
49
50
51
52
53
54
55
56
57
58
59
60

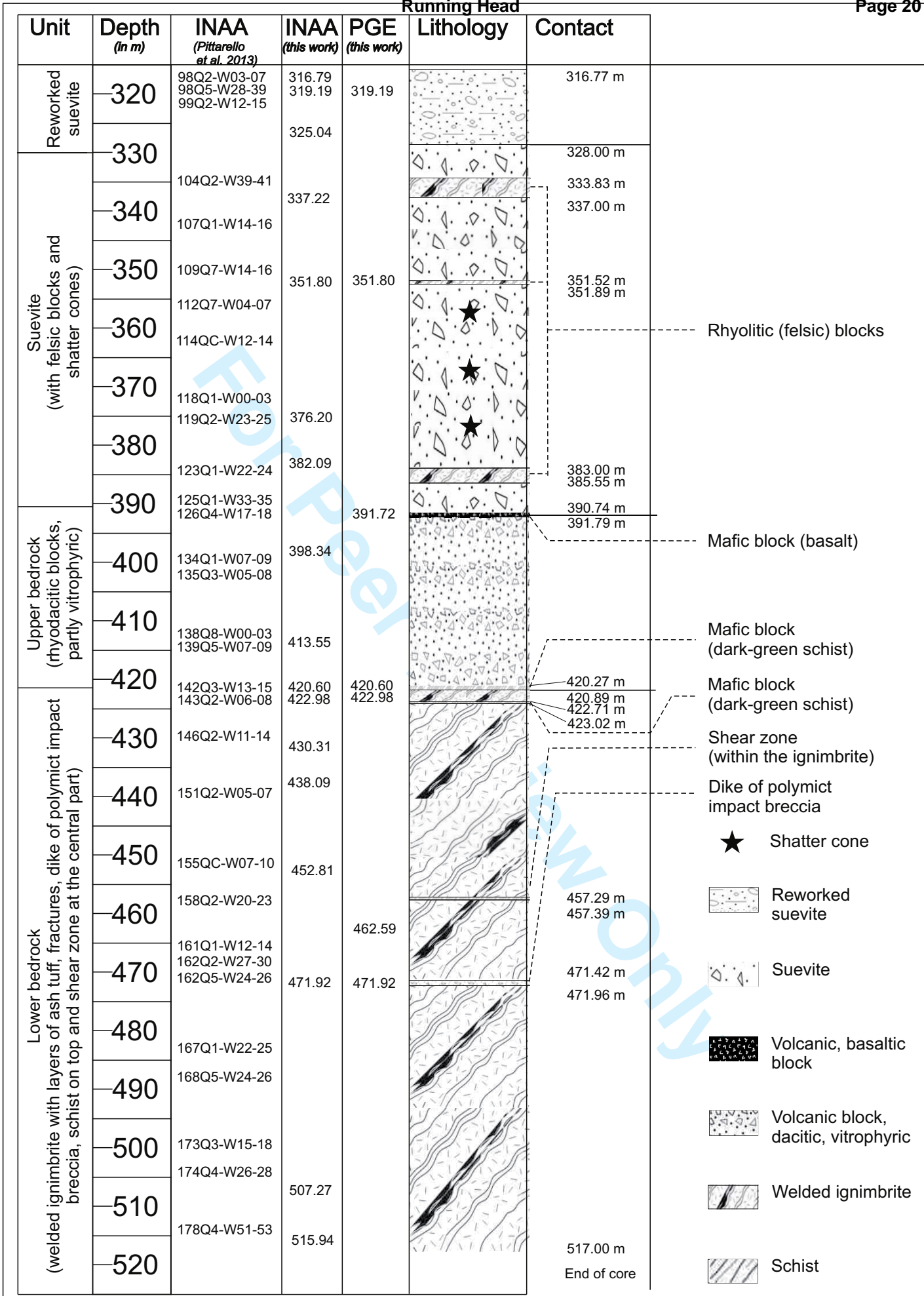


Fig. 3

1
2
3
4
5
6
7
8
9
10
11
12
13
14
15
16
17
18
19
20
21
22
23
24
25
26
27
28
29
30
31
32
33
34
35
36
37
38
39
40
41
42
43
44
45
46
47
48
49
50
51
52
53
54
55
56
57
58
59
60

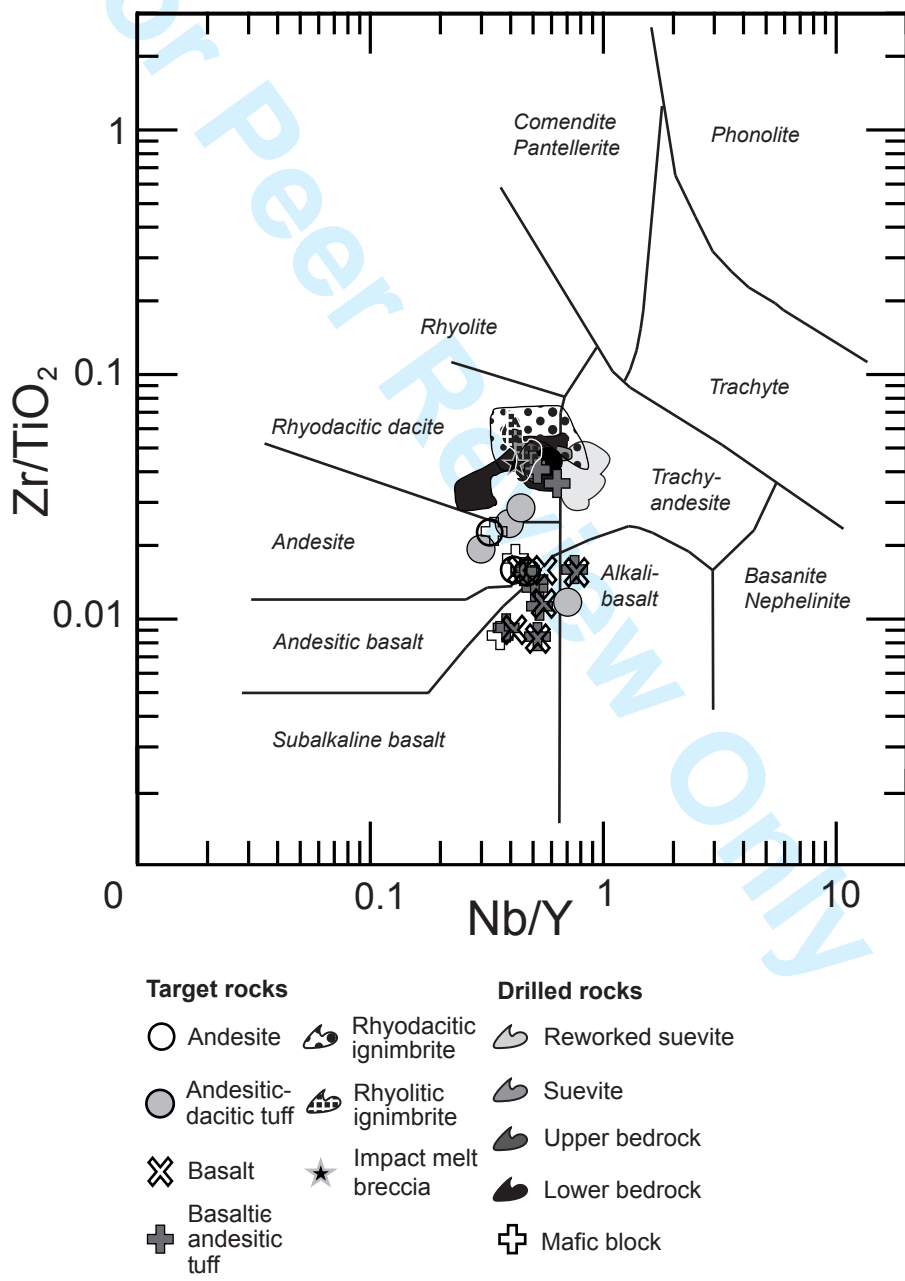


Fig. 4

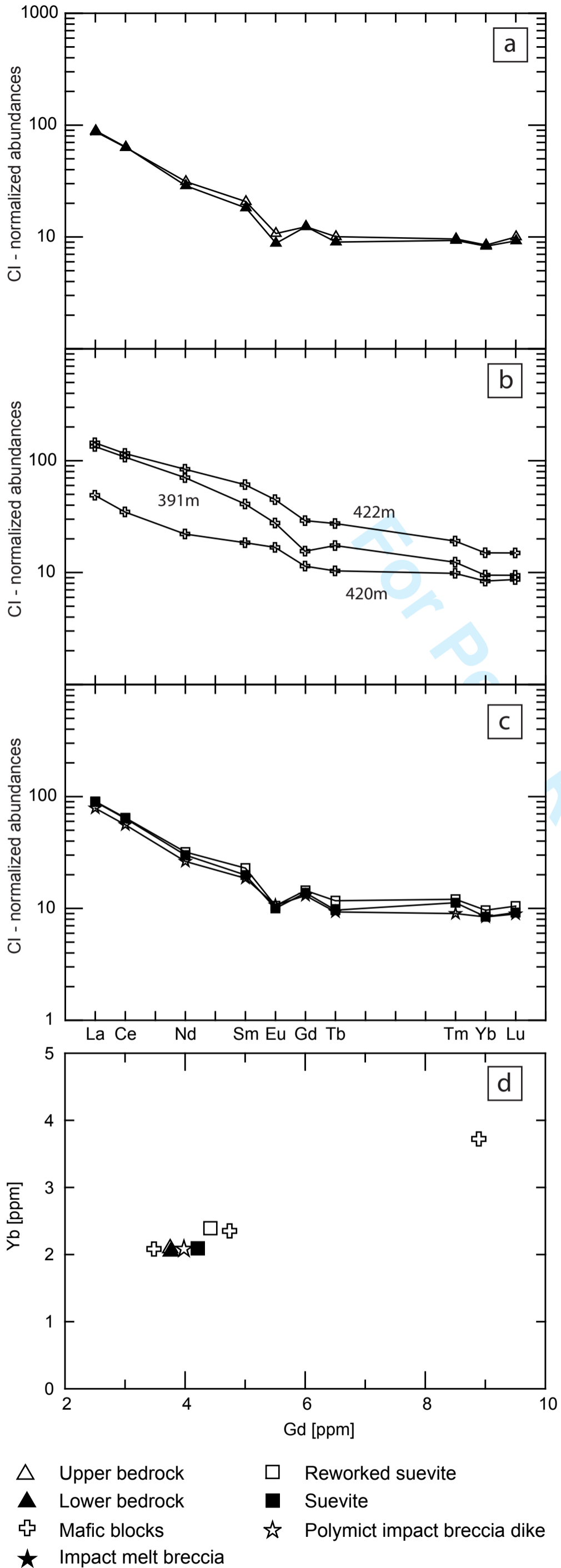
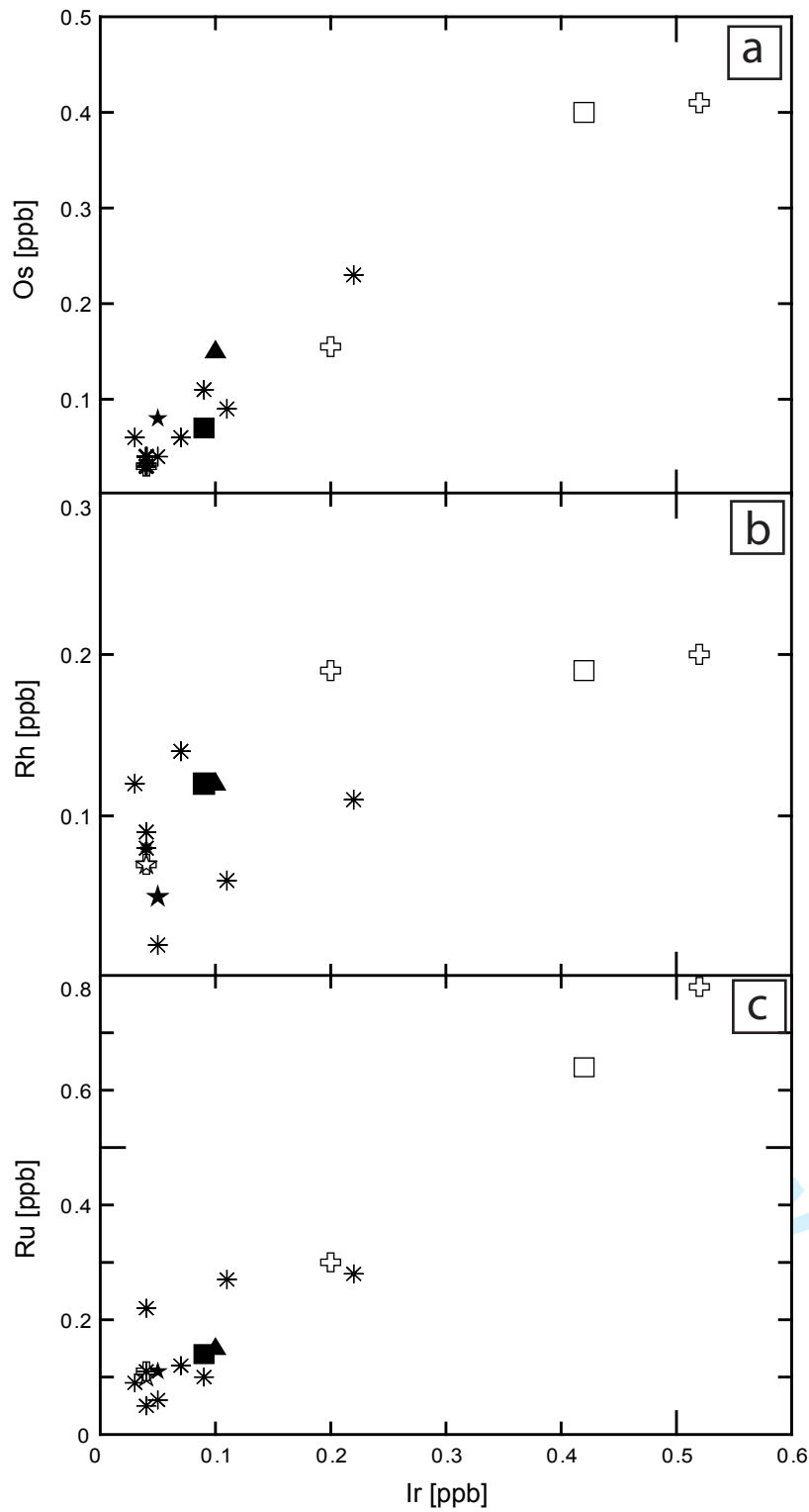


Fig. 5



- △ Upper bedrock
- ▲ Lower bedrock
- ⊕ Mafic blocks
- ★ Impact melt breccia
- Reworked suevite
- Suevite
- ☆ Polymict impact breccia dike
- * Surface volcanic target rocks

Fig. 6

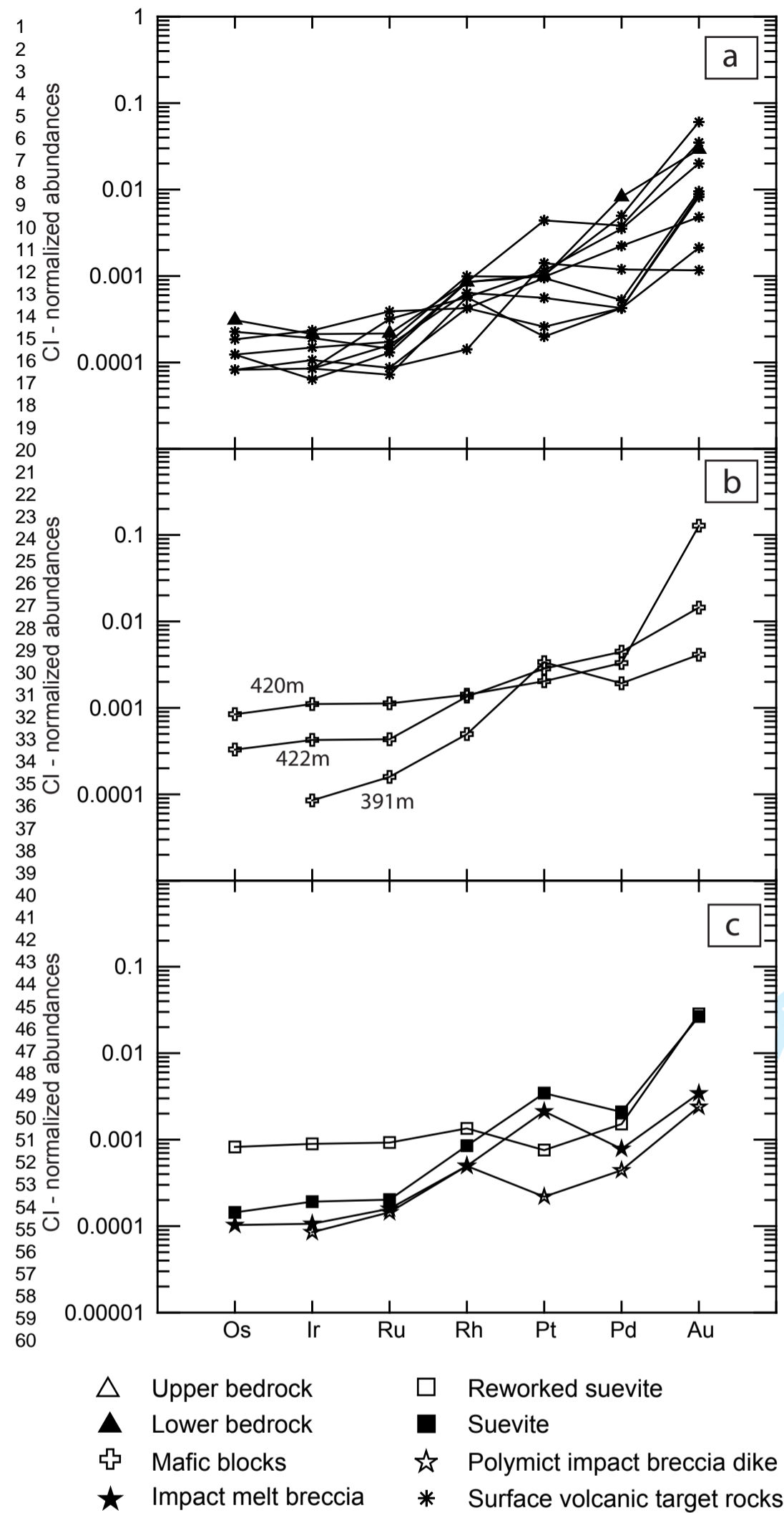


Fig. 7

Table 1. List of ICDP drill core samples for analytical studies.

Sample	ID	Lithology*
<i>(this work)</i>	<i>(by Pittarello et al. 2013)</i>	
UR-ELG_316.79	98Q2-W03-07 (316.80)	rsv
UR-ELG_319.19	98Q5-W28-31 (318.20)	rsv
UR-ELG_325.04	99Q2-W12-15 (319.50)	rsv
UR-ELG_337.22		rsv
UR-ELG_351.80	104Q2-W39-41 (334.70)	sv
UR-ELG_376.20	107Q1-W14-16 (342.70)	sv
UR-ELG_382.09	109Q7-W14-16 (351.40)	sv
	112Q7-W04-07 (355.40)	sv
	114QCC-W02-05 (361.70)	sv
	118Q1-W00-03 (371.30)	sv
	119Q2-W23-25 (374.90)	sv
	123Q1-W22-24 (383.90)	sv
UR-ELG_398.34	125Q1-W33-35 (390.20)	ub
UR-ELG_413.55	134Q1-W07-09 (399.60)	ub
	135Q3-W05-08 (401.80)	ub
	138Q8-W00-03 (412.20)	ub
	139Q5-W07-09 (414.50)	ub
UR-ELG_391.72	126Q4-W17-18 (391.70)	mb
UR-ELG_420.60	142Q3-W13-15 (420.90)	mb
UR-ELG_422.98	143Q2-W06-08 (422.90)	mb
UR-ELG_422.98	146Q2-W11-14 (429.70)	lb
UR-ELG_430.31	151Q2-W05-07 (440.40)	lb
UR-ELG_438.01	155QCC-W07-10(451.40)	lb
UR-ELG_452.81	158Q2-W20-23 (456.90)	lb
UR-ELG_462.59	161Q1-W12-14 (465.10)	lb
	162Q2-W27-30 (468.30)	lb
	162Q5-W24-26 (470.20)	lb
	167Q1-W22-25 (483.10)	lb
	168Q5- W24-26 (487.40)	lb
UR-ELG_507.27	173Q3-W15-18 (500.00)	lb
UR-ELG_515.94	174Q4-W26-28 (503.90)	lb
	178Q4W51-53 (514.30)	lb
UR-ELG_471.92		pibd

*Abbreviations: rsv = reworked suevite, sv = suevite, ub = upper bedrock; lb = lower bedrock; mb = mafic block; pibd = polymict impact breccia dike.

Table 2. Selected major and trace element abundances of samples from the ICDP drill core D1c of the El'gygytyn impact structure, as determined by INAA.

Sample	UR- ELG	UR- ELG	UR- ELG	UR- ELG	UR- ELG	UR- ELG	UR- ELG	UR- ELG	UR- ELG	UR- ELG	UR- ELG	UR- ELG	UR- ELG	UR- ELG	UR- ELG	UR- ELG	UR- ELG
Depth (mblf)	316.79	319.19	325.04	337.22	351.8	376.2	382.09	398.34	413.55	420.6	422.98	430.31	438.09	452.81	471.92	507.27	515.94
Lithology	rsv	rsv	rsv	sv	sv	sv	sv	ub	ub	mb	mb	lb	lb	lb	pibd	lb	lb
ppm																	
Na (wt%)	2.48	2.02	2.00	2.21	2.57	2.22	2.41	2.81	2.09	1.25	0.53	2.76	3.03	2.52	1.89	3.19	1.28
K (wt%)	3.09	3.09	2.88	3.00	3.26	3.17	3.03	2.99	2.50	<0.7	0.91	2.94	3.36	3.30	2.66	4.07	2.04
Fe (wt%)	2.41	2.87	2.26	1.51	1.90	1.78	1.91	1.99	2.01	7.97	7.16	1.75	1.70	1.62	2.42	1.72	2.27
Sc	9.17	11.1	7.95	4.48	6.09	5.38	6.07	6.45	6.14	36.9	42.0	5.46	5.59	4.22	9.04	4.65	2.1
Cr	43.2	92.7	26.1	10.7	11.8	19.8	12.4	18.7	13.6	544	872	13.7	10.2	10	58.6	10.5	11.6
Co	6.03	8.38	5.45	2.92	3.52	3.23	3.54	3.69	3.30	30.8	42.7	3.44	3.23	2.71	6.99	3.24	1.01
Ni	28	41	<26	<20	12	4	11	17	3	98	276	6	<24	<21	24	<22	<20
Zn	67	66	56	49	53	47	56	56	57	348	121	50	50	46	58	51	13
Ga	4.9	3.6	6.9	3.4	5.2	4.1	2.8	4.8	6.1	232	19	3.2	<3.7	5.5	5.3	4.5	3.1
As	15.4	16.2	11.4	4.42	9.57	5.96	7.87	8.96	5.72	42.2	91.5	34.9	19.1	8.91	18.3	14.3	47.8
Se	0.03	<1.7	<1.5	<1.3	<1.4	<1.3	<1.4	<1.4	<1.4	<2.8	7.72	<1.4	<1.5	<1.3	<1.6	<1.4	<1.3
Rb	146	118	115	131	132	138	128	136	107	9.51	75.6	115	129	137	121	174	119
Sr	171	318	153	122	283	122	161	196	165	401	172	178	518	217	235	169	76
Zr	241	241	172	193	218	208	222	234	217	166	678	220	229	187	237	221	127
Sb	1.57	1.55	0.98	0.92	1.39	1.17	1.46	1.69	1.54	1.01	4.09	1.55	1.61	0.97	1.84	1.15	2.37
Cs	9.84	7.64	9.22	8.67	4.55	10.3	8.26	6.64	6.85	3.54	30.1	3.83	4.17	4.94	9.18	5.21	5.76
Ba	464	596	299	431	469	417	439	495	481	65	260	481	569	493	419	592	235
La	34.3	28.3	35.7	30.8	30.1	31.5	31.4	30.4	27.8	19.2	51.1	31.7	34.3	41.6	28.9	37.3	24.6
Ce	64.8	56.1	64.7	57.6	56.7	58.9	60.8	58	53.3	34	111	58.8	64.5	72.9	53.5	68.5	39.8
Nd	25	22.6	22.9	19	20.5	20.4	21.9	20.8	19.2	15.6	59	18.7	21.5	22.2	18.7	22.1	13.1
Sm	5.79	5.08	5.03	3.79	4.3	4.06	4.25	4.2	4.13	4.39	13.7	4.05	4.68	4.43	4.3	4.8	3.05
Eu	0.98	0.89	0.84	0.75	0.88	0.77	0.9	0.93	0.77	1.70	3.84	0.81	0.86	0.76	0.93	0.80	0.44
Gd	5.03	4.54	3.86	3.8	4.26	3.25	3.54	3.67	2.99	4.66	10.2	3.11	4.55	4	3.99	4.67	2.61
Tb	0.78	0.71	0.63	0.46	0.55	0.49	0.56	0.52	0.54	0.7	1.59	0.50	0.59	0.5	0.54	0.57	0.48
Tm	0.4	0.34	0.3	0.27	0.3	0.27	0.27	0.29	0.3	0.34	0.61	0.31	0.33	0.26	0.32	0.34	0.38
Yb	2.54	2.35	2.27	1.78	2.04	1.94	2.10	1.98	1.98	2.17	3.65	1.98	2.13	2.01	2.08	2.13	3.19
Lu	0.43	0.41	0.38	0.3	0.34	0.34	0.35	0.33	0.34	0.32	0.56	0.33	0.34	0.33	0.34	0.38	0.57
Hf	5.28	4.71	4.14	4.17	4.59	4.48	4.57	4.74	4.89	2.04	12.3	4.55	4.46	4.11	4.5	4.88	2.17
Ta	0.89	0.67	0.64	0.69	0.65	0.70	0.64	0.66	0.66	0.33	1.1	0.68	0.79	0.73	0.65	0.93	0.63
Au (ppb)	13	<1.3	0.6	0.6	<1.3	<1.1	<1.5	<1.4	<1.2	<1.5	1.7	<1.5	<1.7	<1.3	<0.9	<1.6	<1.3
Th	15.8	11.9	11.8	15.1	13.3	15.1	13.9	13.3	13.1	1.51	5.32	14.9	16.5	17.8	12.7	19.4	12.6
U	4.69	3.33	3.38	3.55	3.69	3.40	3.16	3.49	3.25	0.36	5.98	3.77	5.15	4.00	3.66	5.04	4.81

Abbreviations: rsv = reworked suevite, sv = suevite, ub = upper bedrock; lb = lower bedrock; mb = mafic block; pibd = polymict impact breccia dike.

Table 3. Concentrations of platinum group elements and Au in impactites and target lithologies from the ICDP drill core D1c and surface outcrops.

Sample	Lithology	Os ppb	Ir ppb	Ru ppb	Rh ppb	Pt ppb	Pd ppb	Au ppb
ICDP drill core								
UR-ELG 319.19 mblf	reworked suevite	0.40	0.42	0.64	0.19	0.76	0.89	4.15
UR-ELG 351.8 mblf	suevite	0.07	0.09	0.14	0.12	3.46	1.23	3.87
UR-ELG 391.72 mblf	mafic block	<0.03	0.04	0.11	0.07	3.38	1.13	0.60
UR-ELG 420.6 mblf	mafic block	0.41	0.52	0.78	0.20	2.04	1.94	18.65
UR-ELG 422.8 mblf	mafic block	0.16	0.20	0.30	0.19	2.86	2.62	2.11
UR-ELG 462.59 mblf	lower bedrock	0.15	0.10	0.15	0.12	1.00	4.84	4.26
UR-ELG 471.92 mblf	polym. impact breccia dike	<0.03	0.04	0.10	0.07	0.22	0.26	0.35
Surface outcrops								
UR-2011_1.1	andesitic-dacitic tuff	<0.03	<0.03	0.06	0.06	0.26	0.25	1.20
UR-2011_3.7	basalt	0.06	0.07	0.12	0.14	0.98	1.31	0.70
UR-2011_4.4	basaltic andesite	0.09	0.11	0.27	0.06	0.95	0.31	1.39
UR-2011_4.5	rhyodacitic tuff	0.11	0.09	0.10	0.12	4.41	2.25	5.10
UR-2011_5.3	rhyolitic ignimbrite	0.04	0.05	0.06	0.02	1.41	0.70	0.17
UR-2011_7.2	andesite	0.04	0.04	0.11	0.09	0.56	0.25	1.29
UR-2011_9.2	basaltic-andesitic tuff	0.06	0.03	0.09	0.12	1.04	2.93	8.82
UR-2011_9.11b	impact melt	0.05	0.05	0.11	0.07	2.13	0.46	0.50
UR-2011_9.12a	rhyodacitic ignimbrite	<0.03	0.04	0.22	0.08	0.20	0.25	0.31
UR-2011_10.1a	rhyodacitic ignimbrite	<0.03	0.04	0.05	0.08	1.16	2.07	2.93

Table 4. Compilation of the average REE contents, their standard deviations, and the Eu/Eu* and La_N/Yb_N ratios of the ICDP El'gygytgyn drill core lithologies.*

	Reworked suevite avg.	Suevite avg.	Polymict impact breccia dike ~471 mblf	Upper bedrock avg.	Lower bedrock avg.	Mafic block ~391 mblf	Mafic block ~420 mblf avg.	Mafic block ~422 mblf avg.
ppm	n = 6	n = 12	n = 1	n = 7	n = 17	n = 1	n = 2	n = 2
La	33.2 ± 3.5	32.9 ± 2.3	28.9	31.8 ± 3.0	32.5 ± 4.1	49.2	18.0 ± 1.4	53.0 ± 1.9
Ce	61.8 ± 5.8	61.0 ± 3.9	53.5	60.0 ± 5.6	60.5 ± 7.7	103.0	33.3 ± 1.0	111 ± 1
Nd	22.7 ± 1.6	21.3 ± 1.9	18.7	22.1 ± 2.9	20.3 ± 2.2	50.2	15.7 ± 0.1	59.6 ± 0.8
Sm	5.30 ± 0.66	4.50 ± 0.56	4.30	4.76 ± 0.72	4.18 ± 0.47	9.47	4.27 ± 0.18	14.1 ± 0.6
Eu	0.92 ± 0.08	0.87 ± 0.13	0.93	0.93 ± 0.18	0.76 ± 0.11	2.41	1.46 ± 0.35	3.90 ± 0.09
Gd	4.43 ± 0.59	4.22 ± 0.66	3.99	3.76 ± 0.67	3.77 ± 0.66	4.75	3.49 ± 1.67	8.90 ± 1.90
Tb	0.68 ± 0.11	0.56 ± 0.08	0.54	0.58 ± 0.09	0.52 ± 0.05	1.01	0.60 ± 0.14	1.59 ± 0.00
Tm	0.43 ± 0.11	0.40 ± 0.10	0.32	0.34 ± 0.05	0.33 ± 0.03	0.44	0.35 ± 0.01	0.68 ± 0.01
Yb	2.39 ± 0.26	2.09 ± 0.22	2.08	2.09 ± 0.20	2.04 ± 0.33	2.35	2.08 ± 0.13	3.72 ± 0.10
Lu	0.40 ± 0.04	0.35 ± 0.03	0.34	0.38 ± 0.04	0.35 ± 0.06	0.36	0.33 ± 0.01	0.57 ± 0.01
Eu _{avg} /Eu _{avg} *	0.58	0.61	0.69	0.67	0.59	1.10	1.16	1.06
La _{avgN} /Yb _{avgN}	9.39	10.64	9.39	10.28	10.77	14.15	5.85	9.63

*Based on data of this work and from Pittarello et al. (2013); n = number of samples, normalization values from Taylor and McLennan, 1985.

Table 5. Compilation of the average Cr, Co, and Ni contents, their standard deviations, and their ratios for the ICDP El'gygytgyn drill core lithologies*; for comparison data for impact spherules from the El'gygytgyn crater are also reported**.

	Reworked suevite avg.	Suevite avg.	Polymict impact breccia dike ~471 mblf	Upper bedrock avg.	Lower bedrock avg.	Mafic block ~391 mblf	Mafic block ~420 mblf avg.	Mafic block ~422 mblf avg.	Impact spherules**
ppm	n = 6	n = 12	n = 1	n = 7	n = 17	n = 1	n = 2	n = 2	n = 13
Cr	34.8 ± 31.4	12.3 ± 13.8	58.6	13.2 ± 7.4	8.1 ± 4.0	95.1	499 ± 64	1061 ± 267	329 ± 267
Co	5.89 ± 1.48	3.97 ± 1.52	6.99	4.49 ± 1.62	3.10 ± 0.60	29.2	32.4 ± 2.2	54.6 ± 16.7	44.4 ± 26.2
Ni	24.9 ± 13.7	10.7 ± 5.2	24	11.8 ± 6.4	10.7 ± 3.4	76.9	100 ± 3	331 ± 77	564 ± 467
Ni _{avg} /Cr _{avg}	0.72	0.87	0.40	0.89	1.32	0.81	0.20	0.31	1.71
Ni _{avg} /Co _{avg}	4.22	2.70	3.43	2.63	3.45	2.63	3.09	6.06	12.70
Cr _{avg} /Co _{avg}	5.91	3.10	8.38	4.00	2.61	3.26	15.40	19.43	7.41

*Based on data of this work and from Pittarello et al. (2013); **based on data from Wittmann et al. (2013); these impact spherules originate from the reworked suevite of the ICDP El'gygytgyn drill core and from outside of the crater; n = number of samples.

1
2
3
4
5
6
7
8
9
10
11
12
13
14
15
16
17
18
19
20
21
22
23
24
25
26
27
28
29
30
31
32
33
34
35
36
37
38
39
40
41
42
43
44
45
46
47
48
49

Table 6. Calculated average PGE composition of the El'gygytgyn target in comparison to the reworked suevite and calculated impactites.*

PGE	Average El'gygytgyn target	Reworked suevite (319 mblf)	Average target + 0.12 % ureilite	Average target + 0.10 % LL chondrite	Average target + 0.07 % CI chondrite
ppb					
Os	0.04	0.40	0.40	0.40	0.39
Ir	0.05	0.42	0.34	0.39	0.38
Ru	0.07	0.64	0.51	0.59	0.57
Rh	0.03	0.19	n.a.**	0.14	0.12
Pt	1.38	0.76	1.87	2.09	2.05
Pd	0.72	0.89	0.79	1.22	1.11
Au	0.27	4.15	0.30	0.39	0.37

*Data based on the average El'gygytgyn target with an admixture of 0.12 % average ureilite, 0.10 % average LL chondrite and 0.07 % average CI chondrite, respectively. Data for ureilite (based on 24 samples) by Warren et al. (2006) and for LL and CI chondrites by Tagle and Berlin (2008). **n.a. = not available.

For Peer Review Only