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1 **Multi-criteria correlation of tephra deposits to source centres applied**  
2 **in the Auckland Volcanic Field, New Zealand**

3  
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17  
18 ***Keywords***

19 Basalt; correlation; tephra; Auckland Volcanic Field; monogenetic volcanic field;  
20 tephrochronology;

21 ***Highlights***

- 22 • Outlines a method to correlate proximal whole rock samples with distal tephra  
23 deposits
- 24 • Uses this correlation with new age data to reconstruct an age order for AVF  
25 eruptions
- 26 • Discusses the spatial, temporal, and geochemical evolution of the AVF

27 **Abstract**

28 Linking tephra back to their source centre(s) in volcanic fields is crucial not  
29 only to reconstruct the eruptive history of the volcanic field but also to  
30 understand tephra dispersal patterns and thus the potential hazards posed by a  
31 future eruption. Here we present a multi-disciplinary approach to correlate  
32 distal basaltic tephra deposits from the Auckland Volcanic Field (AVF) to their  
33 source centres using proximal whole-rock geochemical signatures. In order to  
34 achieve these correlations, major and trace element tephra-derived glass  
35 compositions are compared with published and newly obtained whole-rock  
36 geochemical data for the entire field. The results show that incompatible trace  
37 element ratios (e.g.  $(\text{Gd}/\text{Yb})_N$ ,  $(\text{La}/\text{Yb})_N$ ,  $(\text{Zr}/\text{Yb})_N$ ) vary widely across the AVF  
38 (e.g.  $(\text{La}/\text{Yb})_N = 5$  to 40) but show a more restricted range within samples from a  
39 single volcanic centre (e.g.  $(\text{La}/\text{Yb})_N = 5$  to 10). These ratios are also the least  
40 affected by fractional crystallisation and are therefore the most appropriate  
41 geochemical tools for correlation between tephra and whole rock samples.  
42 However, findings for the AVF suggest that each volcanic centre does not have a  
43 unique geochemical signature in the field as a whole, thus preventing  
44 unambiguous correlation of tephra to source centre using geochemistry alone.  
45 A number of additional criteria are therefore combined to further constrain the  
46 source centres of the distal tephra including age, eruption scale, and location (of  
47 centres, and sites where tephra were sampled). The combination of  
48 tephrostratigraphy,  $^{40}\text{Ar}/^{39}\text{Ar}$  dating and morphostratigraphic constraints allow,  
49 for the first time, the relative and absolute ordering of 48 of 53 volcanic centres  
50 of the Auckland Volcanic Field to be resolved. Eruption frequencies are shown to

51 vary between 0.13-1.5 eruptions/kyr and repose periods between individual  
52 eruptions vary from <0.1 to 13 kyr, with 23 of the 48 centres shown to have pre-  
53 eruptive repose periods of <1000 years. No spatial evolutionary trends are  
54 noted, although a relationship between short repose periods and closely spaced  
55 eruption locations is identified for a number of centres. In addition no temporal-  
56 geochemical trends are noted, but a relationship between geochemical signature  
57 and eruption volume is highlighted.  
58

## 59 **Introduction**

60           The eruptive histories of basaltic volcanic fields can be reconstructed by  
61 the dating of lava and scoria deposits. These reconstructions are critical for  
62 understanding the temporal, geochemical and spatial evolution of the fields in  
63 order to better understand their potential future behaviour. However, within  
64 young fields the errors associated with current dating techniques (e.g.  $^{40}\text{Ar}/^{39}\text{Ar}$   
65 or  $^{14}\text{C}$ ) are often larger than the repose periods, and thus hinder establishment of  
66 a definitive stratigraphic age order of the centres (e.g. Briggs et al. 1994; Cook et  
67 al. 2005; Fleck et al. 2014; Leonard et al. 2017). Similarly, due to the restricted  
68 subaerial distribution of scoria and lavas from small monogenetic centres, field-  
69 wide stratigraphic relationships are often difficult to establish, and cannot  
70 resolve ambiguities that arise from the dating techniques. In these circumstances  
71 distal airfall deposits (tephras) can more reliably resolve the chronological  
72 uncertainties due to their higher preservation potential, and often  
73 stratigraphically restricted relationships.

74           Tephra correlation is used on a number of levels from simply correlating  
75 tephra deposit across cores or outcrops (Hopkins et al. 2015), to defining  
76 stratigraphic marker horizons (e.g. Molloy et al. 2009), or matching horizons to  
77 volcanic source or provenance through comparison of distal and proximal tephra  
78 deposit characteristics (e.g. Alloway et al. 2004; Allan et al. 2008; Zawalna-Geer  
79 et al. 2016). Linking tephras to their source volcanic centre can be  
80 straightforward where the potential number of sources is limited, the eruptive  
81 episodes (and tephras) are precisely dated, stratigraphic successions are  
82 established in proximal tephra layering, and/or the tephras (and sources) have

83 distinctive geochemical signatures (Lowe 2011). Where these criteria are not  
84 met, however, difficulties arise in accurately linking distal tephras to their  
85 sources. In cases where there are multiple potential sources and where proximal  
86 deposits are poorly characterised, or poorly preserved, there is currently no  
87 established method to resolve the origin of identified distal tephras.

88         There are a number of processes and features that should be taken into  
89 account when attempting to correlate tephra deposits. The key ones important  
90 for this study are those that can potentially produce differences in the  
91 geochemistry of glass shards in distal tephra horizons. For example, these could  
92 include atmospheric sorting of components during transportation (e.g. Lirer et  
93 al. 1973), or geochemical variation of magma produced during single eruptions  
94 (e.g. Shane et al. 2008), or the presence of micro-inclusions within individual  
95 glass shards (Lowe 2011). In addition, post-eruption processes such as  
96 reworking of deposits can produce repeated sequences (Hopkins et al. 2015),  
97 whereas poor preservation can result in inconsistent deposit thicknesses; both  
98 make the record harder to interpret (e.g. Davies et al. 2001; Pyne O'Donnell  
99 2011). Methodological discrepancies also need to be considered. In general  
100 different sample types and size fractions are not compared (e.g. Larsson 1937),  
101 nor are analyses using different analytical methods. Many of these issues can be  
102 resolved through methodological, statistical or technical practises that we  
103 discuss below. Overall, if distal deposits could be confidently linked to their  
104 source(s), the chronology of a volcanic region could be better resolved.

105         The Auckland Volcanic Field (AVF) is an example of a volcanic region  
106 where climate and urbanization have resulted in the loss or obscuration of  
107 proximal tephra deposits. The spatial density of centres (53 centres distributed

108 over an area of ca. 600 km<sup>2</sup>; **Fig. 1**) adds further complexity because a given  
109 tephra deposit could have come from a number of possible sources (e.g. Shane  
110 and Smith 2000). In addition, because of the rapid thinning of basaltic tephra  
111 away from source, evidence of stratigraphic successions is often limited to well  
112 preserved basinal deposits, for example in the maar crater infillings (e.g.  
113 Hopkins et al. 2015). The tephrostratigraphy of six cores from the maar craters  
114 in the AVF (Pupuke, Onepoto, Orakei Basin, Glover Park, Hopua, Pukaki;  
115 highlighted in red on **Fig. 1**) has been extensively assessed (e.g. Sandiford et al.  
116 2001; Shane and Hoverd 2002; Molloy et al. 2009; Shane et al. 2013; Hopkins et  
117 al. 2015; Zawalna-Geer et al. 2016). The tephrostratigraphic framework  
118 developed by the careful cross correlation of the tephra deposits between  
119 individual cores, and the geochemistry of the tephra-derived glass is used as a  
120 basis for this study (e.g. Molloy et al. 2009; Hopkins et al. 2015).

121 Proximal lava and coarse-grained scoria cone-forming deposits in the AVF  
122 (defined here as whole-rock samples) have a higher preservation potential than  
123 proximal airfall tephra, and therefore the sources of these materials can be more  
124 easily defined (e.g. Hayward et al. 2011). In addition a large number of whole-  
125 rock analyses already exist for the AVF centres, characterising their geochemical  
126 signatures (**Table 1**). Traditional tephrochronology links distal to proximal  
127 tephra deposits, but in the AVF this process is not possible due to the lack of  
128 unambiguously sourced proximal tephra beyond the cones themselves. Here we  
129 therefore develop and present a method for correlating distal tephra (from  
130 cored maar-lake deposits, represented by glass geochemical analyses) to  
131 proximal deposits (represented by whole-rock geochemical analyses of lava or  
132 large fragments), in order to better constrain the relative and absolute eruption

133 history of the AVF. Here we define “tephra” as the bulk airfall deposits of  
134 material explosively erupted from the volcanoes, now found as unconsolidated  
135 pyroclastic horizons within the maar-lake cores (cf. Lowe 2011). Geochemical  
136 analyses for this study were undertaken on the juvenile glass shards derived  
137 from within these tephra horizons. The term “whole rock” is used here to refer  
138 to analyses of individual pieces of solid rock, from lava flows or from individual  
139 bombs or lapilli.

140

## 141 **Methodology**

142 To provide the most complete basis for tephra-to-source correlations a  
143 critical requirement is an extensive database of characteristics for all volcanic  
144 centres and tephra deposits in the field. For the AVF a large dataset already  
145 exists, including geochemistry of proximal whole-rock samples (e.g. McGee et al.  
146 2013) and geochemistry of distal tephra-derived glass samples (e.g. Hopkins et  
147 al. 2015), ages of eruptive centres (e.g. Leonard et al. 2017), and scale of  
148 eruptions (e.g. Kereszturi et al. 2013). Currently lacking, however, is a collated  
149 field-wide suite of geochemical data of whole-rock compositions, up-to-date  
150 estimates of the ages of the tephra horizons in the maar-lake cores, and  
151 estimates of tephra volumes for the individual centres. Below we present the  
152 methods by which these pre-existing data were collated, and our new data  
153 collected.



154 ***Collation of pre-existing data***

155 *Whole-rock geochemistry for individual centres*

156 A large amount of unpublished whole-rock geochemical data exists for the  
157 AVF. This includes datasets from MSc theses (Bryner 1991; Miller 1996; Franklin  
158 1999; Hookway 2000; Spargo 2007; Eade 2009; see **Table 1**), and the  
159 unpublished data of I.E.M. Smith and co-workers at the University of Auckland.  
160 We also include here data from McGee (2012), the majority of which is published  
161 in McGee et al. (2011, 2012, 2013). For the newly discovered centres of Puhinui  
162 Craters and Cemetery Hill (B. Hayward *pers. comm.*), no geochemical or age data  
163 exist and therefore these centres are not included in this study. The collated  
164 whole-rock major and trace-element dataset can be found in the **supplementary**  
165 **material**.

166 *Glass geochemistry for individual tephra horizons*

167 Hopkins et al. (2015) analysed major and trace element geochemistry for  
168 glass shards from tephra horizons found in the lacustrine maar cores using  
169 Electron Microprobe Analysis (EMPA) and Laser Ablation-ICP-MS (LA-ICP-MS) at  
170 Victoria University of Wellington (VUW). Glass shards from only forty-nine  
171 basaltic horizons from five maar cores could be analysed for trace element  
172 concentrations because glass shard sizes were too small or the samples were no  
173 longer available. These data are combined with previously published major  
174 element data (Sandiford et al. 2001; Shane and Hoverd 2002; Hoverd et al. 2005;  
175 Molloy et al. 2009) reported in Hopkins et al. (2015) and outlined in the  
176 **supplementary material**.

177 *Compatibility of pre-existing and new data*

178           To ensure compatibility between the data sets, and as a quality control  
179 measure, we assessed the analytical methods, accuracy and precision of all data  
180 used in this contribution. For all pre-existing whole-rock analyses (outlined  
181 above), the methods and standardisation procedures were the same. XRF  
182 analyses for major elements were undertaken at the University of Auckland  
183 (UoA), and (where applicable) trace elements were analysed using laser ablation  
184 (LA)-ICP-MS on the XRF glass discs at the Australian National University (ANU).  
185 For XRF methods in-house rock standards were used (see **Supplementary**  
186 **Material**), and the Si concentrations obtained from XRF analysis were used for  
187 the trace element calibration. In addition duplicate analyses were undertaken by  
188 this study to ensure compatibility of the old and new data sets (see  
189 **Supplementary Material**).

190           For tephra-derived glass chemistry all sample preparation followed the  
191 same standard procedures. Major-element geochemistry presented in Sandiford  
192 et al (2001) was acquired at VUW on an older instrument than that used by  
193 Hopkins et al (2015); both of these studies however used wavelength dispersive  
194 X-ray spectroscopy (WDS) techniques. Data presented in Molloy et al (2009),  
195 Shane and Hoverd (2002) and Hoverd et al (2005) were obtained by EMPA at  
196 University of Auckland (UoA), using energy dispersive X-ray spectroscopy (EDS)  
197 techniques. No previous trace-element analysis had been undertaken on these  
198 samples prior to work by Hopkins et al (2015). Accuracy and precision of these  
199 methods is detailed in the **Supplementary Material**. Duplicate analyses from  
200 the same horizons, and from the same shards, were run in order to compare the  
201 newly acquired data with the existing data sets (example reported in

202 **Supplementary Material**). All aspects of these methods for both glass and  
203 whole-rock analyses, and the accuracy and precision reported for the standards  
204 are comparable to the methods used by this study.

#### 205 *Ages for individual centres*

206 To maximise the amount of available age data from individual eruptive  
207 centres, data from three methods have been collated. These methods include  
208 morphostratigraphic evidence (e.g. Hayward et al. 2011),  $^{40}\text{Ar}/^{39}\text{Ar}$  dating of  
209 groundmass material (e.g. Cassata et al. 2008; Leonard et al. 2017), and  $^{14}\text{C}$   
210 dating of organic materials contained within or bounding the volcanic deposits  
211 (compiled in Lindsay et al. 2011). These are detailed in **Table 2**. Modelled ages  
212 for the AVF centres suggested by Bebbington and Cronin (2011) are excluded  
213 from this study, as they are based on tephra horizon ages given by Molloy et al.  
214 (2009), which are superseded by those in Lowe et al. 2013 (for rhyolitic tephra  
215 ages) and Hopkins et al. 2015 (for basaltic tephra horizon thicknesses and  
216 depths).

217 Morphostratigraphy is here defined as the inter-relationships exhibited by  
218 the surface landforms, for example where tephra or lava deposits from one  
219 centre overlie another. Due to the proximity of the centres to one another within  
220 the field (cf. **Fig. 1**), 35 of 53 centres have morphostratigraphic constraints  
221 associated with them (outlined in **Table 2**). These morphostratigraphic  
222 constraints give optimum relative ages, which need to be combined with the  
223 absolute ages derived from  $^{40}\text{Ar}/^{39}\text{Ar}$  or  $^{14}\text{C}$  dating. In all cases the  
224 morphostratigraphic constraints are consistent with the absolute radiometric  
225 age ranges.

226 The  $^{40}\text{Ar}/^{39}\text{Ar}$  ages presented in Leonard et al. (2017) are here given as age  
227 ranges (the 2sd error on the age, reported in **Table 2**). This is because any age  
228 within the range is considered appropriate for the centre, with no extra  
229 emphasis given to the mean ages. For the 20 centres with no  $^{40}\text{Ar}/^{39}\text{Ar}$  or  $^{14}\text{C}$   
230 ages, the relative ages of 14 centres were derived by morphostratigraphy (see  
231 **Table 2**). For the remaining six centres (Otuataua, Pigeon Mt., Robertson Hill,  
232 Boggust Park, Cemetery Hill, and Puhinui Craters) no radiometric ages or  
233 morphostratigraphic relationships are evident. As previously mentioned  
234 Cemetery Hill and Puhinui Craters are not considered in this study, and therefore  
235 Otuataua, Pigeon Mt., Robertson Hill, and Boggust Park are still included as  
236 possible correlatives for any dated horizon during the correlation process.

237

### 238 ***New data acquisition***

#### 239 *Geochemical whole rock data*

240 Prior to this study, 28 of the 53 AVF centres had three or more pre-existing  
241 major and trace element analyses, fifteen centres had less than three, and ten  
242 had no geochemical data at all (see **Table 1**). Volcanic centres with less than  
243 three existing whole rock analyses were targeted in this study. Seventeen centres  
244 had sufficient exposure to be sampled including: Boggust Park, Little Rangitoto,  
245 Mt Albert, Mt Cambria, Mt Hobson, Mt Roskill, Mt Smart, Onepoto, Otuataua,  
246 Pigeon Mt, Pukaki, Pukeiti, Pupuke, Mt Robertson, St Heliers, Taylors Hill and Te  
247 Pou Hawaiki (**Fig. 1**). For an additional seven centres major element data existed  
248 (Miller 1996), but no trace element data were reported. Thus, for these seven  
249 centres (**Fig. 1**; Green Mt, Hampton Park, Mangere Mt, McLaughlins Mt,

250 Mclennan Hills, Mt Victoria, and Otara), samples collected by Miller (1996) were  
251 re-analysed for both major and trace elements by this study. For six centres (Ash  
252 Hill, Kohuora, Mangere Lagoon, Styaks Swamp, Cemetery Hill, and Tank Farm;  
253 **Fig. 1**), there are currently no exposures suitable for sampling (due to  
254 urbanisation and erosion), and therefore, no geochemical data exists.

255 Whole rock samples were crushed to <15 mm in a Rocklabs Boyd crusher,  
256 then powdered using a Rocklabs tungsten-carbide TEMA swing mill at VUW.  
257 Powders were made into fused lithium metaborate glass discs and analysed for  
258 major element oxide concentrations at the Open University, Milton Keynes, UK  
259 using X-ray Fluorescence (XRF) analysis following the methods of Ramsey et al.  
260 (1995). Internal standards WS-E (Whin Sill Dolerite) and OU-3 (Nanhoron  
261 microgranite) were analysed to monitor precision and accuracy. Major element  
262 oxides were accurate to within 2.0% of the recommended values for the internal  
263 standards and analytical precision ( $2\sigma$ ) was 1.5% or better for all elements.

264 For trace element analysis, 50 mg of whole rock powder was treated using  
265 conventional methods of HF-HNO<sub>3</sub> digestion, and analysed on an Agilent 7500CS  
266 ICP-MS (VUW) in solution mode. Trace element abundances were calculated  
267 using the reduction program Iolite (Paton et al. 2011), using BHVO-2 as a  
268 bracketing standard, and BCR-2 as a secondary standard. <sup>43</sup>Ca was used as an  
269 internal standard using CaO contents measured by XRF. Trace element analyses  
270 were accurate to within <6% of the recommended values for the secondary  
271 standard (BCR-2) and precision ( $2\sigma$ ) was <6.5 % with the exceptions of Cr ±10.4  
272 %, Nb ±22 %, Cs ±12.2 %, Ba ±11.8 %, Ta ±20.9 % and Pb ±31 %.

273

274 *Tephra horizon ages*

275        Within the Auckland maar cores as well as the locally derived basaltic tephra  
276 horizons, there are also distal andesitic and rhyolitic tephra deposits from  
277 various other sources within North Island (**Fig. 1B**). These “foreign” tephra can  
278 be used as stratigraphic marker horizons to aid both the absolute and relative  
279 dating of the basaltic deposits. The ages of the basaltic horizons within the cores  
280 are modelled by interpolating ages as a function of deposit depth, with the mean  
281 time interval per millimetre of core (**Fig. 2**). This principle assumes that tephra  
282 represent instantaneous events (Shane 2005), and therefore, their thicknesses  
283 are subtracted from the total sediment thickness. We use the most recent  
284 published ages for the rhyolitic marker horizons (RMHs; e.g. Lowe et al. 2013),  
285 and couple them with the most recent published thicknesses for the basaltic,  
286 andesitic and rhyolitic deposits in the maar cores. For basaltic deposits at Orakei  
287 and Glover Park we use data from Hopkins et al. (2015), and for the Onepoto  
288 core, all tephra thicknesses and depths are adapted from Shane and Hoverd  
289 (2002). Rhyolitic and andesitic deposit thicknesses at Orakei, Hopua, Pupuke,  
290 and lower Pukaki cores (below the Kawakawa/Oruanui RMH (Kk)) are from  
291 Molloy (2008) and in the upper Pukaki core (above Kk) from Sandiford et al.  
292 (2001).

293        Ages and uncertainties for all deposits found above the Maketu RMH are  
294 obtained by Monte Carlo simulation as follows. One thousand simulated sets of  
295 measured ages were found by adding the age’s Gaussian noise with the standard  
296 deviations of the determined ages. Any resulting set of ages out of stratigraphic  
297 order were rejected, that is, the 1000 simulations were conditional on the ages  
298 produced being in decreasing order. The simulations were then used to produce

299 1000 sets of interpolations with the lower 5 and upper 95 percentiles of the  
300 distribution giving the interpolated age uncertainties in  $2\sigma$  (see **Table 3**).

301 Sedimentation rate calculations are used to estimate the ages of the  
302 basaltic deposits found below the Rotoehu RMH (AVF3 to AVFc; no basaltic  
303 deposits are found between Maketu and Rotoehu RMHs). The age of the Rotoehu  
304 RMH itself is currently contentious, with published estimates ranging from ca. 40  
305 to ca. 70 ka, associated with a range of different dating techniques (e.g. Lowe and  
306 Hogg 1995; Lian and Shane 2000; Charlier et al. 2003; Wilson et al. 2007; Danišík  
307 et al. 2012; Flude and Storey 2016). Here, we use an age estimate of  $52 \pm 7$  ka  
308 (D.J. Lowe *pers. comm.*), in order to accommodate the most likely range. In  
309 addition, because there are no dated RMHs below the Rotoehu, these calculations  
310 often assume constant sedimentation rates for a large proportion of the cores,  
311 which is probably unrealistic, and thus they are taken as a guide only (**Table 3**).

312 The basaltic deposit AVFd, was used as a lower constraint for the  
313 sedimentation rate between the Rotoehu and the base of the Onepoto core. This  
314 deposit contains lava and scoriaceous blocks interpreted to represent the  
315 Onepoto maar crater floor (Shane and Hoverd 2002). Although no age exists  
316 from the Onepoto eruption, morphostratigraphy suggests that it is just younger  
317 than Pupuke (Hayward et al. 2011), and we therefore use the mean age  
318 measured for Pupuke ( $193.2 \pm 2.8$  ka by  $^{40}\text{Ar}/^{39}\text{Ar}$  dating: Leonard et al. 2017) as  
319 a maximum age for the eruption of Onepoto. The respective calculated  
320 sedimentation rate of 0.19 mm/yr is comparable to those recorded previously  
321 for younger core sections (0.18 mm/yr: Shane and Hoverd 2002). In addition,  
322 the calculated basaltic tephra horizon ages are comparable to those calculated

323 for the correlated horizons AVF2 and AVF1 in the Orakei Basin core, suggesting  
 324 that the assumptions made to calculate these values are realistic (**Table 3**).

325 In the Glover Park core, for the horizons correlated to other cores (AVF2 and  
 326 AVF1), ages are assigned from an average of the values calculated from these  
 327 core deposits. For horizon AVFa, which is only found at Glover Park, an age  
 328 estimate was obtained through calculating the sedimentation rate between the  
 329 bounding basaltic horizons, AVF1 and AVFb. The ages for these horizons were  
 330 assigned based on the ages calculated for these deposits in Orakei Basin (AVF1)  
 331 and Onepoto cores (AVF1 and AVFb). Calculated ages based on sedimentation  
 332 rate for all basaltic tephra horizons and their associated errors are outlined in  
 333 **Table 3**.

334 *Estimated tephra volumes*

335 Previous studies have estimated total eruptive volumes for the centres of  
 336 the AVF (Allen and Smith 1994; Kereszturi et al. 2013) although, distal tephra  
 337 volumes were not reported due to limited measurable material. Other studies  
 338 (e.g. Kawabata et al. 2015) suggest that tephra volumes for small-scale eruptions  
 339 can be estimated from the volumes of the tuff and scoria cones using the

340 following equation:

341 
$$V_{DRE} = 0.5V_{Tuff} + 1.5V_{Scoria}$$

342 where V is volume, and DRE is dense rock equivalent values (where volumes are  
 343 corrected for void spaces, detailed in Kereszturi et al. 2013). In order to estimate  
 344 tephra volume we use the most recently published DRE values for tuff and scoria  
 345 from Kereszturi et al. (2013). Volume estimates are detailed in **Table 2**.



## 346 **Results**

### 347 ***Whole-rock and glass geochemistry***

#### 348 *Whole-rock geochemistry*

349           Following the rock classification of LeMaitre et al. (2002), the AVF  
350 samples range from basanitic/nephelinitic to basaltic in composition (e.g. SiO<sub>2</sub> =  
351 39-49 wt.%; Mg# = 50-72. Broad positive trends exist between wt.% MgO and  
352 wt.% CaO, and wt.% MgO and wt.% Al<sub>2</sub>O<sub>3</sub>. Although less obvious, there are  
353 discernable broad negative trends exhibited in the AVF data between wt.% MgO  
354 and wt.% SiO<sub>2</sub>, TiO<sub>2</sub>, Fe<sub>2</sub>O<sub>3</sub><sup>tot</sup> and P<sub>2</sub>O<sub>5</sub> (not shown). These elements are more  
355 variable within a single centre than are MgO vs. CaO or Al<sub>2</sub>O<sub>3</sub>. For example the  
356 eruptive products of Motukorea show an almost flat trend for wt.% MgO vs. wt.%  
357 TiO<sub>2</sub>, whereas the Crater Hill samples show a strong positive trend. Although all  
358 samples from the AVF seem to follow the overall major element trends on  
359 variation diagrams, samples from individual AVF centres can define separate  
360 trends (c.f. **Fig. 3**) within this, as previously described by McGee et al. (2013).

361           Trace-element contents in the AVF samples vary substantially, for  
362 example, La 10-90 ppm, Nb 10-80 ppm and Sr 300-1000 ppm (see  
363 **Supplementary Material**). Similar to the major elements, some of the trace  
364 elements show overall general trends for the field, as well as trends specific to  
365 each centre (**Fig. 3**). There is a strong positive trend for wt.% MgO and ppm Cr  
366 and Sc, and a general negative trend of variable slope exists between wt.% MgO  
367 and ppm Th, Nb, Sr, and La (**Fig. 3**).

368           Mantle-normalised trace-element data for near primitive AVF samples  
369 (e.g. Mg# ≥ 60) are broadly similar and are characterised by a positive Nb

370 anomaly and a negative sloping light to heavy rare earth element profile (e.g.  
371 La/Yb range 4 to 40; **Fig. 4**), characteristics that are similar to ocean island  
372 basalts (OIBs). Some centres (e.g. Rangitoto 2 and Te Pou Hawaiiiki) have  
373 geochemical signatures that are less enriched in trace elements than others,  
374 characterised by a shallower rare earth element (REE) pattern gradient (e.g.  
375 La/Yb  $\leq 7.5$ ), and a positive Sr anomaly (e.g. Sr\*  $\geq 1.2$ ). In contrast, samples from  
376 trace element-enriched centres (e.g. Mt Cambria, Mt Hobson, St Heliers) have a  
377 relatively steep REE pattern gradient (e.g. La/Yb  $\geq 20$ ), show a small trough at  
378 Zr-Hf, exhibit no Sr anomaly (e.g. Sr\*  $\leq 1.0$ ), and display a negative K anomaly  
379 (e.g. K\*  $\leq 0.7$ ; **Fig. 4**). These major and trace element signatures for the field are  
380 discussed in detail by McGee et al. (2013), and are attributed to mixing during  
381 ascent of magma from three mantle sources.

#### 382 *Glass geochemistry*

383 The geochemical composition of glass shards found in the AVF tephras  
384 are discussed in detail in Hopkins et al. (2015; see Fig 4 therein). In general they  
385 show a consistent range in MgO (ca. 2 to 7.5 wt.%), CaO (ca. 7 to 15 wt.%), FeO  
386 (ca. 9 to 15 wt.%), K<sub>2</sub>O (ca. 1 to 4 wt %), and TiO<sub>2</sub> (ca. 2 to 4.5 wt.%) between  
387 samples from across all cores. Al<sub>2</sub>O<sub>3</sub> concentrations are shown to be consistently  
388 lower at given MgO values in the Orakei and Onepoto cores, and SiO<sub>2</sub> is  
389 consistently lower at given MgO values in the Onepoto core. Glass shards from  
390 individual horizons have mostly similar major element concentrations with  
391 variations within <1 wt. % for MgO, SiO<sub>2</sub>, FeO, and TiO<sub>2</sub>, and <3 wt. % for CaO,  
392 Al<sub>2</sub>O<sub>3</sub>, Na<sub>2</sub>O, and K<sub>2</sub>O, with minor numbers of horizons showing bimodal or

393 systematic ranges in concentrations of major elements (as discussed in Hopkins  
394 et al. 2015).

395 In addition to major oxides, Hopkins et al. (2015) analysed trace elements  
396 on individual  $\geq 30$   $\mu\text{m}$  diameter glass shards. Their results showed (similar to  
397 whole-rock analyses) high variability in concentrations for trace elements, for  
398 example La ca. 5-100 ppm, Nb ca. 20-175 ppm, and Sr 140-1500 ppm. In general,  
399 glass shard primitive-mantle normalised multi-element plots show comparable  
400 signatures to the whole rock geochemical patterns (**Fig. 4**). Glass shards from  
401 individual tephra horizons have a more limited range in trace-element  
402 concentrations when compared to the whole field, and in many cases show  
403 relatively distinct trace element patterns for each individual tephra horizon (**Fig.**  
404 **4**).

405

#### 406 ***Tephra horizon ages***

407 Age estimates for all tephra horizons used in this study are outlined in  
408 **Table 3** and summarised in **Figure 2**. Basaltic tephra horizons found within 6  
409 cores span a large age range in the field from 0.54 to ca. 143 ka (AVF24 in  
410 Pupuke core and AVFc in Onepoto core respectively). Fourteen horizons have  
411 ages calculated at  $< 28$  ka, nine horizons are found between ca. 28 and 35 ka, and  
412 only 6 horizons have ages of ca. 59-143 ka. Overall the estimated ages are in  
413 good agreement where multiple deposits are correlated across cores (**Fig. 2**).  
414 Two discrepancies, however, arise (highlighted in **Table 3**): 1) The calculated  
415 age for AVF17 appears too young within the AVF number sequence, and 2) the  
416 calculated age of AVF16 appears too old for the AVF number sequence and

417 suspiciously similar to the age of AVF13. These results are potentially  
418 problematic, and are therefore discussed below.

419         The age of AVF17 when estimated using only the Orakei Basin core (23.35  
420 ka), rather than averaging all ages across the cores, is not chronologically out of  
421 place (e.g. AVF18 is 23.2 ka and AVF15 is 24.5 ka). However, using the average  
422 age for AVF18, which is calculated as the average of correlated units from  
423 multiple cores (deposits from within Hopua 25.2 ka, Pukaki 24.6 ka and Orakei  
424 23.35 ka cores) it appears too old (**Table 3**). This is because the ages for the  
425 deposits in the Pukaki and Hopua core are slightly older than those estimated  
426 for just the Orakei Basin core. But, within this section (Okareka to Te Rere), all  
427 of the horizon ages calculated are within error of each other, and therefore  
428 stratigraphic constraints in the cores are required to resolve the absolute  
429 ordering. AVF19 is found above the andesitic horizon Eg36 (**Fig. 2**; Molloy  
430 2008), which is found in all the cores, and therefore acts as a marker horizon to  
431 place AVF19 as the youngest horizon. AVF18 is found above AVF17 within the  
432 Orakei Basin core, further restricting the ordering of these two horizons. The  
433 ordering and correlation of these horizons will therefore be maintained,  
434 however, the errors on the ages must be taken into account during the  
435 correlation process.

436         The ages calculated for AVF16 (Pukaki core only) and AVF13 (Orakei core  
437 only) are identical ( $25.23 \pm 0.86$  ka and  $25.23 \pm 0.31$  ka respectively). The age  
438 estimate for AVF16 implies that it is older than suggested by the original  
439 position in the AVF nomenclature sequence, and there is a strong possibility that  
440 the horizons represent the same deposit. Stratigraphically, there are limited  
441 constraints on the relationship of AVF16 with the other deposits from other

442 cores. The andesitic deposit Eg34 is found below AVF16 but is not found in any  
443 other cores and therefore provides no further regional stratigraphic constraints.  
444 The Te Rere and the Kawakawa/Oruanui RHMs stratigraphically constrain  
445 horizon AVF16 (above and below respectively), but there are no other age  
446 constraints (Te Rere tephra is not found in the Orakei Basin core). In addition  
447 there are limited geochemical data for the deposit AVF16 to confirm or deny its  
448 relationship with AVF13 (Sandiford et al. 2001; Hopkins et al. 2015). Therefore  
449 due to the lack of distinct evidence to suggest these deposits are not the same,  
450 and the overwhelming similarity in the ages, we assume AVF16 and AVF13  
451 record the same event and will be referred to as 'AVF13' with an age of  $25.23 \pm$   
452  $0.86$  ka in the following discussion.

453

## 454 **Discussion**

### 455 *Discriminatory geochemical elements for the AVF*

456 Previous studies on the petrogenesis of AVF eruptive products have  
457 shown that each magma batch feeding a single centre is generated by mixing of  
458 contributions from differing degrees of partial melting of multiple mantle  
459 sources at different depths (Huang et al. 1997; McGee et al. 2013, 2015; Hopkins  
460 et al. 2016). The resulting geochemical signatures of the erupted volcanic  
461 products demonstrate that although there is overlap for many elements,  
462 combinations of some major element ( $\text{SiO}_2$ ,  $\text{MgO}$ ,  $\text{CaO}$ ,  $\text{FeO}$ ,  $\text{P}_2\text{O}_5$ ) and trace  
463 element (Sc, Sr, Zr, Gd, La, Sm, Nd, Nb, Ce) concentrations or ratios (e.g.  $(\text{La}/\text{Yb})_N$   
464 or  $(\text{La}/\text{Y})_N$ ) can be used to discriminate single trends for individual centres (**Fig.**

465 3). The selected elements also show the widest range in concentrations in  
466 eruptive products from the AVF.

467 The rare-earth elements (REEs) are especially useful because fractional  
468 crystallisation of the common silicate phases has only a minor effect on their  
469 concentrations. They can therefore be used to discriminate between melts from a  
470 deep (garnet-bearing mantle = high light REE/heavy REE) or shallow (spinel-  
471 bearing mantle = low light REE/heavy REE) source (e.g. McKenzie and O’Nions  
472 1991; Robinson and Wood 1998; McGee et al. 2013, 2015; Hopkins et al. 2016;  
473 McGee and Smith 2016). As a result of these variations, and of the discriminatory  
474 nature of certain elements and element ratios within the AVF, we show that  
475 geochemical fingerprinting can be used as a method to correlate distal tephra  
476 deposits to their source centre. Below we discuss the techniques by which this  
477 method was tested and developed.

478

### 479 *Geochemical correlation*

480 A key issue in correlating the geochemistry of glass shards in distal tephra  
481 to whole-rock geochemistry of proximal lavas and pyroclastic particles is that  
482 most whole-rock samples contain mineral inclusions (e.g. olivine), whereas small  
483 volcanic glass shards (in tephra) do not. Hence, the concentration of elements  
484 that strongly partition into mineral phases (e.g. Mg, Ni or Cr into olivine) in  
485 whole-rock samples will not be comparable to the respective element contents in  
486 the glass shards (e.g. **Fig. 5A**). Conversely, elements that preferentially remain in  
487 the melt (e.g. those that are incompatible with mineral phases commonly found  
488 in alkali basalts, such as the REE) are likely to have comparable concentrations in  
489 whole-rock and glass shards. In addition, mineral-free groundmass glass from

490 whole-rock samples is likely to have a comparable geochemical signature to the  
491 glass shards forming distal tephra deposits (e.g. Lowe 2011; Allan et al. 2008;  
492 Lowe and Alloway 2015; **Fig. 5B**).

493         These hypotheses were tested initially on samples from a known source by  
494 comparing the geochemical composition of (a) a proximal whole-rock sample  
495 and (b) the matrix-derived glass from that sample to (c) glass shards from a  
496 distal tephra deposit. The whole-rock lava sample Mt. Wellington AU62394 was  
497 chosen for two reasons, 1) it has a fresh, glassy groundmass and, 2) distal tephra  
498 from Mt. Wellington has been unambiguously identified in the Hopua core based  
499 on age and thickness (Molloy et al. 2009). The lava sample was processed first as  
500 a whole-rock sample (XRF and ICP-MS, see methodology). It was also processed  
501 to produce a 'matrix-derived glass' sample by crushing the rock and separating  
502 shards of matrix glass that were of comparable size (30-100  $\mu\text{m}$ ) to the glass  
503 shards found in the tephra horizon from the Hopua core (Molloy et al. 2009).  
504 These separated matrix-derived glass shards were then analysed by EMPA and  
505 LA-ICP-MS using methods outlined in Hopkins et al. (2015).

506 *Geochemical correlation of glass shards from distal tephra deposits with matrix*  
507 *derived glass*

508         **Figure 6** shows MgO vs. Al<sub>2</sub>O<sub>3</sub> (in wt. %. [**Fig. 6A**]) and Gd vs. Zr (in ppm  
509 [**Fig. 6B**]) for matrix-derived glass and the glass from its known distal correlative  
510 from the Hopua core, the overlap in the data demonstrates that their  
511 compositions are comparable. This is the case for a wide range of both major and  
512 trace elements (including, MgO vs. full major element suite plus trace elements  
513 Rb, Zr, Cs, Ni, Cr, Y, and Er; SiO<sub>2</sub> vs. Al<sub>2</sub>O<sub>3</sub>, Na<sub>2</sub>O, K<sub>2</sub>O, and CaO vs. Al<sub>2</sub>O<sub>3</sub>, Na<sub>2</sub>O).

514 Limited variability exists between trace elements (e.g. Rb, Zr, Ni, Cr and Y, and  
515 the REE) when plotted against each other, or against Al<sub>2</sub>O<sub>3</sub> or MgO.

516 For some elements, however, the glass from the distal tephra has larger  
517 variations than does the matrix-derived glass (**Fig. 6A**). This is attributed to  
518 either 1), the matrix-derived glass being made from a single clast and thus  
519 having minimal compositional variation, and/or 2) glass shards from the distal  
520 tephra showing a higher variability due to initial differences in composition of  
521 the erupted magma creating variability in the glass shard composition  
522 throughout the eruption (e.g. McGee et al. 2012). This test proves that matrix-  
523 derived glass from proximal samples can be successfully correlated with glass  
524 shards in distal tephra using trace elements and trace element ratios (**Fig. 6B**).

525 Geochemical analysis using EMPA and LA-ICP-MS techniques are for  
526 individual glass shards, ensuring phenocrysts and microlites are not analysed.  
527 Accordingly matrix-derived glass from proximal samples can be correlated with  
528 glass shards from within distal tephra deposits using both elements that are  
529 highly compatible and elements that are incompatible. Compatible elements are  
530 preferentially incorporated in key crystallising minerals within the whole rock  
531 (e.g. olivine) and therefore result in comparable glass chemistries between  
532 matrix-derived glass and tephra-derived glass. The incompatible trace elements  
533 can also be used because they are not preferentially taken into the crystal  
534 phases. We therefore conclude that matrix-derived glass from whole-rock  
535 samples can be correlated to glass shards from the distal tephra deposits, with  
536 some minor caveats. For example, this method relies on the existence and ability  
537 to extract glass from the groundmass of proximal whole-rock samples, which is  
538 not always possible.



539 *Correlation of glass shards from distal tephra with whole-rock samples*

540 In general, when the entire suite of whole-rock and glass geochemical  
541 datasets are compared, MgO, Cr, and Ni all show distinctly higher concentrations  
542 in whole-rock samples than in the glasses (e.g. MgO in whole rock range from ca.  
543 6-16 wt.%; in glass ca. 2-6 wt.%: **Fig. 5A**). Compared to whole-rock analyses, all  
544 glasses contain higher (but slightly overlapping) wt.% SiO<sub>2</sub>, Al<sub>2</sub>O<sub>3</sub>, Na<sub>2</sub>O, and K<sub>2</sub>O  
545 contents (e.g. SiO<sub>2</sub> in whole rock ca. 38-50 wt.%; glass ca. 42-52 wt.%). CaO, FeO,  
546 TiO<sub>2</sub>, and P<sub>2</sub>O<sub>5</sub> have comparable ranges between whole rock and glass, as do the  
547 trace elements, including REEs (**Fig. 5B**). The REEs in general do show  
548 comparable but slightly wider ranges in concentrations in the glass than in the  
549 whole rock (e.g. Sr in glass = 140-1500 ppm vs. Sr in whole-rock = 300-1000).

550 In addition to the presence of phenocrystic material combined into a bulk  
551 rock analysis, correlating major-element compositions of proximal whole-rock  
552 samples to those of glass shards in distal tephra has proved difficult, due to the  
553 effect that fractional crystallization has on the concentrations of some elements  
554 (e.g. Pearce et al. 2008; Ukestins Peate et al. 2008; Dunbar and Kurbatov 2011;  
555 Óladóttir et al. 2012). Plotting element concentrations (for whole-rock samples  
556 from a single centre or glass shard analyses from one tephra horizon) against  
557 other elements that are compatible with certain crystals (e.g. MgO for olivine,  
558 CaO and Al<sub>2</sub>O<sub>3</sub> for pyroxene or plagioclase) can be used to monitor the effect of  
559 crystal removal on these elements in the glass. If an element shows a positive or  
560 negative correlation ( $r^2 \geq 0.6$ , where no single point is responsible for the trend),  
561 with key compatible major elements (MgO, CaO, Al<sub>2</sub>O<sub>3</sub>) then that element is  
562 significantly affected by crystal removal and therefore not useful for correlation  
563 purposes. In addition to key major elements, trace elements with high partition

564 coefficients for olivine and pyroxene (e.g. Ni, Cr, Sc) are also affected. For  
565 example, **Fig. 7** shows that for MgO vs. Ni, the whole rock  $r^2 = 0.75$ , and for  
566 tephra-derived glass  $r^2 = 0.61$ . Conversely, high field strength elements (HFSE),  
567 such as Nb, Zr, and REE, show no trend with elements tracing fractional  
568 crystallisation (e.g. for MgO vs. La;  $r^2 = 0.02$  for tephra-derived glass, and  $r^2 =$   
569  $0.11$  for whole rock. This exercise discussed above was repeated for all glass-  
570 shard analyses from all tephra horizons and for all whole-rock samples from all  
571 centres using MgO, CaO, Al<sub>2</sub>O<sub>3</sub>, Ni, Mn, and Sr on the x-axis (and all other major  
572 and trace elements on the y-axis). These results suggest that HFSEs are  
573 incompatible with major crystallising phases and are therefore well suited for  
574 geochemical fingerprinting (e.g. **Fig. 6E-F**; **Fig. 7**). Respective trace element  
575 ratios (e.g. (La/Yb)<sub>N</sub>, (Gd/Yb)<sub>N</sub>, (Zr/Yb)<sub>N</sub>, (Ce/Yb)<sub>N</sub>, (Nb/Yb)<sub>N</sub>, and (Nd/Yb)<sub>N</sub>) also  
576 showed no correlation with any of the x-axis elements. Therefore, these ratios  
577 are considered best for geochemical correlation between glass shards and whole  
578 rocks. Such ratios show a broad range in the AVF as a whole, but a relatively  
579 restricted range in samples from each single centre, and no relationship with  
580 indices of fractional crystallisation.

581         When applied to the known Mt Wellington samples, a comparison of  
582 proximal whole rock, matrix-derived glass (of the same whole rock sample), and  
583 distal tephra-derived glass show the expected results. **Figure 6C** shows an  
584 example of element combinations that are comparable for glass-glass  
585 correlations but not for glass-whole rock correlations (e.g. MgO vs. Al<sub>2</sub>O<sub>3</sub>, K<sub>2</sub>O, Ni,  
586 Cr, and the REE). In contrast, some major element combinations do appear to  
587 correlate the whole-rock with glass of the distal tephra (**Fig. 6D**; including SiO<sub>2</sub>  
588 vs. TiO<sub>2</sub> and FeO, and CaO vs. TiO<sub>2</sub>, FeO and Al<sub>2</sub>O<sub>3</sub>). In these cases, however, the

589 strong correlation is mainly due to the small variability observed in the Mt  
590 Wellington samples; it may not be applicable for other centres within the AVF.  
591 **Figure 6E** illustrates an example of incompatible trace elements in glasses that  
592 show slightly more variability than the whole-rock samples do; this discrepancy  
593 is, however, minimised when trace element ratios for the two sample types are  
594 compared (see **Fig. 6F**). The incompatible trace element ratios are sufficiently  
595 distinctive to allow independent correlation to be made between the field-wide  
596 suite of proximal whole-rock and distal glass data, especially  $(La/Yb)_N$ ,  $(Gd/Yb)_N$ ,  
597 and  $(Zr/Yb)_N$ , all of which show a wide range of values in the field as a whole. It  
598 is therefore concluded that by using incompatible-element and LREE/HREE  
599 ratios, it is possible to geochemically correlate individual glass shards from distal  
600 tephra deposits with proximal whole-rock samples. There are, however, some  
601 additional limitations for the AVF.

#### 602 *Limitations on geochemical correlations*

603 Previous studies have demonstrated that the geochemical composition of  
604 the erupted products within some of the AVF centres (e.g. Crater Hill: Smith et al.  
605 2008; Motukorea: McGee et al. 2012), change as the eruptions progress from  
606 initially phreatomagmatic to magmatic eruption styles (**Table 2**). These centres  
607 consistently show, for example, initially low wt.% SiO<sub>2</sub> and Mg/Fe ratios and  
608 higher incompatible element contents that evolve to final products with higher  
609 wt.% SiO<sub>2</sub>, Mg/Fe ratios and lower incompatible element abundances (e.g.  
610 Reiners 1998; Smith et al. 2008; McGee et al. 2012). Such variability may  
611 complicate correlation of proximal units to their related distal tephra deposits

612 because directions and distances of eruptive dispersal may not be constant  
613 through an eruption.

614 For AVF centres, most of the eruptive phases are explosive (**Table 2**), and  
615 therefore, if centres show geochemical evolution through an eruption (e.g.  
616 Motukorea, Crater Hill), there is the potential for tephra deposits (from early  
617 phreatomagmatic phases) to have higher trace element ratios (LREE/HREE)  
618 than their subsequent lava or scoria deposits (from later magmatic phases). This  
619 bias may hinder correlation of some distal tephra to their source centre.

620 To address this issue, **Fig. 8** shows the geochemical progression through  
621 the eruption of Motukorea (data from McGee et al. 2012), compared with the  
622 correlated Motukorea tephra horizon found in the Orakei Basin core. Distal  
623 tephra-derived glass shards appear to show slightly higher SiO<sub>2</sub> concentrations  
624 at given Zr concentrations (due to fractional crystallisation processes), but do  
625 show the full evolutionary geochemical trend for the entire eruption. For the  
626 incompatible trace element ratios the glass shards appear to be geochemically  
627 comparable and again have signatures that are the same as all phases of the  
628 eruption from tuff (explosive early phases), to lava and scoria (less-explosive  
629 later phases) (**Fig. 8**). Although these results generally validate our method, we  
630 still cannot discount the possibility of a mismatch, due to the limited geochemical  
631 data available for the evolution of individual centres.

632 Another limitation of using geochemistry to correlate tephra to their  
633 source centres is that not all the 53 AVF centres show distinct geochemical  
634 signatures. Geochemical composition alone cannot unambiguously fingerprint a  
635 centre if there are either a large number of centres with relatively similar  
636 geochemical compositions, or a general lack of geochemical data (either whole

637 rock or glass). It is therefore essential to include additional criteria (discussed  
638 below) to allow confident correlations to be made.

639

#### 640 ***Multi-criteria correlation of tephra horizons to source centres***

641 We combine four key factors to correlate distal tephra deposits to their  
642 source centres: age, geochemistry, scale of eruption, and location of sources.

643 Where applicable, wind direction is also taken into account.

644 A shortlist of potential source centres (**Table 2**) is created based primarily on  
645 the restrictions provided by the age estimates of the tephra deposits and the age  
646 estimates of the centres. For those shortlisted centres, the major, trace, and trace  
647 element ratios of the proximal whole rock analyses are compared to the distal  
648 tephra derived-glass compositions, focussing on incompatible trace element  
649 ratios (**Fig. 9**). To strengthen potential correlations, other criteria such as the  
650 eruption scale and styles, and the location of the relevant source centre(s), and  
651 the relevant core(s) are also taken into account, as discussed below.

652 Because fall deposits thin systematically with distance (Pyle 1989; Lowe  
653 2011), eruptions with a large estimated tephra volume (ETV) and a dominant  
654 phreatomagmatic component are likely to produce a larger tephra output and  
655 hence a greater dispersal footprint and deposit. Therefore, very thick (primary)  
656 tephra deposits (>100 mm) in a core (Hopkins et al. 2015) require a source  
657 centre that is either 1) close to the deposition site (less than a few kilometres:  
658 Brand et al. 2014), and/or 2) has a predominantly phreatomagmatic eruption  
659 style, and/or 3) has a large magma supply and thus a long eruption duration.

660           Due to the relatively small size of the AVF volcanoes, the tephra dispersed  
661 by single eruption is not thought to cover the entire field for any single event  
662 (Kermode 1992). Therefore, the distribution and thickness of tephra deposits  
663 can be indicative of the region within the field where the source centre is located.  
664 For example, tephra deposits that are only found in the northern maar sites  
665 (Onepoto, Pupuke, Orakei, Glover Park) are inferred to indicate sources in the  
666 north or central AVF (based on the dominant wind direction, discussed below).  
667 Conversely a deposit only found in the southern maar site (Pukaki) is suggestive  
668 of sources in the south of the field. Tephra deposits found in both northern and  
669 southern maar sites are likely to have been derived from the central part of the  
670 field, and/or reflect an eruption large enough to widely disperse tephra from any  
671 source site within the field.

672           Wind direction is also considered, where possible, when making source  
673 correlations, because it has a controlling influence on tephra dispersal. For the  
674 Auckland region, evidence of prevailing past wind directions can be inferred  
675 from the morphology of the volcanic centres, for example, asymmetric tuff rings  
676 or scoria cones (e.g. Motukorea, Hayward et al. 2011). Such morphological  
677 indications are not however definitive for the majority of centres because there  
678 has often been post-depositional erosion, so present cone morphology is not  
679 seen as a definitive wind-direction indicator for an individual eruption. The  
680 dominant prehistoric wind patterns (westerly/south-westerly) are, however,  
681 still the dominant patterns for today (Houghton et al. 2006). This wind direction  
682 generally has resulted in more frequent tephra deposition in the northeast and  
683 east of the field, confirmed by the high number of deposits found within the  
684 Orakei Basin core, situated north-east of most centres (**Fig. 1**). Tephra deposits

685 are therefore more readily traced back to sources to the west and southwest.  
686 Conversely, centres found to the east or north east of the maar sites (e.g. Pigeon  
687 Mt., Hampton Park, Otara, Green Mt., and Styaks Swamp; **Fig. 1**) are less likely to  
688 be represented in the maar-lake tephra record.

689 Hopkins et al. (2015) detailed twenty-eight tephra horizons within six cores.  
690 Eleven of the horizons are cross-correlated between cores, linking two or more  
691 deposits, and seventeen tephra horizons are single deposits found only within  
692 single cores. We here have reduced the number of single horizons to sixteen, and  
693 increased the number of cross-correlated horizons to twelve based on the  
694 correlation of horizons AVF16/AVF13 as previously discussed.

695 For correlation purposes, we assess each tephra horizon individually; all  
696 potential sources are accounted for and discussed, without bias from any other  
697 correlations made (see **Supplementary Material**). A 'confidence value' is  
698 assigned for each correlation based on the number of supporting criteria that are  
699 satisfied (i.e. age, geochemistry, scale and location). In general, if all four criteria  
700 are satisfied a confidence level of 1 is given, when three are satisfied a  
701 confidence level of 2 is given, and if only two are satisfied a confidence level of 3  
702 is given (detailed in **Table 4**). Each of these criteria is variably weighted in  
703 importance with age  $\geq$  geochemistry  $\gg$  locality  $\geq$  eruptive scale. In some cases  
704 the confidence level is skewed to reflect this weighting of criteria, and this skew  
705 is detailed for each individual case in the supplementary material.

706 Discussion of the correlation of all 28 horizons to their proposed source can  
707 be found in the **supplementary material**, with an example of the discussion  
708 outlined below for a single representative tephra horizon (AVF5). For each of the  
709 horizons the proposed source centres are given in **Table 4** along with

710 alternatives that were considered. Of the twenty-eight horizons, eight have been  
711 given a correlation with confidence level of 1, eleven have been given a  
712 confidence level of 2, and seven have been given a confidence level of 3, with two  
713 horizons remaining uncorrelated (**Table 4**).

714 *Example of multi-criteria discussion for a level 3 correlation*

715 **AVF5** is a thick (110 mm) geochemically homogeneous deposit found only  
716 in the Orakei Basin core at a depth of 57.44 m. The bulk tephra sample contains  
717 coarse glass shards (>250  $\mu\text{m}$ ) and abundant country-rock lithic grains. The  
718 source is thus inferred to be relatively close to Orakei Basin in the north of the  
719 field. Its modelled sedimentation rate age is of  $34.2 \pm 0.9$  ka (**Table 2**). Mt.  
720 Cambria is the only candidate with the appropriate age and location, however it  
721 is one of the smallest centres in the field with an estimated tephra volume (ETV)  
722 of  $0.44 \times 10^6$  m<sup>3</sup> (**Table 2**). It is located ca. 5 km away from Orakei Basin, and  
723 therefore, it is highly improbable that it would have produced a 110 mm thick  
724 tephra deposit within the basin. Several other centres have appropriate locations  
725 and eruption scales, but are older than 35 ka (<sup>40</sup>Ar/<sup>39</sup>Ar age ranges [95 %  
726 confidence] from Leonard et al. 2017): Mt. Hobson (45.3-68.5 ka), Mt. St John  
727 (71.9-78.7 ka), Mt. Victoria (AVF4) (42.8-72.4 ka), and North Head (72.3-102.7  
728 ka), or conversely, too young; Little Rangitoto (AVF14) (16.3-25.1 ka), Taylors  
729 Hill (AVF10) (24.2-30.6 ka), and Panmure Basin (AVF13) (>17.5 ka). Of these  
730 centres only Mt. Victoria and Mt. Hobson have a similar (overlapping within  
731 error) geochemical signature to the tephra-derived glass within the AVF5  
732 horizon. Mt. Victoria has an ETV of  $3.9 \times 10^6$  m<sup>3</sup> and is located 4.7 km to the  
733 northwest of Orakei. In comparison Mt. Hobson has an ETV of  $1.8 \times 10^6$  m<sup>3</sup> and is



734 2.5 km downwind to the south west of Orakei basin. Based on this, Mt. Hobson is  
735 more likely than Mt. Victoria to have produced a thick deposit with large shards  
736 in Orakei Basin. The  $^{40}\text{Ar}/^{39}\text{Ar}$  age for Mt. Hobson (44.9-66.9 ka) is older than the  
737 modelled AVF5 tephra horizon age, but the only morphostratigraphic constraint  
738 is that **Mt. Hobson** is older than Three Kings (consistent with this correlation).  
739 We therefore discount the age constraints, which are separated by 9.8 kyr  
740 beyond error bounds. This correlation is predominantly based on the locality  
741 and scale of eruption and the deposit, with inconclusive geochemistry; it is  
742 therefore given a confidence level of 3.

743

#### 744 ***Tephra dispersal in the AVF***

745 Using confident correlations (level 1 and 2 only, which depend primarily on  
746 age and geochemistry) of tephra horizons from cores to their source centres,  
747 inferences can be made about the dispersal distances and thickness of the  
748 deposits from the AVF eruptions. **Table 5** outlines the distance (from source to  
749 depositional core site), thickness (primary horizon thickness identified by  
750 Hopkins et al. 2015), and (where applicable) the estimated shard sizes (based on  
751 grain sieving during glass shard extraction) for each of the centres that have  
752 been assigned a correlation with confidence level 1 or 2. There are no  
753 contemporaneous subaerial deposits in Auckland (cf. Hopkins et al. 2015), and  
754 the recorded thicknesses are here considered to be minima due to potential  
755 post-depositional compaction and erosion (Óladóttir et al. 2012).

756 For all correlations with a confidence level of 1, the maximum dispersal is  
757 of 13.5 km, for the Three Kings eruption recorded in Pupuke maar in a deposit 2

758 mm thick with shards of 50-100  $\mu\text{m}$ . For both confidence level 1 and 2  
759 correlations, the thickest deposits ( $\geq 100$  mm) are all found within 6 km from  
760 source, with a sharp decrease in deposit thickness (all  $< 80$  mm) at distances  $> 6$   
761 km (**Fig. 10A**). The maximum tephra thickness recorded in the cores is 510 mm;  
762 the tephra is from the One Tree Hill eruption in Orakei Basin, 4.6 km from the  
763 core site, suggesting that for a relatively large eruption ( $\text{DRE}^{\text{tot}} = 0.26 \text{ km}^3$   
764 Kereszturi et al. 2013) tephra deposits can be  $> 500$  mm thick at distances of  $> 4$   
765 km. The correlation results also show that shard size decreases with distance  
766 from source (**Fig. 10B**), with 60% of deposits  $< 6$  km from source having shards  
767  $> 200 \mu\text{m}$ , which reduces to 45% of deposits 6-12 km away and 0%  $> 12$  km from  
768 source. These findings are particularly applicable as inputs for tephra dispersal  
769 model simulations, evacuation and 'clean-up' forecasting, planning, and  
770 management (e.g. Tomsen et al. 2014; Wilson et al. 2014; Hayes et al. 2015).

771 Tephra horizon AVF12 correlates to Mt. Eden (**Fig. 1**), and is one of the  
772 most widely dispersed (and thus best preserved) tephra horizons;  $> 10$  mm thick  
773 in both Pupuke and Pukaki cores, which are 11 km and 12 km from source  
774 respectively. The Mt. Eden event also correlates with some of the thickest tephra  
775 deposits in the cores; 410 mm in Orakei (4.5 km from source), and 460 mm in  
776 Hopua (6 km from source). **Figure 11A** shows the decrease in tephra thickness  
777 away from source, coupled with the decrease in tephra shard size. Mt. Eden is  
778 also used as an example to show how the core-to-core and core-to-source centre  
779 correlations can be used to build isopach maps for the dispersal pattern of the  
780 eruption (**Fig. 11B**). The impact of the prevailing westerly winds (Hayward et al.  
781 2011) is considered and therefore produces an inferred elliptical tephra  
782 dispersal footprint. With a calculated total DRE volume of  $0.086 \text{ km}^3$ , the

783 eruption of Mt. Eden was one of the largest in the AVF, and therefore illustrates  
784 the impacts of a more extreme tephra dispersal event from a larger scale  
785 eruption.

786 Smaller eruptions produce more-restricted tephra dispersal; thirteen of the  
787 twenty-nine tephra horizons (45%) are only identified within single cores. Small  
788 eruptions can nevertheless result in near-source tephra horizons of substantial  
789 thickness. For example AVF10, now correlated to the eruption of Taylors Hill  
790 (DRE volume of  $0.0051 \text{ km}^3$ ), is restricted to the north of the field with cross-  
791 correlated deposits found in Orakei Basin (407 mm at ca. 5 km away), Onepoto  
792 (15 mm at ca. 12 km away) and Pupuke (3 mm at ca. 13 km away).

793 Deposits are not necessarily found in all maars along a dispersal pathway.  
794 For example AVF4 is found in Orakei Basin (41 mm) and Pupuke (15 mm) but is  
795 absent in Onepoto, which lies directly between the two. These dispersal patterns  
796 are most likely indicative of either discontinuous preservation and/or complex  
797 distal fallout (Molloy et al. 2009).

798 **Table 6** lists tephra dispersal information from selected basaltic volcanic  
799 fields worldwide together with those for some AVF centres. Monogenetic basaltic  
800 eruptions that show comparable total eruptive volumes, dispersal distances and  
801 thicknesses to some of the larger AVF centres include Mt. Gambier (Newer  
802 Volcanics, Australia) with an estimated  $\text{DRE}^{\text{tot}} = 0.20 \text{ km}^3$  (van Otterloo and Cas  
803 2013) and measured tephra thicknesses  $\leq 5 \text{ cm}$  at 10-12 km distance (Lowe and  
804 Palmer 2005). In comparison One Tree Hill ( $\text{DRE}^{\text{tot}} = 0.26 \text{ km}^3$ ) of the AVF has a  
805 measured tephra thickness of 6 cm at 10 km from source (**Table 6**). Marcath  
806 Volcano (Lunar Crater volcanic field, Nevada, USA) is of a similar eruptive scale  
807 to the mid-range AVF volcanoes, with a  $\text{DRE}^{\text{tot}} = 0.06 \text{ km}^3$  (Johnson et al. 2014).

808 Its tephra is 2 cm thick 7 km from vent, comparable to many AVF eruptions of  
809 similar scale, e.g. Mt. Wellington and Three Kings (**Table 6**). It is difficult to find  
810 global comparisons for the smaller AVF eruptions, but some of the latter show  
811 equivalent values to the larger global examples, for example, Orakei Basin, with a  
812  $DRE^{tot}$  of  $0.0067 \text{ km}^3$  depositing tephra 4 mm thick at 5 km from vent. A number  
813 of factors could potentially contribute to the apparent wider dispersal of tephra  
814 from the smaller AVF centres, including the high proportion of phreatomagmatic  
815 eruptions seen within the field (**Table 2**), the consistent prevailing wind  
816 directions, or the more favourable preservation conditions provided by the maar  
817 sites.

818

#### 819 ***Eruption age order resolution for the AVF***

820 The correlation of tephra deposits to their source centres, coupled with  
821  $^{40}\text{Ar}/^{39}\text{Ar}$  ages and morphostratigraphy, enables us to construct a relative age  
822 model for 48 of the 53 centres, thus allowing us to re-assess the absolute ages for  
823 all centres. As previously outlined, although the  $^{40}\text{Ar}/^{39}\text{Ar}$  age data (Leonard et  
824 al. 2017) provide improved age constraints for many of the AVF centres, the  
825 associated errors preclude ordering eruptive events. We reconstruct the relative  
826 temporal eruptive history for the AVF by combining; 1) the mean  $^{40}\text{Ar}/^{39}\text{Ar}$   
827 (Cassata et al. 2008; Leonard et al. 2017) and  $^{14}\text{C}$  ages (Lindsay et al. 2011;  
828 Needham et al. 2011), 2) the modelled sedimentation rate ages assigned based  
829 on tephra horizon correlations and, 3) the relative positions based on  
830 morphostratigraphic (cf. **Table 3**) or paleomagnetic constraints (Shibuya et al.  
831 1992; Cassidy 2006; Leonard et al. 2017; **Fig. 12**). For five centres there is not

832 enough information to assign absolute or relative ages, and these centres are  
833 therefore not included in the following evaluations. **Table 7** and **Figure 13**  
834 present a new relative age order and absolute ages for 48 of the AVF centres as  
835 defined by this study. A full discussion of the proposed relative and absolute age  
836 order can be found in the **supplementary material**.

837         Two previous studies have attempted reconstructions using statistical  
838 methods. Bebbington and Cronin (2011) reconstructed the temporal history of  
839 the entire field through age simulations based on tephra horizon correlations,  
840 stratigraphic constraints, and radiometric ages. Kawabata et al. (2016) made  
841 improvements to this statistical approach but focussed solely on correlating the  
842 tephra horizons to sources. The input for the original model simulations of  
843 Bebbington and Cronin (2011) included deposit thicknesses and age estimates  
844 for basaltic tephra within maar cores (from Sandiford et al. 2001; Shane and  
845 Hoverd 2002; Molloy et al. 2009), and age estimates for the AVF centres (from  
846 Lindsay et al. 2011). In order to improve on Bebbington and Cronin (2011),  
847 Kawabata et al. (2016) used newly refined ages for the rhyolitic and andesitic  
848 marker horizons from Lowe et al. (2013) as tie points within their  
849 reconstruction, and added wind direction and estimated tephra volumes. This  
850 improved modelling showed only 3 correlations that were consistent with the  
851 previous research, suggesting how easily new data inputs can dramatically  
852 impact the outputs of statistical modelling.

853         When we compare our tephra correlations to those outlined by Kawabata  
854 et al. (2016; **Table 4** and **Fig. 13**), there are three common correlations; AVF1  
855 and Domain, AVF2 and One Tree Hill, and AVF12 and Mt Eden. There are  
856 however a large number of discrepancies that we attribute to differences in

857 input data, in most cases linked to differing tephra horizon characteristics and  
858 the improved age constraints provided by Leonard et al. (2017).

859 **Figure 14** shows a comparison of our field-wide absolute and relative  
860 chronology results to those of Bebbington and Cronin (2011). There is significant  
861 scatter around the 1:1 line, indicating the data sets, and thus the relative orders  
862 are significantly different (**Fig. 14A**). For example, Bebbington and Cronin  
863 (2011) model 21 centres as older, 18 as younger, and 9 in the same positions as  
864 our results show. There are, however only a few large discrepancies (>20  
865 positions) between the two studies. Little Rangitoto, Motukorea, and Te Pou  
866 Hawaiki were all given much older positions (42<sup>nd</sup>, 35<sup>th</sup>, 43<sup>rd</sup> respectively) than  
867 those inferred in this study (13<sup>th</sup>, 12<sup>th</sup>, 16<sup>th</sup> respectively), and McLaughlins Mt.,  
868 Mt. Mangere and Mangere Lagoon are given much younger positions (4<sup>th</sup>, 9<sup>th</sup>, 12<sup>th</sup>  
869 respectively from Bebbington and Cronin, 2011) than those inferred in this study  
870 (30<sup>th</sup>, 33<sup>rd</sup>, 34<sup>th</sup> respectively).

871 For absolute age estimates (**Fig. 14B&C**), variation between the data sets is  
872 apparently greater than for the relative age estimates. Only twenty centres show  
873 offsets of <5 kyr between the modelled ages and our inferred ages, with the  
874 remaining 28 showing larger offsets of between 6.1 kyr (Mt Hobson) up to 124  
875 kyr (Te Pou Hawaiki). In addition, the modelled absolute ages (from Bebbington  
876 and Cronin 2011) cluster around 30 ka, whereas this study infers a broader  
877 spread between 20 and 35 ka for the same centres. The Bebbington and Cronin  
878 (2011) model is heavily weighted towards tephra horizons in the 30 ka age  
879 range, and this may impart a bias on the age constraints of their model's output.  
880 For all centres modelled by Bebbington and Cronin (2011) with ages between 45  
881 and 75 ka, the ages appear to be younger than inferred in this study (e.g. One

882 Tree Hill, Mt. Albert, and Tank Farm). Conversely, modelled ages for centres  
883 older than 75 ka seem to be over estimates (e.g. Little Rangitoto, Orakei Basin  
884 and Onepoto). The conflicting results for both relative and absolute age  
885 estimates between the two studies (e.g. for Onepoto, Pupuke, and Tank Farm), is  
886 likely to reflect differences in the data inputs.

887

## 888 ***Implications for the spatial, temporal and geochemical evolution of the Auckland*** 889 ***Volcanic Field***

### 890 *Spatial and temporal evolution*

891 The newly estimated ages for 48 of the 53 centres suggest that 18 centres  
892 erupted in the first ca. 140 kyr of the AVF's history (190 – ca. 50 ka), with 30  
893 erupting from ca. 50 ka to 0.5 ka. By using the rhyolitic marker horizons (RMHs)  
894 as definitive age constraints the number of eruptions per 1000 years (erup/kyr)  
895 can be calculated: present to Rerewhakaaitu (Rk) (0-17.5 ka) 0.3 erup/kyr; Rk to  
896 Okareka (Ok) (17.5–21.5 ka) eruption rate of 1.0 erup/kyr; Ok to  
897 Kawakawa/Oruanui (Kk) (21.5–25.4 ka) eruption rate of 1.5 erup/kyr; Kk to  
898 Rotoehu (Re) (25.4–52 ka) eruption rate of 0.6 erup/kyr and Re to inception (52-  
899 193 ka) eruption rate of 0.13 erup/kyr. These results suggest that in general  
900 there was an increase in the eruption frequency through time until ca. 21.5 ka  
901 (Okareka RMH; **Table 7**), followed by a decrease since 21.5 ka. Field-repose  
902 periods show a wide range from <0.1–13 kyr (**Table 7**), however eruptions are  
903 not evenly distributed within this range. Only six centres show field-repose  
904 periods of 10-13 kyr, whereas, twenty-three centres erupted after field-repose  
905 periods of 1000 years or less (all except four of which are younger than 50 ka),

906 and eighteen of these twenty-three have field-repose periods of 500 years or  
907 less. In general the longer field-repose periods occur at the beginning of the  
908 field's history, with all of the six centres with field-repose periods of 10-13 kyrs  
909 appearing between 193–86 ka.

910 The distance between successive eruptions (**Table 7**) varies from <0.5 km  
911 to 14 km with two outliers events taking place 21 and 19 km from sites of  
912 preceding events. There is spatial but not temporal alignment of some centres  
913 for example McLaughlins Hill – Wiri Mt. – Ash Hill (**Fig. 1**); these alignments have  
914 previously been attributed to pre-existing crustal fractures and faults (Magill et  
915 al. 2005; von Veh and Németh 2009; Kereszturi et al. 2014). In general there is  
916 no obvious spatial progression or pattern in location of vents through time.

917 Previous studies (Bebbington 2013; Le Corvec et al. 2013) have suggested  
918 that the location of each centre is independent of that of the previous centre, and  
919 for the most part the results presented in this study support this suggestion.  
920 When centre location is linked with the temporal evolution, however, a number  
921 of centres appear to have erupted very closely in space *and* time. These 'coupled'  
922 centres are here defined as having a field-repose period of 1000 years or less and  
923 with centres erupting <1 km away from each other. For example Mt. Wellington  
924 and Purchas Hill are dated to 10.5 ka and 11 ka respectively and are located ca.  
925 0.5 km apart. The other centres include Rangitoto 1 and 2 (Needham et al. 2011),  
926 Styaks Swamp and Green Mt., Mt. Eden and Te Pou Hawaiki, Otara and Hampton  
927 Park, and Wiri Mt. and Ash Hill (**Table 7**). It may also be possible to include  
928 Onepoto and Tank Farm, Mangere Mt. and Mangere Lagoon, and Domain and  
929 Grafton, although the age of one or both volcanoes in each of these pairs is poorly  
930 constrained.



931 *Geochemical evolution*

932           The collation of existing, and collection of new, whole rock and tephra-  
933 derived glass geochemical data presented here provides the most  
934 comprehensive geochemical dataset for the AVF to date (see **Table 1**). These  
935 data reveal a more complete view of the field as a whole, and further support the  
936 work of McGee et al. (2013, 2015), Hopkins et al. (2016), and McGee and Smith  
937 (2016) on the mantle source characteristics and the link between geochemical  
938 signatures of the erupted products (e.g. SiO<sub>2</sub> vs. CaO/Al<sub>2</sub>O<sub>3</sub> (**Fig. 17A**), or SiO<sub>2</sub> vs.  
939 (La/Yb)<sub>N</sub> (**Fig. 17B**)) and the eruptive volume for the centres (from Kereszturi et  
940 al. 2013). The new field-wide data set produced by this study shows that for SiO<sub>2</sub>  
941 vs. CaO/Al<sub>2</sub>O<sub>3</sub> the trend in the data is less well defined in comparison to SiO<sub>2</sub> vs.  
942 (La/Yb)<sub>N</sub> (**Fig 17**). This greater scatter is attributed to the impact of minor  
943 amounts of fractional crystallisation on major elements during magma ascent  
944 (e.g. Hopkins et al. 2016). The (La/Yb)<sub>N</sub> ratio shows a much stronger trend  
945 because these two elements are incompatible, thus less effected by fractional  
946 crystallisation, and therefore are more reflective of the mantle source signature.

947           In addition, McGee et al. (2013) highlighted a relationship between the  
948 trends observed in trace element multi-plots and eruptive volumes, suggesting  
949 that K and Sr anomalies (c.f **Fig. 4**) are also linked to eruptive volume. This  
950 conclusion was, however, based on geochemical data for only 10 centres  
951 (spanning a wide range in eruptive volumes). The addition of our new data  
952 suggests that these relationships may be less clear-cut. For example, the  
953 geochemical data for whole-rock samples from Te Pou Hawaiki shows a highly  
954 subdued K anomaly, coupled with a large Sr anomaly. This signature was linked  
955 by McGee et al. (2013) to centres with large eruptive volumes (e.g. Rangitoto

956  $DRE^{tot} = 0.6 \text{ km}^3$ ), yet Te Pou Hawaiki has a relatively small estimated volume  
957 ( $DRE^{tot} = 0.028 \text{ km}^3$ ). Similarly, Mt Cambria has one of the smallest eruptive  
958 volumes ( $DRE^{tot} = 0.00029 \text{ km}^3$ ), much smaller than Purchas Hill ( $DRE^{tot} =$   
959  $0.0017 \text{ km}^3$ ), yet does not have a more extreme geochemical signature than  
960 Purchas Hill (e.g. it lacks a more pronounced K anomaly, or Zr-Hf trough; **Fig. 4**).

961 If the McGee et al. (2013, 2015) correlations are accepted, then a number  
962 of the newly analysed centres exhibit geochemical signatures that are suggestive  
963 of larger magma batches than fit their inferred eruptive volumes (e.g. Te Pou  
964 Hawaiki; **Fig. 15**). There are three possible explanations for these discrepancies:  
965 1) volume estimates are inaccurate, 2) magma volume is 'lost' on ascent, or 3)  
966 the mantle source is heterogeneous.

967 Volume estimates by Kereszturi et al. (2013) are considered more reliable  
968 than those of Allen and Smith (1994), but the same relationships are seen with  
969 either data set (**Fig. 15**). Distal tephra is not accounted for in either model,  
970 potentially leading to volume underestimates (Kereszturi *pers. comm.*). This  
971 underestimate is not, however, enough to account for the observed discrepancies  
972 between the geochemical signatures and the erupted volumes. It is possible that  
973 there is a loss of magma during ascent, due to either or both of 1) fractional  
974 crystallisation of ascending melt, or 2) trapping of magma within the crust as an  
975 intrusion. Losses through fractional crystallisation are supported by the less  
976 well-defined relationship between the major elements and the erupted volumes  
977 as discussed previously. However, because many of the AVF lavas have a very  
978 primitive geochemical signature, there is only evidence of very limited fractional  
979 crystallisation (e.g. Smith et al. 2008; McGee et al. 2013; Hopkins et al. 2016),  
980 which is again not enough to account for the discrepancies. It is therefore most

981 likely that a heterogeneous mantle source, coupled with minor amounts of  
982 fractional crystallisation and retention of magma in the crust, may affect the final  
983 proportion of magma that is erupted. When geochemical data are combined with  
984 the temporal ordering, there are no obvious patterns identifiable through the  
985 history of the field. The lack of systematic change in the geochemical signatures  
986 through time suggests that the mantle source is not evolving in any systematic  
987 manner. Instead, the magma batches for each eruption are formed through the  
988 variable tapping and mixing of these heterogeneous mantle sources.

989

## 990 **Conclusions**

991 The collation of whole rock major and trace element data for the AVF has  
992 (with a few exceptions) facilitated the development and testing of a method to  
993 correlate distal tephra samples to their source volcanic centres. Geochemical  
994 correlation between distal tephra-derived glass and the glassy matrix of whole  
995 rocks at the source volcano is proved to be reliable. The method produces  
996 reasonable results based on major element signatures alone, with correlations  
997 strengthened by the use of trace-element signatures. Furthermore, incompatible  
998 trace elements and their ratios (particularly versus Yb; e.g.  $(\text{Gd}/\text{Yb})_N$ ,  $(\text{La}/\text{Yb})_N$ ,  
999  $(\text{Zr}/\text{Yb})_N$ ) are representative for individual centres and can therefore be used to  
1000 geochemically correlate distal basaltic tephra to proximal whole-rock samples in  
1001 the AVF. Specifically the ratios listed above are proven to be most useful in  
1002 assigning individual geochemical fingerprints because they are highlighted to be  
1003 the most variable across the field, yet the least variable within any given centre,  
1004 and the least affected by fractional crystallisation processes.

1005 This study has demonstrated geochemistry to be an effective tephra  
1006 correlation tool, but we stress that geochemical compositions are not always  
1007 sufficiently distinct to provide a definitive result. To efficiently correlate tephra  
1008 layers to their source centres, a multi-criteria approach is required. For greatest  
1009 correlation confidence, this approach combines age data (of both distal tephras  
1010 and proximal whole rock deposits) and eruption characteristics (e.g. scale and  
1011 locality), to assign the source centre to tephra deposits. Of the twenty eight  
1012 basaltic tephra horizons in the AVF maar-lake cores, all but two (newA and  
1013 newB) are correlated to a source; eight with a confidence level of 1, eleven with a  
1014 confidence level of 2, and seven with a confidence level of 3.

1015 The correlations with confidence levels of 1 and 2 are used to determine  
1016 tephra dispersal and thickness (e.g. footprint) from the AVF eruptions. The  
1017 maximum tephra dispersal distance is 13.5 km with a primary deposit thickness  
1018 within the core of 2 mm, and for all primary core deposits with a thickness >100  
1019 mm the source is <6 km away. In a number of cases the deposits are restricted to  
1020 sites in close proximity to the source centre, suggesting that in the event of a  
1021 future small-scale eruption, damaging thicknesses of tephra will not inundate the  
1022 entire Auckland area.

1023 Our correlations also provide a clearer picture of the temporal evolution of  
1024 the AVF. Using the stratigraphic relationships of the tephra horizons within the  
1025 cores and their association with the rhyolitic marker horizons, the absolute age  
1026 order of the centres can be resolved. Because of the errors associated with dating  
1027 techniques ( $^{40}\text{Ar}/^{39}\text{Ar}$  and  $^{14}\text{C}$ ) a relative sequencing of the AVF centres was  
1028 previously not possible. Using our new method we provide high-confidence  
1029 relative and absolute eruption age estimates for 48 centres, leaving only five

1030 centres with uncertain ages (Pukaki, Pukewairiki, Boggust Park, Cemetery Hill  
1031 and Puhinui Craters). Our reconstruction of the relative ages of the centres also  
1032 allows the temporal, spatial, and geochemical evolution of the AVF to be  
1033 assessed, confirming that there is no simple temporal pattern in the spatial and  
1034 geochemical evolution of the field.

1035

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## 1 **Figure Captions**

2 **Figure 1.** (A) Map of the Auckland Volcanic Field and its eruptive centres (from  
3 Hayward et al. 2011). The locations of maar craters from which cores documented here  
4 were collected are highlighted by red symbols and red font: Pupuke, Onepoto, Glover  
5 Park, Orakei, Hopua and Pukaki. Although the Glover Park core is from St Heliers  
6 volcano, to avoid confusion here the core location will continue to be called Glover Park.  
7 (B) General location of the AVF within the North Island, New Zealand. Highlighted are  
8 other key volcanic centres including the South Auckland Volcanic Field (SAVF), and the  
9 key rhyolitic sources from the Taupo Volcanic Zone (TVZ) (Taupo Volcanic Centre  
10 (TVC), Okataina Volcanic Centre (OVC)) and andesitic (Tongariro Volcanic Centre  
11 (TgVC), Mt. Taranaki (Tk/Eg)) sources of tephra found in Auckland maar crater cores.

12

13 **Figure 2.** Age-depth profiles for rhyolitic marker horizons (RMHs) and basaltic tephra  
14 deposits within the cores, and individual sedimentation rate profiles for each core.  
15 Abbreviations, errors, and references for the RMHs ages are in **Table 2**. Age envelopes  
16 are highlighted in light grey based on the errors associated with the RMH ages. AVF  
17 basaltic deposits are plotted as red triangles at the appropriate depth in the core, and  
18 horizon Eg36, an andesitic marker horizon from Mt Taranaki, is plotted in green.

19

20 **Figure 3.** Representative major and trace element variation diagrams (in wt%) for AVF  
21 volcanic rocks (n=744; data in **supplementary material**). Highlighted are those which  
22 show examples of the distinct patterns seen within individual centres, grey symbols  
23 show all other data.

24

25 **Figure 4.** Primitive mantle-normalised trace element plots for whole rock (shaded grey)  
26 and glass from selected tephra horizons (coloured lines) from a range of cores showing  
27 a range of geochemistries and ages (high AVF#s = young, low AVF# = old). Values are  
28 normalised to primitive mantle after McDonough and Sun (1995).

29

30 **Figure 5.** Comparison plot for concentrations of major and trace elements for whole  
31 rock and glass for the full sample suite (all data in **supplementary material**). (A) MgO  
32 vs. SiO<sub>2</sub> indicating an example of elements that do not correlate, and (B) (Zr/Yb)<sub>N</sub> vs.  
33 (Gd/Yb)<sub>N</sub> indicating an example of trace element ratios that do correlate for glass and  
34 whole rock samples.

35

36 **Figure 6.** Multi-element plots to show geochemical comparison between glass from a  
37 known Mt. Wellington tephra deposit from the Hopua maar core, a simulated glass  
38 (matrix-derived glass) made from Mt. Wellington whole rock sample AU62394, and  
39 whole rock analyses from Mt Wellington. (A) Glass comparison of MgO vs. Al<sub>2</sub>O<sub>3</sub>, (B)  
40 glass comparison for Gd vs. Zr, (C) example of glass and whole rock concentrations for  
41 major elements which are not comparable (MgO vs Al<sub>2</sub>O<sub>3</sub>), (D) example of glass and  
42 whole rock major element concentration that are comparable (CaO vs FeO), (E) example  
43 of glass and whole rock trace elements that are comparable (Tm vs. Gd), (F) example of  
44 glass and whole rock trace element ratios that are comparable ((Zr/Yb)<sub>N</sub> vs. (Gd/Yb)<sub>N</sub>).  
45 Individual analyses are shown by symbols, field-wide geochemical concentrations of  
46 glass are outlined by orange dashed area and field-wide geochemical whole rock  
47 concentrations are shown by black dashed area.

48

49 **Figure 7.** Selected whole rock and glass sample concentrations to show the effects of  
50 crystal removal. (A) MgO vs. Ni for glass and, (B) whole rock. Both show a high r<sup>2</sup> value  
51 suggesting a statistically significant relationship between the two elements. In  
52 comparison (C) MgO vs. La for glass and, (D) for whole rock. Both show r<sup>2</sup> values near  
53 zero, indicating no statistically significant relationships between the elements.

54

55 **Figure 8.** Graphs showing the variations and comparability of the geochemical  
56 signatures observed through the eruptive products of Motukorea volcano (from McGee  
57 et al. 2012), coupled with the geochemical signatures for the distal glass composition  
58 found within the Orakei Basin core (horizon AVF15). (A) SiO<sub>2</sub> (wt%) vs. Zr (ppm) and  
59 (B) (Zr/Yb)<sub>N</sub> vs. (Gd/Yb)<sub>N</sub> normalised to primitive mantle values (McDonough and Sun  
60 1995). Similar relationships are also seen for (La/Yb)<sub>N</sub>, (Ce/Yb)<sub>N</sub>, (Nb/Yb)<sub>N</sub>, and  
61 (Nd/Yb)<sub>N</sub> (data from **supplementary material**).

62

63 **Figure 9.** Example plots of geochemical correlations. Glass values are shown in coloured  
64 symbols that indicating different cores, whole rock values are shown by coloured  
65 triangles for each centre, and the grey field shows the geochemical spread for the entire  
66 AVF, both whole rock and glass compositions. (A) Example of a confidence level 1  
67 correlation for the Three Kings centre with tephra layer AVF7, showing selected major  
68 element and normalised trace element ratios. (B) Example of an ambiguous geochemical  
69 correlation for Crater Hill centre and tephra horizon AVF8 due to limited trace element  
70 geochemistry for some centres. (C) Example of centres that are of an appropriate age  
71 but show no geochemical correlation to the tephra horizon AVF13.

72

73 **Figure 10.** Data for all correlations with a confidence rating of level 1 or 2 (data in  
74 **Table 6**). (A) Horizon thickness vs. distance from source, showing the thinning of  
75 deposits increases away from source. Grey shaded area marks <6 km, within which all  
76 the deposits >100 mm thick are found. (B) % Shard size vs. distance from source,  
77 indicating the fining of away from source.

78

79 **Figure 11.** Example of the correlation of Mt. Eden eruption to tephra deposit AVF12. (A)  
80 Graph to show change in deposit thickness away from source, note the extreme decline  
81 in thickness after ca. 6 km distance. Also shown on (A) are backscatter electron images  
82 of the glassshards from each core site taken on EMPA. All pictures are at the same scale  
83 with the bar at the base of the images representing 200  $\mu\text{m}$ . (B) Inferred isopach map of  
84 the tephra dispersal from Mt. Eden based on the deposit thicknesses found in the cores.  
85 Dispersal is skewed to the east to reflect the westerly winds likely to have been present  
86 at the time of eruption (Hayward et al. 2011).

87

88 **Figure 12.** Age range chart for all centres (data from **Table 5.1**). Those in red are  
89  $^{40}\text{Ar}/^{39}\text{Ar}$  (from Leonard et al. 2017 or Cassata et al. 2008) (2 sd error) or  $^{14}\text{C}$  ages (from  
90 Lindsay et al. 2011). Markers show the mean ages measured by these techniques with  
91 lines showing the age ranges measured. Lines in orange have their ages based only on  
92 morphostratigraphy, and those in grey have no ages associated with them. Of note is the  
93 number of centres which, based on errors, could have erupted at a given time. For  
94 example there are 18 potential centres whose age ranges include 50 ka (Mt. Cambria,  
95 McLaughlins Hill, Hopua, One Tree Hill, Mt. Victoria, Mt. Hobson, Waitomokia, Onepoto,  
96 St Heliers, Tank Farm, Domain, Grafton, Otuaataua, Puhinui Craters, Mt. Robertson,  
97 Cemetery Hill, Boggust Park, and Pigeon Mt.).

98

99 **Figure 13.** Figure to show the combined age data that allow the centres to be put in  
100 order. Core correlations are from Hopkins et al. (2015), AVF horizon correlations from  
101 this study, Ar-Ar ages and ranges from Leonard et al. (2017), and morphostratigraphic  
102 relationships from Allen and Smith (1994); Affleck et al. (2001); and Hayward et al.  
103 (2011). Key rhyolitic marker horizons are shown in colours, and highlight the  
104 chronostratigraphic age limits for the basaltic horizons. Age ranges depicted by error  
105 bars are not to scale, the ranges are drawn to the associated ages in the cores.

106

107 **Figure 14.** A comparison of relative and absolute age orders for 45 AVF centres from  
108 statistical modelled results (Bebbington and Cronin 2011) versus new data from this  
109 study. (A) Relative age order, (B) absolute age estimates, and (C) 10-50 ka for absolute  
110 age. The 1:1 ratio lines are shown in red on each chart for comparison purposes.

111

112 **Figure 15.** Comparison plots for whole rock geochemistry vs. eruptive volume for all  
113 data available from the AVF. Data are plotted versus eruptive volume estimates from  
114 both Kereszturi et al. (2013) and Allen and Smith (1994) for comparison. All data are  
115 shown in light grey symbols, with mean values for each centre highlighted for pre-  
116 existing data in grey triangles, and for new data in red triangles.

117 **Table Captions**

118 **Table 1.** Catalogue of geochemical whole rock data (pre-existing and additions from this  
119 study) available for the AVF, ordered by the number of analyses, including those centres  
120 without any current data. After the addition of data in this paper, 44 centres now have 3  
121 or more geochemical data points.

122

123 **Table 2.** Details of all 53 centres in the AVF, their eruption type; the current age  
124 estimate and method by which the ages are calculated, the relative age relationships  
125 where known including, and the morphological features which give age constraints.  
126 Sources are: a. Hayward et al. (2011); b. Allen and Smith (1994); c. Affleck et al. (2001);  
127 d. Sandiford et al. (2002); e. Lowe et al. (2013); f. Lindsay et al. (2011); g. Kermode  
128 (1992); h. Newnham et al. (2007); i. Agustín-Flores et al. (2015); j. Leonard et al. (2016);  
129 k. Hayward et al. (2016); the estimated dense rock equivalent (DRE) volumes for the  
130 total, tuff ring and scoria cone from Kereszturi et al. (2013); and the calculated tephra  
131 volumes using the equation reported in Kawabata et al. (2015). For the eruption types,  
132 (A) phreatomagmatic wet explosive eruption which produces maar craters and tuff  
133 rings, (B) dry magmatic eruptions including fire fountaining creating scoria cones, and  
134 (C) effusive eruptions resulting in lava flows, and shield building.

135

136 **Table 3.** The ages and associated errors calculated for each basaltic horizon using  
137 either, Monte Carlo simulations for those younger than the Maketu RMH, and  
138 sedimentation rate calculations for those older than the Maketu RMH (*italicised*).  
139 References: a. Needham et al. (2011); b. Lowe et al. (2013); c. Molloy (2008); d. D.J. Lowe  
140 *pers comm* (2016); and e. Leonard et al. (2016). AVF24 is split into Rangitoto (Ra)1 and  
141 2 identified and dated ( $^{14}\text{C}$  in cal. yr. BP) by Needham et al. (2011), \*indicates  
142 nomenclature from Molloy et al. (2009) for the tephra horizons found in the Pupuke  
143 core. The ages for the rhyolitic marker horizons (shaded grey) are outlined in cal. yr. BP.  
144 The age of AVF17 is shown in grey text as an outlier, and the position of AVF16 also  
145 shown in grey text as out of sequence, both of these are discussed in the text. The age of  
146 deposit AVFd in the base of the Onepoto core is taken from the minimum  $^{40}\text{Ar}/^{39}\text{Ar}$  age  
147 estimation for Pupuke centre, see text for details. All errors are reported as 2 s.d., and  
148 the 95% confidence limits are also reported.

149

150 **Table 4.** Outline of correlations for individual tephra horizons to their source centre.  
151 Average age are calculated by this study (**Table 3**). Proposed centre is given in bold with

152 certainty value (scale 1-3). Ticks indicate where correlation satisfies the criteria of age  
153 (within error of radiometric age), chemistry, scale, and location, '?' indicated where  
154 centre ages are unknown. Alternative possible centres are outlined with their certainty  
155 value and criteria. See supplementary material for explanation of ambiguities in the  
156 table in relation to rating given.

157

158 **Table 5.** For those deposits with a correlation certainty of 1 or 2, the distance to the  
159 deposition site (core) (km), thickness of the deposit within the core (mm) and the  
160 average shard size of the tephra ( $\mu\text{m}$ ) are shown.

161

162 **Table 6.** Comparative global values for tephra dispersal, thickness and total dense rock  
163 equivalent (DRE) volume (in  $\text{km}^3$ ; from Kereszturi et al. 2013 to allow global  
164 comparisons) for monogenetic basaltic volcanoes. \* Cerro Negro is a polygenetic scoria  
165 cone, however it has a comparable total volume estimate from the 1995 basaltic  
166 eruption, and is therefore deemed applicable for comparison. In bold are examples from  
167 this study to allow a direct comparison.

168

169 **Table 7.** Relative order of eruptions with calculated mean ages, time and distance  
170 relationship between the  $n^{\text{th}}$ ,  $n+1$  and  $n+2$  centre. References include a. tephra horizon  
171 ages from this study; b.  $^{14}\text{C}$  from Lindsay et al. 2011; c. Ar-Ar from Leonard et al. 2016 or  
172 Cassata et al., 2008 (see **Table 3**); d. morphostratigraphic constraints (references in  
173 **Table 3**) and/or paleomagnetic constraints ( from Shibuya et al. 1992). Absolute ages  
174 evaluated by this study are discussed in detail in the supplementary material. Note that  
175 for centres where morphostratigraphy suggests contemporaneous eruptions (e.g., no  
176 material between successive volcanic deposits) an arbitrary difference of 500 years is  
177 assigned based on a minimum time taken to form soil horizons.

Figure 1

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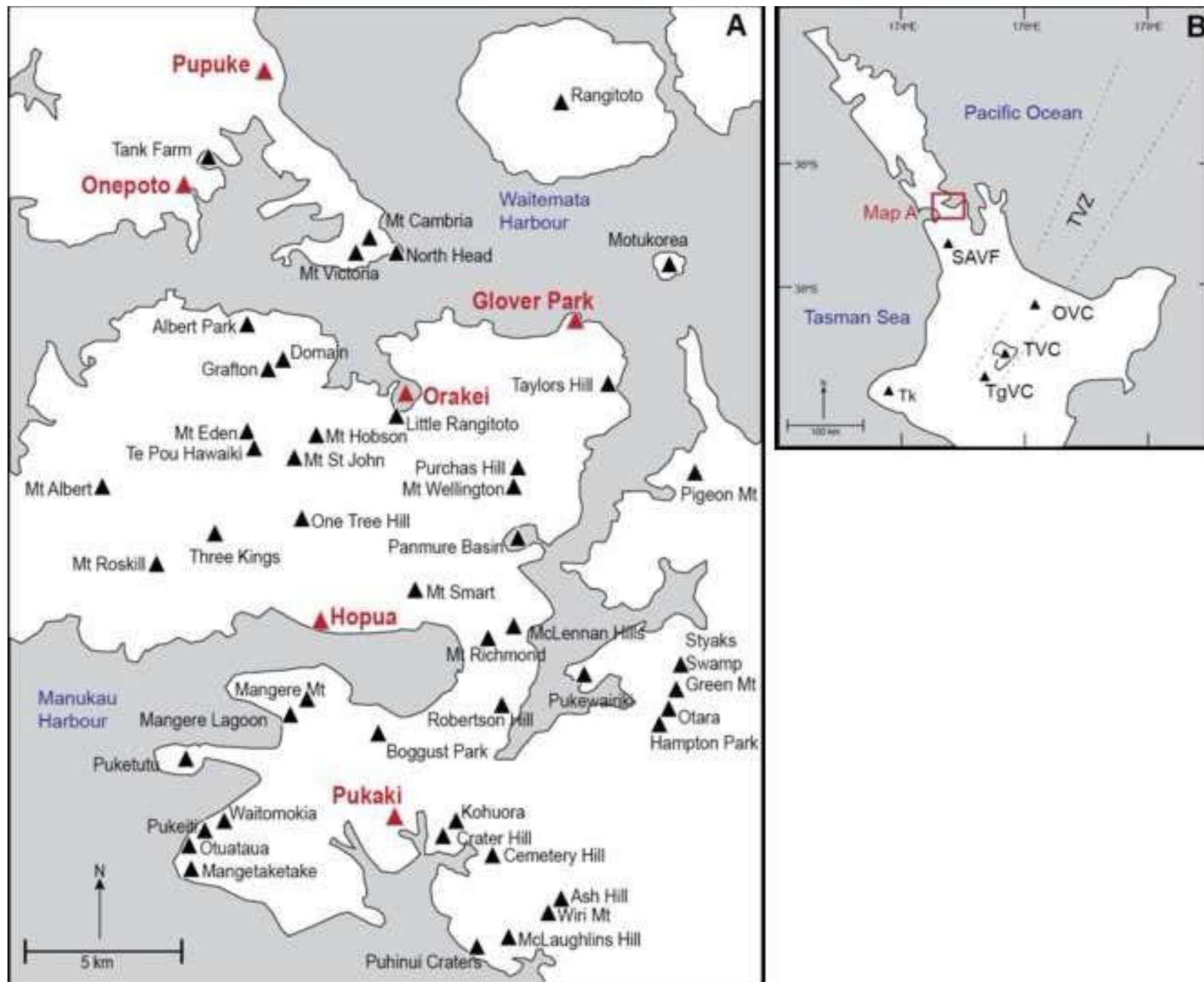
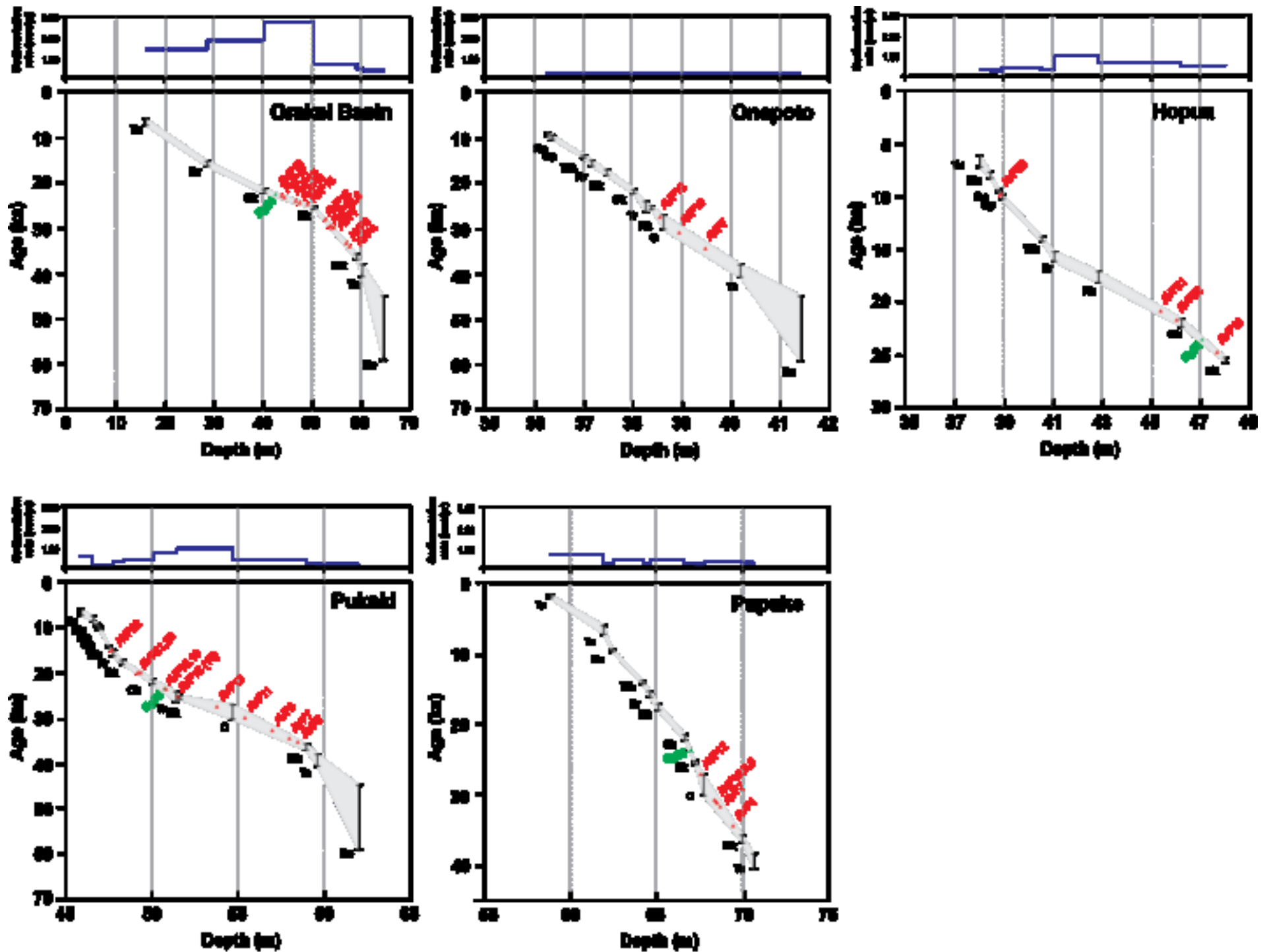


Figure 2

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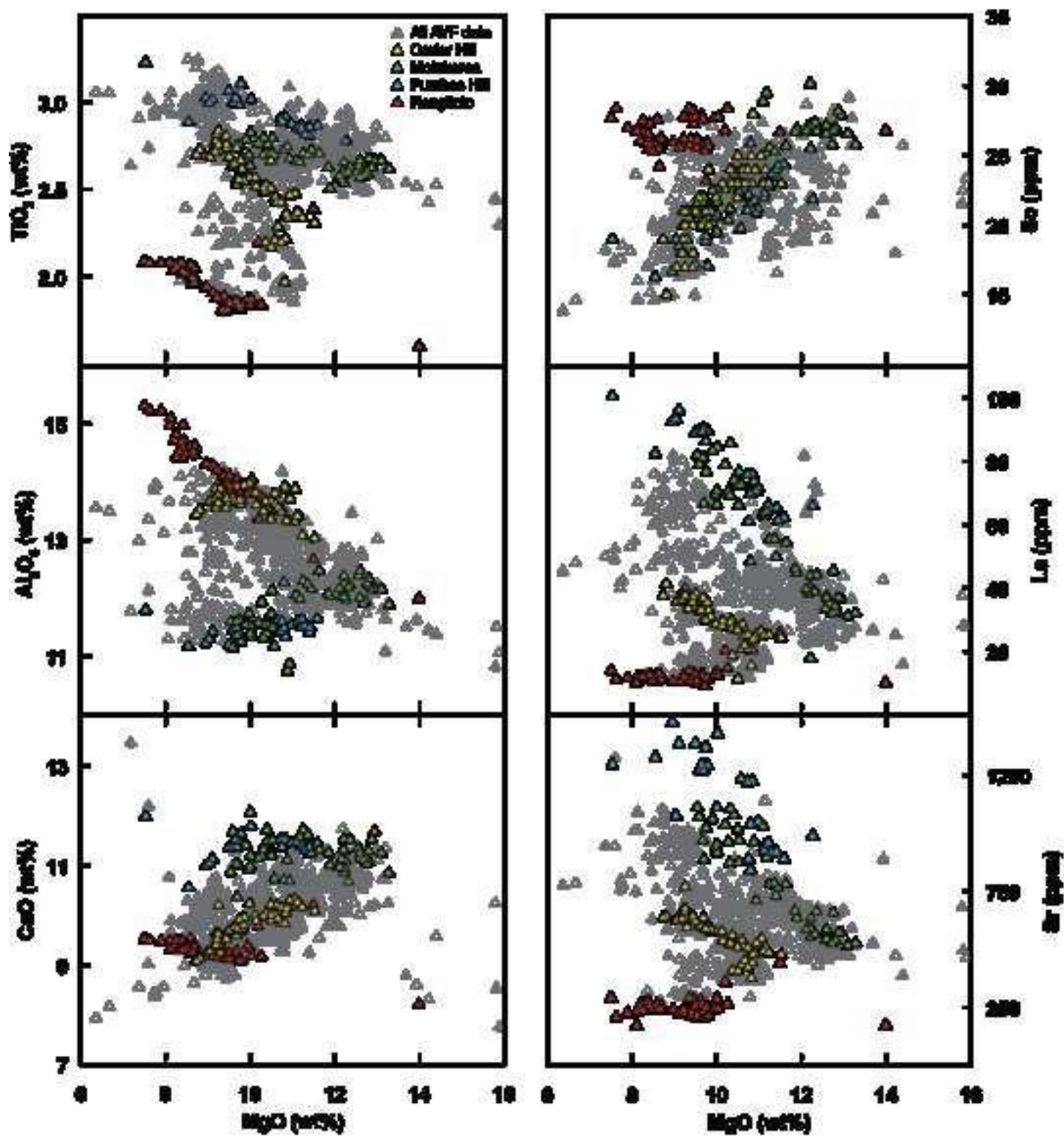
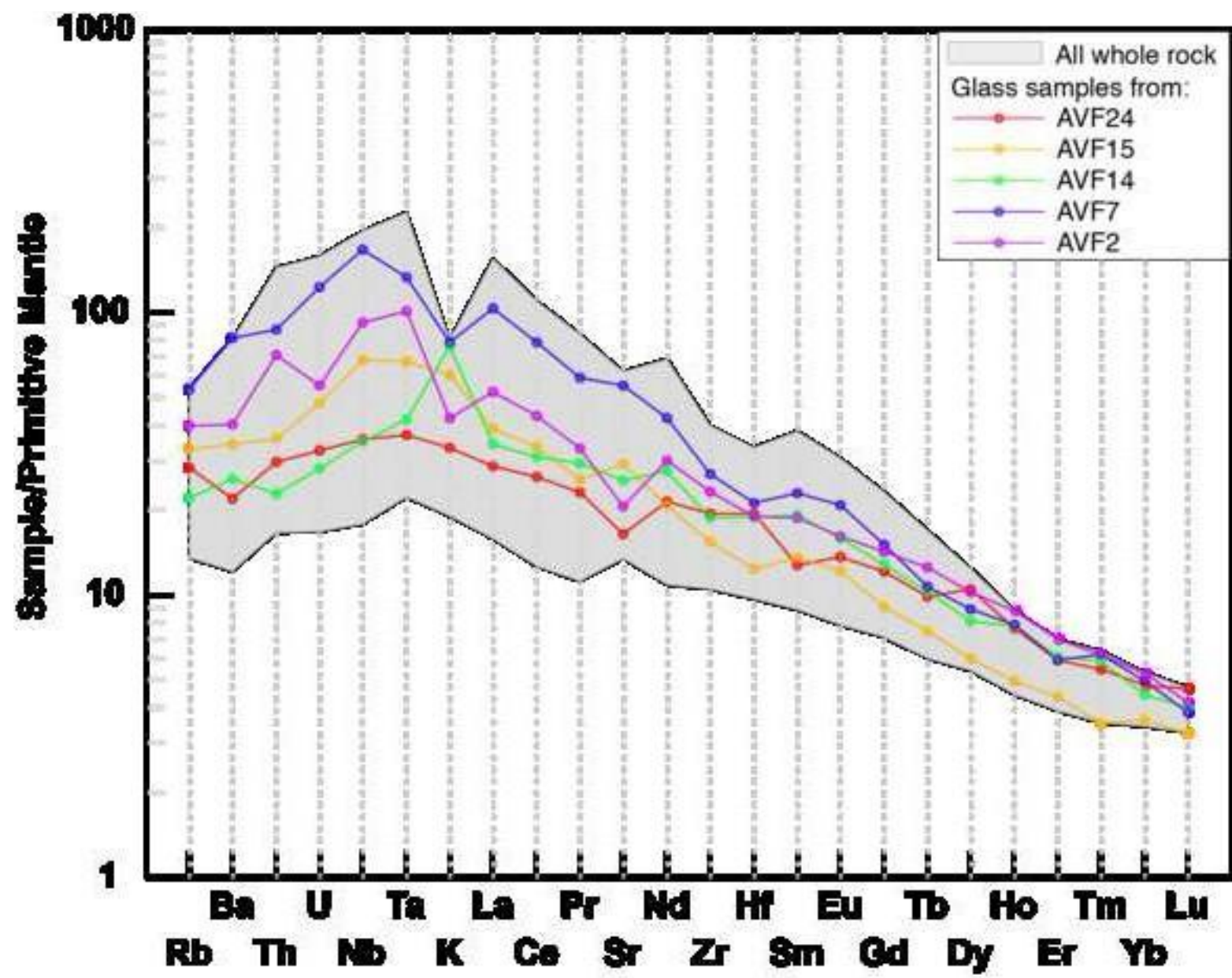
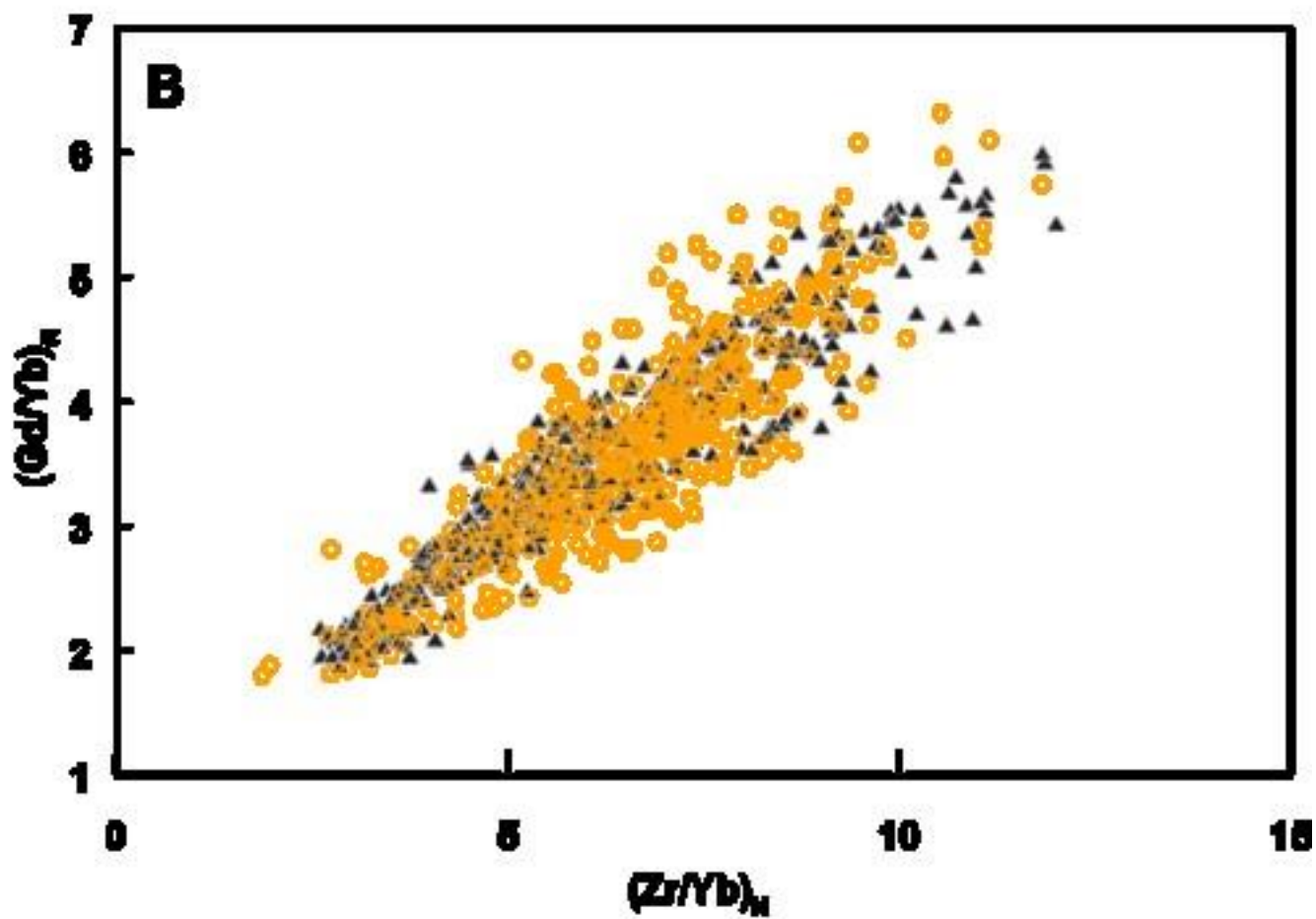
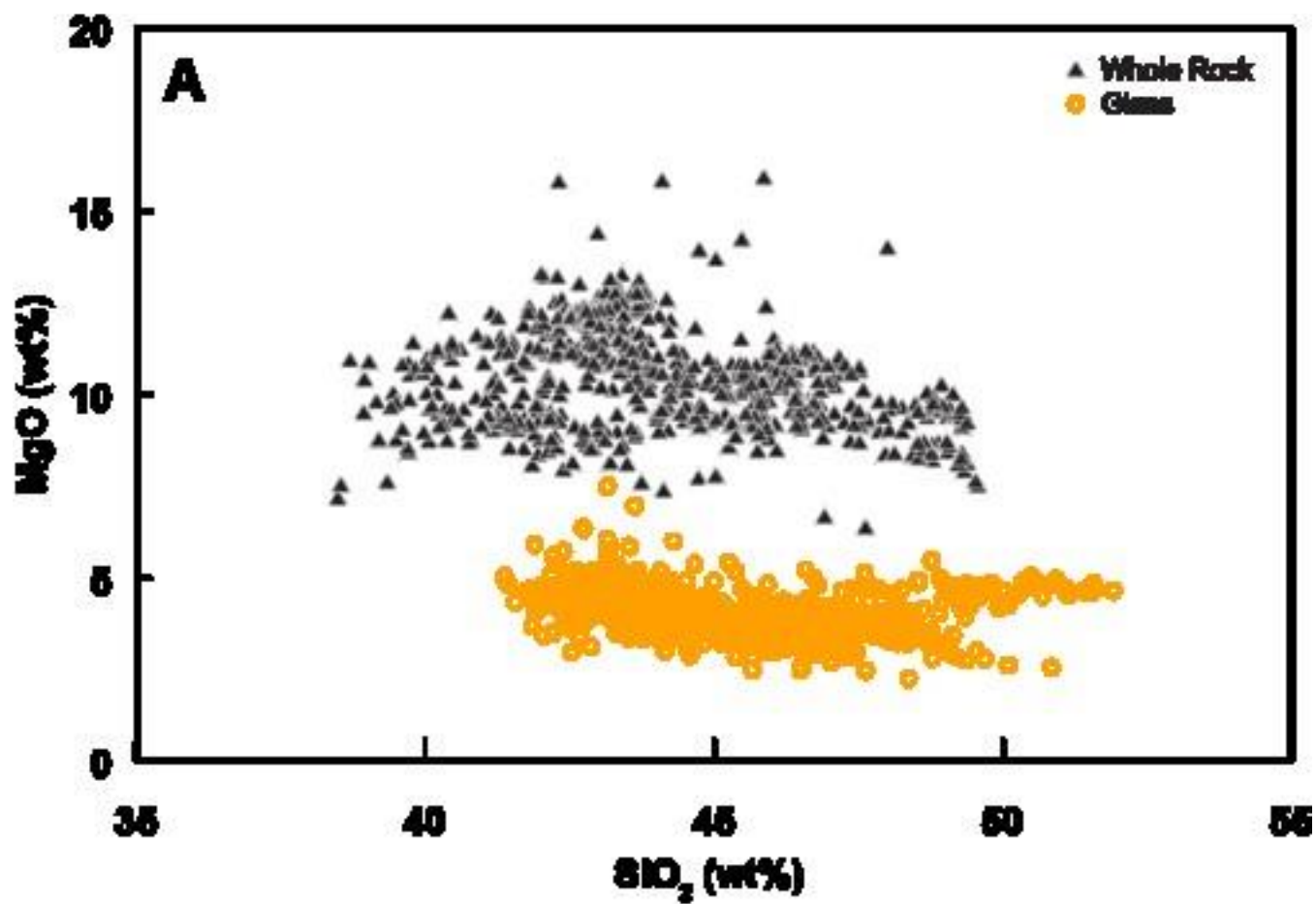
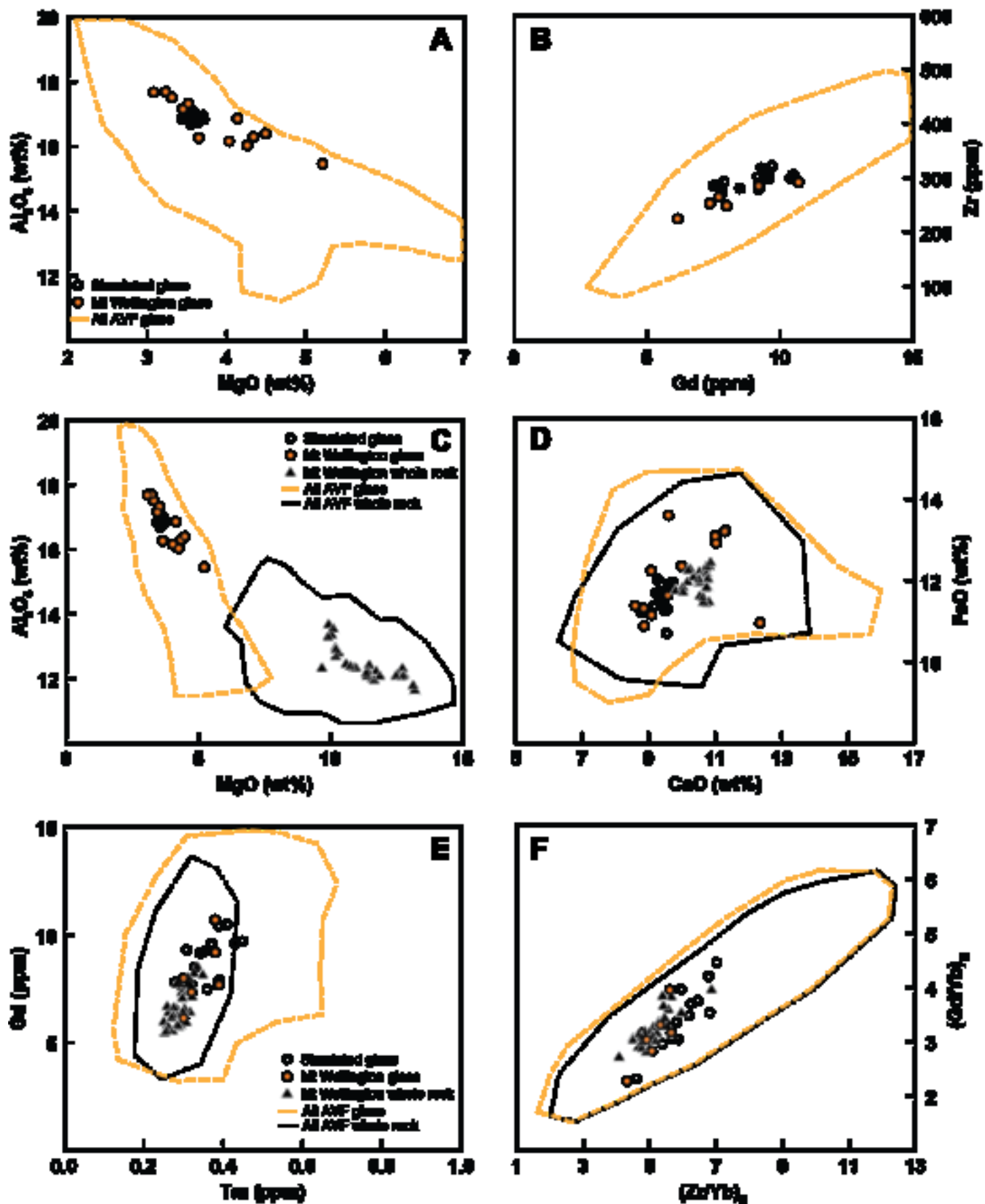
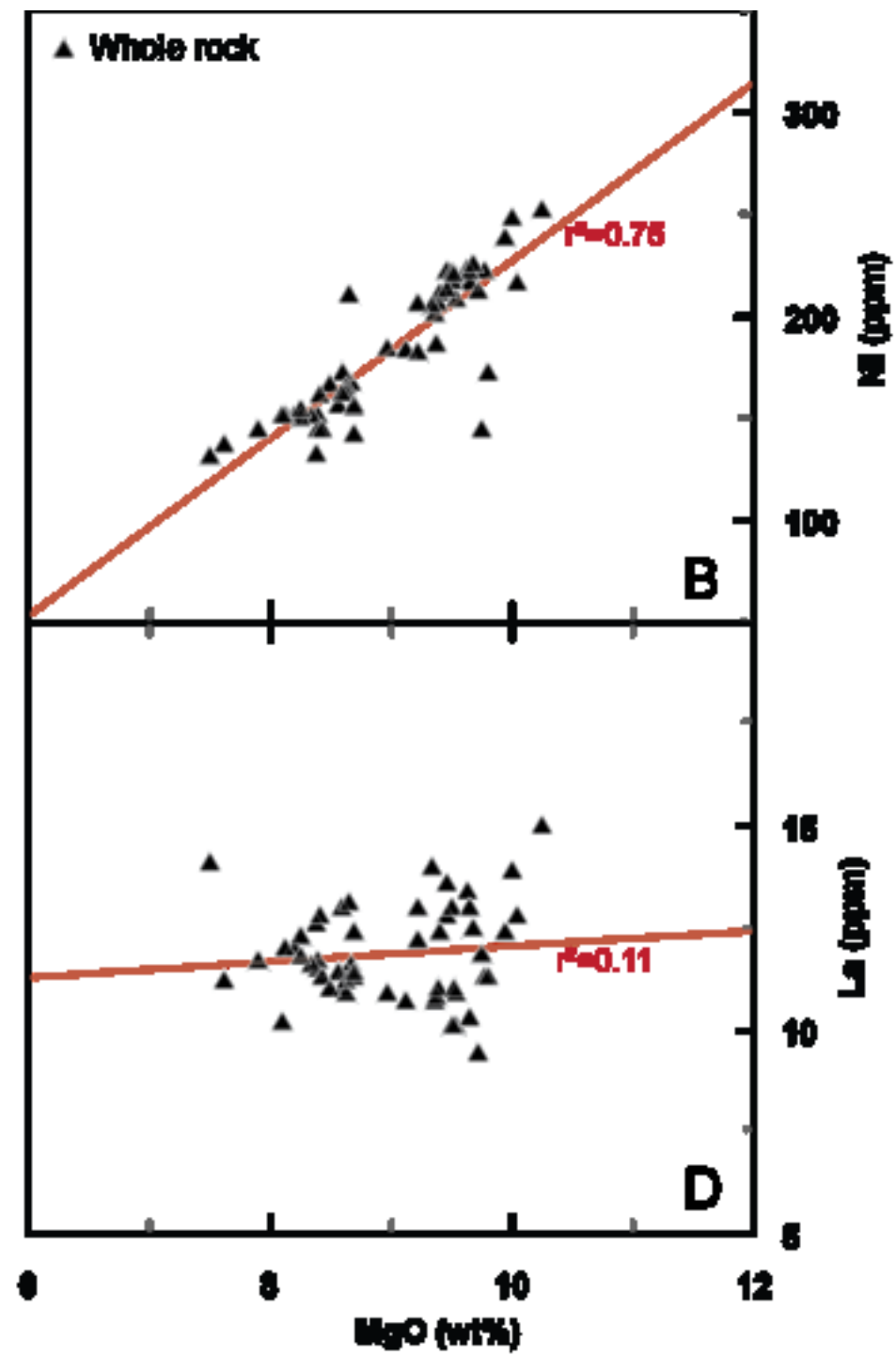
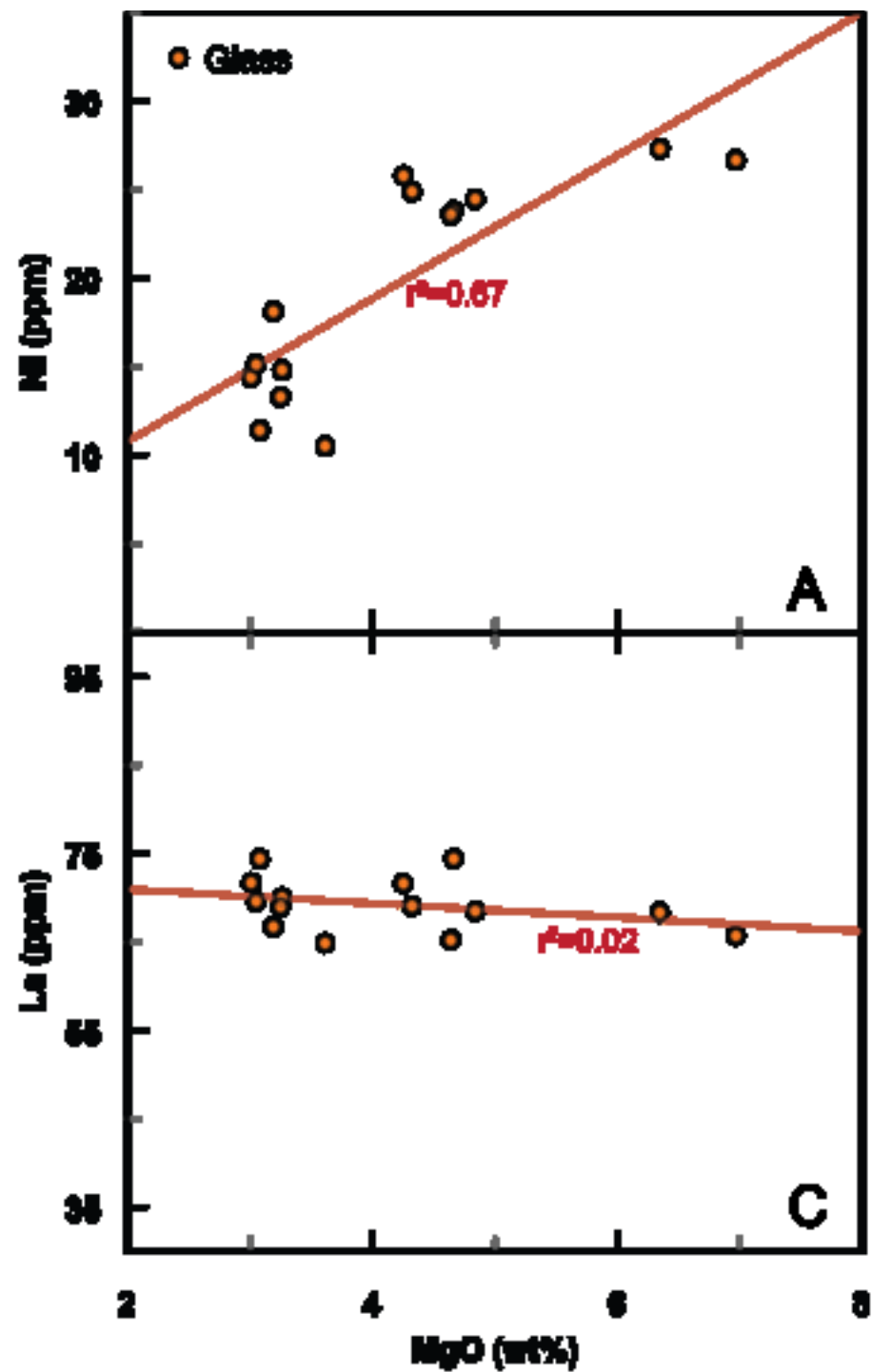


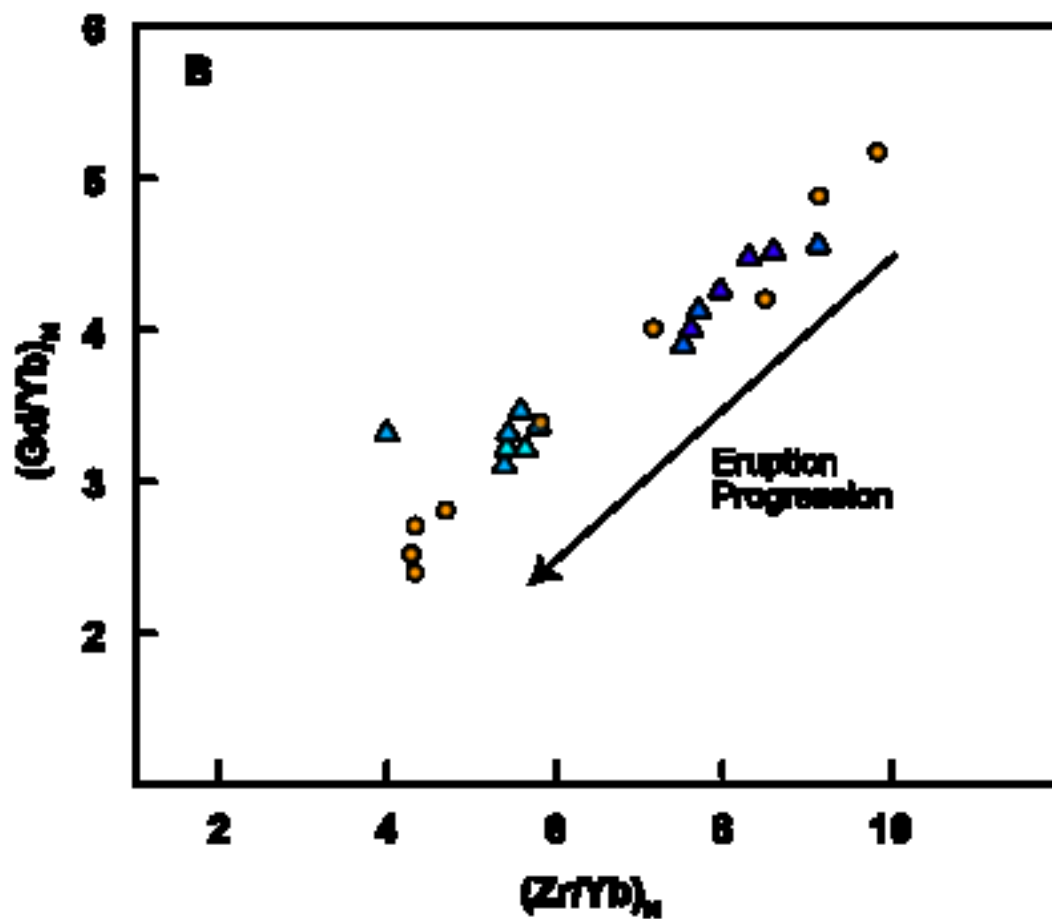
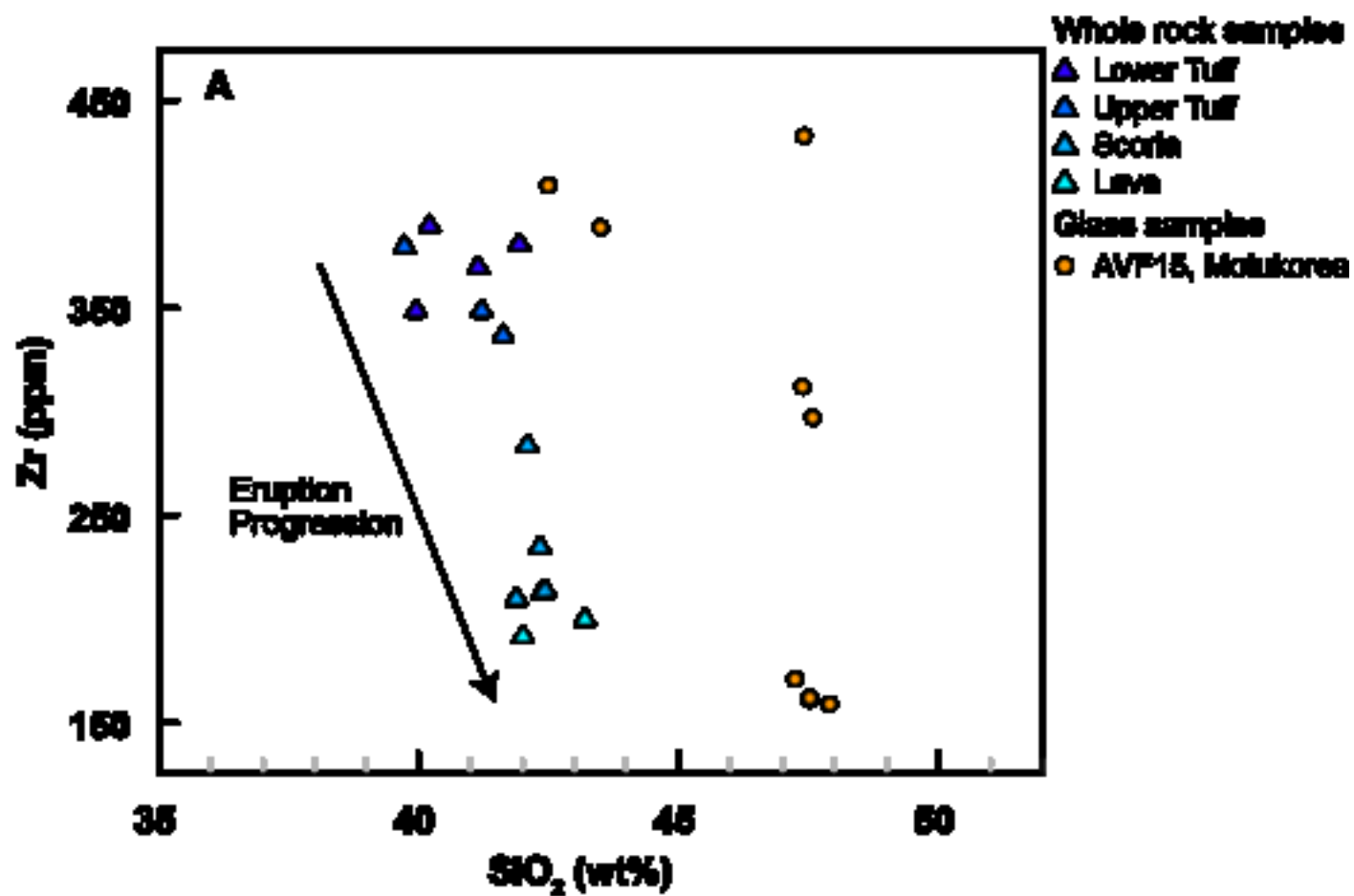
Figure 4



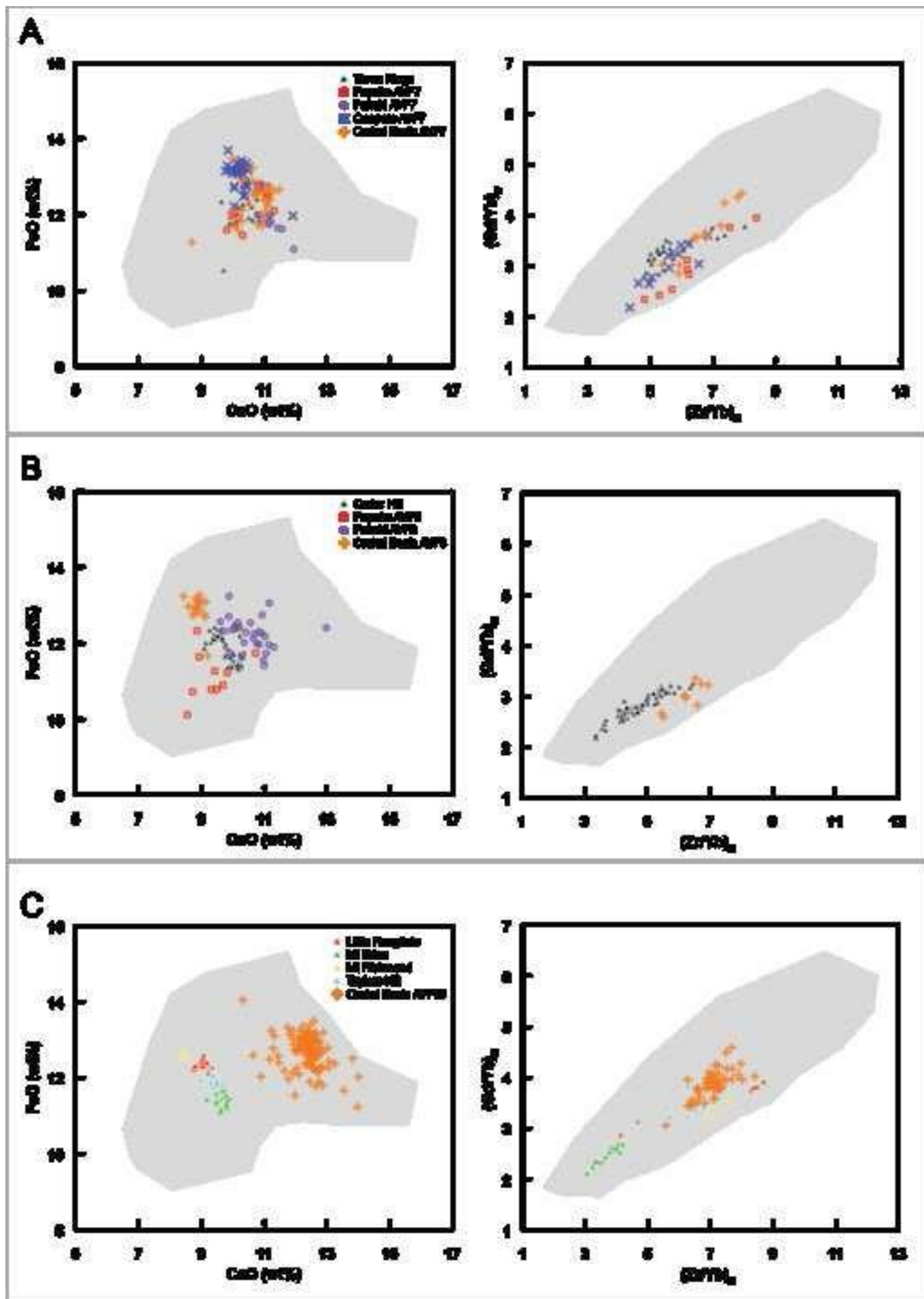


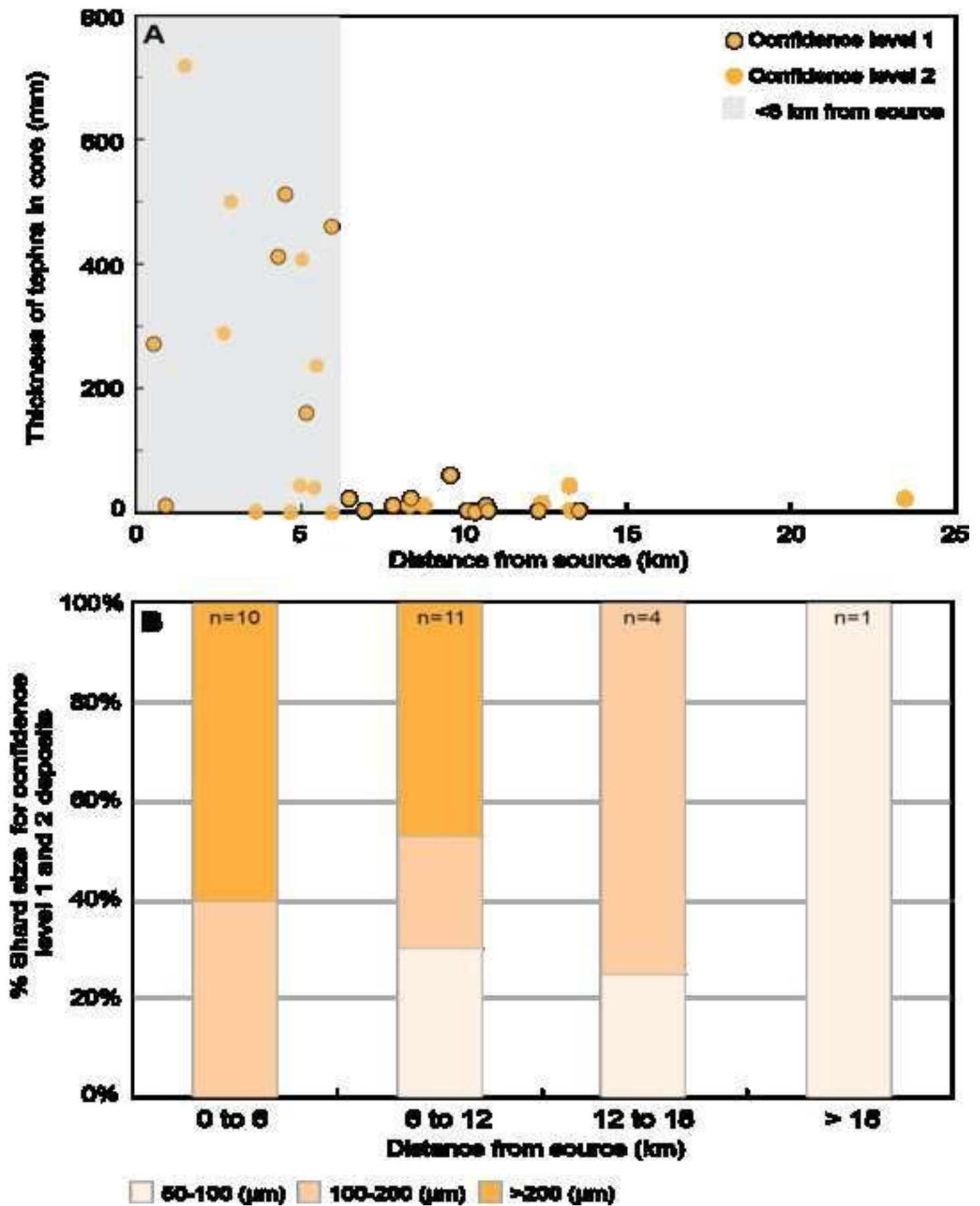














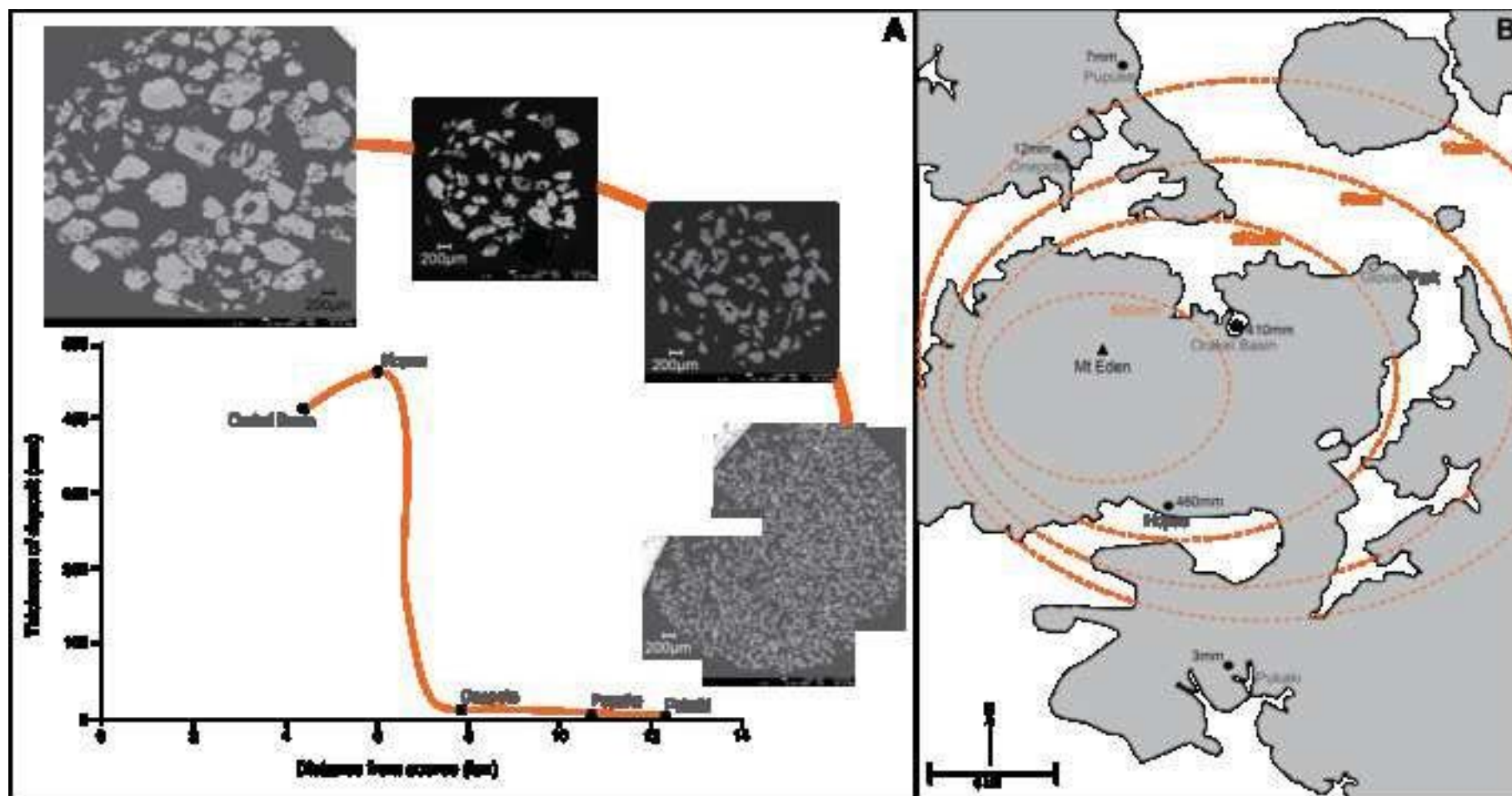
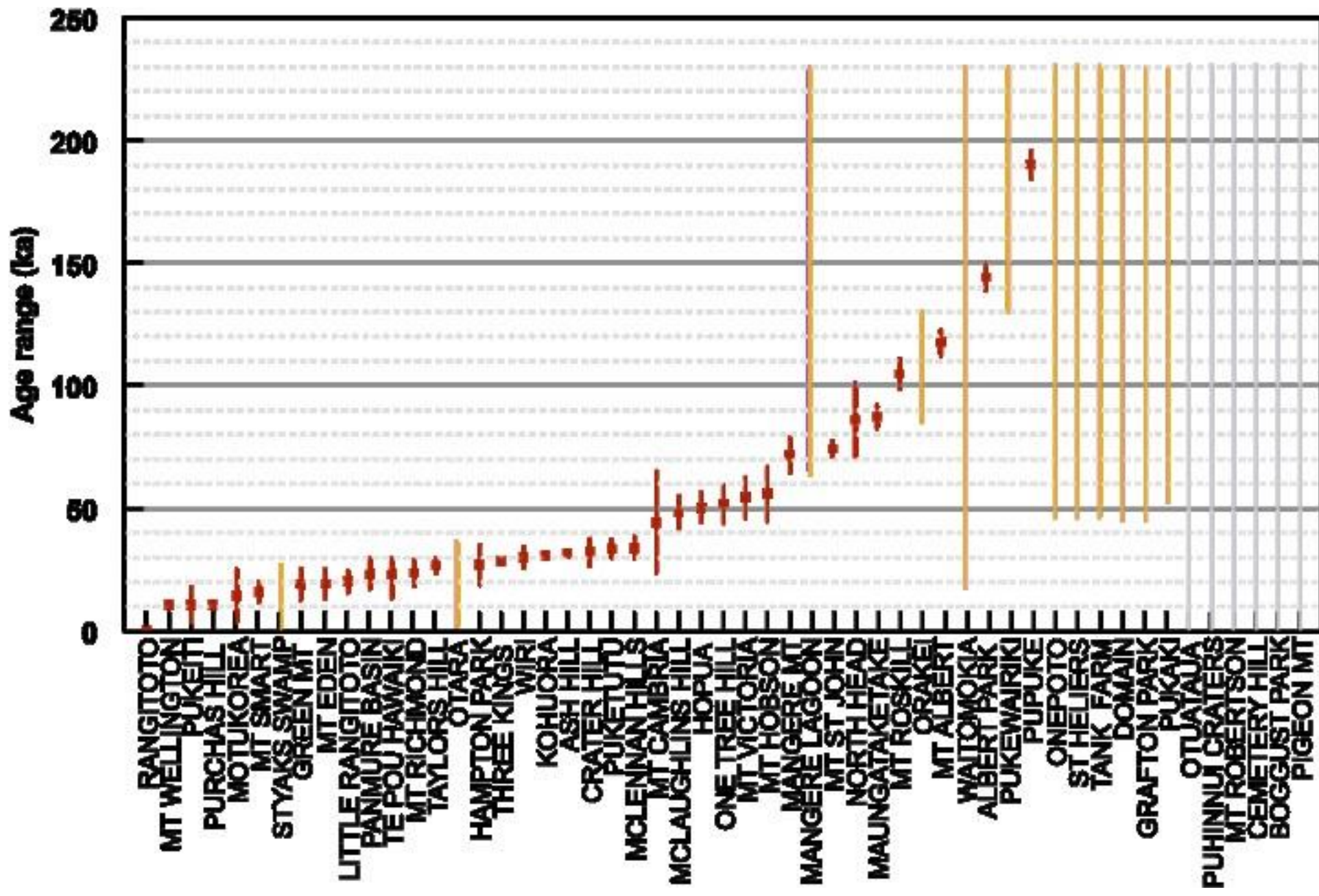
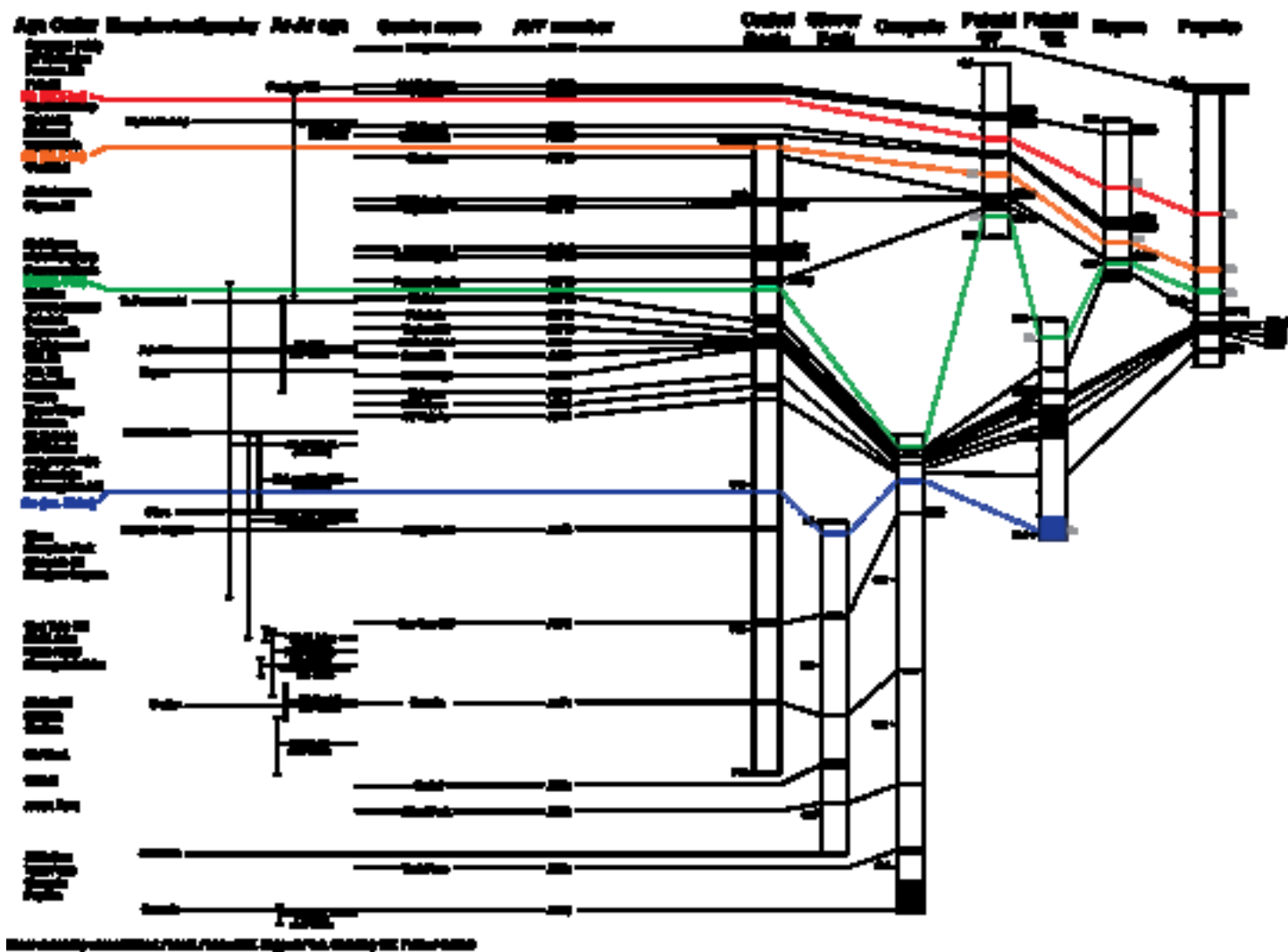
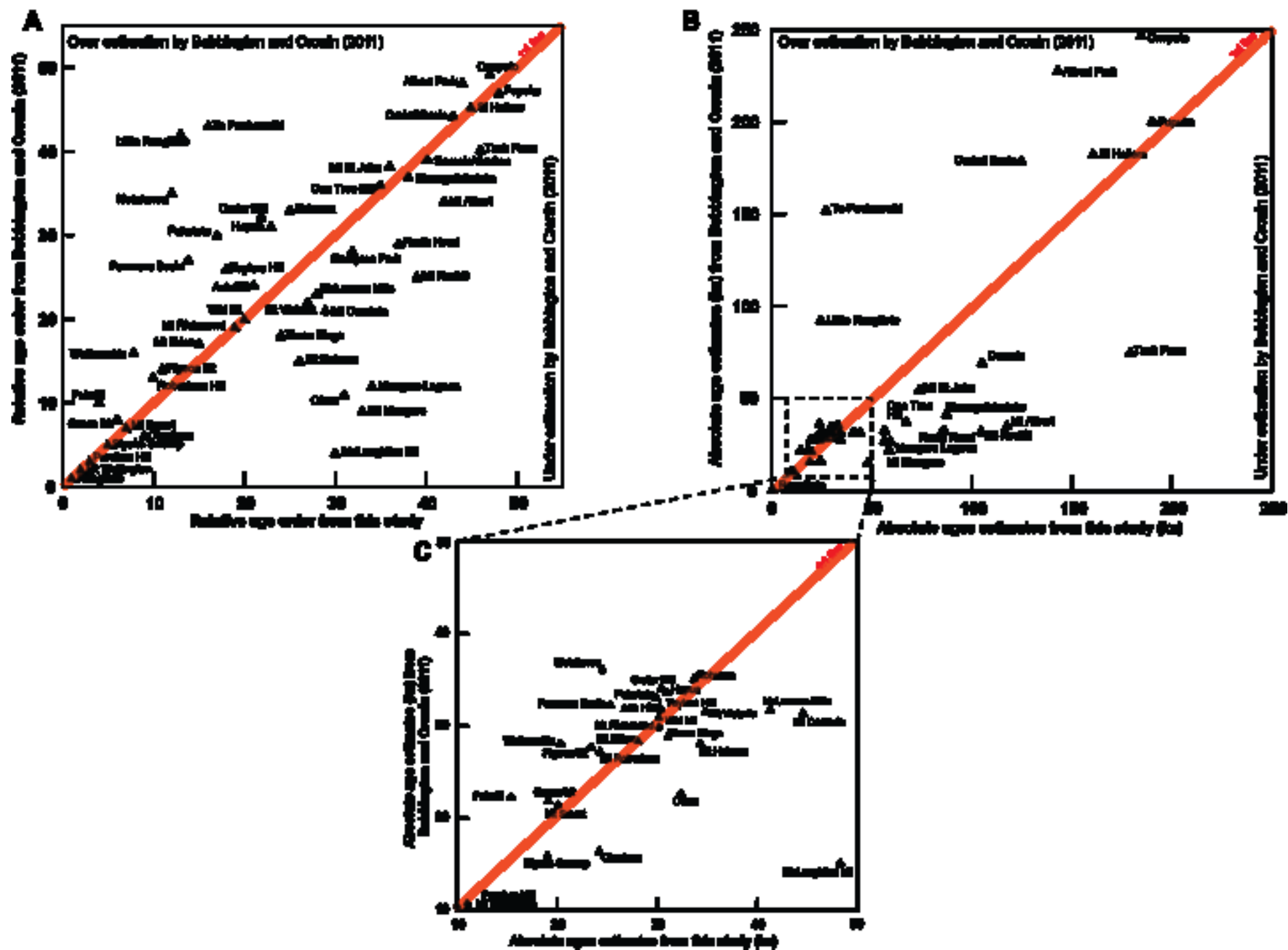


Figure 12







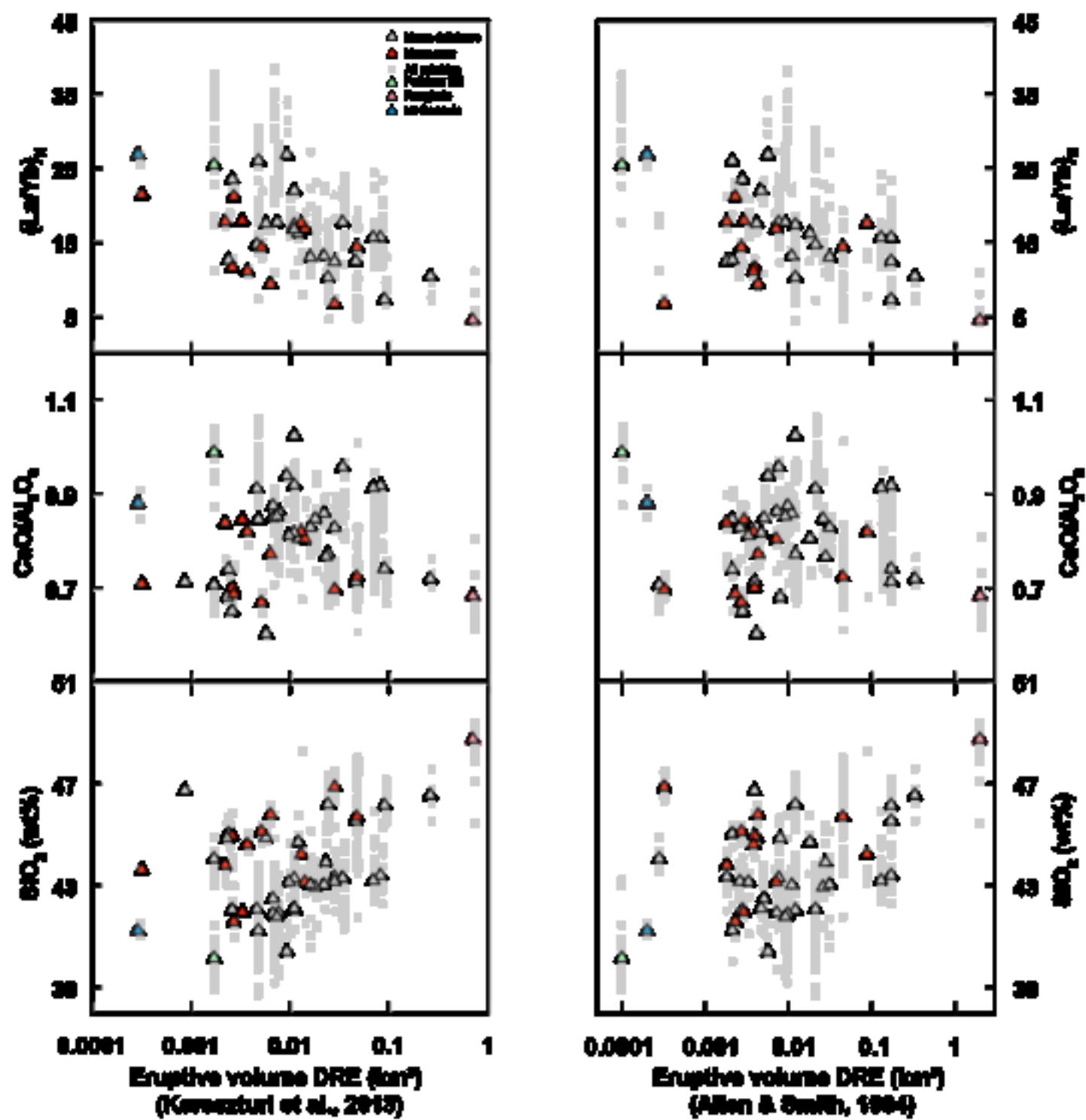


Table 1

Centre	Current whole rock data		References
	Major	Trace	
ALBERT PARK	4	4	McGee, 2012; Smith unpub data
ASH HILL	0	0	
BOGGUST PARK	0	0	
CEMETERY HILL	0	0	
CRATER HILL	61	61	Smith et al. 2008
DOMAIN	19	7	Smith unpub data
GRAFTON PARK	10	10	DEVORA group unpub data
GREEN HILL	3	1	Miller, 1996
HAMPTON PARK	4	0	Miller, 1996
HOPUA	1	1	Smith unpub data
KOHUORA	0	0	
LITTLE RANGITOTO	17	1	Franklin, 1999; Smith unpub data
MANGERE LAGOON	0	0	
MANGERE MT	7	2	Miller, 1996
MAUNGATAKETAKE	23	23	Smith unpub data
MCLAUGHLINS HILL	1	0	Heming and Barnet, 1986
MCLENNAN HILLS	6	3	Miller, 1996
MOTUKOREA	53	53	Bryner, 1991; McGee, 2012, McGee et al. 2012
MT ALBERT	2	4	Smith unpub data
MT CAMBRIA	1	1	Smith unpub data
MT EDEN	29	17	Eade, 2009; McGee, 2012
MT HOBSON	10	2	Smith unpub data
MT RICHMOND	6	3	Eade, 2009; McGee, 2012; Smith unpub data
MT ROSKILL	3	2	McGee, 2012
MT SMART	2	2	McGee, 2012; Smith unpub data
MT ST JOHN	22	13	Franklin, 1999; Eade, 2009
MT VICTORIA	4	2	Smith unpub data
MT WELLINGTON	34	34	McGee, 2012, McGee et al. 2013
NORTH HEAD	6	5	Smith unpub data
ONE TREE HILL	8	4	Eade, 2009; Smith unpub data
ONEPOTO	0	0	
ORAKEI	41	21	Franklin, 1999; Smith unpub data
OTARA	12	0	Miller, 1996; McGee, 2012
OTUATAUA	1	1	Heming unpub data
PANMURE BASIN	22	21	Smith unpub data
PIGEON MT	1	1	Smith unpub data
PUHINUI CRATERS	0	0	
PUKAKI	2	2	Zawalna-Geer, 2012
PUKEITI	1	1	Smith unpub data
PUKEKIWIIRIKI	4	3	Smith unpub data
PUKETUTU	23	13	Miller, 1996; McGee, 2012
PUPUKE	51	51	Spargo, 2007

PURCHAS HILL	27	27	McGee 2012; McGee et al. 2013
RANGITOTO	55	55	Hookway, 2000; Needham et al. 2011
ROBERTSON HILL	0	0	
ST HELIERS	1	1	Smith unpub data
STYAKS SWAMP	0	0	
TANK FARM	0	0	
TAYLOR'S HILL	3	3	McGee, 2012; Smith unpub data
TE POU HAWAIKI	13	0	Franklin, 1999
THREE KINGS	36	35	Eade, 2009; Smith unpub data
WAITOMOKIA	9	9	McGee, 2012
WIRI	12	12	McGee, 2012; McGee et al. 2013
<b>TOTALS</b>	<b>650</b>	<b>511</b>	

This study whole rock data	
Major	Trace

Surface exposure currently non-existent

3	3
---	---

new centre

2

4

Surface exposure currently non-existent

Surface exposure currently non-existent

6	6
---	---

3

5	5
---	---

3

1	1
---	---

5	5
---	---

5	5
---	---

6	6
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6	6
---	---

1

2

2	2
---	---

5

5	5
---	---

5	5
---	---

new centre

4	4
---	---

4	4
---	---

2

1	1
---	---



2	2
4	4
2	2

Surface exposure currently non-existent

Surface exposure currently non-existent

6	6
5	5

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**77**

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**99**

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Table 2

Table\*. Over view of current volcanoes identified in the Auckland Volcanic Field, their age, relative age and Dense Rock Equivalent (DRE) values and geochemical analyses.

Centre name	Eruption types <sup>a</sup>	Age estimate (ka)			Method	Method reference	Relative ages and relationships based on morphology	DRE volumes x10 <sup>6</sup> m <sup>3</sup>			
		min 2sd	mean	max 2sd				Total	Tuff	Scoria Cone	Tephra
ALBERT PARK	A,B,C	141.3	146.9	152.5	Ar-Ar	Leonard et al. 2017		27.8	0.82	0.01	0.43
ASH HILL	A	31.4	31.8	32.2	14C	Hayward, 2008	older than Wiri Mt <sup>b</sup>	0.076	0.05	0.00	0.03
BOGGUST PARK	A? (new)							0.32	0.18	0.00	0.09
CEMETERY HILL	(new)							0.24	0.14	0.00	0.07
CRATER HILL	A,B,C	26.7	32.1	37.5	Ar-Ar	Cassata et al. 2008	Mono Lake' p.mag excursion <sup>l</sup> , younger than Kohuora <sup>a</sup>	24.5	5.88	0.76	4.09
DOMAIN	A,B	52.0					younger than Grafton Park <sup>h</sup> one of the older centres in the AVF <sup>g</sup>	11.4	4.06	0.06	2.11
GRAFTON PARK	A,B	52.0					older than Domain <sup>h</sup> one of the older centres in the AVF <sup>g</sup>	11.4	4.06	0.06	2.11
GREEN MT	A,B,C	13.0	19.6	26.2	Ar-Ar	Leonard et al. 2017	older than Styaks Swamp <sup>a</sup>	12.2	0.36	1.50	2.43
HAMPTON PARK	A,B,C	37.0	55.0	73.0	Ar-Ar	Cassata et al. 2008	unusual p.mag orientation <sup>l</sup> , just older than Otara <sup>a</sup>	2.41	0.11	0.40	0.65
HOPUA	A	45.2	51.6	58.0	Ar-Ar	Leonard et al. 2017	younger than One Tree Hill <sup>a</sup>	0.86	0.31	0.00	0.15
KOHUORA	A	32.0	33.0	34.0	14C	Lindsay et al. 2011	older than Crater Hill <sup>a</sup> , contains Kawakawa/Oruanui tephra (>25.4 ka) <sup>h</sup>	7.24	5.10	0.00	2.55
LITTLE RANGITOTO	B,C	16.3	20.7	25.1	Ar-Ar	Leonard et al. 2017	younger than Orakei <sup>a</sup>	1.71	0.00	0.50	0.75
MANGERE LAGOON	A,B	63.1					just older than Mangere Mt <sup>k</sup>	2.04	0.71	0.01	0.37
MANGERE MT	B,C	63.1	70.3	77.5	Ar-Ar	Leonard et al. 2017	just younger than Mangere Lagoon <sup>k</sup> , younger than One Tree Hill <sup>a</sup>	46.2	0.00	15.01	22.51
MAUNGATAKETAKE	A,B,C	84.1	88.9	93.7	Ar-Ar	Leonard et al. 2017	sea cut platform from last interglacial <sup>a</sup>	33.6	4.40	0.87	3.51
MCLAUGHLINS HILL	A,B,C	41.8	48.2	54.6	Ar-Ar	Leonard et al. 2017	older than Wiri Mt <sup>a,b</sup>	7.58	0.51	0.43	0.90
MCLENNAN HILLS	A,B,C	29.9	34.7	39.5	Ar-Ar	Leonard et al. 2017	Laschamp p.mag excursion <sup>l</sup> , older than Mt Richmond <sup>a</sup>	21.9	0.42	3.79	5.90
MOTUKOREA	A,B,C	2.3	14.3	26.3	Ar-Ar	Leonard et al. 2017		4.56	0.66	1.31	2.30
MT ALBERT	A,B,C	113.6	119.2	124.8	Ar-Ar	Leonard et al. 2017	older than Mt Eden and Mt Roskill <sup>a</sup>	22.9	0.35	3.03	4.72
MT CAMBRIA	B,C	20.1	42.3	64.5	Ar-Ar	Leonard et al. 2017		0.29	0.00	0.29	0.44
MT EDEN	B,C	14.6	21.2	27.8	Ar-Ar	Leonard et al. 2017	much younger than Mt St John, younger than Three Kings, Mt Hobson <sup>a</sup> , One Tree Hill and Domain <sup>g</sup>	89.8	0.00	5.94	8.92
MT HOBSON	B,C	45.3	56.9	68.5	Ar-Ar	Leonard et al. 2017	older than Three Kings <sup>a</sup>	6.68	0.00	1.20	1.80
MT RICHMOND	A,B	24.7	34.3	43.9	Ar-Ar	Leonard et al. 2017	Mono Lake' p.mag excursion <sup>l</sup> , younger than McLennan Hills <sup>a</sup> , older than Okaia tephra (28.6 ka) <sup>d</sup>	5.67	1.17	3.04	5.14
MT ROBERTSON	A,B							2.72	1.01	0.24	0.87
MT ROSKILL	A,B,C	99.1	105.3	111.5	Ar-Ar	Leonard et al. 2017	post-Blake p.mag excursion <sup>l</sup> , younger than Mt Albert <sup>a</sup>	14.4	0.02	1.37	2.07
MT SMART	A,B,C	12.8	16.4	20.0	Ar-Ar	Leonard et al. 2017	younger than One Tree Hill <sup>a</sup>	13.4	0.00	2.34	3.52
MT ST JOHN	B,C	71.9	75.3	78.7	Ar-Ar	Leonard et al. 2017	much older than Mt Eden and Three Kings <sup>a</sup>	28.1	0.00	0.40	0.60
MT VICTORIA	B,C	42.8	57.6	72.4	Ar-Ar	Leonard et al. 2017		4.81	0.00	2.58	3.87
MT WELLINGTON	B,C	9.3	10.3	11.3	14C	Lindsay et al. 2011	just younger than Purchas Hill <sup>a</sup>	82.3	1.93	3.02	5.49
NORTH HEAD	A,B	72.3	87.5	102.7	Ar-Ar	Leonard et al. 2017	raised sea levels ca. 128-116 ka <sup>l</sup>	2.65	1.12	0.04	0.61
ONE TREE HILL	B,C	45.2	52.8	60.4	Ar-Ar	Leonard et al. 2017	older than Hopua, Mt Hobson, Mt Eden, Mt Smart, Three Kings, One Tree Hill <sup>a</sup>	260	0.00	5.70	8.56
ONEPOTO	A	52.0					similar age to Pupuke and Tank Farm <sup>a</sup>	2.62	1.54	0.00	0.77
ORAKEI	A	85.0		130.0			sed. rate ages of tephra horizons	6.70	3.77	0.00	1.89
OTARA	A,B,C	0.0		73.0			morphostratigraphy	2.30	0.11	0.70	1.10
OTUATAUA	A,B,C							6.30	0.00	0.99	1.49
PANMURE BASIN	A,B	17.5					Rerewhakaaitu tephra in drill core	7.44	4.65	0.30	2.77
PIGEON MT	A,B,C							3.31	1.33	0.28	1.08
PUHINUI CRATERS	A? (new)							-	-	-	-
PUKAKI	A	52.0					Core extent	9.19	7.10	0.00	3.55
PUKEITI	B,C	4.2	11.4	18.6	Ar-Ar	Leonard et al. 2017	younger than Otuaataua <sup>a</sup>	3.70	0.00	0.44	0.66
PUKETUTU	B,C	29.8	33.6	37.4	Ar-Ar	Cassata et al. 2008	paleomag excursion (32.4±0.3ka)	11.0	3.00	2.15	4.72
PUKEWAIRIKI	A,C	130.0					morphostratigraphy	17.5	2.29	0.00	1.15
PUPUKE	C,B,A	187.6	193.2	198.8	Ar-Ar	Leonard et al. 2017	similar age to Tank Farm and Onepoto <sup>a</sup>	46.7	20.11	0.00	10.06
PURCHAS HILL	A,B	10.7	10.9	11.1	14C	Lindsay et al. 2011	just older than Mt Wellington <sup>a</sup>	1.68	0.21	0.03	0.14
RANGITOTO 2	A,B,C	0.494	0.504	0.514	14C	Needham et al. 2011	youngest in the field <sup>a</sup>	699	4.65	41.60	64.73
RANGITOTO 1	A,B,C	0.539	0.553	0.567	14C; Needham et al. 2	Needham et al. 2011					
ST HELIERS	A,	52.0					Rotoehu Tephra in drill core	2.20	1.23	0.00	0.62
STYAKS SWAMP	A,	0.0		24.5			morphostratigraphy	0.37	0.25	0.00	0.12
TANK FARM	A	52.0					morphostratigraphy	5.87	4.13	0.00	2.06
TAYLORS HILL	A,B,C	24.2	27.4	30.6	Ar-Ar	Leonard et al. 2017	Mono Lake' p.mag excursion <sup>l</sup>	5.07	0.47	0.18	0.51
TE POU HAWAIKI	B	14.6					morphostratigraphy	28.1	0.00	0.08	0.12
THREE KINGS	A,B,C	27.7	28.7	29.7	14C	Lindsay et al. 2011	younger than One Tree Hill, Mt St John, Mt Hobson, older than Mt Eden <sup>a</sup>	69.3	0.00	3.00	4.51
WAITOMOKIA	A,B	15.6					morphostratigraphy	9.79	2.30	0.11	1.31
WIRI	A,B,C	25.6	30.2	34.8	Ar-Ar	Cassata et al. 2008	Mono Lake' p.mag excursion <sup>l</sup> , younger than Ash Hill <sup>b</sup>	16.4	0.08	0.86	1.34

Table 3

Source	abv	Age	error	ref	interpreted age (yr)	error	95% confidence limits	
AVF24 [P48]*	Ra2	504	5	a				
AVF24 [P49]*	Ra1	553	7	a				
Taupo	Tp	1,718	30	b				
Tuhua	Tu	6,577	547	b				
Mamaku	Ma	7,940	257	b				
Rotoma	Ro	9,423	120	b				
AVF23					9,950	300	9,650	10,240
Opepe	Op	9,991	160	b				
Waiohau	Wh	14,009	155	b				
AVF22					15,310	650	14,660	15,960
Rotorua	Rr	15,635	412	b				
Rerewhakaiaitu	Rk	17,496	462	b				
AVF21					20,080	100	19,080	21,070
AVF20					20,310	142	18,890	21,740
Okareka	Ok	21,858	290	b				
AVF19					24,200	880	23,320	25,090
AVF18					24,260	400	23,860	24,650
AVF17					23,350	350	23,000	23,700
AVF15					24,410	290	24,120	24,700
AVF14					24,550	290	24,270	24,840
Te Rere	Tr	25,171	964	b				
AVF16					25,230	860	24,370	26,090
AVF13					25,230	310	24,920	25,540
Kawakawa/Oruanui	Kk	25,358	162	b				
AVF12					28,030	260	27,760	28,290
Okaia	O	28,621	1428	b				
AVF11					29,770	2240	27,530	32,010
AVF10					30,200	120	30,080	30,320
AVF9					30,200	2080	28,120	32,280
AVF8					30,400	400	30,000	30,810
AVF7					31,040	900	30,140	31,940
AVF6					33,710	1160	32,550	34,870
AVF5					34,200	860	33,340	35,070
AVF4					34,780	2000	32,780	36,780
Maketu	Mk	36,320	575	c				
Tahuna	Ta	39,268	1193	c				
Rotoehu	Re	52,000	7000	d				
AVF3					59,230	10,230	49,000	69,460
AVF2					67,200	6,250	60,950	73,450
AVF1					106,170	4,300	101,870	110,470
AVFa					126,150	3,320	122,830	129,470
AVFb					144,870	2,400	142,470	147,270
AVFc					181,430	580	180,850	182,010
AVFd		193,200	2,800	e				

Table 4

Core	Sample name	New Horizon	Depth (m)	Thickness (mm)	Average age (ka)	Proposed Correlate(s)	Confidence level	Age	Chem	Scale	Location	Alternative(s)	Confidence level	Age	Chem	Scale	Location
<b>East Rarohakaiti (17.8ka)</b>																	
Pupuke	T21-1485829	24	57.90	22	0.6	Rangitoto	1	✓	✓	✓	✓	-	-	-	-	-	-
Hopua	T42-H1-258839	23	38.95	3	9.95 ± 0.3	Mt Wellington	1	✓	✓	✓	✓	-	-	-	-	-	-
Pukaki	T14.47.72m	22	47.72	1.0	15.31 ± 0.65	Pukaki	2	✓	✓	✓	✓	-	-	-	-	-	-
<b>Rarohakaiti to Okarua (17.5 - 21.9ka)</b>																	
Pukaki	AT202 43.15m	21	43.15	3.0	20.08 ± 0.1	Mt Smart	2	✓	✓	✓	✓	Mt Eden	3	✓	✓	✓	✓
Hopua	T5-2-H1-185885(1-5885)	21	45.17	290								PalmuraBasin	2	✓	✓	✓	✓
Pukaki	AT210 49.17m	20	49.17	2.0	20.3 ± 0.14	Wakomoko	2	✓	✓	✓	✓	BoggettPark	3	?	✓	✓	✓
Hopua	T6-H1-1205807(1-5885)	20	45.51	255								Mt Robertson	3	?	✓	✓	✓
Pukaki	T43 51.05	19	51.05	1.0	24.2 ± 0.88	Okarua	3	?	✓	✓	✓	Okarua	3	?	✓	✓	✓
												Win Mt	3	✓	✓	✓	✓
												BoggettPark	3	?	✓	✓	✓
												Mt Robertson	3	?	✓	✓	✓
Pukaki Basin	PNB 476.44.8.99	18	44.22	8								PalmuraBasin	3	?	✓	✓	✓
Hopua	T5-H1-32-58859	18	47.81	40	24.28 ± 0.4	Mt Robertson	3	?	✓	✓	✓	BoggettPark	3	?	✓	✓	✓
Pukaki	T45 51.19	18	51.19	0.5									?	✓	✓	✓	✓
Pukaki Basin	PNB 476.44.8.99 + 44.8.101	17	44.65	5	23.35 ± 0.35	Pigeon Mt	3	?	✓	✓	✓	Little Rangitoto	3	✓	✓	✓	✓
												Mt Eden	3	✓	✓	✓	✓
												Mt Robertson	3	✓	✓	✓	✓
												PalmuraBasin	3	✓	✓	✓	✓
												Taylor Hill	3	✓	✓	✓	✓
Pukaki Basin	PNB 476.44.7.916	15	47.72	12	24.41 ± 0.29	Mokuroa	2	✓	✓	✓	✓	Pigeon Mt	3	?	✓	✓	✓
												Mt Eden	3	✓	✓	✓	✓
Okarua Basin	OB1 433-2-48-12(48-128)	14	48.13	12	24.55 ± 0.29	Little Rangitoto	1	✓	✓	✓	✓	Mt Robertson	3	✓	✓	✓	✓
Okarua Basin	OB1 433-2-48-19(48-278)	newB	48.28	10		?											
Okarua Basin	OB1 433-4-45-14(49-46)	newA	49.46	45		?											
Okarua Basin	OB1 433-3-50-089(49-554)	13	50.09	160	25.23 ± 0.86	Palmura Basin	1	✓	✓	✓	✓	Mt Eden	2	✓	✓	✓	✓
Pukaki	T42-45 51.52	13 (18)	51.52	50.0								Little Rangitoto	2	✓	✓	✓	✓
												Mt Richmond	2	✓	✓	✓	✓
												Taylor Hill	2	✓	✓	✓	✓
<b>Okarua to Roroahu (25.4 - ca. 52 ka)</b>																	
Okarua Basin	OB1 436-2-52-817(53-029)	12	53.03	410								Three Kings	2	✓	✓	✓	✓
Okarua	OK2 44-39.06	12	38.09	12								Te Pahi Haeaki	3	✓	✓	✓	✓
Pukaki	54.355m	12	54.36		28.03 ± 0.26	Mt Eden	1	✓	✓	✓	✓						
Pupuke	P2358947	12	67.59	7													
Hopua	T6-3-H1-3858876	12	48.80	460													
Pukaki	c. 60.6m	11	55.34	<10	29.8 ± 2.2	Pukakū	2	✓	✓	✓	✓	Kohora	3	✓	✓	✓	✓
												Okarua	3	✓	✓	✓	✓
												Mt Robertson	3	✓	✓	✓	✓
												Win Mt	3	✓	✓	✓	✓
Pukaki Basin	PNB 476.44.2.119(44.716)	10	54.21	407	30.2 ± 0.12	Taylor's Hill	2	✓	✓	✓	✓	Palmura Basin	2	✓	✓	✓	✓
Okarua	OK2 44-39.47	10	39.47	15								Pigeon Mt	3	?	✓	✓	✓
Pupuke	T18-7-P2358951	10	68.09	3								Three Kings	3	✓	✓	✓	✓
												Mt Cambria	3	✓	✓	✓	✓
Pupuke	8976880	9	68.15	6	30.2 ± 2.08	Mt Richmond	3	✓	✓	✓	✓	Mt Cambria	3	✓	✓	✓	✓
												Mt Richmond	3	✓	✓	✓	✓
												Hopua	3	✓	✓	✓	✓
Pukaki Basin	PNB 476.44.7.77	8	54.27	45													
Pupuke	P2358953	8	68.24	20	30.4 ± 0.4	Crater Hill	2	✓	✓	✓	✓	Kohora	2	✓	✓	✓	✓
Pukaki	c. 56.4m	8	56.40	ca. 720													
Pukaki Basin	PNB 476.44.7.104 (44.777)	7	54.34	20													
Okarua	OK2 44-39.90(39.914)	7	39.90	20	31.04 ± 0.9	Three Kings	1	✓	✓	✓	✓						
Pukaki	c. 66.8	7	56.80														
Pupuke	T19-3-P2358954	7	68.49	2													
Pukaki	c. 67.6	6	67.10	ca. 600	33.71 ± 1.16	Kohora	2	✓	✓	✓	✓	Crater Hill	2	✓	✓	✓	✓
												Pukakū	2	✓	✓	✓	✓
												Win Mt	3	✓	✓	✓	✓
												Mt Robertson	3	✓	✓	✓	✓
Pukaki Basin	PNB 476.44.7.147(47.44)	5	57.34	110	34.2 ± 0.86	Mt Hobson	3	✓	✓	✓	✓	Little Rangitoto	3	✓	✓	✓	✓
												Mt Cambria	3	✓	✓	✓	✓
												Mt St John	3	✓	✓	✓	✓
												Mt Victoria	3	✓	✓	✓	✓
												North Head	3	✓	✓	✓	✓
												Palmura Basin	3	✓	✓	✓	✓
												Taylor Hill	3	✓	✓	✓	✓
Okarua Basin	OB1 436-5-58-11(58.07)	4	58.11	41	54.78 ± 7.9	Mt Victoria	1	✓	✓	✓	✓	Little Rangitoto	3	✓	✓	✓	✓
Pupuke	P3358960	4	69.32	15								Mt Hobson	3	✓	✓	✓	✓
												Mt St John	3	✓	✓	✓	✓
												North Head	3	✓	✓	✓	✓
												Taylor Hill	3	✓	✓	✓	✓
<b>Pre Roroahu ca. 52 ka</b>																	
Okarua Basin	OB1 445-5-67-039(61.17)	3	67.04	41	59 ± 10.0	Mangere Mt	2	✓	✓	✓	✓	Mt Cambria	3	✓	✓	✓	✓
												Mt Hobson	3	✓	✓	✓	✓
												Mt Victoria	3	✓	✓	✓	✓
												One Tree Hill	3	✓	✓	✓	✓
Okarua Basin	OB1 450-2-73-055	2	73.56	510													
Glover Park	GP16-10-38(10.8)	2	10.60	60	67 ± 6.0	One Tree Hill	1	✓	✓	✓	✓						
Okarua	OK2 46-2-43.86	2	43.86	4													
Okarua Basin	OB1 454-3-80-047	1	80.05	100								Mt Ruakū	3	✓	✓	✓	✓
Okarua	OK2 46-1-11-33	1	54.30	15	106 ± 4.0	Domains/Grafton	3	?	✓	✓	✓	North Head	3	✓	✓	✓	✓
Glover Park	GP16-17-52(16.18)	1	18.15	12													
Glover Park	GP124 -20.78-21.0	AVFb	21.00	40	126 ± 3.0	Okarua	2	✓	✓	✓	✓	Domains/Grafton	3	✓	✓	✓	✓
												Mt Albert	3	✓	✓	✓	✓
												Mt Ruakū	3	✓	✓	✓	✓
												North Head	3	✓	✓	✓	✓
Okarua	OK2 418-62.36	AVFb	62.00	45	145 ± 2.0	Albert Park	2	✓	✓	✓	✓	Domains/Grafton	2	?	✓	✓	✓
Glover Park	GP140 -23.67	AVFb	23.67	10													
Okarua	OK2 421-66.68	AVFb	66.68	270	181 ± 0.6	Tank Farm	1	✓	✓	✓	✓	Pupuke	3	✓	✓	✓	✓

Table 5

Centre (eruption)	AVF#	Confidence level	DRE km <sup>3</sup> (2sf)	Orakei Basin			Glover Park			Onepoto			Pukaki		Hopua			Pupuke		
				Distance to (km)	Horizon thickness (mm)	shard size (µm)	Distance to (km)	Horizon thickness (mm)	shard size (µm)	Distance to (km)	Horizon thickness (mm)	shard size (µm)	Distance to (km)	Horizon thickness (mm)	Distance to (km)	Horizon thickness (mm)	shard size (µm)	Distance to (km)	Horizon thickness (mm)	shard size (µm)
Rangitoto	AVF24	1	0.70		n/a			n/a		n/a			n/a		n/a		8.4	22	>200	
Mt Wellington	AVF23	1	0.082		n/a			n/a		n/a		10.4	1	7.0	3	>200			n/a	
Little Rangitoto	AVF14	1	0.0017	0.9	12	>200		n/a		n/a			n/a		n/a				n/a	
Panmure Basin	AVF13	1	0.0074	5.2	160	100-200		n/a		n/a			n/a		n/a				n/a	
Mt Eden	AVF12	1	0.090	4.4	410	>200		n/a	7.9	12	100-200	12.3	3	6.0	460	>200	10.7	7	50-100	
Three Kings	AVF7	1	0.069	6.5	20	100-200		n/a	10.6	12	50-100	10.1	2		n/a		13.5	2	50-100	
One Tree Hill	AVF2	1	0.26	4.6	510	>200	9.6	60	>200	10.8	4	100-200		n/a		n/a			n/a	
Tank Farm	AVFc	1	0.0059		n/a			n/a	0.6	270	>200		n/a		n/a				n/a	
Pukeiti	AVF22	2	0.0037		n/a			n/a		n/a		4.7	1		n/a				n/a	
Mt Smart	AVF21	2	0.013		n/a			n/a		n/a		7.0	3	2.7	290	>200			n/a	
Waitomokia	AVF20	2	0.010		n/a			n/a		n/a		3.7	2	5.5	235	100-200			n/a	
Motukorea	AVF15	2	0.0046	8.4	12	50-100		n/a		n/a			n/a		n/a				n/a	
Puketutu	AVF11	2	0.018		n/a			n/a		n/a		6.0	<10		n/a				n/a	
Taylor's Hill	AVF10	2	0.0051	5.1	407	>200		n/a	12.4	15	100-200		n/a		n/a		13.2	3	100-200	
Crater Hill	AVF8	2	0.024	13.2	45	100-200		n/a		n/a		1.5	720		n/a		23.5	20	<50	
Kohuora	AVF6	2	0.0072		n/a			n/a		n/a		2.9	500		n/a				n/a	
Orakei Basin	AVFa	2	0.0067		n/a		5.4	40	100-200		n/a		n/a		n/a				n/a	
Albert Park	AVFb	2	0.028		n/a		8.8	10	50-100	5.0	45	100-200		n/a		n/a			n/a	

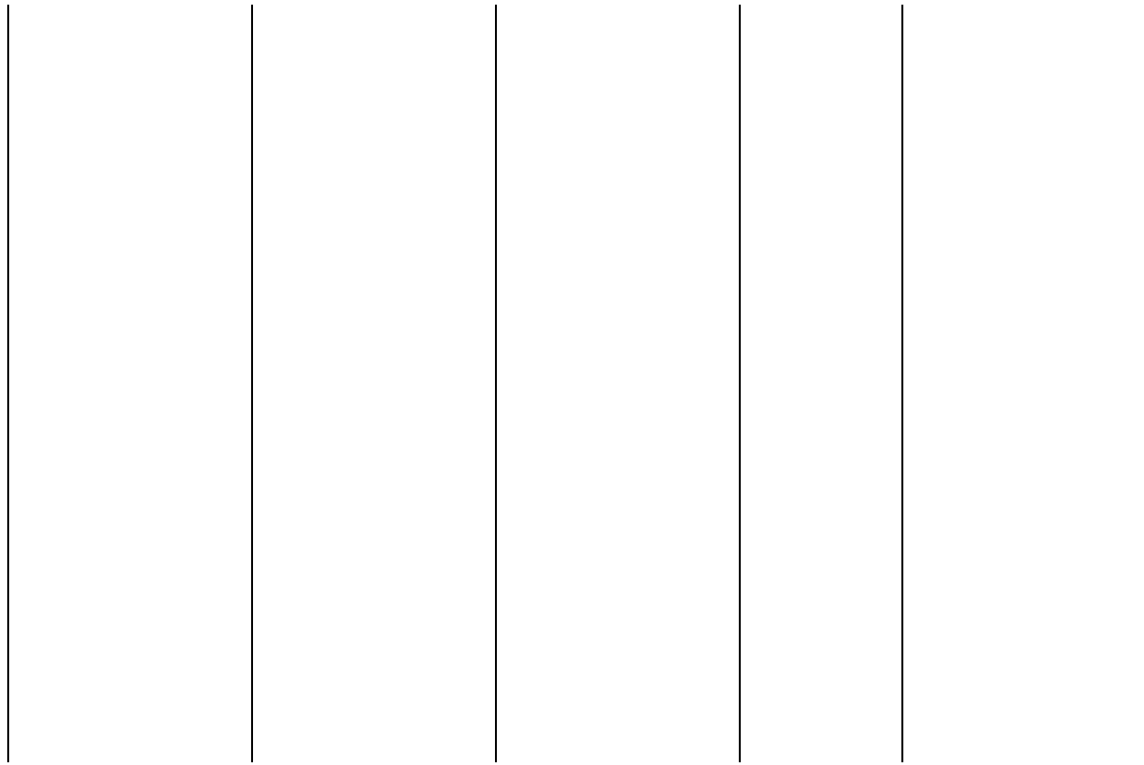


Table 6

Centre Name	Region	Total DRE volume (km <sup>3</sup> )	Tephra thickness (cm)	Dispersal distance (km)
Paricutin	Michoacán-Guanajuato volcanic field, Mexico	2.5	25	7
Sunset Crater	San Francisco volcanic field, Arizona	0.58	10	20
<b>One Tree Hill</b>	<b>Auckland volcanic field, New Zealand</b>	<b>0.26</b>	<b>6</b>	<b>9.6</b>
Mt Gambier	Newer Volcanics province, south-eastern Australia	0.198	≤5	10 to 12
Lanthrop Wells	Southwestern Nevada volcanic field	0.12	1	10
Cerro Negro*	Nicaragua	0.16	0.5	16
<b>Three Kings</b>	<b>Auckland volcanic field, New Zealand</b>	<b>0.069</b>	<b>2</b>	<b>6.5</b>
Marcath Volcano	Lunar Crater volcanic field, Central Nevada	0.06	2	7
<b>Orakei Basin</b>	<b>Auckland volcanic field, New Zealand</b>	<b>0.0067</b>	<b>0.4</b>	<b>5.4</b>

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**Reference**

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Ort et al., 2008

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Ort et al., 2008

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**This study**

Lowe and Palmer, 2005; van  
Otterloo and Cas, 2013

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Valentine et al., 2008

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Hill et al., 1998

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**This study**

Johnson et al., 2014

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**This study**

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Table 7

Relative Order	Centre Name	Mean age (t) in ka	Error (1sd)	Age ref.	Time relationship		Distance r d+1 (km)
					t+1 (ka)	t+2 (ka)	
0	Rangitoto 2	0.50	±0.05	b	-	-	-
1	Rangitoto 1	0.55	±0.07	b	0.05	-	0.1
2	Mt Wellington	10.00	±0.5	a	9.4	9.5	11.7
3	Purchas Hill	10.90	±0.14	b	0.9	10.3	0.6
4	Pukeiti	15.31	±0.65	a	4.4	5.3	13.4
<b>Rerewhakaaitu (ca. 17.5 ka)</b>							
5	Styaks Swamp	19.10	-	d	3.8	8.2	13.9
6	Green Mt	19.60	±3.3	c	0.5	4.3	0.6
7	Mt Smart	20.08	±0.1	a	0.5	1.0	8.1
8	Waitomokia	20.30	±0.14	a	0.2	0.7	7.6
<b>Okareka (ca. 21.9 ka)</b>							
9	Otuataua	24.20	±0.88	a	3.9	4.1	1.7
10	Mt Robertson	24.26	±0.4	a	0.1	4.0	8.8
11	Pigeon Mt	23.35	±0.35	a	-0.9	-0.8	8.5
12	Motukorea	24.41	±0.29	a	1.1	0.1	6.5
13	Little Rangitoto	24.55	±0.29	a	0.1	1.2	9.1
14	Panmure Basin	25.23	±0.86	a	0.7	0.8	4.8
<b>Oruanui/Kawakawa (ca. 25.4 ka)</b>							
15	Mt Eden	28.03	±0.26	a	2.8	3.5	8.3
16	Te Pou Hawaiki	28.53	-	d	0.5	3.3	0.7
17	Puketutu	29.80	±2.2	a	1.3	1.8	9.3
18	Taylor's Hill	30.20	±0.12	a	0.4	1.7	8.0
19	Mt Richmond	30.20	±2.08	a	0.0	0.4	8.0
20	Wiri Mt	30.20	±4.6	c	0.0	0.0	8.5
21	Ash Hill	30.70	-	d	0.5	0.5	1.0
22	Crater Hill	30.40	±0.4	a	-0.3	0.2	4.4
23	Hopua	31.00	-	d	0.6	0.3	6.5
24	Three Kings	31.04	±0.9	a	0.0	0.6	4.6
25	Kohuora	33.71	±1.16	a	2.7	2.7	11.5
26	Mt Hobson	34.20	±0.86	a	0.5	3.2	12.0
27	Mt Victoria	34.78	±2.0	a	0.6	1.1	6.6
28	McLennan Hills	41.30	±1.2	d	6.5	7.1	12.2
29	Mt Cambria	42.30	±11.1	c	1.0	7.5	12.5
30	McLaughlins Hill	48.20	±3.2	c	5.9	6.9	21.5
<b>Pre Rotoehu (ca. 52 ka)</b>							
31	Otara	56.5	-	d	8.3	14.2	8.7
32	Hampton Park	57.0	±16.0	c/d	0.5	8.8	0.5
33	Mangere Mt	59.0	±10.0	a	2.0	2.5	5.4
34	Mangere Lagoon	59.5	-	d	0.5	2.5	0.9
35	One Tree Hill	67.0	±6.0	a	7.5	8.0	8.6
36	Mt St John	75.3	±1.7	c	8.3	15.8	8.2
37	North Head	87.5	±7.6	c	12.2	20.5	6.8
38	Maungetaketake	88.9	±2.4	c	1.4	13.6	19.4

39	Mt Roskill	105.3	±3.1	c	16.4	17.8	9.1
40	Domain	106.0	±4.0	a	0.7	17.1	6.7
41	Grafton	106.5	-	d	0.5	1.2	0.3
42	Mt Albert	117.6	±5.2	c	11.1	11.6	5.6
43	Orakei	126.0	±3.0	a	8.4	19.5	8.6
44	Albert Park	145.0	±2.0	a	19.0	27.4	4.4
45	St Heliers	161.0	±18.0	d	16.0	35.0	8.8
46	Tank Farm	181.0	±1.0	a	20.0	36.0	11.4
47	Onepoto	187.6	-	d	6.6	26.6	0.6
48	Pupuke	193.2	±2.8	c	5.6	12.2	3.3

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**Undated centres**

Pukaki	>52.0
Pukewairiki	>130
Boggust Park	?
Cemetery Hill	?
Puhinui Craters	?

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**relationship**  
**d+2 (km)**

-  
11.8  
10.7  
12.8



13.4  
13.5  
8.1  
12.3



9.1  
7.1  
16.8  
13.9  
8.5  
9.1




4.0  
7.8  
9.9  
9.6  
9.2  
15.8  
8.5  
3.7  
11.5  
10.2  
7.5  
3.2  
17.4  
7.8  
0.4  
9.4



16.0  
6.5  
14.0  
6.3  
3.3  
7.3  
14.7  
12.7

11.5  
15.1  
6.4  
6.0  
3.6  
6.3  
5.4  
5.5  
11.4  
2.7





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