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1 **Trace element perspective into the ca. 2.1-billion-year-old shallow-marine microbial**
2 **mats from the Francevillian Group, Gabon**

3
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20 **Abstract**

21 The sedimentary fabrics of Precambrian mat-related structures (MRS) represent some of the
22 oldest convincing evidence for early life on Earth. The *ca.* 2.1 billion-year (Ga) old MRS in the
23 FB₂ Member of the Francevillian basin in Gabon has received considerable attention not only
24 because they contain remnants of microbial mats that colonized large areas in oxygenated,
25 shallow-marine settings, but they also contain evidence for ancient multicellular organisms that
26 thrived on these microbial mats using them as a food source. Despite these insights, what
27 remains lacking is a full characterization of the geochemical composition of the MRS to test
28 whether the bulk composition of fossilized MRS is distinct from the host sediments (sandstones
29 and shales). Here, we show that the trace element (TE) content of microbial textures belonging
30 to pyritized MRS, poorly pyritized MRS, and “elephant-skin” textures (EST) is highly variable
31 and differs from that of the host sediments. The poorly pyritized MRS contain a unique matrix
32 with embedded Ti- and Zr-rich minerals and syngenetically enriched in TE. The EST, some of
33 which are developed along the same stratigraphic horizon as the poorly pyritized MRS, display
34 a distinct distribution of TE-bearing heavy minerals, suggesting a local difference in physical
35 conditions during sedimentation. Similarly, high chalcophile-element (CE) content in pyritized
36 MRS relative to the host sediments of the FB₂ Member further points to local bacterially
37 influenced enrichments with high rates of microbial sulfate reduction during early diagenesis.
38 The geochemical relationship between the MRS and the Francevillian sediments (*e.g.*, FB, FC,
39 and FD formations) indicates that specific biological pathways for CE enrichments (*i.e.*,
40 microbially controlled accumulation) are not apparent. Our findings highlight bulk-rock TE
41 distinction between the 2.1-billion-year-old MRS and their host sediments, but also indicate
42 that environmental conditions, such as hydrodynamic regime and water-column redox
43 chemistry, may simply overwhelm any potential biological signal. Our data suggest that the
44 microbial impact may have only passively influenced TE enrichment in the studied sediments,
45 implying that TE concentrations in MRS are a poor biosignature. Importantly, this work
46 indicates that bulk TE geochemistry does not unveil specific microbiological processes in the
47 rock record, which is consistent with the observed patterns in modern analogues.

48

49 **Keywords**

50 Mat-related structures, elemental geochemistry, biological signal, Palaeoproterozoic, Gabon.

51 1. Introduction

52 The *ca.* 2.1 billion-year-old Francevillian sedimentary rocks record significant changes in the
53 Earth's ocean and atmosphere chemistry as they were deposited in the immediate aftermath
54 of the first substantial rise of atmospheric oxygen (Bekker, 2015a; Holland, 2002) and at the
55 time of Earth's longest and most extreme marine positive carbon isotope excursion (Bekker,
56 2015b; Canfield et al., 2013; El Albani et al., 2010). These sediments record the presence of
57 cm-sized macroscopic fossils with a wide range of characteristics, including lobate, elongated,
58 rod-shaped, and disc-shaped morphotypes (El Albani et al., 2014, 2010), as well as diverse
59 string-shaped structures that have been interpreted as evidence for ancient organism motility
60 (El Albani et al., 2019). In addition, benthic microbial mat communities prevailed in the
61 Francevillian sedimentary deposits (Aubineau et al., 2018; Reynaud et al., 2017) and are now
62 preserved as mat-related structures (MRS) with an exceptional diversity of mat growth patterns
63 (*e.g.*, "elephant-skin" texture, "fairy rings") and mat-protected features (*e.g.*, wrinkle structures)
64 (Aubineau et al., 2018). Based on the observed tufted microtextures and comparison with
65 modern analogues it has been suggested that shallow-marine conditions within the photic zone
66 allowed the development of phototrophic microbial mats composed of oxygenic
67 photosynthesizers such as cyanobacteria (Aubineau et al., 2018; Reynaud et al., 2017). Lastly,
68 these MRS have been shown to record an ancient metabolic activity through a potassium (K)-
69 controlled mineral transformation that occurred during diagenesis (Aubineau et al., 2019).

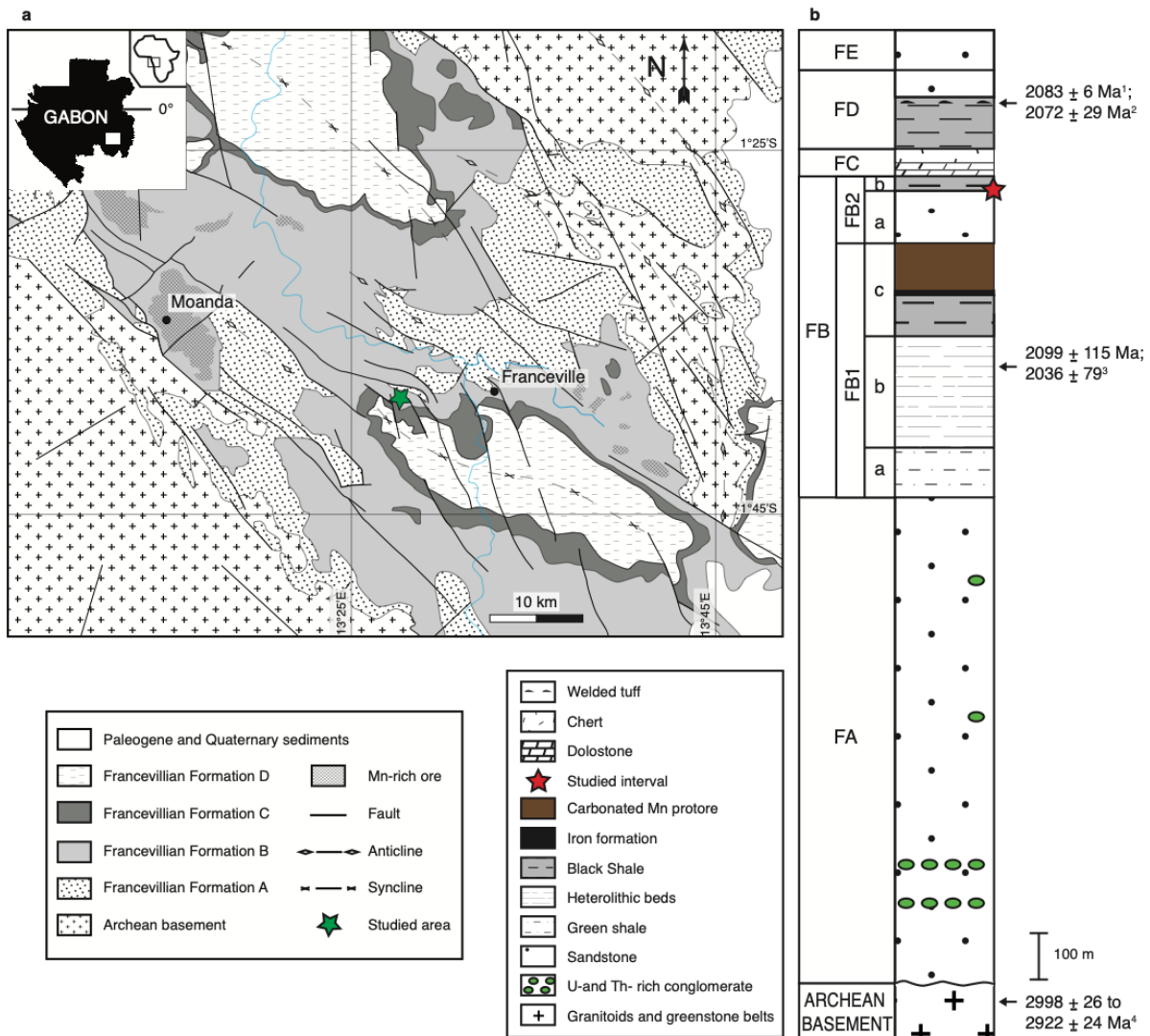
70 The observation of MRS in the Francevillian Group offers us the opportunity to further
71 ascertain whether specific chemical "biosignatures" unique to MRS might also be present. This
72 is crucial because organic microbial remnants, preserved as graphite, are common features in
73 early Archean rocks (Dodd et al., 2017; Mojzsis et al., 1996; Nutman et al., 2016; Rosing, 1999;
74 Schidlowski et al., 1979), but these have been subjected to deep burial diagenesis and
75 metamorphism, and as a consequence, without clear physical features of MRS, it is difficult to
76 assess their biogenicity. Indeed, it has been suggested that carbonaceous matter may have
77 been generated abiotically under hydrothermal conditions *via* a process similar to the Fischer–
78 Tropsch synthesis (Brasier et al., 2002; Lindsay et al., 2005; Van Zuilen et al., 2002).
79 Therefore, seeking chemical biosignatures for ancient microbial mats is a sensible way forward
80 because it is well documented that modern microbial mats, comprised of various bacterial
81 species and their extracellular polymeric substances (EPS), are frequently enriched in metal
82 cations at the nmol/g (Huerta-Diaz et al., 2012, 2011; Petrash et al., 2015) to µg/g (Sancho-
83 Tomás et al., 2018; Webster-Brown and Webster, 2007) range. This is a function of the EPS
84 containing a variety of organic functional groups (*e.g.*, carboxyl, phosphate, and sulfhydryl)
85 that deprotonate at circum-neutral pH conditions, thus making them anionic and prone to metal
86 adsorption and potential authigenic mineral nucleation (Konhauser, 1998, 1997). The same
87 functional groups are also responsible for the trapping and binding of detrital particles onto the

88 surface of mats as they grow upward, including heavy mineral grains, such as titanium oxides
89 and zircons (Homann et al., 2015), or clays (Konhauser et al., 1998; Newman et al., 2017,
90 2016; Playter et al., 2017).

91 There is no reason to believe that Precambrian mats behaved differently from the
92 modern so it stands to reason that fossilized mat fabrics should similarly show trace element
93 (TE) enrichments relative to host sediments. In this regard, we measured the bulk TE
94 concentrations of the MRS and host sediments of sandstone and black shale from the *ca.* 2.1
95 Ga Francevillian FB₂ Member of the Francevillian Group in Gabon to ascertain whether
96 potential biological TE enrichments are recorded. These sedimentary rocks were chosen
97 because previous work has already demonstrated that they contain diverse microbial mat
98 morphotypes (Aubineau et al., 2019, 2018), previously described as pyritized MRS, poorly
99 pyritized MRS, and unpyritized “elephant-skin” textures (EST). The rocks also possess
100 geochemical record that tracks Earth’s oxygenation history and MRS with associated K
101 enrichment (Aubineau et al., 2019; Canfield et al., 2013; Chi Fru et al., 2019).

102 **2. Geological context of the Francevillian Group**

103 The unmetamorphosed and minimally deformed *ca.* 2.1 Ga old Francevillian Group in
104 southeastern Gabon is a 1.0 to 2.5 km-thick succession that overlies an Archean crystalline
105 basement (Fig. 1a) (Gauthier-Lafaye and Weber, 1989; Weber et al., 2016).
106 Lithostratigraphically, the Francevillian Group has been divided into five formations, labelled
107 FA to FE from the oldest to youngest (Fig. 1b) (Weber, 1968). The Archean crystalline
108 basement is mainly composed of granitoids and greenstone belts (Thiéblemont et al., 2009)
109 with ages *ca.* 3.0–2.9 Ga (Mouélé et al., 2014). The basal FA Formation unconformably
110 overlies the Archean basement and is predominantly composed of fluvial to fluvio-deltaic,
111 coarse- to fine-grained quartz arenites, and conglomerates with minor interlayered mudstones.
112 Collectively, this sedimentary sequence indicates a progressive transition to marine conditions
113 (Bankole et al., 2015; Gauthier-Lafaye and Weber, 1989).



114
 115 **Figure 1.** Geological map of the Paleoproterozoic Franceville sub-basin (a) modified from Bouton et al.
 116 (2009), showing Gabon in Africa and the Franceville sub-basin in Gabon, and general lithostratigraphic
 117 column of the Francevillian Group (b). The studied area of the Moulendé Quarry is shown with the green
 118 star, while the red star indicates the stratigraphic position of the studied interval with MRS for which
 119 synthetic lithostratigraphic column observed in the quarry is shown on Figure 2a. Sources of ages: 1 -
 120 Horie et al. (2005); 2 - Thiéblemont et al. (2009); 3 - Bros et al. (1992); 4 - Mouélé et al. (2014).

121 The predominantly marine FB Formation sedimentary rocks conformably overlie the FA
 122 Formation and are characterized by fine-grained sediments with interbedded debris flows,
 123 dolostones, thin iron-rich intervals, manganese ore deposits, and sandstones (Gauthier-
 124 Lafaye, 1986; Gauthier-Lafaye and Weber, 1989). The FB Formation is classically subdivided
 125 into FB₁ (including a, b, and c units) and FB₂ (with a and b units) members based on lithological
 126 variations (Azziley Azzibrouck, 1986; Weber, 1968). The FB₁ Member is dominated by
 127 greenish, silty to sandy shales, heterolithic beds that consist of shales alternating with
 128 interbedded dolomite-cemented siltstones, and manganese-rich black shales (Azziley
 129 Azzibrouck, 1986; Pombo, 2004; Reynaud et al., 2017). The FB₂ Member is essentially
 130 dominated by coarse-grained sandstones rapidly deposited from high-density gravity currents

131 (FB_{2a} unit) that formed largely structureless beds with some intervals showing soft-sediments
132 deformation structures due to dewatering (Reynaud et al., 2017). The sandstones are overlain
133 by black shales with thin interlayered siltstones (FB_{2b} unit), that were deposited during a sea-
134 level fall. Overall, the FB₂ Member sediments were deposited in a shallow-marine environment
135 with an oxygenated water column (Canfield et al., 2013; Ossa Ossa et al., 2018). The overlying
136 FC Formation consists of massive dolostone beds, and stromatolitic cherts interbedded with
137 black shales (Amard and Bertrand-Sarfati, 1997; Bertrand-Sarfati and Potin, 1994; Pr  at et al.,
138 2011; Weber, 1968). An exceptionally well preserved, Gunflint-type microfossil assemblage
139 was recently documented in cherts of the FC Formation, giving additional support to the diverse
140 ecosystem in the Franceville sub-basin (Lekele Baghekema et al., 2017). The FD Formation
141 is mainly dominated by black shales alternating with volcanic tuffs and fine- to medium-grained
142 sandstones (Thi  blemont et al., 2014). The topmost FE Formation consists of medium-grained
143 arkose with interlayered shales (Gauthier-Lafaye and Weber, 1989; Thi  blemont et al., 2014).

144 A number of geochronologic dates has been obtained for the Francevillian Group
145 sediments, but the timing of deposition of the Lower Francevillian Group (FA and FB
146 formations) is still controversial. Diagenetic illites in the FB_{1b} unit yielded a Sm-Nd isotope ages
147 of 2099 ± 115 Ma and 2036 ± 79 Ma on $<0.4 \mu\text{m}$ and $<0.2 \mu\text{m}$ clay fractions, respectively,
148 which constrained the age of early diagenesis in this unit (Bros et al., 1992). Finally, the
149 depositional age of the FD Formation is well constrained with ignimbrite tuffs and sandstones
150 that provided U-Pb ages of 2083 ± 6 Ma (Horie et al., 2005) and 2072 ± 29 Ma (Thi  blemont
151 et al., 2009), respectively. Based on carbonate and organic carbon isotope compositions, the
152 FB and FC formations record the end of the Lomagundi carbon isotope excursion (Canfield et
153 al., 2013; El Albani et al., 2010; Pr  at et al., 2011), dated elsewhere between 2.11 and 2.06
154 Ga (Karhu and Holland, 1996; Martin et al., 2013). Given these age constraints, deposition of
155 the Francevillian Group likely occurred *ca.* 2.1 Ga ago, during a critical juncture when Earth's
156 atmosphere underwent a critical transition towards persistently oxygenated conditions.

157 **3. Methods**

158 **3.1. Sampling**

159 Highly diverse microbial mats and their host sediments were sampled from the Moulend  
160 Quarry, Franceville sub-basin, Gabon. The mat layers were carefully separated from the
161 underlying sediments with a blade, avoiding contamination with the underlying sediments.
162 Directly from below the MRS, the host sediments were also sampled for geochemical analyses.

163 **3.2. Microscopy**

164 Mineral composition and textural relationships were studied in polished and carbon-coated thin
165 sections, using a FEI Quanta 200 scanning electron microscope (SEM) equipped with an
166 energy-dispersive X-Ray Spectrometer (EDX) as previously described (Aubineau et al., 2018).

167 **3.3. Major and trace element analyses**

168 Whole-rock geochemical analyses of major and trace elements were performed on 7 pyritized
169 MRS, 7 poorly pyritized MRS, 4 unpyritized “elephant-skin” textures, 7 sandstones, and 15
170 black-shale samples. Major element (Si, Al, Fe, Mn, Mg, Ca, Na, K, Ti, and P) concentrations
171 were measured by inductively-coupled plasma optical emission spectrometry (ICP-OES), while
172 trace and rare earth element (As, Ba, Bi, Cd, Ce, Co, Cr, Cs, Cu, Dy, Er, Eu, Gd, Hf, Ho, La,
173 Lu, Mo, Nb, Nd, Ni, Pb, Pr, Rb, Sb, Sm, Sn, Sr, Ta, Tb, Th, Tm, U, V, W, Y, Yb, Zn, Zr) contents
174 were analyzed by inductively coupled plasma mass spectrometry (ICP-MS) at *Service*
175 *d’Analyse des Roches et Minéraux* (SARM) of the *Centre de Recherches Pétrographiques et*
176 *Géochimiques* (CRPG), Nancy, France. Total sulfur content of the MRS and host sediments
177 was measured by infrared absorption with a carbon-sulfur EMIA 320-V2 (HORIBA) at SARM-
178 CRPG.

179 Prior to analysis, each sample was powdered in an agate mortar. The glass beads were
180 then prepared by fusion with ultra-pure lithium metaborate (LiBO_2) at 980 °C followed by a
181 dissolution in a 1 mol/L nitric acid, 0.5% hydrogen peroxide, and 10% glycerol mixture. The
182 analytical accuracy and precision were found to be better than 5% for major elements, with the
183 exception of Ca (10%) and P (>25%). According to the analyzed trace element concentrations,
184 the uncertainty is lower than 5–10% for elemental concentrations > 50 ppm, less than 5–20%
185 for elemental concentrations between 10 and 1 ppm, and higher than 25% for elemental
186 concentrations near the detection limit (Carignan et al., 2001).

187 The enrichment factor (EF; cf. Tribovillard et al., 2006) relative to the average shale
188 was used to assess whether an element was authigenically enriched or depleted. The EF is
189 considered to be detectable and substantial when it is >3 and >10, respectively (Algeo and
190 Tribovillard, 2009). EFs were calculated as $(X/Al)_{\text{sample}}/(X/Al)_{\text{average shale}}$, where X stands for the
191 element concentration. Descriptive statistics for trace elements was performed with R software
192 (version 3.5.1; R Core Team, 2018) to enhance data resolution relative to the number of
193 individual samples within each lithology.

194 **3.4. X-ray fluorescence (XRF) microscopy**

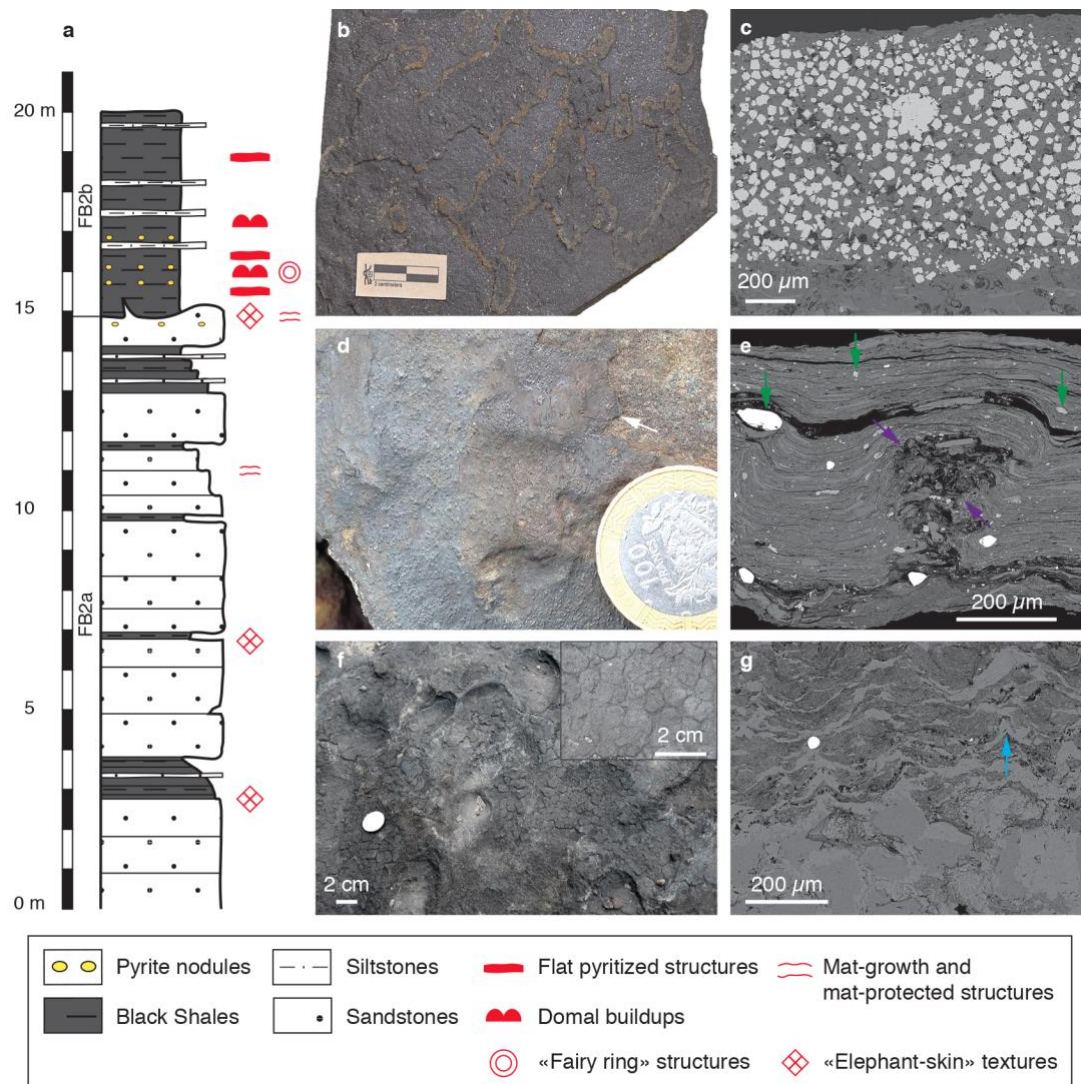
195 Scanning X-ray fluorescence microscopy analysis was performed on polished slab sections of
196 representative MRS samples. The XRF elemental distribution maps were collected on several
197 areas within the investigated samples at the Nanoscopium hard X-ray nanoprobe beamline

198 (Somogyi et al., 2015) of Synchrotron Soleil (L'Orme des Merisiers Saint-Aubin, France). The
199 monochromatic X-ray beam with 12 keV energy was focused on the sample by a Kirckpatrick-
200 Baez nano-focusing mirror. We used the fast continuous scanning (FLYSCAN) technique to
201 obtain micrometer resolution for elemental maps of mm²—sized sample areas (Medjoubi et
202 al., 2013). We have increased the solid angle of detection by the means of two Si-drift detectors
203 (VITUS H50, KETEK GmbH) to collect full XRF spectra in each pixel of the scans. We provide
204 the distribution of Fe, As, and Ni with micron-scale spatial resolution and high analytical
205 sensitivity. Each elemental map was normalized to 10 ms/pixel exposure time.

206 **4. Results**

207 **4.1. Textural diversity of MRS**

208 Mat-related structures, marine shales, and sandstones collected from across the FB₂ Member
209 (Fig. 2a) are different in their macroscopic and microscopic textures, mineralogy, and
210 geochemistry (Figs. 2b-g; Tables S1, S2). Pyritized MRS are developed on top of thin bedded,
211 parallel-laminated black shales. They range from 0.8 to 3 mm in thickness and contain
212 euhedral pyrite crystals, 20 to 30 μm in size (Figs. 2b, c). In this study, pyritized MRS refer to
213 flat, pyritized structures, domal buildups, and “fairy ring” features that were previously
214 described by Aubineau *et al.* (2018) in the same rocks. The poorly pyritized MRS from the
215 sandstone unit reflect mat growth or mat-protected (resulting from mat preservation of
216 structures formed independently from mat growth) patterns (cf. Aubineau et al., 2018; Sarkar
217 et al., 2008). They vary between 0.2 and 0.5 mm in thickness (Figs. 2d, e), and contain a large
218 number of detrital minerals (Fig. 2e; Fig. S1) (Aubineau et al., 2018). Among them, randomly
219 arranged zircons and titanium oxides, and clay particles up to 100 μm in size, are common.
220 Numerous sub-euhedral to euhedral pyrite crystals with varying size, but generally smaller than
221 those in the pyritized MRS, were observed at deeper levels in the poorly pyritized MRS (Fig.
222 S2). In addition, interpreted gas escape and bubble-like structures, having clay minerals
223 disorganized at the edges, but aligned around them, were likely formed before compaction
224 (Fig. 2e). The unpyritized EST, some of which developed on the same bedding planes in
225 sandstones as the poorly pyritized MRS, is one of the most common mat morphologies,
226 characterized by tufted microbial fabrics (Figs. 2f, g) (Aubineau et al., 2018; Reyes et al., 2013;
227 Shepard and Sumner, 2010). They reach up to 2 mm in thickness and lack embedded heavy
228 minerals and putative gas escape structures.



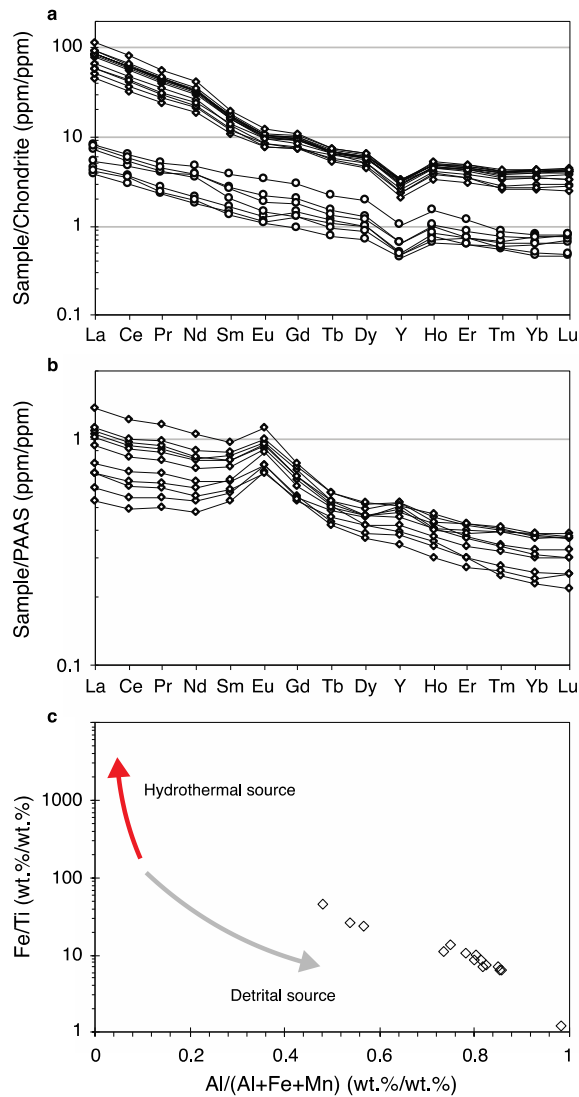
229

230 **Figure 2.** Lithostratigraphic column and representative mat-related structures (MRS) with their
 231 associated petrographic textures from both sandstone and black shale facies. **a**, Composite
 232 lithostratigraphic section of the FB₂ Member in the studied area showing stratigraphic levels where
 233 pyritized MRS (flat pyritized, domal buildup, and ‘fairy ring’ structures), poorly pyritized MRS (mat-growth
 234 and mat-protected structures), and unpyritized MRS (EST: “elephant-skin” texture) developed. **b**, Flat
 235 pyritized MRS from the FB_{2b} unit. **c**, SEM image of pyritized MRS in cross-section perpendicular to the
 236 bedding plane. Note that pyrite grains and clay particles are developed within biofilms. **d**, Mat-layer
 237 structures (white arrow) on the bedding surface of sandstone from the FB_{2a} unit. **e**, SEM image of
 238 micrometer-thick poorly pyritized MRS in cross-section perpendicular to the bedding plane. Green and
 239 purple arrows show trapped and bound heavy minerals and a gas escape structure, respectively. **f**, EST
 240 from the FB_{2a} unit. Inset box shows reticulate pattern. **g**, SEM image of tufted microbial fabrics from the
 241 EST above the poorly sorted quartz sandstone. Blue arrow points to tufted microbial fabrics. Images in
 242 b and c, and f and g are modified from Aubineau *et al.* (2018) and Aubineau *et al.* (2019), respectively.

243 4.2. Provenance characterization

244 The host sediments have similar chondrite-normalized rare earth elements (REE) + Y patterns
 245 (Fig. 3a; Table S2). The steep light-REE pattern with a flat heavy-REE profile in the post-

246 Archean shales reflects REE abundances in the upper continental crust exposed to weathering
 247 (Large et al., 2018; Taylor and McLennan, 2001). REE contents decrease with increasing SiO₂
 248 content, reflecting quartz dilution. The similar REE + Y patterns of sandstone and black shale
 249 lithologies suggest that the sediments on which the MRS developed had the same provenance.
 250 Furthermore, sandstones and black shales are interlayered (Fig. 2a), indicating that they were
 251 deposited in the same general setting, but at different energy levels (cf. Reynaud et al., 2017).



252
 253 **Figure 3.** Geochemistry of the host rocks, black shales (diamonds) and sandstones (circles). **a**,
 254 Chondrite-normalized REE + Y patterns. Chondrite values are from Schmitt et al. (1964) and Evensen
 255 et al. (1978). **b**, PAAS-normalized REE + Y patterns. PAAS values are from Taylor and McLennan
 256 (1985). **c**, Binary plot of Fe/Ti vs. Al/(Al+Fe+Mn) after Pecoits et al. (2009).

257 The REE abundances have been normalized to the Post-Archean Australian Shale
 258 (PAAS), hereafter referred to as REE_(SN) (Fig. 3b). The latter is a widely used reference for
 259 normalizing sedimentary lithologies, allowing easy comparison with previous studies. The
 260 PAAS-normalized REE + Y patterns of black shales from the FB₂ Member show moderately
 261 positive Eu anomalies (Eu/Eu*_{SN}; see Table S2 for method of calculation), ranging from 1.29 to

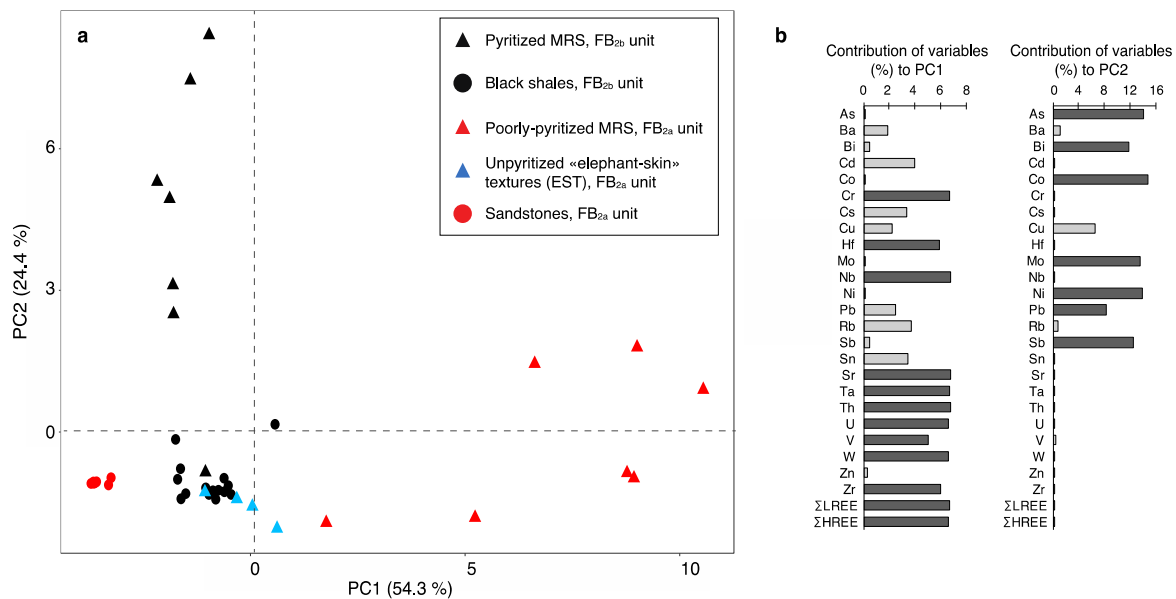
262 1.44. A strong positive Eu anomaly in sediments is usually interpreted to indicate high-
263 temperature hydrothermal fluxes to the water column (Bolhar et al., 2004; Hiebert et al., 2018;
264 Planavsky et al., 2010). Moreover, the REE + Y patterns reveal a moderately steep trend with
265 light-REE enrichments, as shown by Pr/Yb_(SN) ratios ranging between 1.45 and 3.54 (Fig. 3b;
266 Table S2). In addition, the binary plot of Fe/Ti vs. Al/(Al+Fe+Mn), which resolves the influence
267 of a hydrothermal vs. detrital source (Fig. 3c; cf. Pecoits *et al.* 2009), suggests that the studied
268 black shales predominantly reflect a detrital source, with limited hydrothermal influence.
269 Similarly, the average Y/Ho value of 31.4 (Table S2) is lower than the modern seawater Y/Ho
270 ratio that is >44, but above the invariable Y/Ho ratio of continental crust (~26; Bolhar et al.,
271 2004).

272 **4.3. Chemical composition of host sediments and MRS**

273 Petrographic analysis showed that Fe-(oxyhydr)oxide coatings were not found around the
274 pyrite crystals (Figs. 2c, S2). In addition, the X-ray diffraction patterns of the Francevillian MRS
275 and host sediments revealed the occurrence of pyrite as the main iron sulfide, while Fe-
276 (oxyhydr)oxides were lacking (Aubineau et al., 2019). Combined petrographic and XRD
277 analyses suggest that chemical composition of our samples was not appreciably affected by
278 recent oxidative weathering.

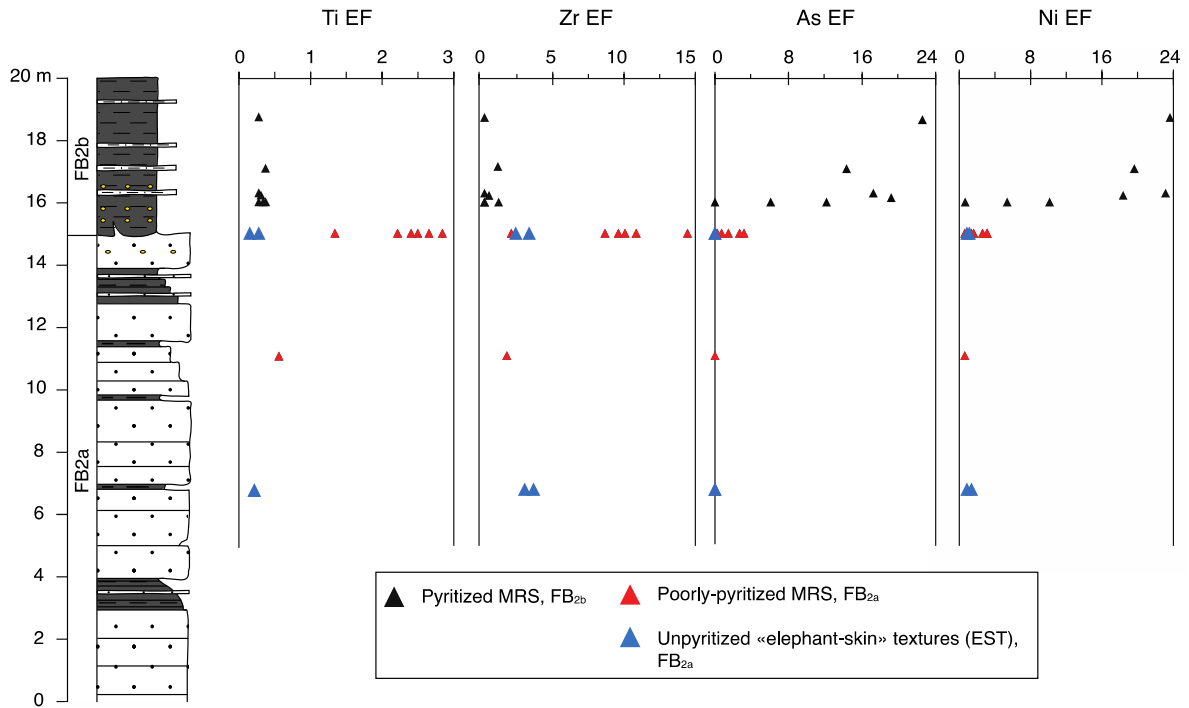
279 Descriptive statistical relationships among independent TE variables of MRS and
280 individual sediment lithologies show the largest variation along the first and second Principal
281 Components (PC) (Fig. 4a). Collectively, both PC explain more than 78% of the total variance
282 in the dataset (Table S3). According to the contribution of each variable along the PC1, the
283 variation is induced by TE that are usually taken to be proxies of the detrital flux (Fig. 4b;
284 McLennan 1989, 2001; Tribovillard *et al.* 2006). Along the PC2, it appears that the metals
285 enriched in pyrite, apart from Cd and Zn, are responsible for differences amongst the
286 lithologies. The PC1 accounts for 54.3% of the total variance, with the poorly pyritized MRS
287 being different from the pyritized MRS, EST, and host sediments (Fig. 4a). Interestingly, EST,
288 pyritized MRS, and black shales do not cluster into distinctly separate groups. Their spread
289 along the PC1 axis seems largely determined by the elements that are abundant in the heavy
290 mineral fraction, as indicated by the large contribution of Zr and REE (Fig. 4b). This is
291 consistent with petrographic observations that indicate the presence of detrital zircon and Ti-
292 oxides in the poorly pyritized MRS. The PC2 contributes 24.4% to the total variance, with the
293 pyritized MRS plotting above the poorly pyritized MRS, EST, and host sediments. The pyrite
294 content likely controls their spread along the PC2 axis. One pyritized MRS sample has a
295 signature similar to that of black shales, suggesting either a large contribution of the host

296 sediment during sample separation or a depletion in metals enriched in pyrite compared to the
 297 other pyritized MRS (Fig. 4a).



298
 299 **Figure 4.** Descriptive statistical treatment. **a**, Principal component analysis of 40 individuals
 300 representing the relationships among 3 MRS morphotypes and 2 host sediments (black shales and
 301 sandstones) and 26 independent variables (trace element concentrations). **b**, Contribution of each
 302 variable (%) along PC1 (left) and PC2 (right). The stronger the contribution of TE, the darker the bar is.

303 Stratigraphic geochemical profiles of the FB₂ Member reveal dramatic difference in bulk
 304 composition among the three MRS morphotypes even on the same stratigraphic level, but also
 305 heterogeneous composition of the poorly pyritized and pyritized MRS (Fig. 5). In the poorly
 306 pyritized MRS samples, EFs have a mean of 2.1 for Ti (EF_{Ti}) and 6.6 for Zr (EF_{Zr}), suggesting
 307 moderate to strong enrichment, respectively. These elements are predominantly depleted to
 308 moderately enriched in the EST (mean EF_{Ti} = 0.2; mean EF_{Zr} = 2.5) and in the pyritized MRS
 309 (mean EF_{Ti} = 0.3; mean EF_{Zr} = 0.5). Chalcophile elements (CE) are depleted to weakly
 310 enriched (mean EF_{As} = 0.1; mean EF_{Ni} = 1.1) and weakly enriched (mean EF_{As} = 1.4; mean
 311 EF_{Ni} = 2.0) in the EST and the poorly pyritized MRS, respectively. On the other hand, the
 312 pyritized MRS are strongly enriched in CE (mean EF_{As} = 13.2; mean EF_{Ni} = 14.5). The host
 313 sediments (both sandstone and black shale) show no enrichment to weak TE enrichment for
 314 most samples, with mean EFs for Ti, Zr, As, and Ni of 0.4, 1.5, 0.5, and 1.2, respectively (Fig.
 315 S3).

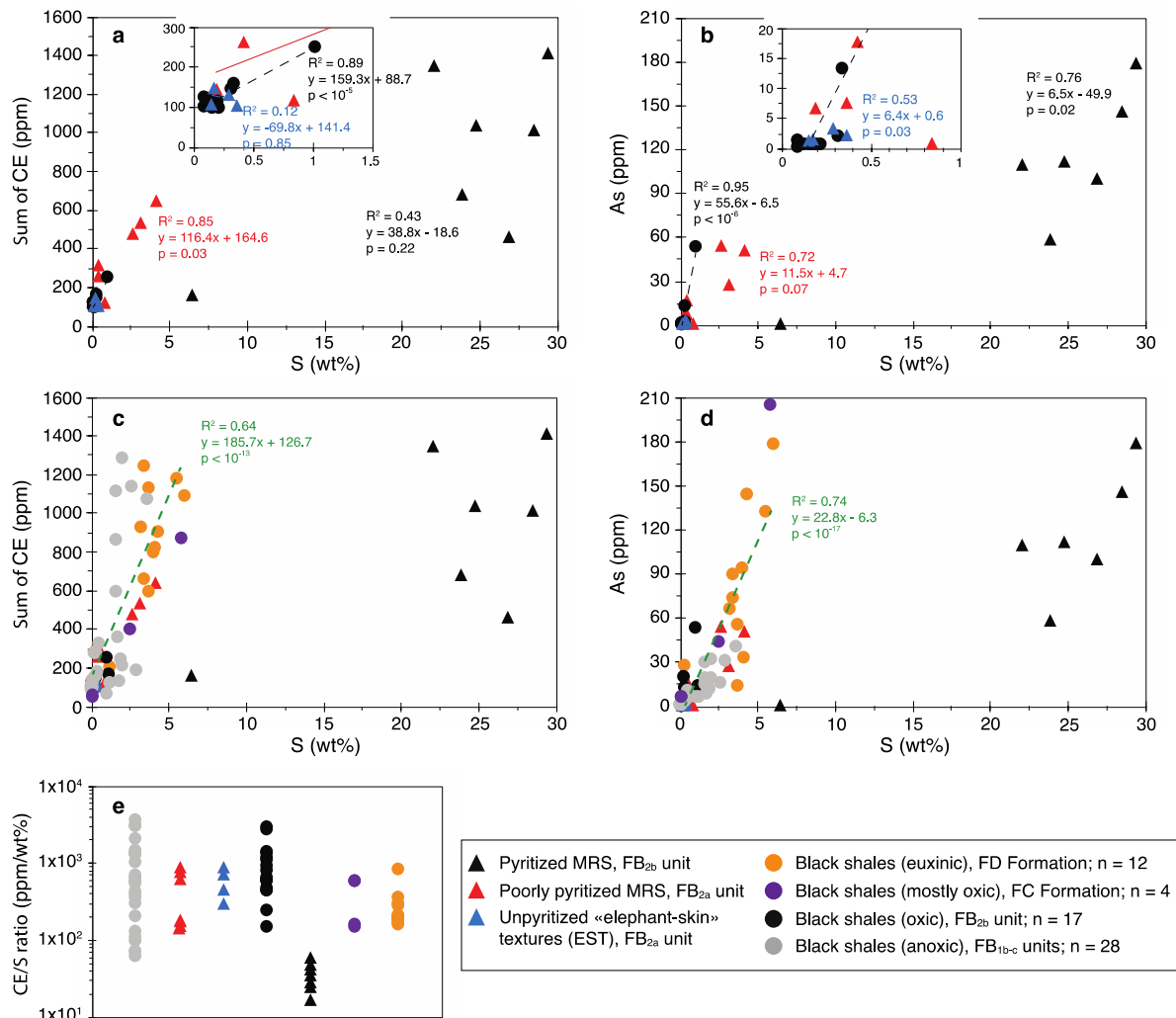


316

317 **Figure 5.** Stratigraphic profile of enrichment factors (EF) of selected detrital and chalcophile elements
 318 within MRS. EF are normalized to the average shale (Ti, Zr, and Ni concentrations for the average shale
 319 are from Taylor and McLennan (2001) and As concentration for the average shale is from Li and
 320 Schoonmaker, 2003).

321 Generally, CE are enriched within the sediment in association with S-bearing minerals,
 322 and when plotted against each other, show positive covariation. The trends found within the
 323 MRS and black shales from the FB₂ Member show more or less the expected pattern for the
 324 preferential incorporation of CE into sulfide minerals (Figs. 6a, b). The mixing trends associated
 325 with the poorly pyritized MRS, EST, and host shale reflect the variable content of CE-bearing
 326 pyrite. Nonetheless, the correlation trends of pyritized MRS differ from those of pyrite-poor
 327 MRS and host sediments, indicating that either some intrinsic properties (*e.g.*, mineralogy,
 328 organic matter, or microbial sulfate reduction rate) or the dynamic fluctuations in supply of CE
 329 controlled incorporation of pyrite-hosted metals. For instance, organic substrate may have
 330 controlled the availability of CE as the carbonaceous material of MRS and host rocks was
 331 derived from different sources (*i.e.*, benthic vs. planktonic origin). Minor to no contribution of
 332 benthically produced organic matter to the host sediments, which is suggested by the lack of
 333 microbial mat structures (Aubineau et al., 2018) and K enrichment (Aubineau et al., 2019),
 334 argues for two distinct carbon sources. Specifically, K incorporation occurred during the living
 335 stage of benthic MRS through a bacterially mediated process, whereas the sinking flux of
 336 particulate organic material did not incorporate and deliver K to the sediments to the same
 337 extent. Moreover, 43% ($R_2=0.43$) of total co-variation in the pyritized MRS between CE and S
 338 (Fig. 6a) implies that CE are also associated with other components. However, this relationship
 339 might be biased as it seems to be controlled by the one pyritized MRS sample with lower CE

340 content, by a factor of 10 to 100, and moderate S concentration. It reveals high heterogeneity
 341 in CE contents in pyritized MRS.



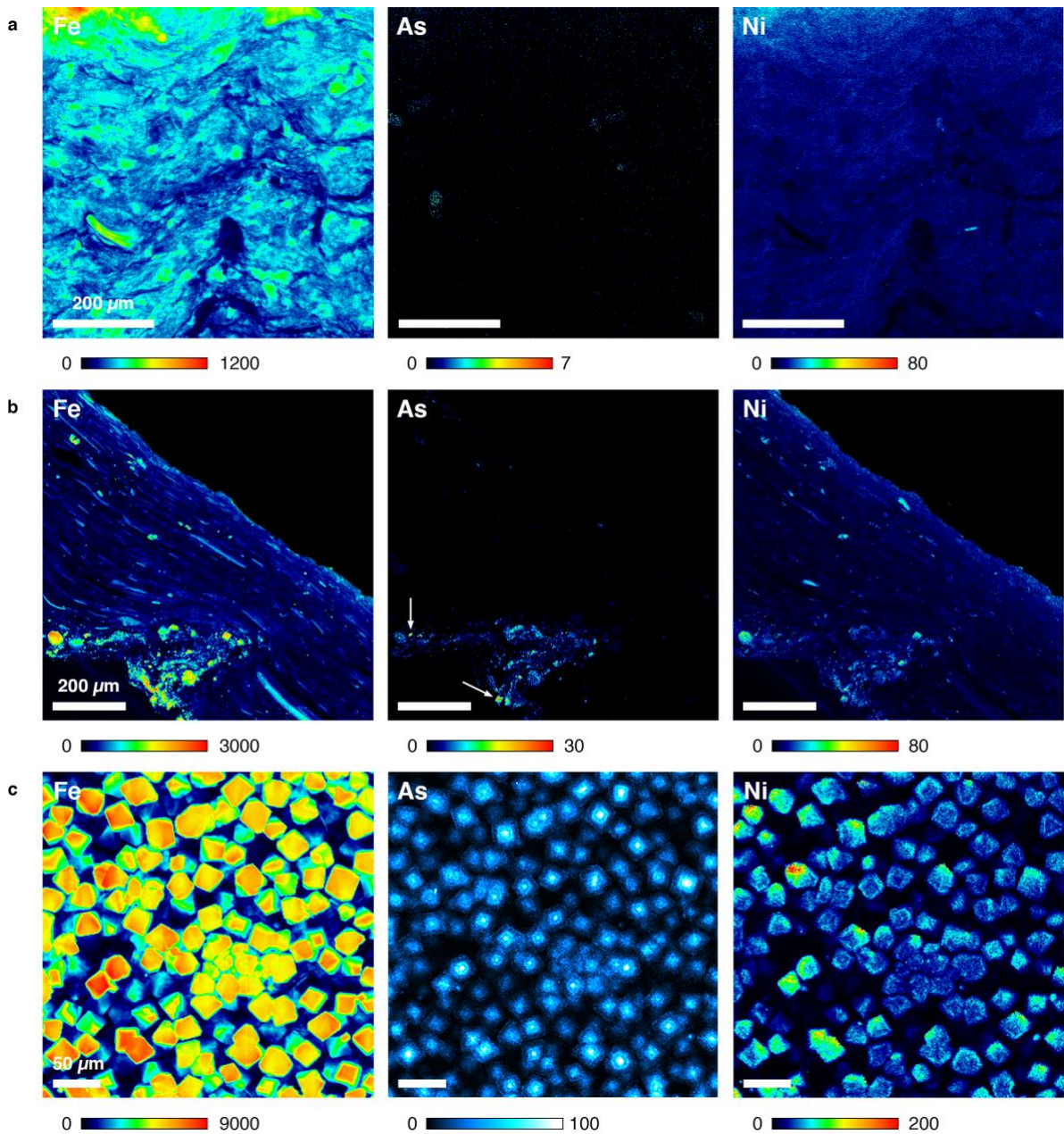
342
 343 **Figure 6.** Relationship between chalcophile elements and bulk-rock sulfur in pyritized MRS, poorly
 344 pyritized MRS, unpyritized «elephant-skin» textures (EST), and Francevillian Group black shales (from
 345 the FB_{1b} and FB_{1c} units and the FD Formation). **a**, Cross plot of Σchalcophile element (As, Bi, Co, Cu,
 346 Mo, Ni, Pb, and Sb) contents vs. bulk-rock S concentrations for MRS and host black shales. **b**, Cross
 347 plot of As content vs. bulk-rock S concentration for MRS and host black shales. Insets (a and b) show
 348 co-variation at low CE and S contents. **c**, Cross plot of Σchalcophile (As, Bi, Co, Cu, Mo, Ni, Pb, and
 349 Sb) contents vs. bulk-rock S concentrations for MRS and all the Francevillian Group black shales (green
 350 trend). **d**, Cross plot of As contents vs. bulk-rock S concentrations for MRS and all the Francevillian
 351 Group black shales (green trend). **e**, CE/S ratios for 3 MRS morphotypes and 2 host sediments (black
 352 shales and sandstones). CE and S contents for black shales of the Francevillian Group are from Canfield
 353 *et al.* (2013). p-value (p) is provided with a confidence level of 95%. Data for some samples are below
 354 the detection limit; in that case a value of half of the detection limit was used.

355 Compiled data for CE and S contents of the Francevillian Group black shales (from the
 356 FB_{1b} unit to FD Formation) illustrate an overall linear relationship (Fig. 6c), suggesting that
 357 sulfide production controlled the enrichment in CE. The strong positive co-variation ($R^2 = 0.64$)
 358 for black shales of the Francevillian Group suggests that CE contents are only moderately
 359 controlled by other mineral phases. The pyritized MRS show a different trend with higher S

360 contents for the same CE concentrations. The same trend is also observed for As and S (Fig.
361 6d), indicating that the As content was affected by the same processes that controlled CE
362 enrichments. Finally, the CE/S ratios of pyrite-poor MRS and the Francevillian Group black
363 shales are mostly within the same range from 10^2 to 10^3 , whereas the pyritized MRS show the
364 lowest ratios down to 17 (Fig. 6e).

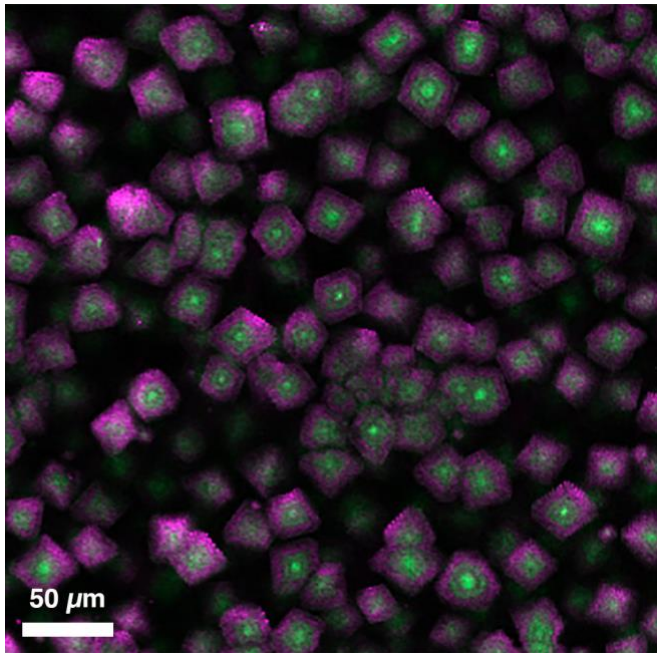
365 **4.4. High-resolution metal distribution in the MRS**

366 Millimeter-sized areas of each MRS morphotype were analyzed by synchrotron-based
367 scanning X-ray fluorescence microscopy revealing spatial variations in S, K, Ca, Ti, Ce, Ba,
368 Cr, Mn, Fe, Ni, Cu, Zn, Ga, Ge, Hf, and As. The distribution maps of Fe, As, and Ni are shown
369 in Figures 7 and 8. Pyrite crystals are highlighted by Fe distribution and location of CE by Ni
370 and As maps. Although some Fe is hosted in phyllosilicates, the highest abundance of iron is
371 in pyrite grains, highlighting the difference amongst the MRS. EST have Ni and As
372 concentrations close to the detection limit (Fig. 7a). These elements are unlikely to be
373 associated with pyrite, since pyrite is not present at a detectable level in the same spots. This
374 also applies to the absence of a strong relationship between CE and S in the EST. The poorly
375 pyritized MRS contain higher CE concentrations relative to the EST that are heterogeneously
376 distributed in pyrite grains (Fig. 7b). A few small, euhedral pyrite crystals display As-rich cores.
377 Arsenic is preferentially located within the pyrite cores in the pyritized MRS with a variation in
378 As content of $\pm 10\%$ (1 SD), whereas Ni is mostly concentrated within the overgrowths (Fig.
379 7c). This pattern suggests high As content in early pore-waters from which pyrite cores
380 nucleated and higher Ni content during later diagenesis when the overgrowths were formed.
381 Arsenic distribution within individual pyrite crystals is heterogeneous, resulting from
382 micrometer-thick As-rich rims around the cores and a decrease in As content toward the crystal
383 surface (Fig. 8). Ni also exhibits a highly variable abundance pattern often with a micrometer-
384 thick Ni-rich zone at the pyrite crystal surface. CE appear to be enriched by a factor of 15 (As)
385 and 6 (Ni) in pyritized MRS compared to poorly pyritized MRS.



386

387 **Figure 7.** Synchrotron-based XRF microscopy showing distribution of Fe, As, and Ni. **a**, unpyritized
 388 EST. **b**, poorly pyritized MRS. White arrow indicates As-rich pyrite core. **c**, pyritized MRS. Colour bars
 389 indicate intensity scale of the metal XRF distribution maps in counts/10 ms (black colour indicates
 390 concentrations below the detection limit). Higher XRF intensities correspond to higher metal contents.



391

392 **Figure 8.** Composite image: As (green) and Ni (magenta) XRF maps for pyritized MRS.

393 **5. Discussion**

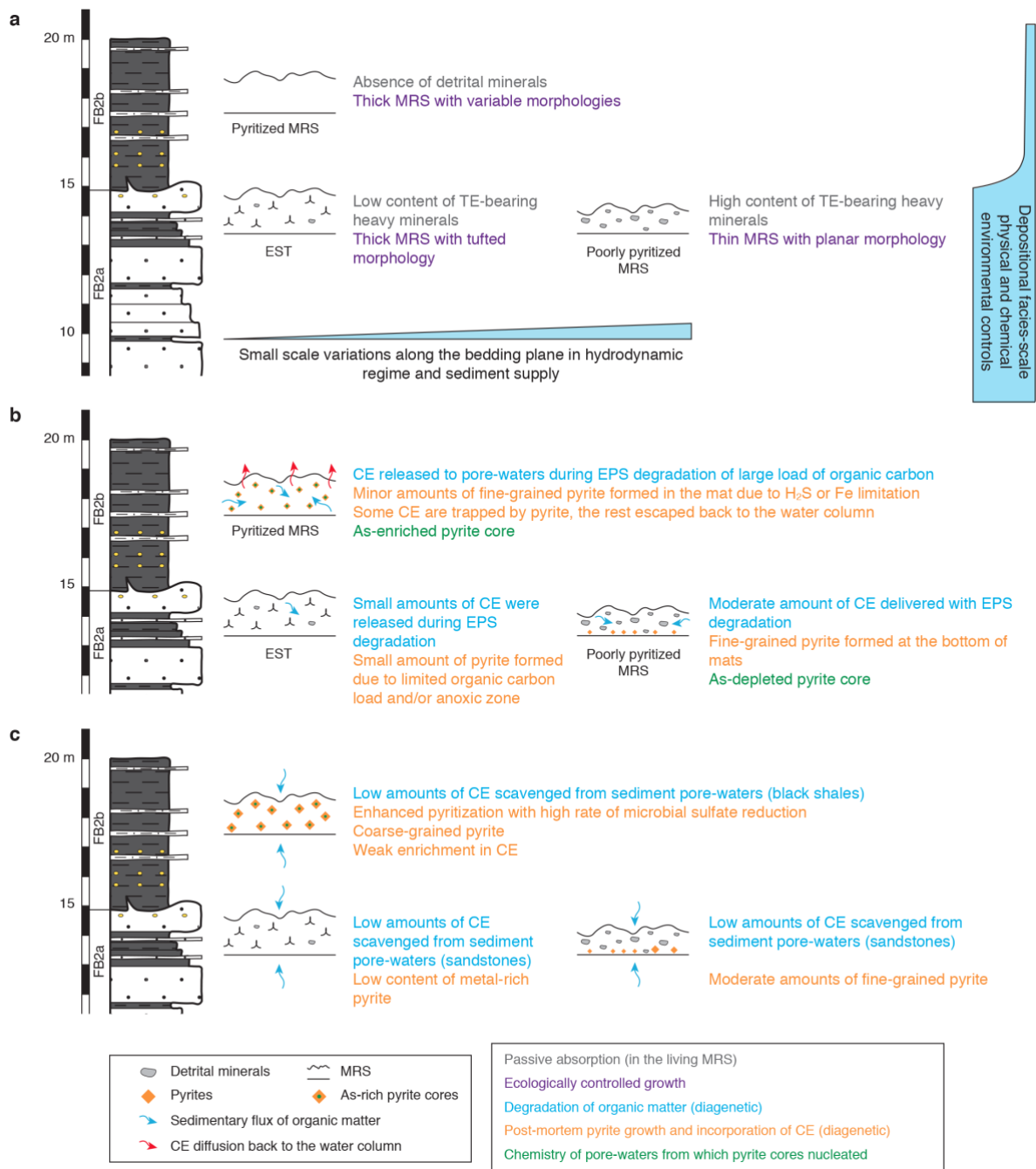
394 **5.1. Microbial fabric and response to the environment**

395 At high concentrations of suspended material in the water column, laboratory-cultured
396 cyanobacteria can become rapidly coated with clay minerals, which means that they may
397 preserve the TE and mineralogical context of detrital input to the depositional setting in which
398 they lived (Newman et al., 2017, 2016; Playter et al., 2017). The Francevillian poorly pyritized
399 MRS and EST contain in a high abundance TE-bearing heavy minerals relative to the host
400 sandstone facies. The absence of heavy mineral enrichment in the overlying pyritized MRS
401 reflects a significant decrease in sedimentation rate and energy in the depositional setting (Fig.
402 9a), consistent with a change in sediment supply from sand-rich high density flows to fine-
403 grained black shales (Reynaud et al., 2017). As such, the distribution of heavy minerals in the
404 Francevillian Group MRS, or for that matter any MRS, is best explained by passive trapping
405 and binding processes (Sforna et al., 2017).

406 Importantly, poorly pyritized MRS and EST developed along the same uppermost
407 stratigraphic horizon in coarse-grained sediments (Fig. 9a). Although the transition from
408 sandstone to black shale might imply a transgressive surface, the sandstones of the FB
409 Formation were deposited from pulses of high density flows into a depositional setting with a
410 background mud sedimentation (Reynaud et al., 2017). It is thus reasonable to assume that
411 the poorly pyritized MRS and EST were formed under exactly the same depositional
412 conditions. The variable abundance of Ti- and Zr-bearing heavy minerals in these mats only
413 few meters apart likely reflects minor changes in environmental conditions under which the

414 MRS flourished (e.g., hydrodynamic regime or sediment delivery pattern) (Fig. 9a). These
415 changes appear to have impacted microbial growth patterns, resulting in a shift to different mat
416 morphology (Fig. 9a). Indeed, numerous observations on stromatolite morphology have shown
417 active adaptation by the bacterial community to environmental conditions, such as water
418 energy and the availability of light and nutrients (e.g., Allwood et al., 2006; Jahnert and Collins,
419 2012; Petrash et al., 2012). Similarly, distribution patterns of distinct microbial mat
420 morphotypes across ancient tidal flats constrain palaeoenvironmental conditions (Bose and
421 Chafetz, 2009; Homann et al., 2015; Sarkar et al., 2014).

422 The poorly pyritized MRS typically display the smallest mat thickness and planar
423 morphologies that likely reflect a morphological adjustment to high energy hydrodynamic
424 regime, while the EST and pyritized MRS show larger mat thicknesses and well-developed
425 mat-growth fabrics (*i.e.*, tufted features, domal buildups) that likely indicate optimal growth
426 conditions (e.g., Gerdes et al., 2000). Differences in physical environmental conditions would
427 have also led to highly variable abundance of TE-bearing heavy minerals in the Francevillian
428 Group MRS and may have triggered ecologically controlled growth (Fig. 9a). While mat
429 morphology and thickness are linked to external environmental factors, it has been recently
430 documented that the mat thickness may similarly influence the community structure (Suarez
431 et al., 2019). This leads to different redox profiles and internal carbon recycling styles amongst
432 the microbial communities since thick mats contain a well-expressed anoxic area in the mat
433 structure, whereas thin mats create only a limited anoxic region (Suarez et al., 2019). However,
434 effects of thickness on microbial architecture are difficult to quantify because the hydrodynamic
435 regime (Battin et al., 2003; Flemming et al., 2016), water depth (Lydmark et al., 2006),
436 temperature (He et al., 2016), and dissolved carbon to nitrogen ratios (Okabe et al., 1996) also
437 impact on the microbial community structure. Thus, in combination, these environmental
438 parameters might have promoted the lack or poor expression of an anoxic zone in the thick
439 unpyritized EST (Fig. 9b).



440

441 **Figure 9.** Conceptual model for the spatial and temporal TE variations in the Francevillian Group FB₂
 442 unit MRS. **a**, processes that took place during the living stage of mats. Depositional, small-scale to
 443 facies-scale environmental factors controlled the distribution of heavy minerals in the Francevillian MRS.
 444 The external mat morphologies provide information on local ecologic conditions. **b**, short-lived, first stage
 445 of pyrite formation on the earliest phase of diagenesis. The EPS degradation resulted in the release of
 446 CE to pore-waters. The amount of CE delivered with EPS decay should be proportional to the organic
 447 carbon content that is buried. Depending of sulfur and Fe availability, CE were trapped by fine-grained
 448 pyrite or diffused back to the water column. **c**, long-lived, second stage of pyrite formation as diagenesis
 449 progressed. The remaining organic matter was degraded and, thus, larger amounts of H₂S and Fe were
 450 released to pore-waters. High concentrations of biogenic H₂S were locally produced, but only small
 451 amounts of CE were scavenged into coarse-grained pyrite.

452

5.2. Microbial mineralization

453 Recent studies on transition metals and metalloids associated with present-day hypersaline
454 mats (Huerta-Diaz et al., 2012; Petrash et al., 2015; Sancho-Tomás et al., 2018; Sforza et al.,
455 2017) have revealed that TE can be trapped by and bound to the organic matrix, progressively
456 accumulating with depth via incorporation into sulfides. The latter form as the original mat
457 organic material is degraded by the activity of anaerobic heterotrophs (e.g., sulfate-reducing
458 microorganisms), eventually diminishing the cation-binding capacity of the EPS matrix (Dupraz
459 et al., 2009; Dupraz and Visscher, 2005). Then, as authigenic pyrite forms by the reaction of
460 reactive Fe and biogenic H₂S, metals liberated from mats to pore-waters are scavenged into
461 sulfidic microenvironments where pyrite crystals grow (Fig. 9b), resulting in high spatial
462 variability of metal distribution (Huerta-Diaz et al., 2012; Sforza et al., 2017). The spatial
463 restriction of pyrite grains to deeper levels in the mat in the poorly pyritized MRS or their
464 development over the entire mat structure in the pyritized MRS likely reflects the extent of the
465 sulfidic region (Fig. 9b).

466 The pyrite overgrowths in the MRS that formed later during diagenesis precipitated from
467 pore-waters that were no longer enriched in As, but in Ni (Fig. 9b). Alternatively, the preferential
468 incorporation of As into the early pyrite phase and Ni in the later phase may simply suggest
469 that these two elements have different kinetic behavior whereby As is more reactive towards
470 H₂S than Ni (cf. Huerta-Diaz and Morse, 1992). Hydrothermal fluids affecting sediments during
471 diagenesis could produce CE-rich pyrite rims (Large et al., 2009), but we have no evidence for
472 migration of such fluids or late-stage remobilization that could account for strongly enriched
473 CE in pyrite rims during late diagenesis. Variability in concentration of pyrite-hosted metals in
474 the three microbial mat morphotypes is mainly linked to the sulfide production and
475 concentration in pore-waters (Figs. 9b, c). The relationship between CE and total S is,
476 however, complicated since whole-rock data seem to reflect several host minerals. Some of
477 these CE might be incorporated into clays, scavenged by organic matter, and/or absorbed onto
478 the surface of Fe oxides (Chappaz et al., 2014; Chi Fru et al., 2019; Nachttegaal et al., 2005),
479 which might explain weak to moderate correlations within each MRS morphotype.

480 **5.3. Implications of weak enrichment in CE in the pyritized MRS**

481 The CE enrichment is not unique to the MRS, but instead typical for the depositional setting of
482 the Francevillian Group where seawater and pore-water sulfide levels controlled their
483 incorporation into sediments. Although a moderately positive Eu anomaly is observed, low
484 Y/Ho and Fe/Ti ratios in combination with high Al/(Al+Fe+Mn) ratios (see Fig. 3) do not support
485 a strong influence of hydrothermal fluids on the depositional site, and rather suggests that CE
486 were supplied by continental weathering. Consistent with this interpretation, large variations in
487 CE/S ratios for the Francevillian Group shales (FB_{1b-c} units, and FC and FD formations) were

488 not observed. However, absence of substantial CE enrichment in the pyritized MRS is
489 unexpected because they contained appreciable amounts of carbonaceous materials to
490 maintain extensive pyritization. The pyritized MRS samples with high bulk-rock S
491 concentrations have lower CE/S ratios than the Francevillian Group sediments, which requires
492 an explanation for the weak enrichment in CE.

493 Weak enrichment in CE in the pyritized MRS might reflect a change in aqueous fluxes
494 of CE to the ocean during the living stage of the MRS, but this would not explain the similar
495 range of CE content in the pyrite-poor MRS and the Francevillian Group black shales (FB_{1b-c}
496 units, and FC and FD formations). Similarly, different scavenging capabilities of planktonic (in
497 the Francevillian Group black shales) and benthic (in MRS) organic matter is not likely to
498 explain weak enrichment in CE in the pyritized MRS. Rather, low CE/S ratios recorded by the
499 pyritized MRS could be related to profuse sulfide production in pore-waters in
500 microenvironments with high content of easily degradable benthic organic matter. Basinal
501 restriction from the global ocean could have also diminished the reservoirs of metals and
502 metalloids, which could have resulted in low CE/S ratios in the pyritized MRS. However, it is
503 unlikely that the depletion of the basinal inventory of TE would have occurred at the time when
504 the development of pyritized MRS took place, but not during deposition of the Francevillian
505 Group black shales. Importantly, restricted basins with near-quantitative drawdown of the
506 sulfate reservoir often show positive $\delta_{34}\text{S}$ values for early diagenetic pyrite (Diamond et al.,
507 2018). By contrast, highly negative $\delta_{34}\text{S}$ values of the pyritized Francevillian Group macrobiota
508 specimens point to bacterial sulfate reduction during early diagenesis with open connection to
509 the overlying water column (El Albani et al., 2019, 2010). In this regard, we suggest that the
510 amount of CE delivered to the sediments was largely controlled by both the mat thickness and
511 sedimentary flux of organic matter (Figs. 9b, c). With high concentration of biogenic H_2S in the
512 studied MRS, CE were not as enriched as in black shales from the Upper Francevillian Group,
513 which were deposited under euxinic conditions where pyrite formed in the water column had
514 ready access to a much larger reservoir of CE.

515 We propose that the pyrite formation and CE trapping in the pyritized MRS followed a
516 two-stage process, resulting in weak enrichment in CE (Figs. 9b, c). First, the EPS degradation
517 released to pore-waters CE, which were subsequently either incorporated into microcrystalline
518 pyrite or returned to the water column (Fig. 9b). As EPS degradation takes place immediately
519 after deposition (Decho et al., 2005), CE trapping at this stage was limited by sulfur and/or iron
520 availability and, specifically, their reduction and release from the sediments. Hence, CE that
521 were not trapped by pyrite diffused back to the water column, resulting in loss of co-variance
522 between CE and S contents. During the second, protracted stage, most of remaining organic
523 matter in sediments was degraded under anoxic and, locally, massive sulfidic conditions which
524 resulted in the formation of coarse-grained, crystalline pyrite and overgrowths with low CE

525 contents (Fig. 9c). As this stage lasted longer, larger amounts of sulfide and iron were released
526 to pore-waters, whereas pyrite formed was not enriched in CE since CE were either already
527 trapped or lost from sediments during EPS degradation on the first stage.

528 The observed trend of weak enrichment in CE in the pyritized MRS thus reflects: (1)
529 pyrite formation in pore-waters in the MRS and incorporation of CE delivered to the sediments
530 with organic matter, specifically with EPS, and (2) the released to pore-waters during EPS
531 degradation rather than a specific, biologically controlled pathway for accumulation of CE.
532 Diagenetic pyrite production within the MRS likely also enhanced the potential for preservation
533 of the Francevillian Group macrobiota.

534 **5.4. TE geochemistry: a biosignature of benthic microbial activity?**

535 As discussed above, the TE content is greater in MRS than in host sediments of the FB₂
536 Member, a pattern that likely reflects variations in hydrodynamic regime and water-column
537 redox chemistry. Specifically, the TE data for the Francevillian Group MRS suggest that the
538 relationship between CE and sulfide concentrations are not consistent with a biological control
539 over metal enrichment. These observations differ from Sforza et al. (2017) who suggested
540 biologically controlled accumulation of TE in organic-rich globules in modern microbialites was
541 a consequence of specific metabolic pathways (Sforza et al., 2017). By contrast, our results
542 are similar to Rico et al. (2020) who observed that living microbial mats from the Middle Island
543 Sinkhole (located 23 m below the water level of Lake Huron) do not significantly influence bulk
544 TE geochemistry at the sediment–water interface, although highly heterogeneous redox
545 conditions were observed. Considering that the water column redox conditions (e.g., oxygen,
546 sulfate, and iron concentrations) in the Middle Island Sinkhole could be analogous to those
547 that were in Proterozoic shallow-marine environments (cf. Biddanda et al., 2012; Rico and
548 Sheldon, 2019; Ruberg et al., 2008), we believe that these microbial mats provide a close
549 analogue to the MRS of the Francevillian Group FB₂ Member. Together, TE data from benthic
550 microbial communities in modern and ancient settings indicate that TE geochemistry is not a
551 straightforward biosignature to be utilized in studies of the geological rock record.

552 **6. Conclusion**

553 Textural relationship and combined with TE geochemistry suggest distinct biological signals in
554 the Francevillian Group MRS relative to their host sediments. However, no conclusive example
555 of biologically controlled TE enrichment was found. Instead, our results, summarized in
556 Figure 9, indicate highly variable TE contents with environmental factors overprinting
557 biologically controlled TE distribution.

558 A lack of accumulation of heavy minerals and pyrite-hosted CE characterizes the EST.
559 The poorly pyritized MRS contain in a high abundance of Ti- and Zr-bearing minerals
560 concentrated through passive absorption. External morphology of these two MRS could be a
561 potential biogenicity indicator and provides information on local environmental and ecological
562 conditions, largely controlled by hydrodynamic regime and sediment supply. The CE
563 concentrations in the poorly pyritized MRS co-vary ($R_2 = 0.85$) with pyrite content, which
564 suggests that the CE are largely hosted by pyrite. Abundant organic matter in the now pyritized
565 MRS did not result in similar CE/S ratios as in the Francevillian Group black shales deposited
566 in euxinic settings due to the limited trapping capacity of pore-waters close to the sediment-
567 water interface immediately after sediment deposition when CE-enriched EPS degraded.
568 Arsenic-rich pyrite cores indicate high As concentrations in the early diagenetic pore-waters,
569 where it was likely supplied via decay of EPS.

570 In summary, bulk-rock TE data provide additional evidence for biological activity during
571 deposition of the *ca.* 2.1 Ga Francevillian Group, but there is no definitive biological signal
572 since physical controls and water column redox chemistry outweighed microbial influences on
573 the TE distribution in MRS. Recognition of the overprinting power of physical and chemical
574 environmental conditions over microbially imprinted TE patterns in MRS further contributes to
575 our understanding of interaction among ancient microbial life and the environmental and
576 biological controls affecting TE contents in MRS, and their relationship to TE sinks in deep
577 time.

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587 **Data availability.** The authors declare that the data supporting this study are available within
588 the paper and its supplementary information files.

589 **Declaration of competing interests.** The authors declare no competing interests.

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APPENDIX A: SUPPLEMENTARY DATA

Trace element perspective into the ca. 2.1-billion-year-old shallow-marine microbial mats from the Francevillian Group, Gabon

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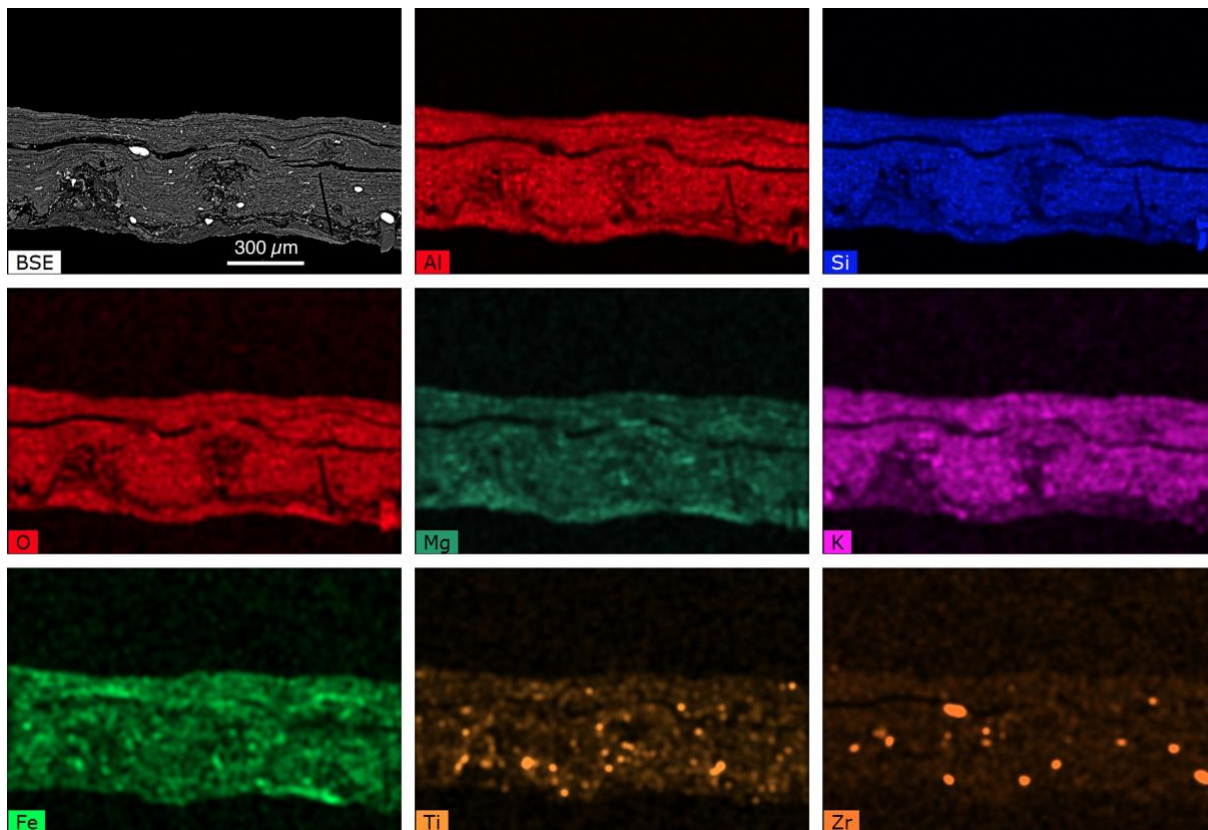
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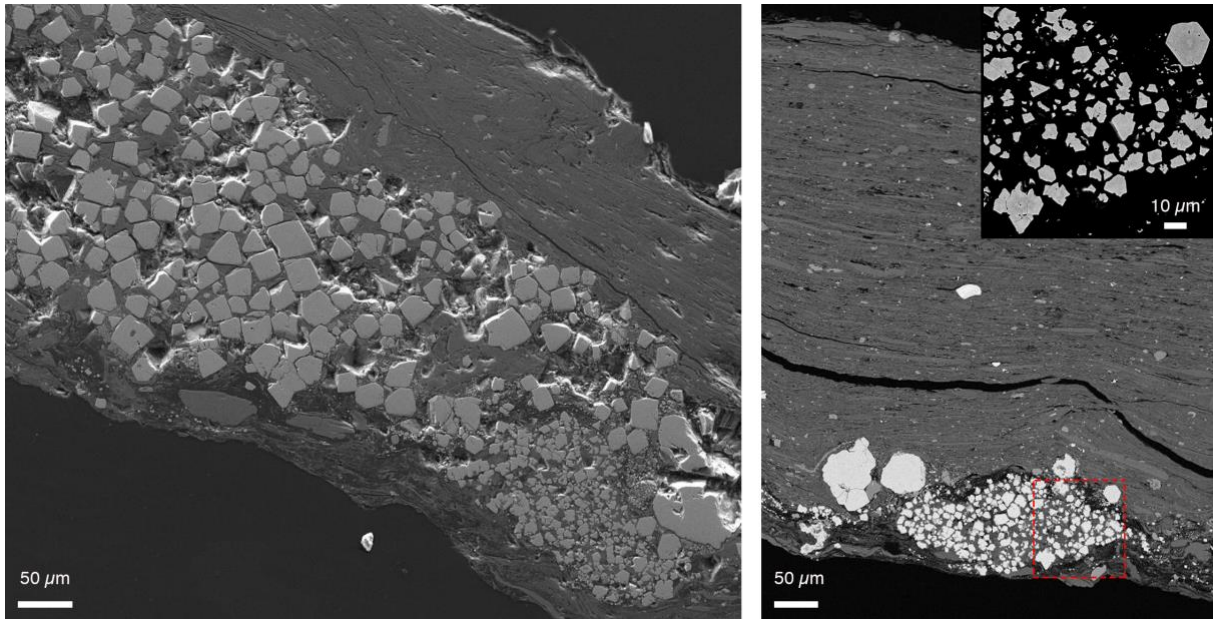


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955 **Figure S1.** Petrography, SEM, and EDX of a poorly pyritized MRS in cross-section
 956 perpendicular to the bedding plane. Back-scattered electron (BSE) and composite (Al, Si, O,
 957 Mg, K, Fe, Ti, and Zr) elemental maps show mineral composition of the poorly pyritized MRS.
 958 A number of Ti- and Zr-rich heavy mineral grains are embedded within these mats.

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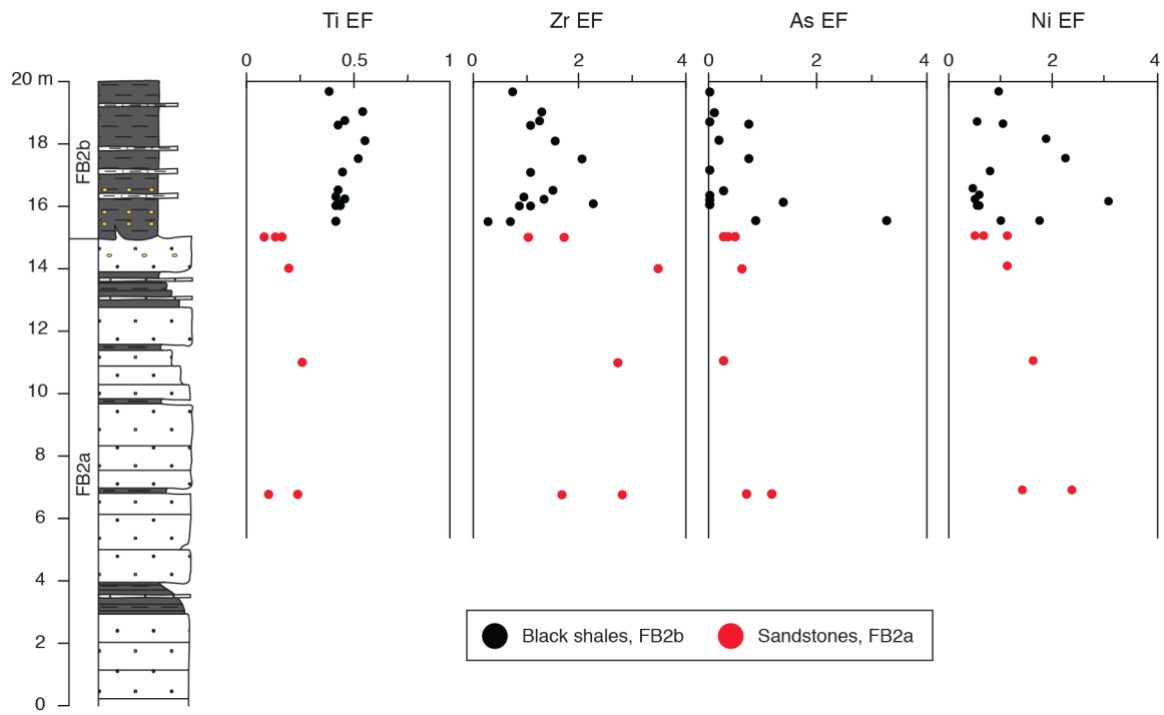
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962 **Figure S2.** SEM image of poorly pyritized MRS in secondary electron (left) and BSE (right)
963 modes. Pyrite-rich layers observed at the bottom of the MRS are highly variable in thickness.
964 Inset shows numerous, tiny sub-euhedral to euhedral pyrite crystals.

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967 **Figure S3.** Stratigraphic profile with enrichment factor (EF) of selected immobile and
 968 chalcophile elements within host sediments. EF are normalized to the average shale (Ti, Zr,
 969 and Ni concentrations for the average shale are from Taylor and McLennan (2001), and As
 970 concentration for the average shale is from Li and Schoonmaker, 2003).

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973 **TABLE S1.** Whole-rock composition of major and trace elements and enrichment factors (EF)
 974 of Ti, As, Ni, and Zr.

Formation	Lithology	Sample ID	Height m	S _r wt%	Si wt%	Al wt%	Fe wt%	Mn wt%	Mg wt%	Ca wt%	Na wt%	K wt%	Ti wt%	P wt%	EF _{Ti}	As ppm	Ba ppm	Bi ppm	Cd ppm	
FB2b	Pyritized MRS	FPS_s5-mat	18.7	29.37	10.83	5.35	26.88	0.08	0.81	0.42	0.14	1.56	0.08	0.02	0.27	179.20	754.50	4.24	0.64	
		DB_s4-mat	17.1	22.02	15.34	5.14	20.82	0.12	0.98	0.91	0.13	1.50	0.11	0.02	0.36	110.10	693.20	5.76	0.29	
		FPS_s4-mat	16.3	24.81	9.91	4.37	29.04	0.07	1.09	0.58	0.10	1.11	0.07	0.02	0.28	111.90	526.50	4.01	0.21	
		FPS_s3-mat	16.2	28.45	12.46	5.09	24.68	0.09	1.12	0.61	0.14	1.34	0.09	0.02	0.30	146.40	680.70	2.71	0.05	
		DB_s2-mat	16.0	23.88	12.99	6.26	14.57	0.06	0.87	0.36	0.16	1.89	0.12	0.02	0.32	58.37	983.70	0.05	0.05	
		DB_s3-mat	16.0	26.90	11.28	5.56	20.90	0.06	0.86	0.61	0.14	1.63	0.09	0.02	0.28	100.30	885.10	0.22	0.10	
		FR_s1-mat	16.0	6.46	23.94	8.33	7.23	0.19	1.46	1.49	0.23	2.52	0.20	0.02	0.42	0.80	1234.00	0.25	0.57	
		WS_s1-sed	19.7	0.09	27.50	8.83	2.58	0.38	2.06	1.61	1.61	0.25	2.45	0.20	0.02	0.38	0.36	1131.81	0.11	0.23
		KS_s1-sed	19.0	0.09	28.86	8.02	2.67	0.18	1.81	1.48	1.48	0.22	2.35	0.25	0.09	0.53	1.41	1136.75	0.15	0.11
		FPS_s5-sed	18.7	0.22	26.30	8.85	1.67	0.37	1.82	1.89	0.25	2.63	0.29	0.23	0.07	0.45	0.80	1231.00	0.20	0.52
FB2a	Pyrite-poor MRS	AFBSO-12	18.6	N.D.	9.87	2.41	0.05	1.18	0.26	0.29	3.01	0.25	0.07	0.43	10.93	1505.00	0.31	0.34		
		MLS_s7-sed	18.1	0.32	28.72	7.00	4.89	0.17	2.86	1.60	0.16	1.31	0.22	0.02	0.55	2.06	648.41	0.15	0.12	
		AFBSO-9	17.5	N.D.	25.55	6.38	4.55	0.57	3.33	3.72	0.17	1.22	0.19	0.05	0.52	6.99	651.30	0.16	0.39	
		DB_s4-sed	17.1	0.13	27.11	9.04	1.96	0.34	1.84	1.90	0.25	2.75	0.23	0.06	0.44	0.80	1294.00	0.21	0.42	
		AFBSO-7	16.5	N.D.	36.25	7.03	0.25	0.01	0.32	0.07	0.16	2.36	0.17	0.02	0.42	2.94	1204.00	0.19	0.19	
		FPS_s4-sed	16.3	0.16	30.88	6.37	6.21	0.09	1.85	0.17	1.61	0.16	1.12	0.16	0.07	0.41	0.80	1482.00	0.59	0.66
		FPS_s3-sed	16.2	0.17	26.98	9.85	1.54	0.20	1.64	1.48	0.26	3.10	0.24	0.07	0.46	0.80	1125.00	0.18	0.80	
		AFBSO-6	16.1	N.D.	26.91	8.51	1.64	0.24	1.79	1.85	0.27	2.48	0.25	0.23	0.07	0.42	13.16	587.60	0.24	0.54
		DB_s2-sed	16.0	0.16	26.18	9.45	1.50	0.23	1.72	1.60	0.26	2.98	0.24	0.24	0.02	0.43	0.81	1318.60	0.09	0.15
		FR_s1-sed	16.0	0.20	27.21	9.98	1.65	0.19	1.72	1.55	0.28	3.14	0.24	0.24	0.06	0.41	0.80	1448.00	0.32	0.27
MLS_s8-sed	15.5	1.02	25.34	10.97	2.71	0.37	1.86	1.50	0.29	3.57	0.27	0.25	0.02	0.41	53.11	1633.94	1.35	4.48		
MLS_s9-sed	15.5	0.34	26.27	10.13	2.12	0.28	1.80	1.50	0.27	3.21	0.25	0.25	0.02	0.42	13.23	1467.75	0.31	0.82		
FB2a	Pyrite-poor MRS	MLS_s1-mat	15.0	4.16	15.75	11.96	9.17	0.25	2.48	0.72	0.26	3.46	1.66	0.24	2.39	50.84	1539.00	1.52	2.61	
		LP_s3-mat	15.0	0.19	19.30	14.66	3.52	0.08	2.10	0.24	0.44	4.43	1.14	0.15	1.34	6.78	2035.00	0.86	0.65	
		MLS_s2-mat	15.0	3.08	15.44	12.34	6.61	0.08	2.47	0.31	0.21	3.36	2.03	0.16	2.83	28.01	1409.00	2.33	3.35	
		MLS_s6-mat	15.0	2.61	17.87	10.94	9.33	0.12	2.67	0.57	0.21	2.87	1.40	0.28	2.20	54.46	1225.00	1.43	1.84	
		MLS_s4-mat	15.0	0.36	16.62	12.99	3.98	0.08	2.61	0.49	0.21	3.68	1.92	0.22	2.62	7.71	1512.00	1.39	2.54	
		MLS_s5-mat	15.0	0.42	16.59	12.99	4.48	0.07	2.80	0.43	0.22	3.56	1.88	0.21	2.48	17.75	1488.00	1.64	2.59	
		EST_s3-mat	15.0	0.36	14.78	14.76	2.43	0.03	3.37	0.02	0.29	2.31	0.22	0.04	0.26	2.33	1453.62	0.40	0.14	
		EST_s4-mat	15.0	0.29	14.77	15.18	2.97	0.04	4.06	0.02	0.25	1.91	0.14	0.02	0.16	3.31	1203.89	0.28	0.12	
		MLS_s3-mat	11.0	0.84	20.68	16.82	1.74	0.02	1.74	0.12	0.36	4.76	0.52	0.20	0.53	1.05	2542.63	0.41	0.11	
		EST_s1-mat	6.9	0.17	16.93	16.41	1.90	0.01	3.34	0.07	0.37	2.96	0.20	0.02	0.21	1.56	2350.62	0.53	0.17	
EST_s2-mat	6.9	0.15	17.99	16.63	1.74	0.01	2.42	0.15	0.46	4.16	0.20	0.07	0.20	1.36	3253.14	0.34	0.16			
Sandstones		MLS_s1-sed	15.0	N.D.	41.65	2.47	0.68	0.09	0.58	0.47	0.08	0.50	0.02	0.02	0.13	1.90	281.40	0.18	0.27	
		LP_s3-sed	15.0	N.D.	43.70	1.30	0.46	0.05	0.38	0.43	0.03	0.26	0.07	0.02	0.08	0.80	145.20	0.05	0.05	
		MLS_s2-sed	15.0	N.D.	41.65	1.70	0.56	0.21	0.70	0.87	0.12	0.22	0.02	0.02	0.16	0.80	126.20	0.05	0.05	
		LP_s2-sed	14.0	N.D.	44.35	1.10	0.39	0.01	0.24	0.04	0.03	0.16	0.01	0.01	0.20	1.01	95.36	0.02	0.04	
		MLS_s3-sed	11.0	N.D.	44.09	1.51	0.47	0.02	0.35	0.01	0.02	0.18	0.02	0.02	0.26	0.70	97.90	0.02	0.03	
		EST_s1-sed	6.9	0.02	44.21	1.05	0.40	0.01	0.12	0.01	0.07	0.01	0.24	0.01	0.02	1.14	120.74	0.08	0.05	
EST_s2-sed	6.9	N.D.	44.79	0.78	0.44	0.01	0.10	0.05	0.05	0.02	0.15	0.01	0.03	1.35	82.64	0.02	0.05			

978 *Note:* FPS: flat, pyritized structure; DB: domal buildup; FR: “fairy-ring” structure; WS: wrinkle
979 structure; KS: “Kinneyia” structure; MLS: mat-layer structure; LP: linear pattern; EST:
980 “elephant-skin” texture as previously described by Aubineau et al. (2018).

981 EF_x refers to enrichment factor of X, where X stands for any chemical element.

982 Italicized numbers refer to samples for which half or more of the data were below the detection
983 limit. The detection limit is calculated as 6 times the standard deviation of the mean, plus the
984 mean, on 100 blank measurements.

985 N.D. = not determined.

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992 *Note:* FPS: flat, pyritized structure; DB: domal buildup; FR: “fairy-ring” structure; WS: wrinkle
993 structure; KS: “Kinneyia” structure; MLS: mat-layer structure; LP: linear pattern; EST:
994 “elephant-skin” texture as previously described by Aubineau et al. (2018).
995 Σ LREE refers to the sum of REE from La to Sm.
996 Σ HREE refers to the sum of REE from Gd to Lu.
997 REE_T includes all REE.
998 Eu/Eu* refers to Eu anomaly where Eu* is the interpolated value for Eu based on Sm and Gd.
999 Eu/Eu* (ppm/ppm) is calculated using the following equation: $Eu_{(SN)}/(0.66*Sm_{(SN)}+0.33*Tb_{(SN)})$.
1000 The subscript ‘SN’ indicates that the REE are normalized to shale composite (PAAS).
1001 N.D. = not determined.
1002
1003

1004 **TABLE S3.** Weight of each principal component.

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	eigenvalue	variance %	cumulative variance %
PC1	14.12	54.32	54.32
PC2	6.35	24.44	78.76
PC3	1.87	7.18	85.94
PC4	1.36	5.22	91.15
PC5	0.73	2.82	93.98
PC6	0.47	1.8	95.78
PC7	0.34	1.32	97.1
PC8	0.24	0.92	98.02
PC9	0.15	0.58	98.6
PC10	0.11	0.43	99.04
PC11	0.09	0.33	99.37
PC12	0.05	0.17	99.54
PC13	0.03	0.13	99.68
PC14	0.02	0.09	99.77
PC15	0.02	0.07	99.84
PC16	0.01	0.06	99.9
PC17	0.01	0.04	99.94
PC18	0.01	0.02	99.96
PC19	0	0.01	99.98
PC20	0	0.01	99.99
PC21	0	0.01	99.99
PC22	0	0	100
PC23	0	0	100
PC24	0	0	100
PC25	0	0	100
PC26	0	0	100

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