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**Oxygen isotopes and volatile contents of the Gorgona komatiites, Colombia: a confirmation
of the deep mantle origin of H₂O**

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Abstract

We report O isotopes in olivine grains (Fo_{89–93}) and volatile contents (CO₂, H₂O, F, S, Cl) in olivine-hosted melt inclusions from one Gorgona picrite and five komatiites with the aim of constraining the origin of H₂O in these magmas. These samples have previously been analysed for major and trace elements and volatile concentrations (H₂O, S, Cl) and B isotopes in melt inclusions. A distinctive feature of the included melts is relatively high contents of volatile components and boron, which show positive anomalies in, otherwise depleted, primitive mantle normalised trace and rare earth element patterns and range in $\delta^{11}\text{B}$ from -11.5 to 15.6‰ . In this study, the olivines were systematically analysed for O isotopes (1) in the centre of grains, (2) near the grain boundaries and, (3) as close as possible to the studied melt inclusions. The majority of olivines ($\sim 66\%$) are “mantle”-like, $4.8\text{‰} \leq \delta^{18}\text{O} \leq 5.5\text{‰}$, with a subordinate but still significant number ($\sim 33\%$) above, and only 2 grains below, this range. There is no systematic difference between the central and marginal parts of the grains. Higher than “mantle” $\delta^{18}\text{O}_{\text{Ol}}$ values are ascribed to low- T ($<300\text{ °C}$) serpentinisation along inner fractures and grain boundaries of olivine phenocrysts. The measured concentrations of volatile components in the melt inclusions corrected for the effects of post-entrapment crystallisation and H₂O-CO₂ exsolution in inclusion shrinkage bubbles are: 286–1748 $\mu\text{g/g}$ CO₂, 0.2–0.86 wt.% H₂O, 48–82 $\mu\text{g/g}$ F, 398–699 $\mu\text{g/g}$ S and 132–198 $\mu\text{g/g}$ Cl. They correspond to a pressure of $86 \pm 44\text{ MPa}$ or $\sim 2.5\text{-km}$ crustal depth of olivine crystallisation. The correlations of S and, to a lesser extent, of H₂O, with highly incompatible lithophile elements and the correlation of F with Cl, but no relationships of H₂O with Cl, rule out shallow depth magma degassing and/or crustal contamination. Our new $\delta^{18}\text{O}$ olivine and volatile component data combined with the existing, highly variable $\delta^{11}\text{B}$ values for melt inclusions also support the deep mantle origin of H₂O (and probably other volatiles) in the Gorgona mafic and ultramafic magmas.

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53 **Keywords**

54 Gorgona Island, Komatiites, Olivine, Oxygen isotopes, Volatile components, Ion microprobe

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1. Introduction

Komatiites are olivine spinifex-textured, ultramafic, mantle-derived rocks formed by high degrees of partial melting and typically having high magnesium (>18 wt.% MgO), with low silicon (40–45 wt.% SiO₂), titanium (<1 wt.% TiO₂) and incompatible trace element concentrations (e.g., [Arndt and Nisbet, 1982](#); [Kerr and Arndt, 2001](#) and references therein). Komatiites were first recognized in the late 1960s in the Barberton greenstone belt, South Africa, and were named from their type locality along the Komati River ([Viljoen and Viljoen, 1969a,b](#)). There has long been a considerable debate as to whether komatiites are derived from dry and hot mantle, and are the products of partial melting of ascending super-hot Archaean plumes (e.g., [Bickle et al., 1977](#); [Campbell et al., 1989](#); [McDonough and Ireland, 1993](#); [Arndt et al., 1997, 1998](#); [Herzberg et al., 1995, 2007](#); [Berry et al., 2008](#) among others) or are formed by hydrous melting of the mantle at much lower temperatures during subduction (e.g., [Allègre, 1982](#); [Parman et al., 1997](#); [Grove and Parman, 2004](#); [Parman and Grove, 2005](#)). Because most komatiites have undergone significant sub-solidus alteration, primary volatile abundances of the magmas from which they crystallised, remain controversial.

In this context, the Late Cretaceous (~90 Ma) ultramafic and mafic lavas of the Gorgona Island, Colombia, first identified as komatiites by [Gansser et al. \(1979\)](#) and then studied in more detail by e.g., [Echeverría \(1980\)](#), [Dietrich et al. \(1981\)](#), [Aitken and Echeverría \(1984\)](#), [Echeverría and Aitken \(1986\)](#), [Kyser et al. \(1987\)](#), [Kerr et al. \(1996\)](#), [Arndt et al. \(1997\)](#), [Kerr \(2005\)](#), [Révillon et al. \(2000, 2002\)](#), [Serrano et al. \(2011\)](#), among others, represent suitable candidates for the study of volatile components because they are generally much fresher compared to most of their Archaean counterparts. In addition, the komatiites from Gorgona remain the only known non-Archaean komatiites and, therefore, studying them in detail may have important implications for our understanding the effects of water on mantle potential temperatures.

Although, overall, olivine-hosted melt inclusions in Gorgona komatiites are depleted in incompatible trace elements, they are distinctive in that they possess relatively high concentrations of volatile components and boron (Kamenetsky et al., 2010). These high concentrations result in positive anomalies (10 to 50 fold enrichment) on primitive mantle normalised trace and rare earth element diagrams and a wide range of $\delta^{11}\text{B}$ values from -11.5 to $+15.6\%$. Based on these data, Gurenko and Kamenetsky (2011) concluded that the inclusions with “mantle”-like $\delta^{11}\text{B}$ (around -10%), also enriched in H_2O and Cl , represent the unmodified composition of their parental magmas. Moreover, they demonstrated that although a range of contaminants (seawater, NaCl saline brines, altered oceanic crust, serpentinised peridotite or marine sediments) could potentially affect the composition of these magmas, the observed volatile enrichment could be better explained by possible input of less than 3 wt.% of a H_2O - and ^{11}B -enriched fluid into the source of the Gorgona komatiites.

The present work is follow-up study (from the work of Kamenetsky et al., 2010); Gurenko and Kamenetsky, 2011) on olivines from Gorgona komatiites and their included melts and primarily focuses on their detailed oxygen isotopic composition, determined by SIMS (= Secondary Ion Mass Spectrometry). The main purpose of the present study is therefore to further constrain the origin of volatile components, and especially H_2O in the Gorgona parental magmas. Furthermore, in order to evaluate the pressure (or depth) of olivine crystallization (with the aim of discriminating between upper and lower crust levels, which presumably have different O-isotope signatures), and to place additional constraints on the behaviour of volatile components during magma origin and consequent crystallization, we re-analysed the same olivine-hosted melt inclusions remaining after the Kamenetsky et al. (2010) and Gurenko and Kamenetsky (2011) SIMS and laser ablation ICP-MS measurements. These previously studied melt inclusions, together with several newly exposed inclusions were analysed for CO_2 , H_2O , F , S and Cl (note that neither CO_2 nor F concentrations were previously reported). Based on the existing

and our new volatile component and O-isotope data, this study will help to determine if the ascending high-Mg melts, currently preserved as inclusions in olivine, have suffered contamination by either crustal rocks at shallow depth (Arndt et al., 1997), or by seawater or seawater-derived components (Shimizu et al., 2009). If such contamination has occurred, one would expect a deviation of $\delta^{18}\text{O}$ measured in the studied Gorgona olivines from that of the mantle (4.8–5.8‰; Matthey et al., 1994; Eiler, 2001) to more ^{18}O -enriched values. In this context, our new oxygen isotope data on olivines, when coupled with the available volatile contents and boron isotope data on melt inclusions will be key to unravelling the origin of wet high-Mg magmas, now preserved as komatiites and picrites on the island of Gorgona.

2. Geological setting and samples

Gorgona is located ~40 km off the Pacific coast of Colombia (Fig. 1). The geology of this small island (approximately 8.3 km long and 2.5 km wide), was first studied by Gansser (1950), who noted the occurrence of “olivine rocks” containing skeletal olivine phenocrysts, which later were recognised as “komatiites” by Gansser et al (1979) and Echeverría (1980). Echeverría (1980) also reported picrites from the south part of the island that had no skeletal olivine crystals. An undeformed peridotite-gabbro complex is located in the central part of the island and surrounded by a sequence of tilted and faulted blocks of mafic and ultramafic rocks and tuff breccias (Dietrich et al., 1981; Aitken and Echeverría, 1984). The first petrographic and geochemical work on these rocks was reported by Dietrich et al. (1981) and Echeverría (1980) and demonstrated that the olivine spinifex (komatiite) lavas contain 13 to 24 wt.% MgO. Aitken and Echeverría (1984) proposed that the parental magma of the komatiites contained ~20 wt.% MgO. Kyser et al. (1987) reported the first stable isotope (δD and $\delta^{18}\text{O}$) data on Gorgona komatiites.

Our present work is based on the same 50 olivine fragments studied by [Kamenetsky et al. \(2010\)](#) and [Gurenko and Kamenetsky \(2011\)](#). These fragments contain small portions of melt entrapped along zones of skeletal growth of the former euhedral to subhedral olivine phenocrysts, as well as 21 additional, randomly selected olivine phenocrysts from a picrite sample GOR94-32. The komatiitic samples were collected from fresh coastal exposures on the north east side of the island (**Fig. 1**; for more detail, see [Kamenetsky et al., 2010](#)).

Sample GOR94-32 was collected from a picrite lava block, whose whole rock geochemistry was reported by [Arndt et al. \(1997\)](#). As first noted by [Echeverría and Aitken \(1986\)](#) and studied subsequently by [Kerr et al. \(1996\)](#), [Révillon et al. \(2000\)](#) and [Kerr \(2005\)](#), the fragmental picritic rocks (with angular-to-rounded blocks ranging in size from a few centimetres to > 1 m) and picritic tuffs are exposed in the southern part of the island and are intruded by a suite of comagmatic dykes (**Fig. 1**). As noted by [Kerr \(2005\)](#), the picrite blocks comprise ~25 vol.% euhedral, non-skeletal olivine phenocrysts (0.5–1.0 mm), set in a groundmass of very fine-grained Cr-spinel microphenocrysts and acicular crystals of clinopyroxene and plagioclase.

3. Analytical methods

Oxygen isotopic composition of olivine grains was systematically analysed by ion microprobe. Several previously studied and newly exposed olivine hosted melt inclusions were also analysed for CO₂, H₂O, F, S and Cl concentrations by SIMS. In order to check that no host olivine was ablated by the ion beam, the spot position was carefully examined with a petrographical microscope after analysis. As an additional check, aluminium (since it is incompatible in olivine but present in the melt inclusions in concentrations >15 wt.% Al₂O₃) was also included in the analytical sequence along with the volatile components in order to monitor and correct for any ablation of olivine (for more detail, see *Supporting online material*). Following these analyses, olivine composition was re-analysed by electron microprobe (hereafter

EPMA = Electron Probe Microanalysis) as close as possible (10–30 μm) to the spots left by SIMS. Most of the studied olivines were analysed for major elements by Kamenetsky et al. (2010). However, here we report and discuss only the newly acquired olivine major element data, except for four olivine grains that were lost during remounting from an epoxy mount into indium metal. For these olivines, we use major element compositions published in Kamenetsky et al. (2010).

4. Results

4.1. Elemental and oxygen isotope variability of olivines

The olivines were systematically analysed for oxygen isotopes: a) in the centre of grain, b) near the grain boundaries and c) as close as possible to the previously studied and newly exposed melt inclusions. As noted by Kamenetsky et al. (2010), olivine phenocrysts from the studied komatiite samples are serpentinised along inner fractures and grain margins. Consequently only smaller fragments of former olivine phenocrysts remained intact for geochemical analysis and the melt inclusion study. For this reason, no direct link exists between internal and external parts of the analysed grains and the true cores and rims of the former olivine crystals.

The analyses of oxygen isotopes were replicated 2 to 28 times (except for 2 of 71 olivine grains that were analysed only in one spot) during three separate analytical sessions (for reproducibility results, see **Fig. S1.1** of *Supporting online material*). In total, 138 oxygen isotope analyses of olivine, each based on an average of up to 14 spot analyses (in total 824 individual $^{18}\text{O}/^{16}\text{O}$ determinations) were completed during the present study.

The olivines from the Gorgona picrite sample (GOR94-32) have a wider range of forsterite contents that extend to higher values ($\text{Fo}_{88.3-93.2}$), as compared to those from the komatiite samples ($\text{Fo}_{88.3-91.5}$), and contain 0.29–0.38 wt.% CaO, 0.34–0.46 wt.% NiO,

0.09–0.19 wt.% Al_2O_3 and 0.07–0.17 wt.% Cr_2O_3). However, the olivines from this picrite have systematically lower NiO concentrations than the komatiite olivines at equivalent Fo contents, resulting in two distinct trends of olivine crystallisation (**Fig. 2a, b**). These trends are in agreement with the observation of [Kerr et al. \(1996\)](#), who noted that Gorgona picrites usually have lower Fe_2O_3 , Ni and Zr and somewhat higher Sc and Y concentrations (at a given MgO) than the komatiites. We think that the observed narrow Fo-ranges within individual samples ($\Delta\text{Fo} = \text{Fo}_{\text{max}} - \text{Fo}_{\text{min}}$ for a given sample is between 0.6 and 3.2 mol.% in komatiites and 4.9 mol.% in picrite; **Fig. 2a**), similar concentrations of NiO (0.36–0.50 wt.%, except for one grain with 0.32 wt.%; **Fig. 2b**), CaO (0.3–0.36 wt.%, except for one grain with 0.4 wt.%), Al_2O_3 (0.07–0.14 wt.%) and Cr_2O_3 (0.07–0.15 wt.%) (**Table S2.1 – Supporting online material**), and almost complete overlap of the compositions of the internal and external parts of the studied olivine fragments imply almost no elemental intra-crystal zoning.

The $\delta^{18}\text{O}_{\text{Ol}}$ span a range of values between 4.7 and 6.0‰ (**Fig. 2c, d**). Because most individual grains were multiply analysed in the centre of the grain and near the grain boundaries, the average external precision of $\delta^{18}\text{O}_{\text{Ol}}$ determinations in the internal and external grain zones is $\pm 0.2\text{‰}$, 2 SE (broad range is 0.07–0.47‰). The majority of $\delta^{18}\text{O}_{\text{Ol}}$ values (~66% i.e., 91 of 138) are mantle-like, with $\delta^{18}\text{O}_{\text{Ol}}$ ranging from 4.8 to 5.4‰. A significant number of $\delta^{18}\text{O}_{\text{Ol}}$ determinations (~33%, i.e., 45 of 138) exceed the “mantle” range towards higher values, and only two of 138 are below 4.8‰. The probability density distribution curves calculated using *Isoplot 3.70* ([Ludwig, 2008](#)) reveal a common maximum of $\delta^{18}\text{O}_{\text{Ol}}$ at ~5.3‰ for most of the studied olivine “cores” and “rims” (**Fig. 2e**), but the “rims” exhibit also an additional, subordinate peak at ~5.6‰. The absence of systematic differences (or zoning) in both Fo and $\delta^{18}\text{O}_{\text{Ol}}$ values between internal and external parts of the grains can be interpreted as evidence for only minor interaction of the crystallised olivines with a second primitive or evolved magma, or interaction of the olivine parental magma with surrounding crustal rocks.

Prior to this study only a limited amount of oxygen isotope data was available for Gorgona mafic rocks: [Kyser et al. \(1987\)](#) reported (from two basalts and one picrite) a range of $\delta^{18}\text{O}$ from 6.0 to 6.3‰ for olivine and 7‰ for one clinopyroxene. The results of this study were interpreted as supporting a mantle origin of the Gorgona komatiites, suggesting that the mantle source of these rocks may have been more fertile, compared to that of Archaean komatiites. Moreover, the δD and $\delta^{18}\text{O}$ values obtained for chrysolite from Gorgona komatiitic samples by [Kyser et al. \(1987\)](#) were interpreted to be consistent with serpentinisation by meteoric water rather than from seawater. A subsequent study by [Révillon et al. \(2002\)](#) reported a $\delta^{18}\text{O}$ range from 5.25 to 5.73‰ for clinopyroxene separates from the Gorgona picrites and komatiites. Accordingly, using the oxygen isotope fractionation factor between olivine and clinopyroxene ($\Delta^{18}\text{O}_{\text{Cpx-Ol}} = 0.33\text{--}0.37$; [Chiba et al., 1989](#)), the $\delta^{18}\text{O}_{\text{Ol}}$ equivalent calculated for 1300–1400°C using the above clinopyroxene O-isotope data is in the range of 4.9–5.4‰, i.e., values that are similar to the O-isotope range obtained in the present study.

4.2. Volatile concentrations

As previously shown by [Kamenetsky et al. \(2010\)](#) and [Gurenko and Kamenetsky \(2011\)](#), melt inclusions from the studied Gorgona olivines are characterised by pronounced enrichment of volatile components (0.2–1.0 wt.% H_2O , 520–770 $\mu\text{g/g}$ S, 220–310 $\mu\text{g/g}$ Cl) and boron (0.61–2.02 $\mu\text{g/g}$ B) which show positive anomalies in, otherwise depleted, primitive mantle normalised trace element and REE patterns. In order to constrain a minimum pressure (or depth) of olivine crystallisation (with the aim of discrimination between upper and lower crustal rocks that have different O-isotope composition; e.g., [Gregory and Taylor, 1981](#)), and to place additional constraints on the behaviour of volatile components during magma origin and consequent crystallization, we re-analysed 6 melt inclusions remaining after SIMS and laser ablation ICP-MS measurements, whose compositions are reported in [Kamenetsky et al. \(2010\)](#)

and [Gurenko and Kamenetsky \(2011\)](#), and analysed 9 additional melt inclusions for CO₂, H₂O, F, S and Cl by SIMS. More detailed information about the selection and preparation of melt inclusions prior to analysis, and their laboratory heating and quenching conditions is given in [Kamenetsky et al. \(2010\)](#). No melt inclusions in olivine from the picrite sample GOR94-32 were analysed during the present study.

Prior to interpreting the existing volatile data it is important to assess whether the original concentrations of water dissolved in the included melts are compromised by possible diffusive gain or loss of H₂O through olivine host. Such diffusion of H₂O was first noted by [Sobolev and Danyushevsky \(1994\)](#) and then experimentally investigated by [Portnyagin et al. \(2008\)](#) and other more recent studies. The problem of possible rapid exchange of water between melt inclusions in olivine and a host magma in relation to the studied Gorgona melt inclusions were briefly discussed in [Kamenetsky et al. \(2010\)](#), who have concluded that “...the enrichment of the studied melt inclusions in the volatile elements is magmatic in origin” (page 1005). Moreover, if H₂O was acquired by the inclusions due to H⁺-diffusion inside olivine hosts, this would imply the existence of a H₂O-rich ambient magma, from which the olivines with “normal mantle” $\delta^{18}\text{O}$ signature have crystallized. The existence of such a H₂O-rich magma would additionally support our previous conclusion about wet origin of the Gorgona komatiites. We thus maintain that the only process that could potentially bias water concentrations is H₂O-loss that could also happen during laboratory heating of the inclusions. However, several lines of evidence argue against extensive loss of H₂O during olivine crystallisation at depth or upon eruption:

1. The first and compelling evidence arguing against possible diffusive gain or loss of H₂O is existence of multiple significant correlations between volatile components, especially H₂O, and major (SiO₂ in the case of H₂O) and multiple incompatible trace elements (see below).

We contend that if the concentrations of water in the melt inclusions have been compromised, it is unlikely that these correlations would be preserved.

2. A main signature of the Gorgona komatiites is a spinifex texture that points towards the exceptionally fast crystallization of the parental magma resulting in formation of skeletal olivine. These olivine crystals contain very few melt inclusions and a search of >20,000 olivine fragments provided us with only a few tens of inclusions from ~ 30 μm to ~60 μm size used in this and our previous studies (see respective phototables in [Kamenetsky et al., 2010](#) and [Gurenko and Kamenetsky, 2011](#)). We contend that the studied fragments of skeletal olivine containing melt inclusions formed during and shortly after emplacement of the parental magma as a result of its rapid cooling, suggesting that there was virtually no time to compromise H_2O concentrations in the included melts. Moreover, the observed wide variations in H_2O in the studied melt inclusions (0.3–1.1 wt.%; see below), in contrast to relatively narrow variations in forsterite content of the olivine hosts ($\text{Fo } 90.7 \pm 0.6$ mol.%, excluding more evolved GOR94-17 komatiite and GOR94-32 picrite samples) would additionally support the absence of post-entrapment H_2O re-equilibration, otherwise no H_2O -range would have been preserved in the inclusions.

3. Formation of a spinifex texture requires not only rapid cooling of MgO -rich parental magma but also the cooling must occur in a strong thermal gradient (7–35 $^{\circ}\text{C}/\text{cm}$; [Faure et al., 2006](#)). If the parental magma ponded in a magma reservoir and had begun to crystallise olivine, the thermal gradient is unlikely to have been strong enough, and the cooling too slow, to form spinifex-textured olivine. All this argues against extensive olivine fractionation in a plumbing system, where diffusive H_2O loss might occur. However unlike komatiites, the studied picrite GOR94-32 is characterized by compositionally diverse olivine phenocrysts ($\text{Fo } 91.4 \pm 1.2$ mol.%), which show no skeletal appearance and contain abundant partially crystallised, glassy

inclusions. It is therefore possible that in contrast to komatiites, H₂O contents in melt inclusions from this picrite sample could be strongly compromised.

4. An additional argument in support of the rapid quenching of komatiite parental magma could be the presence of clusters of Cr-spinel crystals in the groundmass of the studied rocks. Many of these spinel crystals, although extending to more Fe-rich compositions, are compositionally similar to those included in olivine and exhibit no rims enriched in both ulvöspinel and magnetite components (in contrast to what could be observed, if a host magma had a long crystallization history; e.g., [Kamenetsky et al., 2001](#); [Gurenko et al., 2006](#)). Given that the majority of olivine-hosted spinel inclusions are in the same growth zones of skeletal olivine (as the melt inclusions), this suggests rapid crystallization of spinel occurred in the rising and erupted magma along with its nearly simultaneous entrapment by growing olivine.
5. Partial loss of H₂O from inclusions during laboratory heating is possible and thus the concentrations of H₂O in the included melts have to be considered as lower-end estimates. However as noted by [Sobolev et al \(2016\)](#) in Archaean komatiites, melt inclusions with a diameter >30 µm appear to preserve unbiased H₂O concentrations if magma was cooling fast (while forming spinifex), as well as during laboratory heating. Moreover, the existing significant correlations of H₂O with incompatible elements (see below) suggest that this process is unlikely to have had a significant effect.

The present results of volatile analysis in the studied melt inclusions are listed in **Table S2.2** (*Supporting online material*), along with the volatile concentrations, boron contents and isotopic composition obtained by [Kamenetsky et al. \(2010\)](#) and [Gurenko and Kamenetsky \(2011\)](#). The following observations can be made based on these data:

1. The concentrations of volatile components in the analysed melt inclusions (35–297 $\mu\text{g/g}$ CO_2 , 0.26–1.11 wt.% H_2O , 67–107 $\mu\text{g/g}$ F, 556–879 $\mu\text{g/g}$ S and 186–304 $\mu\text{g/g}$ Cl) are in very good agreement with our previously published data set. Correction for post entrapment crystallization of olivine (14–38 wt.% Ol; **Table S2.2 – Supporting online material**) results in the following ranges: 26–222 $\mu\text{g/g}$ CO_2 , 0.2–0.83 wt.% H_2O , 52–86 $\mu\text{g/g}$ F, 431–730 $\mu\text{g/g}$ S and 142–209 $\mu\text{g/g}$ Cl. Comparison of these ranges with those reported for chromite-hosted melt inclusions by Shimizu et al. (2009) (up to 4000 $\mu\text{g/g}$ CO_2 and below 0.4 wt.% H_2O , except for one inclusion with 0.91 wt.% H_2O), reveals that almost all inclusions analysed in the present study are characterised by significantly lower CO_2 concentrations but have higher H_2O at an equivalent CO_2 (**Fig. 3a**).
2. Decompression degassing of magma results in preferential exsolution of CO_2 (as opposed to H_2O) into vapour phase (Dixon and Stolper, 1995). The concentrations of H_2O and CO_2 directly measured by SIMS in the Gorgona melt inclusions reflect the last vapour-melt equilibrium in the magmatic system for the time of eruption and thus represent a minimum pressure of magma fractionation, as the melts could have been partially degassed. According to the *VolatileCalc* solution model (Newman and Lowenstern, 2002), they correspond to H_2O - CO_2 gas pressures between 5 and 30 MPa or to <1-km crustal depth (**Fig. 3a**).
3. Numerous studies (Esposito et al., 2011; Steele-Macinnis et al., 2011; Hartley et al., 2014; Moore et al., 2015; Wallace et al., 2015; Mironov et al., 2015) have demonstrated that 40 to 90% of the original CO_2 dissolved in the melt at the time of inclusion entrapment can be lost to a shrinkage bubble during post-entrapment cooling. A correction algorithm described by Wallace et al. (2015) was therefore used to reconstruct the original CO_2 and H_2O concentrations in the included melts (see *Supporting online material*). These calculations reveal that 79–92% (average 87%) of the original CO_2 and 1–5% (average 2%) of the original H_2O have been lost to the shrinkage bubble, in agreement with the results of the above

studies. The following ranges of volatile concentrations were obtained after correction: 286–1748 $\mu\text{g/g}$ CO_2 , 0.2–0.86 wt.% H_2O , 48–82 $\mu\text{g/g}$ F, 398–699 $\mu\text{g/g}$ S and 132–198 $\mu\text{g/g}$ Cl. According to the *VolatileCalc* solution model (Newman and Lowenstern, 2002), these results correspond to H_2O - CO_2 gas pressures between 20 and 200 MPa (average 86 ± 44 MPa, 1 SD; **Table S2.2** – *Supporting online material*) or to ~2.5-km crustal depth (**Fig. 3b**).

4. The concentrations of sulphur (either corrected for post entrapment crystallisation and shrinkage bubble volatile exsolution or not) are lower than those of MORB and ocean island basalt (OIB) submarine glasses (usually $\text{S} > 800$ $\mu\text{g/g}$ at ~7–10 wt.% FeO; Wallace and Carmichael, 1992), but plot above the “sulphur saturation line” defined by compositions of sulphide saturated experimental melts from Haughton et al. (1974) (**Fig. 4a**). The presence of micron-scale sulphide globules inside several melt inclusions suggests saturation of the included melts with sulphur. It is possible that this saturation might be caused by post-entrapment loss of Fe from the inclusions (Danyushevsky et al., 2002). However, there is good positive correlation of S with Ce ($R^2 = 0.77$ and 0.96 for uncorrected and corrected S concentrations, respectively, except for one outlier; **Fig. 4b**), as well as significant correlations of S with other incompatible elements [e.g., Ti (0.87), Sr (0.71), Zr (0.61), Ba (0.47), La (0.54), Pr (0.67), Nd (0.49), Sm (0.42), Eu (0.64), Tb (0.77), Ho (0.58), Er (0.81), Lu (0.56), Hf (0.53); the respective R^2 values for corrected S concentrations are given in brackets]. Such relationships cannot be explained by either the later, post-entrapment origin of sulphide phase in the inclusions or by magma contamination, but may imply incompatible behaviour of S at an early stage of magma fractionation and probably undersaturation of high-Mg “primary” magma with sulphur, as demonstrated on MORB magmas by Sobolev and Hofmann (1999).
5. Similarly, H_2O concentrations (in particular, post-entrapment crystallisation- and shrinkage bubble-corrected values) have a clear positive correlation with SiO_2 ($R^2 = 0.60$) (**Fig. 4c**) and

still significant negative correlations with some incompatible elements, such as Ti (0.50), Sr (0.43), Y (0.43), Zr (0.54), Ce (0.44), Sm (0.79), Dy (0.49), Ho (0.73), Er (0.58), Tm (0.43), Yb (0.74), Lu (0.46); the respective R^2 values are given in brackets (**Fig. 4d**).

6. There are, however, no correlations of H₂O with fluid mobile incompatible elements (e.g., Ba, K), Pb and Th, also implying that H₂O, along with S, might have not been degassed during magma fractionation at shallow depth. One possible explanation of the observed positive correlation of H₂O with SiO₂ (**Fig. 4c**) could be incompatible behaviour of water during magma fractionation. However, one would expect also a coherent increase in H₂O with increasing concentrations of the other incompatible elements. As this is not the case, we suggest that the correlation between SiO₂ and H₂O, in conjunction with negative trace element-H₂O correlations, could be explained by a gradual increase in the degree of partial melting of a clinopyroxene-depleted mantle source under hydrous conditions (e.g., [Parman and Grove, 2004](#)). The coherent increase in SiO₂ and H₂O contents in the magma can be explained by H₂O-fluxed decompression melting of refractory mantle accompanied by increasing contribution of orthopyroxene to the melting reaction, and consequent progressive depletion of trace element concentrations in the resulting melts.

7. The concentrations of chlorine in the included melts show no relationship with H₂O, CO₂ or S (as well as the ratios of these volatile component concentrations to those of incompatible elements, such as CO₂/Nb, H₂O/Ce, F/Zr, Cl/Nb, and S/Ce) but show a significant linear correlation with fluorine ($R^2 = 0.43$ and 0.46 for uncorrected and corrected concentrations, respectively) (**Fig. 4e**). Chlorine and fluorine usually behave as incompatible elements during partial melting and magma fractionation, if their juvenile concentrations were not affected by later degassing or contamination events (e.g., [Carroll and Webster, 1994](#); [Jambon, 1994](#)). The observed F-Cl correlation may thus suggest that the reported F and Cl concentrations were not

influenced by degassing or contamination at crustal depth and are likely to reflect mantle derived compositions, in agreement with incompatible behaviour of S and H₂O.

5. Discussion

5.1. Message from oxygen isotopes

Oxygen isotopes represent one of the most effective ways of tracing crustal components in the source of mafic magmas and can provide robust constraints on shallow-level crustal contamination because fractionation of oxygen isotopes is much more significant in low-temperature hydrothermal processes, than in the mantle (e.g. [Kyser et al., 1987](#); [Woodhead et al., 1993](#); [Eiler et al. 1995, 1996, 1997](#); [Garcia et al., 1998](#); [Harris et al., 2000, 2015](#); [Widom and Farquhar 2003](#); [Bindeman et al. 2005, 2006, 2008](#); [Wang and Eiler, 2008](#); [Day et al., 2009](#); [Martin et al., 2011](#); [Genske et al., 2013](#)). Oxygen isotope heterogeneity of the oceanic crust is generally caused by seawater and/or seawater-derived fluid alteration (e.g., [Muehlenbachs et al., 1974](#); [Hattori and Muehlenbachs, 1982](#); [Condomines et al., 1983](#); [Muehlenbachs 1986](#); [Alt et al., 1986](#)). The oceanic crust can be affected by both low (<300–400°C) and relatively high (>400°C) temperature fluid-rock interaction. It can vary in oxygen isotopic composition with depth, mostly due to cross-over of oxygen isotopic fractionation factors (and their magnitude) between minerals and water at ~300°C. Indeed, most of the oceanic crustal rocks altered by seawater cover a ~10‰ range of $\delta^{18}\text{O}$ (from ~2–3‰ to ~11–13‰) spanning the “normal”- $\delta^{18}\text{O}$ mid-ocean ridge basalt (MORB) magma value of $\sim 5.6 \pm 0.2\text{‰}$ ([Gregory and Taylor, 1981](#); [Alt et al., 1986](#); [Stakes and Taylor, 1992](#); [Bindeman, 2008](#); [Martin et al., 2011](#); [Yamaoka et al., 2012](#); [Jacques et al., 2013](#) among others). But it is also never altered to $\delta^{18}\text{O}$ values lower than ~2‰ in its middle and lower sections because $\Delta^{18}\text{O}_{\text{rock-water}}$ isotope fractionation is effectively $0 \pm 1\text{‰}$ at high (>400°C) temperatures ([Gregory and Taylor 1981](#); [Stakes and Taylor 1992](#); [Bindeman, 2008](#); [Martin et al., 2011](#)). During subsequent subduction and recycling, the relative proportions

of ^{18}O -enriched and ^{18}O -depleted material has been shown to remain broadly similar to those in the ocean crust prior to subduction, leading to $\delta^{18}\text{O}$ values in erupted volcanic-arc products that are distinct from that of pristine mantle (e.g., [Bebout and Barton 1989](#); [Putlitz et al. 2000](#)).

Previous studies of tholeiitic to alkali basalts from Iceland, Hawaii, the islands of Tristan da Cunha, Gough, Canary and Azores (e.g., [Harris et al., 2000](#); [Bindeman et al., 2008](#); [Wang and Eiler, 2008](#); [Day et al., 2009](#); [Genske et al., 2013](#)) have demonstrated that a coherent decrease of $\delta^{18}\text{O}_{\text{Ol}}$ and Fo is most likely to be a signature of assimilation of low- $\delta^{18}\text{O}$ crustal rocks by ascending magma. This results in crystallization of olivine with lower than typical “mantle” (i.e., 4.8–5.4‰) $\delta^{18}\text{O}$ values from hybrid magma. The slope of such Fo- $\delta^{18}\text{O}$ relationships may vary significantly depending on the particular O-isotope composition of a contaminant, ranging from the most extreme case of Hawaii ($\delta^{18}\text{O}_{\text{Ol}}$ decreases from 5.3‰ to 4.6‰ in the range from Fo₉₀ to Fo₈₈ i.e., 0.35‰/mol.% Fo; [Wang and Eiler 2008](#)) through the Canaries (from 5.0‰ to 4.7‰ in the range from Fo₈₄ to Fo₇₉, i.e. 0.06‰/mol.% Fo; [Day et al., 2009](#)) to the Azores ($\delta^{18}\text{O}_{\text{Ol}}$ = 5.2–4.8‰ in olivines Fo_{90–77}, i.e., 0.03‰/mol.% Fo) (see *contamination trends* in **Fig. 2c**). Conversely, olivine phenocrysts crystallized from mantle-derived magmas can show no clear signs of shallow-depth crustal contamination (for example, Society, Samoa and Canary islands, and the Karoo and Etendeka large igneous provinces). These olivine phenocrysts are characterised by elevated $\delta^{18}\text{O}$ relative to typical upper mantle values and this was ascribed to the presence of an ^{18}O -enriched component in the magma source (e.g., [Eiler et al., 1997](#); [Gurenko et al., 2011](#); [Harris et al., 2015](#)).

To justify such trends, the analytical precision of a “single grain” O-isotope determination must be at ± 0.1 – 0.2 ‰, 2 SD, and such values are typical for laser fluorination analyses (e.g., [Bindeman, 2008](#)). In the case of SIMS measurements, one viable way to maintain the required analytical precision is to increase a number of individual measurements within a single grain. As noted above, most of the studied Gorgona olivine grains were multiply analysed

by SIMS. No substantial compositional difference has been recognised between internal and external parts of the grains, except for the second subordinate maximum in the “rim” probability density distribution curve at ~5.6‰ (**Fig. 2e**). As most of the studied olivine phenocrysts exhibit varying degrees of serpentinisation but no clear relationship of Fo with $\delta^{18}\text{O}_{\text{Ol}}$ (**Fig. 2c**), and since shallow crustal contamination of the studied samples was not previously identified by [Gurenko and Kamenetsky \(2011\)](#) nor by our present volatile data, our preferred explanation for the second maximum and probably for the all $\delta^{18}\text{O}_{\text{Ol}}$ values in excess of the “mantle” range is a diffusive oxygen isotope exchange between olivine (4.8–5.4‰) and products of its low-temperature alteration (for example, with serpentine with $\delta^{18}\text{O}$ up to ~13‰; [Martin et al., 2011](#)).

A less probable possibility is that the shift to more positive $\delta^{18}\text{O}$ olivine compositions (**Fig. 2e**) could be related to admixture of a portion of a second magma derived from an ^{18}O -enriched mantle component, or by analogy with [Martin et al. \(2011\)](#) who reported $\delta^{18}\text{O}$ values of up to 6.1‰ in Mt. Shasta olivine phenocrysts, and related them to “a ^{18}O pre-enriched” peridotite mantle wedge component. However, we consider this explanation unlikely mostly because (a) the rapid, *in-situ* crystallisation (advocated herein) of skeletal olivine in the magma either en route to the surface or later in the lava flow would result in more diffused distribution of $\delta^{18}\text{O}_{\text{Ol}}$ values, with no well shaped maximums, as shown in **Fig. 2e**, and (b) the observed $\delta^{18}\text{O}$ enrichment is related to the edges of the studied olivine grains, which were in direct contact with serpentine filling grain fractures and whose origin can be ascribed to later alteration stage of the lava flows. In addition, a subordinate number of $\delta^{18}\text{O}$ -enriched compositions determined for olivine “cores” (**Fig. 2e**) are ascribed to the so-called “3D-effect”, when ion beam visually placed in the centre of a grain is in fact analysing its remaining outer part oriented parallel to the sample surface.

Calculation the average “single grain” $\delta^{18}\text{O}_{\text{Ol}}$ composition from all individual $\delta^{18}\text{O}_{\text{Ol}}$ measurements performed for each olivine grain, regardless of whether they represented “core” or

“rim” areas, results in the average standard deviation of mean external reproducibility of $\pm 0.12\text{‰}$, 2 SE (grain to grain range from ± 0.06 to $\pm 0.26\text{‰}$ due to the different number of replicates), being comparable with those of single-grain laser fluorination. Figure 5 shows averaged “within-grain” oxygen isotope compositions of the studied Gorgona olivines with no discrimination between their internal and external zones. Like in **Fig. 2c**, no relationship between Fo contents and $\delta^{18}\text{O}_{\text{Ol}}$ values for olivines can be observed for komatiite lavas. In contrast, olivines from picrite GOR94-32 show a subtle positive correlation ($R = 0.53$), best approximated by a polynomial quadratic equation (**Fig. 5a**), which is significant at 95% confidence level (critical value of the Pearson correlation coefficient R is 0.44 at $N = 20$). Similarly, as in **Fig. 2e**, the probability density distribution curves reveal a common maximum of typical mantle $\delta^{18}\text{O}_{\text{Ol}}$ values between 5.2 and 5.3‰ (**Fig. 5b**). A subordinate peak at $\sim 5.6\text{‰}$ for olivines from komatiites and three less-pronounced subordinate peaks at $\sim 5.6\text{--}5.9\text{‰}$ for olivines from picrite are thought to have been inherited from the analyses of rims.

If our assumption about the possible impact of later serpentinisation on the O-isotope composition of olivine is correct, then some of the $\delta^{18}\text{O}_{\text{Ol}}$ values, including those from the “mantle” range, could have been systematically shifted to more ^{18}O -enriched compositions. This would result in more diffused shape of the maximums in the density distribution (**Fig. 2e**) and representing the upper level of $\delta^{18}\text{O}_{\text{Ol}}$ estimates. Moreover, the absence of a clear relationship between Fo content and $\delta^{18}\text{O}_{\text{Ol}}$ in komatiites also could be explained by this olivine-serpentine O-isotope exchange, during which the isotopic composition of olivine was modified to a greater extent than the Fo content (as serpentinisation does not affect forsterite number in olivine, as recently shown by Birner et al., 2016). In this context, the most Fo-rich olivines of the “picrite” trend in **Fig. 5a** having $\delta^{18}\text{O}_{\text{Ol}}$ values outside “mantle” range could be affected by this serpentinisation process, suggesting also that their former unmodified $\delta^{18}\text{O}_{\text{Ol}}$ values originally corresponded to the “normal” mantle $\delta^{18}\text{O}_{\text{Ol}}$ range.

5.2. Evidence against crustal contamination

The parental magmas of komatiites are hotter, up to 1600°C for Archaean komatiites (e.g., Herzberg et al., 2007 and references therein), as compared to MORB and OIB magmas, and therefore are more likely to melt and assimilate surrounding wall rocks en route to the surface. A melt inclusion study of 2.7-Ga-old Belingwe komatiites, Zimbabwe, demonstrated that their parental magmas probably contained up to 1 wt.% H₂O, but the source of water was uncertain (i.e., shallow magma contamination versus melting of a hydrous mantle plume source; e.g., Shimizu et al., 2001; Berry et al., 2008; Kent et al., 2009). Recently, Sobolev et al. (2016) reported marked H₂O enrichment of Archaean komatiites from the Abitibi greenstone belt, Canada, with H₂O/Ce >6000, and proposed the presence of a hydrous, deep mantle reservoir formed early in the Earth's history. However, those melts with high Cl/K₂O \approx 7, Cl/F \approx 100 and Cl/H₂O \approx 0.6 ratios were explained by Sobolev et al. (2016) as being due to assimilation of serpentinites altered by seawater with possible involvement of seawater-derived brines. For comparison, Cl/K₂O from 0.23 to 0.65, Cl/F from 2.4 to 3.5, and Cl/H₂O from 0.02 to 0.09 in the studied Gorgona melt inclusions are well below the above values ascribed to contamination by Sobolev et al. (2016), and the controversy remains as to whether the obtained volatile concentrations preserved in the Gorgona melt inclusions are pristine or they were inherited during assimilation of crustal rocks by the magma (e.g., Shimizu et al., 2001; Berry et al., 2008; Kent et al., 2009; Kamenetsky et al., 2010).

One of the main arguments against a hydrous origin of the Gorgona komatiites is their extreme depletion in incompatible trace elements. This depletion becomes progressively more pronounced on going from basalts through komatiites to picrites. However, it is clear that all these rock types are related to a similar type of depleted mantle source (e.g., Arndt et al., 1997; Révillon et al., 2000, 2002). These authors ascribed the origin of volatiles (and in particular

water) either to sub-solidus alteration processes or to assimilation of hydrated crust in a shallow-level magma reservoir. This latter conclusion was also reached by Shimizu et al. (2009), who in their study of chromites from the beach sands of Gorgona described CO₂-rich (40–4000 µg/g) melt inclusions with highly variable H₂O (0.03–0.9 wt. %) and Cl (6–1056 µg/g) and proposed that the concentrations of volatile components, and especially Cl, were influenced by magma degassing or interaction with seawater or brine prior to entrapment.

Significantly, however, Gurenko and Kamenetsky (2011), who studied boron contents and B isotopic compositions in the same melt inclusions analysed during the present study, argued that it was not possible to unequivocally prove magma contamination only on the basis of Cl/K₂O vs. H₂O/K₂O relationships (or similar, where Cl- and H₂O-concentrations are normalised to a concentration of an incompatible element). These authors showed that although H₂O/K₂O-Cl/K₂O or H₂O/K₂O-B/K₂O systematics can be explained by contamination of magma by seawater or NaCl-rich brines, the relationships between B/K₂O and Cl/K₂O or B/K₂O and $\delta^{11}\text{B}$, for the same inclusions do not support their role, as well as a role of altered oceanic crust, serpentinites or marine sediments as possible contaminants.

The highest obtained H₂O/Ce = 4490 of the studied melt inclusions is significantly lower than that of the Archaean komatiites (H₂O/Ce >6000; Sobolev et al., 2016) and the highest CO₂/Nb = 3320 and Zr/Y = 2.1 are very close to, but F/Zr = 2.6 is significantly lower than, those of the “primary komatiite” of Shimizu et al. (2009) (i.e., CO₂/Nb = 4000 ± 2200, F/Zr = 7.0 ± 1.7, Zr/Y = 1.7 ± 0.3). In addition to not being as high as in this hypothetical “primary komatiite”, the F/Zr ratios (1.5–2.6) in the olivine melt inclusions analysed in this study, are at the lower end of the F/Zr range in the chromite-hosted inclusions of Shimizu et al. (2009). We note that all chromites studied by Shimizu et al. (2009) are from the beach sands and are likely to come from any type of mafic rocks exposed on Gorgona, not necessarily from komatiites. This may suggest that the inclusions analysed during the present study and those of Shimizu et al.

(2009) perhaps represent two similar but quite different magmatic systems, at least with respect to their dissolved volatiles.

The observed significant correlations of S, and to a lesser extent H₂O, with almost all highly incompatible elements (**Fig. 4b, d**), the good correlation between F and Cl (**Fig. 4e**), in addition to there being no clear relationship between Cl and H₂O, suggest that neither shallow magma degassing nor crustal contamination have affected volatile concentrations in the studied magmas during ascent.

Moreover, if rapid crystallization of skeletal olivine occurred at shallow crustal depth (no deeper than ~2.5 km, as follows from CO₂-H₂O data) during magma transport to the surface and probably continued *in-situ* in the lava flow, in contrast to the much slower crystallization in a magma chamber, then H₂O (as well as chlorine and probably other volatiles) must have been supplied to the magma by interaction with near-surface, seawater pre-conditioned crustal rocks. This would imply crystallization of olivine with higher than “normal mantle” $\delta^{18}\text{O}$ composition (i.e., 4.8–5.4‰), because upper-level crustal rocks suffered low-temperature (<300 °C) alteration by seawater are characterized by elevated, up to 13‰, $\delta^{18}\text{O}$ values (see above). Since the majority of the studied olivines are “mantle”-like in respect of $\delta^{18}\text{O}$, we consider the possibility of surface contamination of the komatiite parental magma by variable amount of seawater-preconditioned crustal components as unlikely.

5.4. Boron vs. oxygen isotope relationships

Boron vs. oxygen isotope systematics have been shown to be a very powerful tool in discriminating between pure mantle melting and crust recycling, slab-mantle wedge interaction or shallow depth magma contamination. In particular, [Chaussidon and Jambon \(1994\)](#), [Chaussidon and Marty \(1995\)](#), [Smith et al. \(1995\)](#) have demonstrated on primitive MORB and OIB glasses that positive $\delta^{11}\text{B}$ associated with increasing oxygen isotope ($\delta^{18}\text{O} > 5.5\text{‰}$) and

radiogenic Sr isotope ratios are more likely to result from shallow-level assimilation of oceanic crust, whereas negative $\delta^{11}\text{B}$ values accompanied by low $\delta^{18}\text{O} < 5.5\text{‰}$ point towards interaction of primitive magmas with the Layer 3 gabbroic crust. Moreover, the systematically higher $\delta^{11}\text{B}$, reported for arc lavas by [Peacock and Hervig \(1999\)](#), as compared to exhumed subduction related metamorphic rocks, may suggest that dehydration reactions can significantly decrease $\delta^{11}\text{B}$ of the subducting oceanic crust and sediments, thereby implying significant fractionation of ^{11}B and ^{10}B and enrichment of the fluid in ^{11}B isotope.

Although variable but generally positive $\delta^{11}\text{B}$ values have been identified in mantle wedge serpentinites at the slab-mantle interface (e.g., [Benton et al., 2001](#); [Savov et al., 2005, 2007](#)), it has also been shown that subduction-related serpentinites, eclogites and slab-derived melts may be strongly depleted in ^{18}O , and have significantly lower $\delta^{18}\text{O}$ than that of typical mantle (down to $\sim 2\text{--}4\text{‰}$) (e.g., [Bindeman et al., 2005](#); [Bindeman 2008](#); [Martin et al., 2011](#); [Jacques et al., 2013](#)). In addition, [Martin et al. \(2011\)](#) have demonstrated that slab-derived fluids can be markedly enriched in ^{18}O , if dehydration occurred at low T , i.e., $< 300^\circ\text{C}$. This is because the fluid-rock O-isotope fractionation factor, $\Delta^{18}\text{O}_{\text{fluid-rock}}$, can be as high as 10‰ at this temperature. But O-isotope composition of the fluid can broadly correspond to the composition of the source rock if released at higher temperatures (because $\Delta^{18}\text{O}_{\text{fluid-rock}} \approx 0$ at $T > 400^\circ\text{C}$).

[Gurenko and Kamenetsky \(2011\)](#) previously identified two groups of Gorgona olivine-hosted melt inclusions in B- $\delta^{11}\text{B}$ space. The first group of inclusions has relatively low $\delta^{11}\text{B}$ from -11.5 to -7.3‰ (average $-9.0 \pm 1.5\text{‰}$), spanning the entire range of B concentrations and closely corresponding to the Earth's mantle value of $-10 \pm 2\text{‰}$ ([Chaussidon and Marty, 1995](#)). The second group is characterised by generally higher $\delta^{11}\text{B}$ values than mantle $\delta^{11}\text{B}$ (up to $+15.6\text{‰}$) and increasing B concentrations up to $1.6 \mu\text{g/g}$ (see Fig. 3 in [Gurenko and Kamenetsky, 2011](#)). The first group was interpreted to be of mantle origin, whereas the second was explained by input of $< 3 \text{ wt.}\%$ of ^{11}B -enriched (possibly subduction related) fluid into the

source of the Gorgona komatiites. The main limitation of the proposed mechanism, however, is the absence of a clear geochemical signature of subduction in the composition of the studied melt inclusions such as depletion of high field strength elements (Nb, Zr, Ti) coupled with U-shaped patterns of incompatible element spectra due to the enrichment of large ion lithophile elements (Rb, Ba, K, Sr) and Pb. It is important to note that a similar absence of the subduction signatures in the H₂O-rich olivine-hosted melt inclusions from the 2.7-billion-year-old komatiites from the Abitibi greenstone belt in Canada led [Sobolev et al \(2016\)](#) to propose the presence of a hydrous deep-mantle reservoir in the Archaean mantle.

Figure 6 shows the B isotopic composition of Gorgona melt inclusions from [Gurenko and Kamenetsky \(2011\)](#) linked with new O-isotopic data on the hypothetical melts in equilibrium with olivines, for which $\delta^{18}\text{O}_{\text{Ol}}$ were determined (olivine equivalent). In order to calculate the O isotope fractionation factor we used a coefficient $A = -0.54$ in the general isotope fractionation equation:

$$\Delta^{18}\text{O}_{\text{ol-melt}} = A \times 10^6/T^2 \quad \text{Eq. 1}$$

defined for komatiite by [Matthews et al. \(1998\)](#). The calculated $\Delta^{18}\text{O}_{\text{ol-melt}}$ was found to be between -0.199 and -0.212 in the T -range of 1320 – 1380 °C, i.e., the crystallization temperatures of the studied olivines ([Kamenetsky et al., 2010](#)). Surprisingly, but even with a large range of boron isotopic composition from -11.5 to 15.6 ‰ of $\delta^{11}\text{B}$, the assigned $\delta^{18}\text{O}_{\text{melt}}$ values only slightly exceed the “mantle” interval, ranging from 5.0 to 5.9 ‰, whilst the inclusions of the first, low- $\delta^{11}\text{B}$ group of [Gurenko and Kamenetsky \(2011\)](#) fit almost entirely the “mantle” array in respect of both O and B isotope systematics (**Fig. 6**).

Although the mantle origin of the first, low- $\delta^{11}\text{B}$ group of inclusions can be determined on the basis of combined $\delta^{11}\text{B}$ - $\delta^{18}\text{O}$ isotopic data, the origin of the second group of inclusions

with varying $\delta^{11}\text{B}$ data remains enigmatic. We have calculated several two-component mixing trends in $\delta^{18}\text{O}$ - $\delta^{11}\text{B}$ space (**Fig. 6**), illustrating the effects of possible interaction of fluids released from oceanic crust at different temperatures, with a MORB-like depleted mantle source, as well as the effects of the possible presence of recycled crust and variously dehydrated serpentinite in the magma source cf. [Gurenko and Kamenetsky \(2011\)](#). We used B contents and $\delta^{11}\text{B}$ values of the mixing end-members as defined by [Gurenko and Kamenetsky \(2011\)](#) coupled with the respective oxygen isotope data (for explanation see **Table S2.3** – *Supporting online material*). As shown by the calculations, our new B-O isotopic data also confirm the importance of a H_2O - and B-rich fluid component released at low T (*FL1*) and/or serpentinite (*SERP1*) interacting with mantle source rocks (less than 5% of input is required) and, in contrast, demonstrate the unlikely, but still possible, role of the recycled crust (*REC*), dehydrated serpentinite (*SERP2*) and fluids released at greater depths and higher temperatures (*FL2* and *FL3*) (more than 10% addition to the source is required). However, we note that although the potential role of each of the selected mixing end-members cannot be completely ruled out taking into account the analytical precision of $\delta^{11}\text{B}$ and $\delta^{18}\text{O}$ determinations, their possible admixture to the *MANT* source should not significantly affect the volatile concentrations of the resulting melts as no difference in volatile contents exists between high- $\delta^{11}\text{B}$ and low- $\delta^{11}\text{B}$ groups of inclusions.

However, the question exists as to whether the Gorgona picrites and komatiites came from a distinct or a compositionally similar mantle source. As noted by [Kerr et al. \(1996\)](#), in addition to the difference in major and trace element ratios (e.g., Al/Ti, Ce/Sm, Gd/Yb), they also show a subtle difference in their radiogenic isotope signatures, suggesting that their mantle source is unlikely to be the same. Similar conclusions also can be derived from the two distinct trends of magma fractionation in the Fo-NiO diagram (**Fig. 2b**). As shown by several recent studies, Ni concentration in olivine (along with Mn and Ca) is a very good proxy to access relative contribution of the peridotite (mantle) and pyroxenite (recycled crust) components in the

mantle source of primary magmas, and these relative contributions can be directly related to the radiogenic isotope composition of the host lavas (e.g., Sobolev et al., 2007, 2008; Gurenko et al., 2013, among others). Using the Sobolev et al. (2008) parameterisation for Ni in olivine, we obtain substantially different fractions of anticipated pyroxenite-derived melt in the picritic ($X_{pxm} = 0.27 \pm 0.07$, 1 SD) and “komatiitic” ($X_{pxm} = 0.40 \pm 0.06$, 1 SD) primary melts that also may partially account for the observed subtle difference in oxygen and radiogenic isotope compositions. Further implications of this method for the possible variability in mineral, chemical and isotopic compositions of these types of rocks are beyond the scope of the present paper. However, given that clinopyroxene is a major inventory of H₂O as, for instance, in the depleted, MORB-like mantle (e.g., Danyushevsky et al., 2000), differences in pristine concentrations of H₂O in Gorgona picrite and “komatiite” parental magmas (where komatiites are more H₂O-rich) can be predicted.

6. Conclusions

1. The $\delta^{18}\text{O}$ values in the studied olivines from one picrite and five komatiites of Gorgona, Colombia, range from 4.7 to 6.0‰, with the majority (~66% population) being within the typical “mantle” olivine range of 4.8–5.4‰, but with a subordinate but still significant number (~33% population) above, and only 2 olivine grains below, this range. No zoning or systematic difference in forsterite contents and $\delta^{18}\text{O}_{\text{Ol}}$ values between the internal and external parts of the grains have been recognised.
2. The probability density distribution curves reveal a common maximum of $\delta^{18}\text{O}_{\text{Ol}}$ at ~5.3‰ for both internal and external zones of olivine, but the “rims” exhibit also an additional, subordinate maximum at ~5.6‰ probably due to diffusive oxygen isotope exchange between fresh and low-temperature alteration domains in the olivines. The O-isotope mantle signature of the majority of the studied olivines do not support shallow depth contamination of the

Gorgona high-Mg magmas, and rather favour a deep mantle origin of high volatile concentrations in the olivine-hosted melt inclusions.

3. The concentrations of volatile components in olivine-hosted melt inclusions analysed during the present study agree well with our previously reported data. After correction for post-entrapment crystallization of olivine in the inclusions upon cooling, the concentrations of CO₂ and H₂O reveal minimum pressure estimates of 5–30 MPa, which likely indicate the last vapour-melt equilibrium shortly prior to eruption of magma (<1 km crustal depth). However, ~87% and ~2% of the original CO₂ and H₂O, respectively, have been lost to the inclusion shrinkage bubbles during magma cooling. After correction for this effect, the obtained presumable undegassed ranges of volatile concentrations imply a H₂O-CO₂ gas pressure of 86 ± 44 MPa or ~2.5 km of crustal depth of olivine crystallisation.
4. Significant correlations of S and, to a lesser extent of H₂O, with a number of highly incompatible elements, along with the correlation between F and Cl, but no clear relationship of H₂O with Cl, argue against shallow depth degassing and/or crustal contamination of the studied magmas.
5. The new B-O isotopic data confirm the importance of a H₂O- and B-rich fluid component interacting with the source rocks of the Gorgona magmas, but reduce the likelihood of a role for recycled crust and serpentinite. Therefore, the results of this study suggest that the reported enrichment of volatile components and boron represent a deep mantle signature of the studied Gorgona mafic and ultramafic magmas.

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Figure captions

Fig. 1. Simplified geological map of the Gorgona island (modified after Echeverría, 1980; Kerr et al., 1996; Révillon et al., 2000). Sampling locations are marked by sample numbers referred to in the text.

Fig. 2. Elemental and oxygen isotopic composition of olivine from Gorgona picrite and komatiite samples. Panels (A) and (D) represent inter- and intra-sample variations of forsterite contents and $\delta^{18}\text{O}_{\text{Ol}}$ values, respectively. Shown here *error bars* represent the largest $\pm 2\sigma$ analytical uncertainty attained during replicate measurements of the San Carlos olivine standard by EPMA (± 0.2 mol.% Fo) and SIMS ($\pm 0.4\text{‰}$ of $\delta^{18}\text{O}$). Panels (B) and (C) are variation diagrams of NiO (B) and $\delta^{18}\text{O}_{\text{Ol}}$ (C) vs. Fo in the central parts and near the grain boundaries; (E) cumulative probability density curves of $\delta^{18}\text{O}_{\text{Ol}}$ in *Ol core* and *Ol rim* parts calculated using *Isoplot 3.70* (Ludwig, 2008). The *Gorgona Ol literature data* set (panel B) shows Fo and NiO concentrations of olivines from Kamenetsky et al. (2010), Gurenko and Kamenetsky (2011), as well as authors' unpublished data. The *shaded band* (panels C, D, E) denotes the range of typical upper mantle olivine (4.8–5.4‰; Matthey et al. 1994) and of olivine in equilibrium with N-MORB magmas (assuming olivine-melt fractionation of -0.4‰ and the N-MORB range of 5.2–5.8‰, Eiler, 2001). Dashed, numbered curves in panel (C) represent the Azores (1) and Hawaii (2) *contamination trends* after Wang and Eiler (2008) and Genske et al. (2013), respectively.

Fig. 3. Concentrations of CO_2 and H_2O dissolved in the olivine-hosted melt inclusions from Gorgona komatiites measured by SIMS before correction for the loss in the shrinkage bubble of CO_2 and H_2O (A) and after the correction applied (B) (see text). Isobars of melt compositions in equilibrium with $\text{CO}_2\text{-H}_2\text{O}$ fluid (*solid lines*) and isopleths of fluid composition (dashed lines

with labels returning molecular fraction of H₂O in the fluid) were calculated using *VolatileCalc* solution model (Newman and Lowenstern, 2002). Both isopleths and isobars were calculated for a magma having 47 wt.% SiO₂ at 1360°C, corresponding to the averaged values of the studied melt inclusions (Kamenetsky et al., 2010). Error bars represent average $\pm 2\sigma$ analytical uncertainty attained within three analytical sessions, i.e., $\pm 14\%$ relative for CO₂ and $\pm 12\%$ relative for H₂O, and they are smaller than the size of the symbol, if not shown. The data from Shimizu et al. (2009) are shown for comparison.

Fig. 4. Relationships between concentrations of S, H₂O, Cl and F with FeO, SiO₂, Ce and Y in melt inclusions. (a) FeO vs. S variations. The *MORB* & *OIB* field presents S concentrations in presumably undegassed MORB and Hawaiian submarine tholeiitic and alkali basaltic glasses after Mathez (1976), Clague et al. (1995), Dixon et al. (1997) and Wallace and Carmichael (1992). Dashed curve represents sulfur solubility in basaltic melts based on the sulfide-saturated experiments of Haughton et al. (1974). (b) Ce vs. S, (c) SiO₂ vs. H₂O, (d) Y vs. H₂O and (e) F vs. Cl variations. The observed relationships do not support shallow depth contamination of the Gorgona high-Mg magmas, and rather favour a deep mantle origin of high volatile concentrations in the olivine-hosted melt inclusions, in agreement with Kamenetsky et al. (2010), Gurenko and Kamenetsky (2011) and Sobolev et al. (2016). The positive SiO₂-H₂O correlation may also be interpreted as reflecting gradually increasing degree of source melting at hydrous conditions.

Fig. 5. Averaged, single-grain oxygen isotope compositions of the studied Gorgona olivines. (A) $\delta^{18}\text{O}_{\text{OI}}$ vs. Fo variation diagram, (B) probability density distribution curves of $\delta^{18}\text{O}_{\text{OI}}$ values for olivine from komatiites and picrite GOR94-32.

Fig. 6. Oxygen versus boron isotopic compositions of olivine hosted melt inclusions from Gorgona komatiites. Subdivision of the inclusions on to “low- $\delta^{11}\text{B}$ ” and “high- $\delta^{11}\text{B}$ ” groups and their respective boron isotopic compositions are from [Gurenko and Kamenetsky \(2011\)](#). To infer oxygen isotope composition of the included melts, we used experimental data of [Matthews et al. \(1998\)](#) for komatiite to calculate O isotope fractionation factor between olivine and melt for the temperature range of 1320–1380 °C. Two-component mixing lines between the depleted MORB-type mantle (*MANT*), recycled crust (*REC*), serpentinised peridotite (*SERP1* and *SERP2*) and hypothetical oceanic crust fluids released at different temperatures (*FL1*, *FL2* and *FL3*) are shown (for end-member compositions see **Table S2.3**, *Supporting online material*). Shaded grey field represents the composition of the mantle, with $\delta^{11}\text{B}$ of $-10 \pm 2\text{‰}$ ([Gurenko and Kamenetsky, 2011](#) and references therein) and $\delta^{18}\text{O}$ of $5.5 \pm 0.5\text{‰}$, assuming a calculated oxygen isotope fractionation of -0.2‰ between olivine and high-Mg melt and the olivine range of 4.8–5.8‰, as defined in **Fig. 2**.

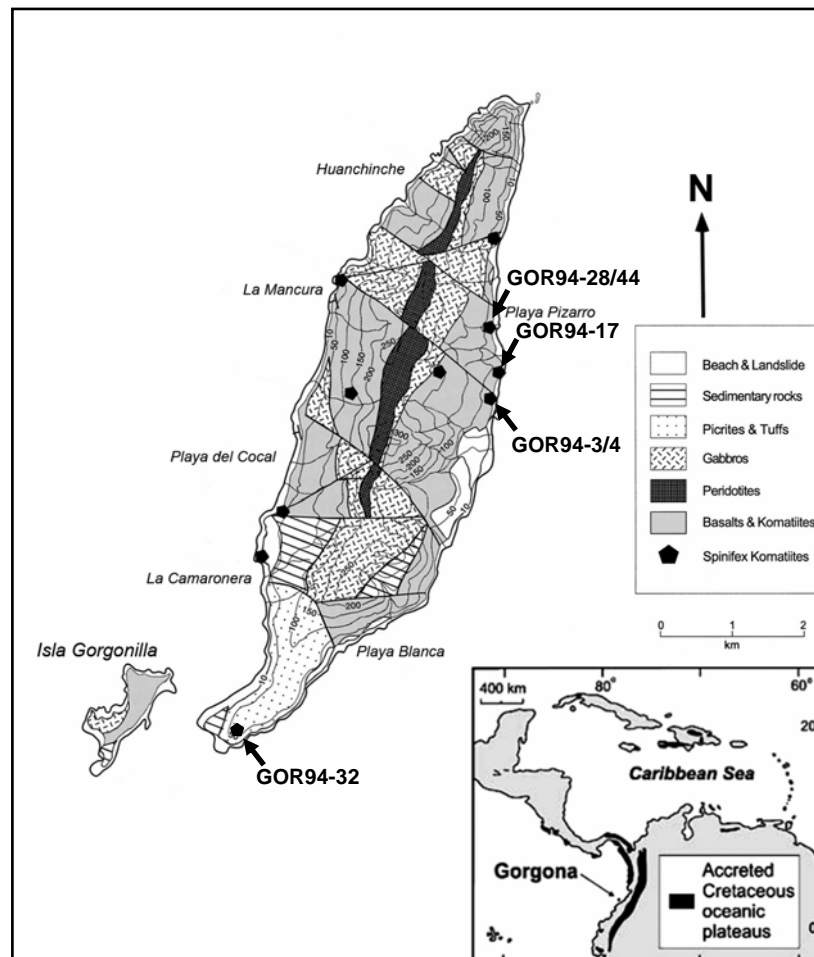


Fig. 1. Gurenko et al.

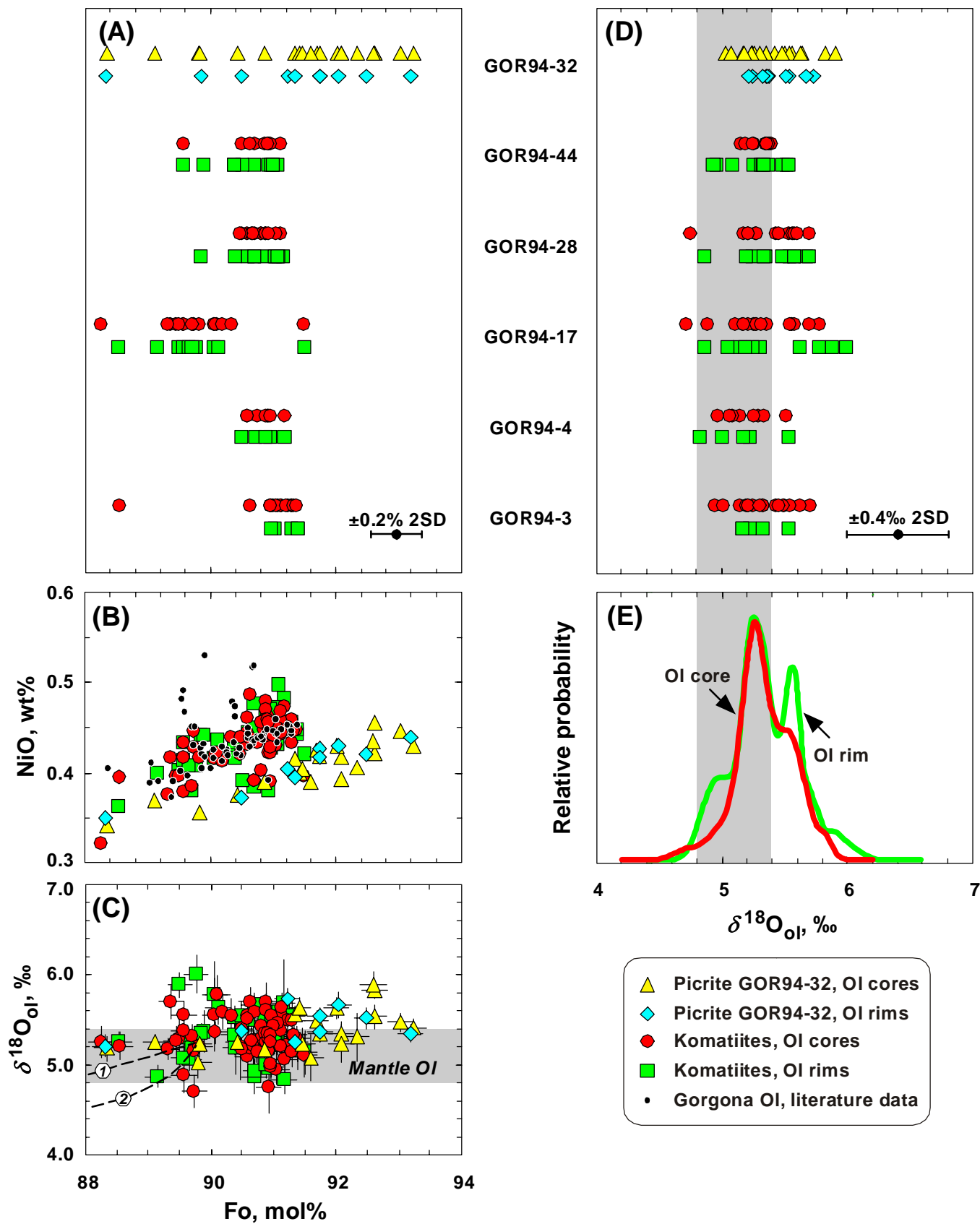


Fig. 2. Gurenko et al.

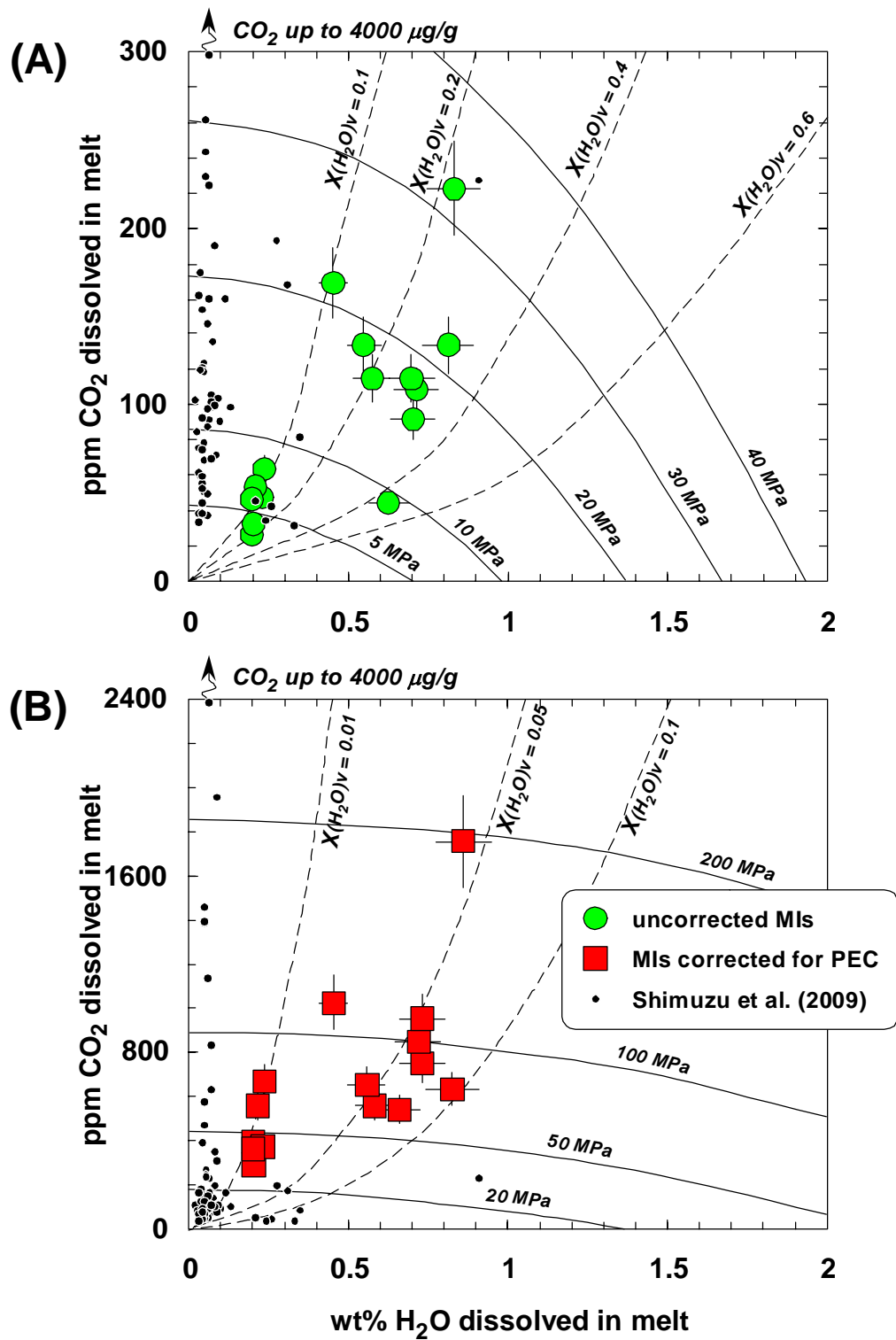


Fig. 3. Gurenko et al.

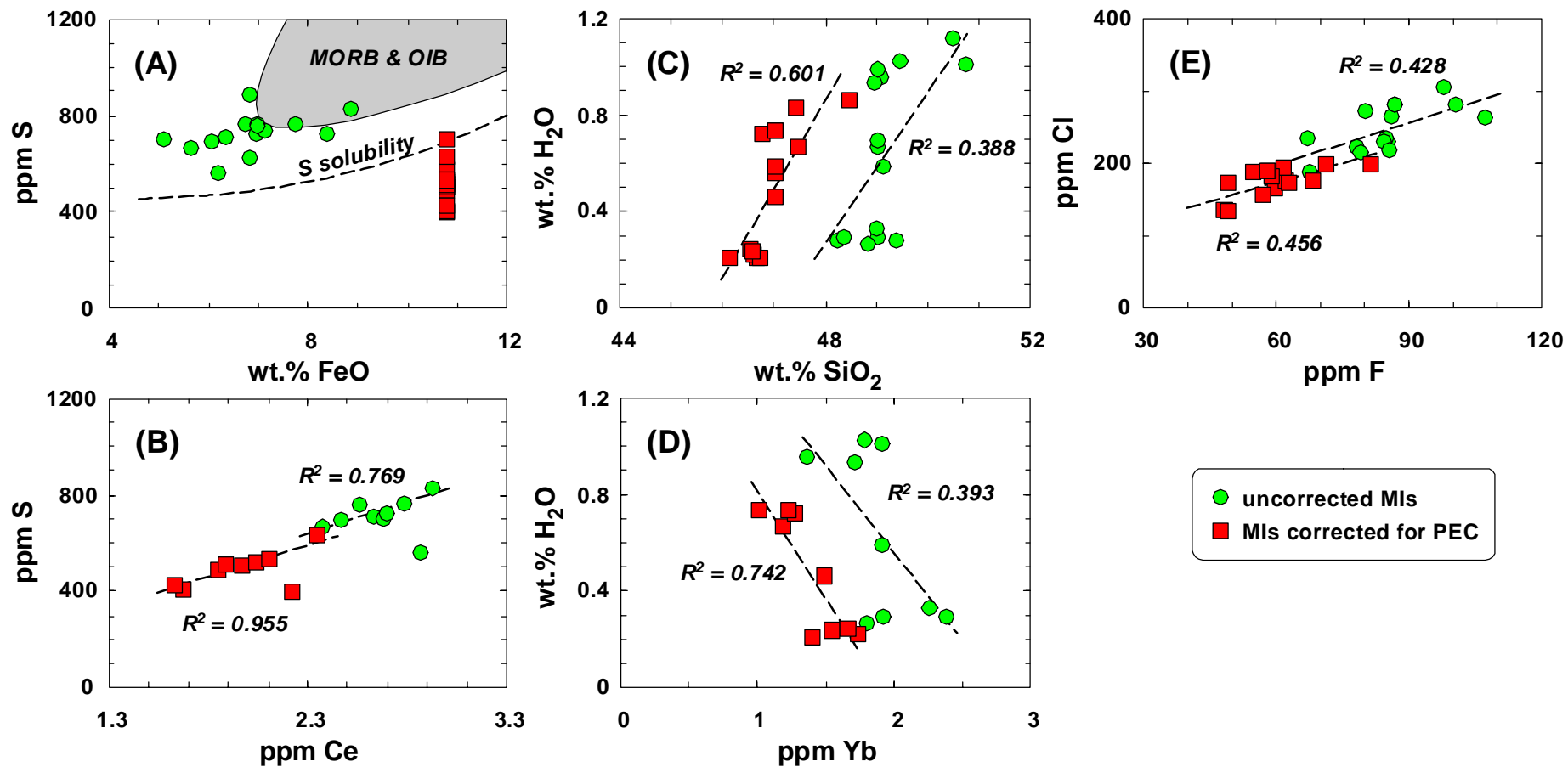


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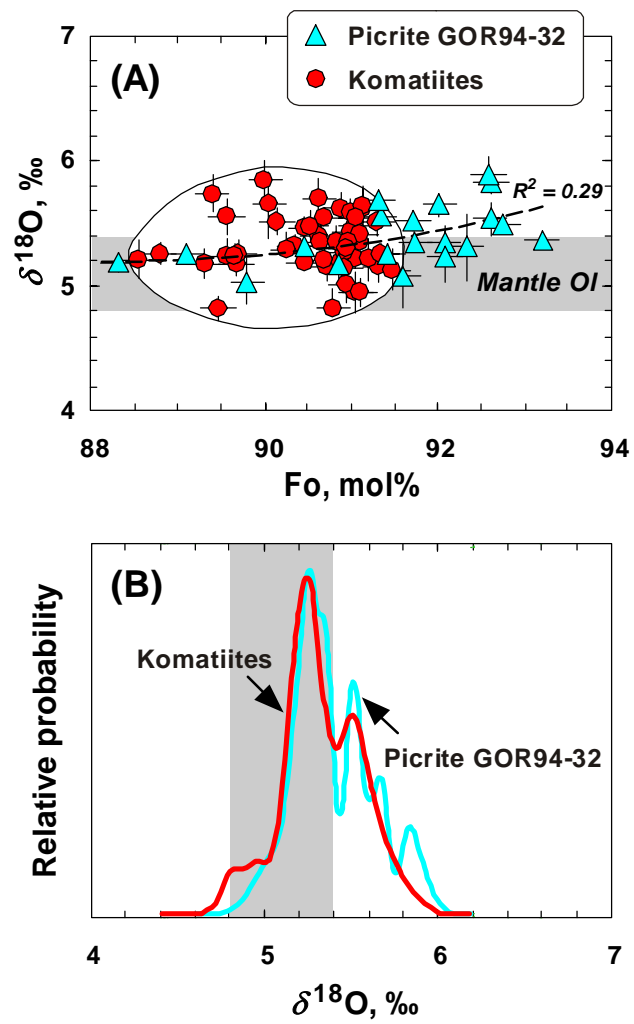


Fig. 5. Gurenko et al.

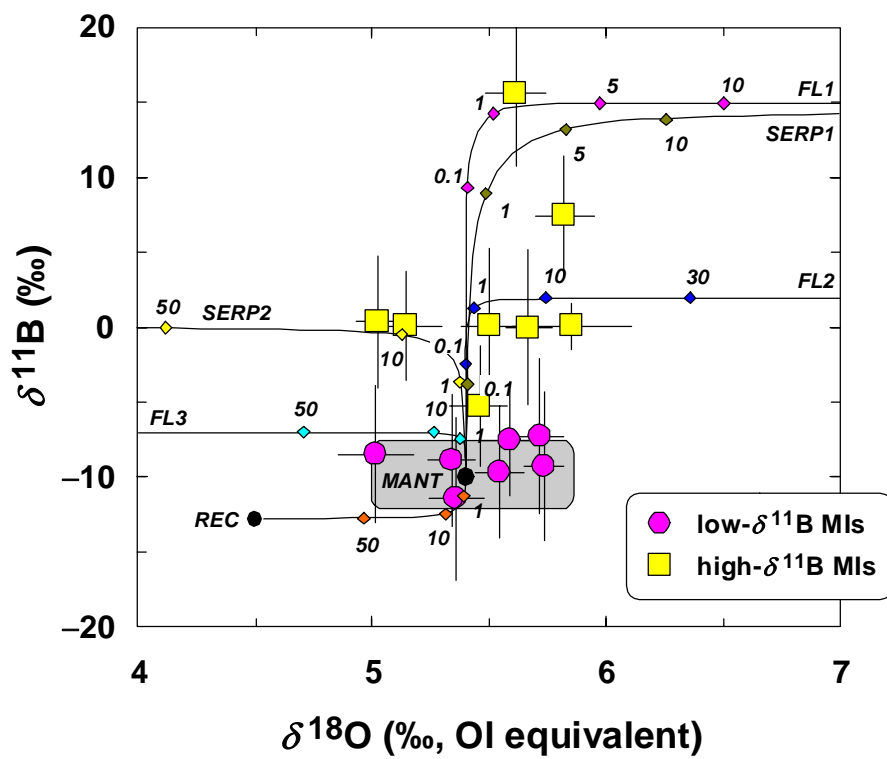


Fig. 6. Gurenko et al.

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