

1 **Geochemical records of the end-Triassic Crisis preserved in**
2 **a deep marine section of the Budva Basin, Dinarides,**
3 **Montenegro.**

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26 **Abstract**

27 The end-Triassic extinction event (~ 201.5 Ma) is one of the five major mass extinction events in
28 Earth's history, however, considerable discussion continues on the exact causes and timing of the
29 event. This is because, whilst certain geochemical data on T-J sections appears to be largely
30 comparable globally, with for example a significant (up to 6‰) negative carbon-isotope ($\delta^{13}\text{C}$)
31 excursion at the extinction horizon, more often than not other geochemical variations are neither
32 uniform nor fully consistent between sections. Critical to this discussion is that the majority of the
33 studied sections containing the end-Triassic extinction event are limited to shallow marine or
34 terrestrial sections, which are prone to discontinuities and hiatuses. In this study, we present
35 carbon isotopes ($\delta^{13}\text{C}_{\text{carb}}$), total organic carbon (TOC), major and trace, mercury (Hg) and highly
36 siderophile elements (HSE), osmium-isotope compositions and paleomagnetic data of a relatively
37 less studied deep-marine T-J succession in the Budva Basin, Čanj, Montenegro. At Čanj, deep-
38 marine Triassic limestones are abruptly interrupted by a ~6 cm finely laminated clay layer, before
39 transitioning to more argillaceous Jurassic red beds. The clay layer is interpreted to represent the
40 end-Triassic extinction interval and is characterized by a negative carbon isotope excursion,
41 relative heavy rare earth element (HREE) enrichment, Hg increase, HSE enrichment and a sharp
42 shift to unradiogenic osmium-isotopic ratios. This establishes the Čanj section as a unique and
43 well-preserved outcrop that exquisitely encapsulates the end-Triassic extinction in the Tethyan
44 marine realm. The distinct geochemical markers recorded at Čanj are consistent with the Central
45 Atlantic Magmatic Province as the main driver behind the end-Triassic extinction.

46

47

48 **1. Introduction**

49 The Late Triassic featured one of the five major mass extinction events of the Phanerozoic Eon.
50 Predating the end of the Triassic by ~140 ky, the onset of the End-Triassic extinction (ETE) event
51 (~ 201.5 Ma; Schoene et al., 2010; Blackburn et al., 2013; Wotzlaw et al., 2014) marked the demise
52 of close to 50% of known genera in both the marine and continental realms (see e.g., Raup and
53 Sepkoski, 1982; Bambach, 2006; Kiessling et al., 2007). The event coincided with an acute carbon-
54 cycle perturbation, as indicated by a large (up to 6‰) negative carbon-isotope ($\delta^{13}\text{C}$) excursion
55 (e.g., Pálffy et al., 2001; Ward et al., 2001; Hesselbo et al., 2002; Whiteside et al., 2010; Lindström
56 et al., 2017, 2021) in both marine and terrestrial sedimentary records, as well as stomatal index
57 and pedogenic carbonate evidence for a rise in atmospheric CO_2 (McElwain et al., 1999; Schaller
58 et al., 2011; Steinthorsdottir et al., 2011). Stratigraphic records of the extinction are also marked
59 by mercury (Hg; Thibodeau et al., 2016; Percival et al., 2017; Lindström et al., 2019; Kovács et
60 al., 2020; Yager et al., 2021) and highly siderophile element (HSE) enrichments (e.g., Ir, Ru, Rh,
61 Pt, Pd; Olsen et al., 2002a, b; Whiteside et al., 2021), as well as a shift towards a relatively
62 unradiogenic $^{187}\text{Os}/^{188}\text{Os}_i$ composition of seawater (Cohen and Coe, 2002; Kuroda et al., 2010).
63 The ETE was contemporaneous with the early onset of volcanism associated with the Central
64 Atlantic Magmatic Province (CAMP), based on radioisotopic dating of volcanic ash layers just
65 above the extinction horizon, CAMP igneous units themselves, and the documentation of CAMP
66 lavas interbedded with sedimentary records of the event (e.g., Schoene et al., 2010; Marzoli et al.,
67 2011; Blackburn et al., 2013; Wotzlaw et al., 2014; Davies et al., 2017). Consequently, increased
68 emissions of CO_2 and other gases such as SO_2 (Bacon et al., 2013; Steinthorsdottir et al., 2018)
69 from CAMP volcanic outpouring and/or sill intrusions are widely implicated as the main trigger
70 of the extinction (e.g., Wignall, 2001; Deenen et al., 2010; Davies et al., 2017; Heimdal et al.,

71 2018). The Hg and platinum-group element (PGE: Ru, Rh, Pd, Os, Ir, Pt) enrichments and shifts
72 towards more unradiogenic osmium-isotope seawater compositions have been explained as a
73 further consequence of, and proxy for, this volcanism (e.g., Kuroda et al., 2010; Thibodeau et al.,
74 2016; Percival et al., 2017; Lindström et al., 2019; Whiteside et al., 2021). Alternative hypotheses
75 have been proposed, such as methane hydrate release as an alternative source of the carbon
76 emissions (Beerling and Berner, 2002), and a meteorite impact event as a source of the HSE
77 enrichment (Olsen et al., 2002a,b; de Graaff et al., 2017), which is similar, though an order of
78 magnitude lower, to the well-documented HSE enrichment associated with the Cretaceous-
79 Paleogene (K-Pg) boundary mass extinction (Alvarez et al., 1980, Smit and Hertogen, 1980;
80 Goderis et al., 2013, 2021). However, the mounting evidence for CAMP volcanism over methane
81 hydrate release from other sources (see e.g., Heimdal et al., 2018, 2019; Lindström et al., 2019;
82 Whiteside et al., 2021) and the lack of additional supportive evidence for an impact event such as
83 impact spherules, Ni-rich spinel crystals, shocked quartz or an impact structure of a correct age
84 (see e.g., Tanner et al., 2004, 2008), leaves CAMP volcanism as the more widely accepted cause.

85 The emplacement of the CAMP initiated the break-up of Gondwana and Pangaea,
86 ultimately forming one of the largest continental flood-basalt provinces on Earth, exceeding 7
87 million km² in aerial extent (Marzoli et al., 1999). The earliest known CAMP magmas formed
88 intrusive bodies in Africa and South America at $\sim 201.635 \pm 0.029$ Ma, roughly 100 kyr prior to
89 the ETE (Davies et al., 2017). The oldest dated extrusives of the main CAMP volcanic activity
90 yields 201.566 ± 0.031 Ma (North Mountain Basalt; Blackburn et al., 2013). This earliest known
91 pulse of extrusive magmatism is thought to have been coeval with the geologically abrupt negative
92 carbon-isotope excursion (named the Initial, or Marshi carbon-isotope excursion: ICIE; Hesselbo
93 et al., 2002; Lindström et al., 2017) and a dramatic turnover in marine fauna (Ward et al., 2001;

94 Hori et al., 2007; Deenen et al., 2010; Ruhl et al., 2011). Additional intrusive sills from the
95 Amazonas Basin (Brazil) have also been dated to that time and could have resulted in the
96 generation of non-magmatic gases via the heating of volatile-rich lithologies (Davies et al., 2017;
97 Heimdal et al., 2018). The initial onset of volcanism in Morocco was rapidly (within a few kyr)
98 followed by a series of magmatic pulses recorded by lavas in both NE North America and Morocco
99 (Deenen et al., 2010), with eruptions also documented in SW Europe. Whilst the exact sequence
100 of environmental and biospheric perturbations remains debated, with, for example, a carbon-
101 isotope excursion preceding the ICIE (termed the pre-cursor CIE; Ruhl and Kürschner, 2011)
102 suggesting that a global carbon cycle disturbance preceded the onset of CAMP, the main CAMP
103 eruptions broadly coincided with a calcification crisis in calcareous nannofossils (van de
104 Schootbrugge et al., 2007), extinction of numerous faunal groups, including the conodonts, and
105 the development of widespread marine anoxia (Wignall and Bond, 2008; Kasprak et al., 2015; Jost
106 et al., 2017a). Furthermore, long-lasting disturbances to the global environment continued long
107 after the onset of major CAMP volcanism, with evidence for a further carbon-cycle perturbation
108 (recorded as the Main CIE), oceanic acidification, marine faunas dominated by bio-siliceous taxa,
109 and at least localized marine anoxia recorded from several locations (e.g., Hesselbo et al., 2002;
110 van de Schootbrugge et al., 2007; Greene et al., 2012; Richoz et al., 2012; Kasprak et al., 2015;
111 Thibodeau et al., 2016; Jost et al., 2017a,b), before the first occurrence of Jurassic ammonites
112 ~100–200 kyr after the onset of the extinction marking the start of the Jurassic (201.36 Ma;
113 Schoene et al., 2010; Wotzlaw et al., 2014). U-Pb geochronology indicates that CAMP
114 emplacement continued for at least 600 kyr after its onset (see Marzoli et al., 2018; and references
115 therein), broadly similar to the durations of other well dated LIPs (e.g., Siberian and Deccan Traps;
116 Burgess et al., 2017; Schoene et al., 2019; Sprain et al., 2019). However, $^{40}\text{Ar}/^{39}\text{Ar}$ ages suggest a

117 considerably longer duration for the CAMP, with some basalts dated as being up to 10 My younger
118 than the oldest erupted products (Jourdan et al., 2009; Marzoli et al., 2011), and potentially
119 suggesting that (minor) CAMP volcanism continued until long after faunal recovery took place in
120 the Jurassic. With a clear synchronicity between the extinction and CAMP eruptions, but possible
121 disconnect between the continued outpouring of CAMP and faunal recovery, it is apparent that
122 questions remain regarding the exact mechanisms behind the extinction, and the respective roles
123 of magmatic degassing, thermogenic emissions, and/or other processes such as methane clathrate
124 destabilization or a possible impact event.

125 Critical to this discussion is that most of the studied sections containing the ETE event are
126 limited to shallow marine or terrestrial sections, which are prone to discontinuities and hiatuses
127 (see e.g., Lindström et al., 2019 for an overview), with comparatively few examples of well-
128 studied deep-marine T-J records, with those that do exist being limited to Panthalassic Ocean
129 sections (see e.g., Hori et al., 2007; Fujisaki et al., 2018). In Čanj, Montenegro, a relatively less
130 studied continuous deep-marine section of the T-J boundary exists; first described by Goričan
131 (1994), the Čanj section has since that time only been further investigated by Črne et al. (2011)
132 (Fig. 1A and B). These studies presented stable-isotope records and radiolarian data, which support
133 a Triassic–Jurassic age and preservation of a relatively complete ETE horizon. In this study, we
134 present carbon isotopes ($\delta^{13}\text{C}_{\text{carb}}$), total organic carbon (TOC), major and trace elements, with a
135 special focus on mercury (Hg), highly siderophile elements (HSE), osmium-isotope compositions
136 and paleomagnetic data of the Čanj section. Through this approach we aim to provide a detailed
137 geochemical record from a deep-marine succession in the Tethyan realm, and present new insights
138 into the causes of the ETE.

139

140 **2. Geological setting**

141 The Čanj section is located in coastal Montenegro near the village of Čanj (42° 9' 40.04" N, 18°
142 59' 29.85" E; Fig. 2). It is part of the Budva tectonic zone of the External Dinarides (Petković,
143 1956; Goričan, 1994), which encompasses several south-west verging thrust units spanning from
144 Budva to Hercegnovi (Goričan, 1994). The Budva Zone separates the Dalmatian Zone (Aubouin,
145 1960) in the southwest from the High Karst Zone (Kossmat, 1924) in the northeast. The Budva
146 Zone is generally subdivided into a lower and upper tectonic unit (Goričan, 1994), with the Čanj
147 section representing part of the lower unit. The subdivision in tectonic units corresponds
148 approximately to the Mesozoic paleogeography, with sediments deposited in the lower Budva
149 Zone tectonic unit thought to be more distal with respect to the upper tectonic unit (Goričan, 1994).
150 Paleogeographically, the Budva Zone records a narrow area of deposition, commonly referred to
151 as the Budva Basin, located between two carbonate platforms, the Adriatic Carbonate Platform in
152 the west and the Dinaric Carbonate Platform in the east (D'Argenio et al., 1971). This intra-
153 platform deep-marine basin was formed by the rifting and separation of Adria from Africa during
154 the Middle and Late Triassic, separating the Adriatic and Dinaric Carbonate Platforms (Čadjenović
155 et al., 2008; Schmid et al., 2020). The basement rock of the Budva Basin consists of volcanic rocks
156 of Middle Triassic age (Fig. 2). The oldest sediments deposited on the volcanic basement are Upper
157 Triassic limestones, which consist of thin beds of pelagic limestone with intercalated chert nodules
158 and infrequent (thin) chert or clay layers (Fig. 2; Goričan, 1994; Črne et al., 2011; Đaković et al.,
159 2018; van Unen et al., 2019). The Triassic succession is conformably overlain by near-continuous
160 Lower Jurassic red radiolarites, limestones and shales and Middle Jurassic to Upper Cretaceous
161 thinly bedded pelagic micritic limestone and radiolarite lithologies (Fig. 2; Goričan, 1994; Črne et
162 al., 2011). The entire stratigraphic succession contains carbonate gravity-flow deposits and

163 turbidites (Črne et al., 2011). Sedimentation stopped with lower Eocene flysch (Fig. 2; Schmid et
164 al., 2020), which provides a time constraint regarding the age of thrusting of the High Karst unit
165 over the Budva Zone (Schmid et al., 2020).

166

167 **3. The Čanj section**

168 The Čanj section exposes at least 20 m of the Upper Triassic Halobia limestone Formation, which
169 consists of thin beds of pelagic carbonates with intercalated chert nodules and chert layers (varying
170 from a few cm to upwards of 15 cm), alternated with shales and/or marls (Fig. 1A). This part of
171 the section is characterized by a near-continuous repetition of grey, fine-grained, and fossiliferous
172 limestone beds (Črne et al., 2011) of ~1 – 20 cm thickness followed by a marl or shale bed, or a
173 combination of both (varying between 0.1 ~ 2 cm in thickness) (Fig. 1A; Appendix A). The
174 Halobia limestone formation is interrupted at Čanj by two slump intervals. The first slump occurs
175 ~ 9 m below the top of the Halobia limestone and has a varying thickness of approximately 1.65
176 m at its maximum. It contains clasts of reworked material originating from the beds below.
177 Underlying limestone beds terminate laterally due to incision of the slump. The second,
178 approximately 1 m thick, slump occurs ~ 4.5 m below the top of the formation. This slump is less
179 brecciated and instead has an internal structure of undulating layers with chert nodules, which are
180 not laterally continuous. These slumps are the only gravity-flow deposits observed in the Halobia
181 limestone formation at the Čanj section, marking the top of the last slump as the start of a
182 continuous sedimentary succession until well into the Jurassic.

183 The Halobia limestone is overlain by the Lower Jurassic Passée Jaspeuse Formation (Črne
184 et al., 2011), which consists of red siliceous limestones, radiolarian cherts, sandstones, siltstones,
185 shales, marls, and clays well over 30 m in thickness (Fig. 1A; Goričan, 1994). The boundary

186 between the Halobia limestone and the Passée Jaspeuse is marked by a clay layer of varying
187 thickness, which can be subdivided into 2 distinct clay beds (Fig. 1B). At the base, a bed of grey
188 to red, platy shale of ~ 3 cm, which is slightly undulating with thin laminae of less than 0.1mm
189 and no visible grains. This bed is overlain by a ~3.5 cm, dark brown to grey black clay. No distinct
190 bedding or laminae are visible and calcite grains, likely secondary in origin, have been observed.
191 Above this clay layer, the basal strata of the Passée Jaspeuse remain relatively argillaceous for tens
192 of centimeters before carbonate-rich lithologies reappear. A switch from Triassic to Jurassic
193 radiolarian fauna somewhere between ~ 90 cm below the clay layer to ~ 40 cm above it (near the
194 top of the carbonate-depleted layers), together with documentation of the ICIE correlative with the
195 clay layer itself (Črne et al., 2011), supports placement of the Triassic–Jurassic boundary slightly
196 above the clay layer interval.

197

198 **4. Methods and materials**

199 *4.1. Sample selection and preparation*

200 The Čanj section was visited and sampled in May 2019. The entire Čanj section was logged
201 from ~ -6.5 m to + 21.8 m (0 m marking the base of the clay layer); at the millimeter scale with
202 continuous sampling every few centimeters (Appendix A). Further sample selection and
203 geochemical analyses were carried out at the Analytical, Environmental and Geo-Chemistry
204 (AMGC) laboratory at the Vrije Universiteit Brussel (Brussels, Belgium; VUB), except for HSE
205 and Re-Os isotopic analyses, which were done at the University of Tokyo at Komaba and the Japan
206 Agency for Marine-Earth Science and technology (JAMSTEC) in Japan, respectively. Eighty
207 samples were selected for carbon-isotope analysis, to construct a continuous curve across the ETE
208 interval, spanning strata from just above the last slump of the Halobia limestone at around -4 m

209 and up to the highest sampled carbonate layer of the Passée Jaspeuse at +21m. Further sample
210 selection for total organic carbon (TOC) contents (n=40), major and trace element analyses (n =
211 40), with S being done separately (n = 36), Hg contents (n = 40) and HSE and osmium-isotopic
212 analyses (n = 21) focused around the clay layer to resolve potential variations in the concentrations
213 and isotope ratios of these elements across the extinction horizon (see Appendix B and III).
214 Sampling for paleomagnetic analyses (n = 24) was done on site using a modified STIHL drill with
215 diamond tipped Hoffmann drill bits. Sample locations were chosen based on the expected
216 occurrences of magnetic reversal intervals using the sedimentation rate estimates in the Halobia
217 limestone and the Passée Jaspeuse (Črnek et al., 2011; de Graaff et al., 2017) and
218 magnetostratigraphy data of the T-J interval (Korte et al., 2019).

219 For carbon-isotope analyses, samples were powdered using a hand-held drill on fresh
220 surfaces and collected in clean glass vials. For major and trace elements, including TOC contents,
221 Hg and HSE concentrations, and osmium isotope analysis, approximately 5 cm³ of sample was
222 crushed to a homogenous powder using an agate mortar and pestle, followed by sieving until a ~
223 <125 μm grain size was achieved. Larger rock samples were cut using a diamond board table saw
224 beforehand and were subsequently washed with ultrapure 18.2 MΩ cm water in an ultrasonic bath
225 before crushing. Samples that required sawing were not selected for HSE or ¹⁸⁷Os/¹⁸⁸Os analyses,
226 to avoid possible contamination.

227

228 *4.2. Carbon-isotope composition analysis*

229 Carbon-isotope ratios of bulk carbonate material ($\delta^{13}\text{C}_{\text{carb}}$) were determined on a Nu
230 Instruments Nu Perspective isotope-ratio mass spectrometer (IRMS) coupled to a gas preparation
231 (GasPrep) automated gas bench at the VUB. Between 1–10 mg of homogenized powder

232 (depending on carbonate content) was weighed into a glass vial, which was then sealed and flushed
233 with helium gas, before addition of phosphoric acid to liberate the carbonate-bound CO₂ over a
234 few hours. Data calibration, and, if necessary, machine drift correction, was carried out using
235 multiple aliquots of Carrara marble ($\delta^{13}\text{C} = +3.41 \text{ ‰}$) on a daily basis for each individual batch,
236 while international reference material IAEA 603 ($\delta^{13}\text{C} = +2.46 \text{ ‰}$) was also used to assess
237 analytical precision. Measured results for IAEA 603 had an average of 2.49 ‰, consistent with the
238 certified value. Measurement uncertainty on these standards was $\pm 0.07 \text{ ‰}$ (1σ , $n = 30$) for Carrara
239 marble and $\pm 0.05 \text{ ‰}$ (1σ , $n = 25$) for IAEA 603.

240

241 *4.3. Total organic carbon analysis*

242 For each sample, between 2–3 grams of homogenized sample powder were decarbonated
243 using 10% HCl following the method outlined in Percival et al. (2022). The organic carbon content
244 of the treated samples was determined on a Nu Instruments Horizon 2 isotope-ratio mass
245 spectrometer (IRMS) coupled to a Eurovector elemental analyzer EuroEA3000 at the VUB. These
246 measurements were then converted to bulk TOC contents by accounting for the mass lost during
247 decarbonation. Data accuracy and reproducibility were monitored through analysis of international
248 standards IVA33802151 (organic-rich sediment) and IVA33802153 (organic-poor soil), yielding
249 9.06 ± 0.11 (1σ , $n = 4$) and 1.57 ± 0.03 (1σ , $n = 3$) respectively.

250

251 *4.4. Major and trace element composition analysis*

252 Approximately 100 mg (± 0.1 mg) of homogenized sample powder was weighed into in
253 trace metal-clean 15 ml polytetrafluoroethylene (PTFE) beakers. A mixture of 3 ml 14 M HNO₃
254 and 1 ml 29 M HF was added to the PTFE beakers, after which these were left to react for 4 days

255 at 120°C. The samples were subsequently dried down, re-dissolved in 2 ml of 14 M HNO₃ and left
256 to react for another 3 days at 90°C until the liquid was completely clear of any residue and 3 ml of
257 ultrapure 18.2 MΩ cm water was added. Exclusively high-purity trace metal grade acids were used
258 in the procedure. After adequate dilution (by factor 2500 to 5000) and addition of indium as an
259 internal standard, selected major and trace element concentrations were determined using the
260 Thermo Scientific Element 2 high-resolution inductively coupled plasma-mass spectrometer (ICP-
261 MS) housed at the AMGC research unit at the VUB. Both the low and medium resolution modes
262 of the instrument were used for one or more monitored isotopes of the element of interest, and
263 mathematical corrections were applied for isobaric interferences where necessary (as commonly
264 applied for the rare earth elements (REE), see e.g., de Graaff et al., 2022). Elemental concentrations
265 were determined versus an external calibration curve prepared from single element standard
266 solutions. Reference basalt BE-N of the Centre de Recherches Pétrographiques et Géochimiques,
267 basalt BIR-1 of the United States Geological Survey, and limestone CCB-1 of Liège Université
268 were digested and analyzed as secondary standards using the same procedure as the samples. All
269 reported values were consistent with the certified standard data. A fraction of randomly selected
270 samples was digested twice to test sample heterogeneity, after which the results were compared
271 and averaged. The average bias between reference values and experimentally obtained values for
272 international reference values was typically 5 to 10%. Based on the repeated analysis of the
273 reference materials and a subset of samples, the external reproducibility for the elements measured
274 is better than 10% relative standard deviation (RSD) depending on the concentration level.

275

276 *4.5. Sulphur content analysis*

277 Sulfur concentrations of 36 samples across the Čanj T-J boundary were measured using micro-X-
278 ray fluorescence (μ XRF). Besides semi-quantitative high-resolution element mapping, this non-
279 destructive technique also allows quantitative spot analysis to determine major and trace element
280 concentrations (de Winter and Claeys, 2016; Kaskes et al., 2021), with the determination of low
281 concentrations of S (< 0.1 wt%) proven to be successful (Gulick et al., 2019). We used an M4
282 Tornado benchtop μ XRF surface scanner (Bruker nano GmbH, Berlin, Germany) equipped with a
283 Rh tube as X-ray source and two XFlash 430 Silicon Drift detectors, available at VUB (de Winter
284 and Claeys, 2016). Homogenized powders were measured under near vacuum conditions (20
285 mbar), without the use of an X-ray source filter, and using repeated spot analysis ($n = 10$) per
286 powder with a spot size of $200 \mu\text{m}$ and an integration time of 120 s per spot (following Vellekoop
287 et al., 2022). This integration time was selected to allow the Time of Stable Reproducibility and
288 Time of Stable Accuracy to be reached, which allows the concentration of a range of elements to
289 be quantified (de Winter et al. 2017). A range of certified geological reference materials ($n = 18$)
290 was measured under the same conditions to perform a matrix-matched multi-standard calibration
291 to correct the pre-determined S-data based on the Fundamental Parameters Method. Limit of
292 detection of S was determined to be in the order of 50 ppm based on repeated spot analysis ($n =$
293 10) of certified carbonate reference material BCS CRM393 (Bureau of Analyzed Samples Ltd.,
294 Middlesbrough, UK).

295

296 *4.6. Mercury concentration analysis*

297 Mercury (Hg) concentrations were determined on an Advanced Mercury Analyzer (AMA)
298 254.7 at the VUB, broadly following the method outlined in Liu et al. (2021). Approximately 100
299 mg of homogenized bulk-rock powder was analyzed for clay-rich samples, whilst ~ 250 mg was

300 used for measurements of carbonate-rich samples. Mercury was volatilized from the untreated rock
301 powder at 750 °C, and collected in an amalgamator containing a gold trap, before being analyzed
302 by atomic absorption spectrometry. Blank measurements on the AMA were better than 0.05 ng.
303 Each sample was analyzed at least twice, with measurement reproducibility typically better than
304 5%, and the mean concentration was taken. Measurement accuracy was confirmed through
305 analysis of multiple aliquots of the international reference materials SRM MESS-3 (89.87 ± 0.6
306 ppb, 1σ , $n = 3$) and JP-1 (5.23 ± 0.29 ppb, 1σ , $n = 9$), consistent with certified values (SRM MESS-
307 3 = 91 ppb Hg) and JP-1 = 5.3 ppb Hg).

308

309 *4.7. Highly siderophile element concentrations and Re-Os isotopic analysis*

310 Highly siderophile element (HSE: Os, Ir, Ru, Pt, Pd, and Re) concentrations and osmium-isotope
311 compositions were determined by isotope dilution mass spectrometry (ID-MS). Sample digestion
312 was done with quartz glass tube ($\phi = 14$ mm, $L = 290$ mm) digestion using inverse aqua regia on
313 powdered samples (~ 0.1 – 0.5 g) and subsequent complete desilicification in 38% HF on the
314 residual solids (used only for Ir, Ru, Pt, Pd and Re). Chemical purification of osmium was
315 performed using standard microdistillation into HBr using $\text{Cr}^{\text{VI}}\text{O}_3\text{-H}_2\text{SO}_4$, following solvent
316 extraction with CCl_4 (Birck et al., 1997; Cohen and Waters, 1996). The detailed methods for
317 sample digestion, chemical purification and mass spectrometry are based on those reported in
318 Ishikawa et al. (2014), following adaptations presented by Sato et al. (2021).

319 The concentrations for Ir, Ru, Pt, Pd, and Re were measured using a Thermo Element XR
320 ICP-MS at the University of Tokyo at Komaba. For sample introduction, a combination of a 100
321 $\mu\text{L}/\text{min}$ PFA self-aspirating nebulizer and dual cyclonic/Scott double-pass spray chamber was used
322 during all measurements, and the oxide level (based on the HfO/Hf) was set to $\sim 1\%$. Sample and

323 standard solutions were interspersed throughout the analytical sessions to monitor and correct for
324 instrumental fractionation. The average total procedural blanks for the analysed elements were
325 0.20 ± 0.05 pg Ir, 0.17 ± 0.08 pg Ru, 8.1 ± 4.0 pg Pt, 14.6 ± 3.6 pg Pd, and 0.39 ± 0.09 pg Re (n=3,
326 1SD). All analyses were blank corrected to account for the variable background contribution for
327 each sample, depending on their HSE concentrations (0.096–12% for Ir, 0.27–7.5% for Ru, 0.22–
328 19% for Pt, 0.45–27% for Pd, and 0.038–26% for Re). The reported uncertainties on each sample
329 were calculated by error propagation of the ICP-MS measurement uncertainty (2 SE) and the blank
330 correction.

331 Osmium concentrations and isotopic compositions were measured by negative thermal
332 ionisation mass spectrometry (N-TIMS, Thermo Triton Plus) at JAMSTEC, Japan. The average
333 total procedural Os blank was 0.12 ± 0.02 pg with a $^{187}\text{Os}/^{188}\text{Os}$ ratio of 0.159 ± 0.008 (n=3, 1SD).
334 Blank corrections were applied for all analyses, although the blank contributions to the measured
335 Os concentrations and $^{187}\text{Os}/^{188}\text{Os}$ ratios were almost negligible for all samples and less than 4%
336 and 3%, respectively. The uncertainties of $^{187}\text{Re}/^{188}\text{Os}$ and $^{187}\text{Os}/^{188}\text{Os}$ were calculated by error
337 propagation of the blank uncertainties. The reproducibility for basaltic reference material BIR-1a
338 yielded 0.4% RSD for $^{187}\text{Os}/^{188}\text{Os}$, 8.6% RSD for Os, 8.0% RSD for Ir, 4.0% RSD for Ru, 3.8%
339 RSD for Pt, 2.7% RSD for Pd, and 1.2% RSD for Re (~0.5 g powdered samples; Goderis et al.,
340 2021), consistent with data for larger sample amounts (~1–2 g; Ishikawa et al., 2014). The
341 measured $^{187}\text{Os}/^{188}\text{Os}$ ratios were corrected for post-depositional decay of rhenium to osmium by
342 utilising the established concentrations of rhenium and osmium in the samples, together with their
343 estimated age (201.3 Ma) to give the initial isotopic composition of the sediment at the time of
344 deposition ($^{187}\text{Os}/^{188}\text{Os}_i$), following standard protocols (Cohen et al., 1999).

345

346 4.8. Paleomagnetic and magnetic property analysis

347 Paleomagnetic analyses were carried out on 24 samples (two specimens per sample) at the
348 Geophysical Centre of the Royal Meteorological Institute of Belgium (Dourbes, Belgium). The
349 cylindrical standard-sized palaeomagnetic samples were subjected to stepwise alternating field
350 (AF) demagnetization in steps from 0 to 70 mT. The remaining remanent magnetization was
351 measured after each step. To do so, an upgraded three-axis, model 760, 2G Enterprises, cryogenic
352 magnetometer with DC SQUIDS and sample access of 7.6 cm was used. The magnetometer is
353 equipped with a Cryomech 4K, model PT405, pulse tube cryorefrigerator; an inline three-axial
354 alternating field demagnetizer and automatic sample handling system.

355 A subset of 6 samples from different stratigraphic intervals were selected for magnetic
356 property analyses with a magnetic property measurement system (MPMS3) from QuantumDesign.
357 For each sample field cooled and zero-field cooled hysteresis loops were acquired within a field
358 range of $\pm 5 T$ at following temperatures 20, 50 and 300 K. In addition, field-cooled direct current
359 susceptibility warming curves were measured to reveal characteristic low temperature transitions
360 such as Verwey and Morin to support the interpretation of the magnetic mineralogy.

361

362 5. Results

363 5.1. Carbon-isotope compositions

364 Our $\delta^{13}\text{C}_{\text{carb}}$ results are consistent with those of Črne et al. (2011). Values are typically
365 between 1–2 ‰ in the Halobia limestone formation, with the lowest $\delta^{13}\text{C}_{\text{carb}}$ ratios occurring as
366 part of a minor fall in values below the base of the clay layer, between -1.7 and -2.6 m (Fig. 3). A
367 significantly more pronounced negative excursion is recorded across the Halobia limestone–
368 Passée Jaspeuse boundary, with values falling to -1.18 ‰ just below the base of the clay layer and

369 remaining at ~ -1 ‰ within that level (Fig. 3). Three clay-layer samples record considerably lower
370 $\delta^{13}\text{C}_{\text{carb}}$ ratios (< -4 ‰), but it should be noted that those intervals featured a very low carbonate
371 content; thus, these low values may reflect poor analytical precision (Appendix B). The $\delta^{13}\text{C}_{\text{carb}}$
372 ratios return to pre-excursion ratios around 25 cm above the base of the clay layer, before a second,
373 smaller, negative excursion of around 0.5 ‰ occurs approximately 75 cm higher up in the
374 succession. Following this second shift, there is a gradual decline of ~ 0.5 ‰ in $\delta^{13}\text{C}_{\text{carb}}$ until 5–6
375 m above the base of the clay layer, which is succeeded by a slight rise of ~ 0.5 ‰ in values across
376 the next 7 meters. The upper 7 meters of the studied strata record two additional transient negative
377 excursions in $\delta^{13}\text{C}_{\text{carb}}$, of between 2 and 3 ‰ in magnitude (Fig. 3).

378

379 *5.2. Total organic carbonate contents*

380 Sedimentary TOC contents are presented in Appendix C and are shown to be very low
381 throughout the boundary interval of the section. The Halobia limestones are marked by TOC
382 contents of < 0.1 wt% on average, with just two thin argillaceous horizons 88cm and 145 cm below
383 the base of the clay layer reaching slightly higher organic content levels (0.13 wt% and 0.16 wt%,
384 respectively). There is a slight increase in TOC within the four clay layer samples, reaching a
385 maximum of 0.29 wt% 2 cm above its base. Above the clay layer, Passée Jaspeuse samples return
386 to very low TOC contents (< 0.1 wt%), similar to the Halobia limestones.

387

388 *5.3. Major and trace element abundances*

389 Major and trace element compositions are presented in Appendix C. Notable major element
390 oxide variations are shown by differences in MgO, Al₂O₃ and TiO₂ contents of samples from the
391 Halobia limestones, Passée Jaspeuse and the clay layer, with the latter featuring clearly elevated

392 MgO, Al₂O₃ and TiO₂ levels (up to 9.96, 14.9 and 0.77 wt%, respectively) compared to an average
393 of 1.17 wt%, 3.07 wt%, and 0.11 wt% for the Halobia limestones and 2.08 wt%, 8.63 wt%, 0.39
394 wt% for the Passée Jaspeuse. These variations follow the dominant lithological changes between
395 both formations and the clay layer. Moreover, they highlight the presence of more clay rich layers
396 in the Halobia limestone at the -265, -247.5, -182 and -145.5 cm sample heights (Appendix C).
397 This is further substantiated with these intervals being relatively enriched in Th (up to 13.8 ppm
398 compared to 2.21 ppm average in calcareous/siliceous samples), an element typically associated
399 with clay content in sediments (e.g., Plank, 2014). Micro-X-ray fluorescence analyses reveal that
400 sulfur concentrations are consistently low (<300 ppm; Appendix D) across the entire interval, with
401 the Triassic Halobia limestone Formation yielding slightly higher values (~100-200 ppm;
402 Appendix D) compared to the Jurassic Passée Jaspeuse Formation, for which all but one sample
403 have sulphur concentrations below the limit of detection (50 ppm for S; Appendix D).

404 The Halobia limestones display considerable variability in absolute trace element content
405 between the samples, yet show generally consistent chondrite normalized trace element patterns
406 (Fig. 4). The latter show an overall trend of relative light rare earth element (LREE) enrichment
407 compared to the heavy rare earth elements (HREE; Fig. 4). Relative to CI-chondritic values (Sun
408 and McDonough, 1989), La is observed to be up to 100 times chondrite compared to <1 times
409 chondrite for Lu, and the limestones are characterized by a distinct negative Ce anomaly. A similar
410 trend is observed for the Passée Jaspeuse, although with higher LREE contents (up to ~270 times
411 chondrite for La) overall, and less variation in absolute concentrations between the samples (Fig.
412 4; Appendix C). All element concentrations surpass 4 times chondrite in the Passée Jaspeuse
413 samples (Fig. 4), and no pronounced Ce anomalies are documented. By contrast, the clay layers
414 show a distinctly different trend with relative enrichment in both LREE and HREE (Fig. 4).

415 5.4. Mercury concentrations

416 Mercury concentrations in Halobia limestone samples are generally very low (median of 0.38 ppb,
417 with most below 0.1 ppb), likely due to the carbonate-rich lithology. For almost all limestone
418 samples, a Hg content below the reliable detection limit of 0.05 ng (i.e., sub-blank level) was
419 measured from 250 mg of analyzed material, indicating a negligible mercury concentration. By
420 contrast, Hg contents in the argillaceous samples range from 0.58–28.5 ppb (Fig. 5). There is a
421 sharp spike in Hg concentrations spanning 20–30 cm across the Halobia limestone–Passée
422 Jaspeuse boundary, beginning just below the clay layer in the uppermost Halobia strata, and
423 continuing 10 cm above the base of the Passée Jaspeuse (Fig. 5). This excursion is recorded by
424 two samples from the clay layer, which have Hg contents of up to 117 ppb. Further up the Passée
425 Jaspeuse stratigraphy, Hg levels return to relatively low values, reaching a maximum of 7.02 ppb,
426 91 cm above the base of the clay layer (Fig. 5). When normalizing the Hg concentrations to Al₂O₃
427 it is noted that the enrichments persist (Fig. 5), even elucidating a peak before the clay layer,
428 suggesting a lack of correlation between Hg enrichment and clay content (Fig. 5).

429

430 5.5. Highly siderophile element concentrations

431 Concentrations of Os, Ir, Ru, Pt, Pd and Re are low for most samples from the Čanj section
432 (Fig. 6; Appendix C). The exception is enrichment of HSEs in the main clay layer, where Os, Ir,
433 Ru, Pt, Pd and Re contents reach up to 3.33 ppb, 0.41 ppb, 0.12 ppb, 7.6 ppb, 6.3 ppb and 2.1 ppb,
434 respectively, in the most enriched interval (+ 1.5 cm) compared to a section average of 0.43 ppb
435 Os, 0.073 ppb Ir, 0.035 ppb Ru, 1.2 ppb Pt, 1.2 ppb Pd and 0.15 ppb Re (Fig. 6). Overall, the
436 Passée Jaspeuse has a higher Ir, Ru and Pd content compared to the Halobia limestones, with
437 values only slightly above average. When normalizing the HSE concentrations to Th it is noted

438 that for most elements the enrichments disappear, suggesting a relation between HSE and clay
439 content (Fig. 6).

440

441 *5.6. Osmium-isotope compositions*

442 Age-corrected whole rock $^{187}\text{Os}/^{188}\text{Os}_i$ (201.3 Ma) ratios for the Čanj section show an
443 overarching gradual trend towards lower, more unradiogenic, values up section: from ~ 0.72 to
444 0.26 . $^{187}\text{Os}/^{188}\text{Os}_i$ values decrease from 0.72 to 0.52 across the Halobia limestone strata, with a
445 further decline up section in the Passée Jaspeuse from 0.50 to 0.26 (Fig. 6; Appendix C). Notable
446 exceptions to this general trend include a minor shift to the most radiogenic values of 0.758 and
447 0.735 , at -1.86 and -1.35 m, respectively, and a more pronounced shift from 0.481 to 0.254 within
448 the clay layer, which records the most unradiogenic $^{187}\text{Os}/^{188}\text{Os}_i$ compositions within the section
449 (Fig. 6). This sharp shift correlates with the overall enrichment in HSEs in the clay layer as
450 described in the previous section; however, there is no correlation between the overarching decline
451 in $^{187}\text{Os}/^{188}\text{Os}_i$ values up section and HSE contents (Fig. 6).

452

453 *5.7. Paleomagnetic results*

454 Results indicated unstable remanence directions during alternating field demagnetization for most
455 of the samples. In general, about 50 to 80 % of the natural remanent magnetization is left at
456 demagnetizing fields of 70 mT. Only a few samples could be demagnetized almost completely,
457 which indicate the presence of high-coercivity minerals. A primary magnetization component
458 could not be determined. The intensity of the natural remanent magnetization varies between 0.34
459 and 9.1×10^{-7} Am²/kg, which is rather low.

460 Magnetic property analyses indicate intermediate to high coercive forces ranging from 12
461 to 116 *mT*. Four of the six samples analyzed show a magnetization change peaking at 116 to 119
462 *K*. These values coincide with the so-called Verwey transition observed in 1.5 *mm* large magnetite
463 single crystals (Özdemir and Dunlop, 1999). Thus, magnetic mineralogy of the Čanj section
464 consists of multidomain magnetite and single domain hematite and paramagnetic minerals of
465 variable contributions. Significant exchange bias fields, that would induce a shift of the hysteresis
466 loop at low temperatures, and which may occur in oxidized magnetite core-shell nanoparticles
467 (Phan et al., 2016) possibly during weathering or diagenesis, are not observed.

468 The magnetic property analyses conducted confirm the presence of high-coercivity
469 minerals such as hematite and the absence of nanometer-sized magnetic minerals. The
470 heterogeneity of the magnetic mineralogy and the rather low concentration of potential remanence
471 carries did not favor the acquisition of a stable primary depositional remanent magnetization.
472 A magnetostratigraphic record of the Triassic–Jurassic interval is likely not preserved at Čanj.
473 Therefore, no further effort was undertaken to further investigate the Čanj section for
474 paleomagnetic analyses.

475

476 **6. Discussion**

477 *6.1. Čanj and the global T-J boundary record*

478 We interpret the abrupt negative $\delta^{13}\text{C}_{\text{carb}}$ excursion recorded at the Čanj section, correlative
479 with the clay layer and turnover in radiolarian fauna, as being stratigraphically equivalent to the
480 ICIE that marks the End Triassic extinction (ETE) horizon globally, as also concluded by Črne et
481 al. (2011). This interpretation is supported by the existence of similar abrupt changes from
482 calcareous to argillaceous lithologies correlative with this CIE at a number of other, more shallow

483 marine NW Tethyan sites, for example at Kuhjoch (Eiberg Basin, Austria; Ruhl et al., 2009), and
484 Val Adrara (Lombardy Basin, Italy; e.g., Bachan et al., 2012) (Fig. 7). Whilst some of these shifts
485 have been interpreted as resulting from lithological changes and/or diagenetic alteration, a similar
486 secondary origin for the recorded negative excursion at Čanj is deemed unlikely. Firstly, the base
487 of the $\delta^{13}\text{C}_{\text{carb}}$ shift takes place in the uppermost 20 cm of the Halobia limestone, and is thus not
488 associated with any change in lithology, suggesting that it records a genuine variation in the
489 isotopic composition of seawater. Secondly, a similar negative shift in the isotopic composition of
490 bulk organic matter ($\delta^{13}\text{C}_{\text{org}}$) has been documented at the same level as the $\delta^{13}\text{C}_{\text{carb}}$ shift at Čanj
491 (Črne et al., 2011). Finally, the interpretation of these shifts as documenting a global carbon-cycle
492 perturbation during the T-J faunal turnover is consistent with the worldwide record of this
493 environmental disturbance (see Korte et al., 2018).

494 By contrast, the radiolarian T-J boundary in Japanese Panthalassic records (Katsuyama and
495 Kurusu) has been interpreted as being stratigraphically correlative with the Main CIE (Fujisaki et
496 al., 2018; Du et al., 2020; or Spelae CIE; Lindström et al., 2020). Thus, it could be argued that the
497 $\delta^{13}\text{C}_{\text{carb}}$ shift at Čanj also records this later excursion, or that it comprises a combination of the
498 Initial and Main CIEs due to the condensed nature of the clay layer. However, if the Čanj shift
499 were indeed equivalent to the Main CIE, it would be expected that a sharp excursion in the
500 uppermost Halobia limestones would mark the Initial CIE, which is not the case. Furthermore, the
501 only other Tethyan site featuring radiolarian biostratigraphic information (albeit poorly
502 constrained) at Csővár (Hungary), shows a $\delta^{13}\text{C}_{\text{carb}}$ shift interpreted as the Initial CIE between the
503 highest known Triassic radiolaria and lowest Jurassic fauna (Pálffy et al., 2007). Thus, our
504 interpretation of the Initial CIE at Čanj is consistent with this Tethyan record, as well as previously

505 studied eastern Panthalassic sites at New York Canyon (Nevada, USA), and the Queen Charlotte
506 Islands (British Columbia, Canada) (see Orchard et al., 2007; Williford et al., 2007).

507 Thus, a stratigraphic correlation between the Main CIE and T-J radiolarian turnover in
508 Japan appears to be the exception, rather than the rule. Moreover, the T-J radiolarian turnover in
509 Japan also stratigraphically correlates with the last appearance of conodonts, which is documented
510 around the ICIE level at almost all other sites. Consequently, either radiolarian and conodont
511 turnover in the Tethyan and Panthalassic realms were asynchronous (see Du et al., 2020), or the
512 faunal extinctions were coeval in these two ocean basins, and the isotopic shifts correlative with
513 radiolarian turnover/conodont extinction at Katsuyama and Kurusu are in fact equivalent to the
514 ICIE, as at Kuhjoch and Csővár. Given these uncertainties in global stratigraphic correlations, we
515 favor correlation between Čanj and the more proximal Tethyan sites such as Csővár and Kuhjoch:
516 i.e., that the $\delta^{13}\text{C}_{\text{carb}}$ shift associated with the lithological change and T-J faunal turnover marks the
517 ICIE in the Budva Basin (Fig. 7).

518 Assuming that the negative $\delta^{13}\text{C}$ excursion associated with the Čanj clay layer is indeed
519 the ICIE, it might indicate that the lower magnitude isotopic shifts 2 m below and 1 m above are
520 equivalent to the Precursor and Main CIEs, respectively (Ruhl and Kürschner, 2011). However,
521 the low magnitude of these excursions and lack of further stratigraphic information means that
522 these interpretations remain speculative. As noted above, it is also possible that the Main CIE is
523 also encompassed (together with the Initial excursion) in the $\delta^{13}\text{C}_{\text{carb}}$ shift around the clay layer.
524 Additionally, the two negative excursions documented at the top of the studied Čanj section (Fig.
525 3; Fig. 7), likely of Sinemurian age (see Goričan, 1994; Črne et al., 2011), may be equivalent to
526 the isotopic shifts reported in lower Sinemurian strata elsewhere (e.g., van de Schootbrugge et al.,
527 2005; Bartolini et al., 2012) (Fig. 7). However, proving these hypotheses is hindered by the limited

528 biostratigraphic information and lack of magnetostratigraphic constraints at Čanj. Consequently,
529 the remainder of the manuscript focusses primarily on the ETE interval recorded by the clay layer
530 and the strata immediately above and below, which can be more reliably correlated with the
531 stratigraphy at other sites.

532

533 *6.2. Implications of the CIE at Čanj*

534 At Čanj, the ICIE occurs at the same stratigraphic level as enrichments in HREEs, Hg, and HSEs,
535 as well as a sharp shift to unradiogenic $^{187}\text{Os}/^{188}\text{Os}_i$ values (Fig. 4, 5, 6). Črne et al. (2011)
536 interpreted the CIE at Čanj to reflect a global carbon-cycle disturbance that caused either
537 accelerated carbonate dissolution, resulting in shoaling of the calcium compensation depth, or a
538 biocalcification crisis that reduced carbonate input (e.g., a shift from tropical to microbally
539 mediated mud-mound factory). Either of these processes could have been amplified in this region
540 by increased tectonic subsidence, relative sea-level rise and/or a possible gap at the lithological
541 boundary resulting in a relatively low magnitude of the CIE compared to other sites (Črne et al.,
542 2011). Črne et al. (2011) suggested that either scenario causing the CIE can be explained with
543 increased CO_2 , SO_2 , and CH_4 fluxes due to CAMP volcanism (Črne et al., 2011; see also Greene
544 et al., 2012; Lindström et al., 2021); however, Črne et al. (2011) did not document specific
545 geochemical markers at Čanj that could have directly linked the CIE to CAMP volcanism. The
546 next step is thus to geochemically discern what happened around the ICIE at Čanj.

547

548 *6.3. Origin of the highly siderophile element enrichment*

549 Specific HSE concentrations can be an indication of either an extraterrestrial signature, (see e.g.,
550 Alvarez et al., 1980, Smit and Hertogen, 1980; Goderis et al., 2021), changes in redox states (Wang

551 et al., 1993), and/or volcanic activity (e.g., Whiteside et al., 2021). The enrichment of all reported
552 HSEs in the extinction interval is apparent. However, when normalizing the HSE content to Th,
553 an element associated with clay dilution, the HSE enrichments mostly disappear, the small peaks
554 of Os/Th, Ir/Th, Pt/Th and Pd/Th that remain present are solely because of the scale used (Fig. 6).
555 These observations suggest the HSE enrichments are closely related to the clay fraction in a given
556 stratum and thus likely reflect terrestrial weathering as the source for HSE enrichment at Čanj.
557 Figure 8 compares the Ir and Pt trends of Čanj and elsewhere in the NW Tethys (Kuhjoch) to that
558 of sections in North America (Partridge Island) and the Panthalassic Ocean (Kurusu). What stands
559 out is that the Ir and Pt enrichments are distinctly inconsistent in both pattern and level of
560 concentrations. These inconsistencies are exemplified by the existence of multiple Ir peaks in the
561 Partridge Island section, which contrasts with other sections either featuring general enrichment
562 (Kuhjoch) or very small enrichments (Kurusu) (Fig. 8). The palynologically determined T-J
563 interval at Partridge Island correlates with the second Ir peak, and, while this is the largest Ir peak,
564 it is not at the same level as the ICIE in that locality (Fig. 8). For Kuhjoch, the comparison with
565 Čanj is more straightforward as the initial enrichment in Ir is synchronous with the ICIE (Tanner
566 et al., 2016), while for Kurusu only a small (~70 ppt) Ir anomaly exists (Hori et al., 2007). Čanj
567 appears to contain the largest Ir enrichment recorded, with only the Partridge Island section
568 matching the Ir enrichment (of up to 450 ppt), whilst lacking a similar peak in Pt content. The
569 exact correlation between these peaks is difficult, as the use of HSEs for correlation remains
570 tentative at best, with so far, no reliable correlation between specific HSE anomalies (see
571 Whiteside et al., 2021 for further details). However, these data do show that Ir and HSE enrichment
572 occur globally for the ETE. With the Čanj section likely indicating a terrigenous derived HSE
573 enrichment, the question that remains is whether this is the case globally. For example, a global

574 HSE enrichment is observed for the K-Pg boundary, although unlike the ETE the HSE enrichment
575 at the K-Pg is present globally (e.g., Goderis et al., 2013).

576 Specific HSE ratios can be employed to distinguish between terrigenous or extra-terrestrial
577 sources for these elements. In Figure 9 the Pd and Pt ratio over Ir is presented: for non-fractionated
578 chondrites these ratios should be lower than for fractionated values in mantle material (see e.g.,
579 Tegner et al., 2020), whilst the composition of iron meteorites encompasses a large range of
580 compositions (Fig. 9). The observed element ratios of the Čanj section highlight the fractionated
581 nature of the PGEs and, for the majority, the data overlaps with known compositions of CAMP
582 volcanism. This is exemplified when comparing the CI-chondrite normalized HSE content of the
583 Čanj section, where one can observe a strongly comparable normalized signature of all units when
584 compared to CAMP, with notably the clay interval overlapping with CAMP compositions (Fig.
585 9C). Furthermore, when comparing the Pd/Ir and Pt/Ir ratios of Čanj to that of other T-J sections
586 (Fig. 9B) it becomes apparent that ratios are comparable between sections, regardless of
587 geographic location and absolute enrichment. In contrast, the Pd/Ir and Pt/Ir ratios of K-Pg
588 boundaries overlap in composition with chondritic values, and do not plot close to compositions
589 observed for coeval Deccan volcanism, suggesting that a chondritic origin of HSE enrichment
590 should be observable over a terrigenous source in these elemental ratios, if a large impact occurred
591 coevally with a LIP. This strongly argues that HSE enrichment at Čanj, and globally, are derived
592 from the CAMP. Whilst the Pd/Ir and Pt/Ir ratios of iron meteorites encompass a large range in
593 composition, overlapping in composition with all reported data sets (Fig. 9B), iron meteorites tend
594 to have subchondritic HSE values, which greatly differs from the compositions observed at Čanj
595 (Fig. 9C). Furthermore, Tegner et al. (2020) noted that iron meteorites are also known to be
596 relatively enriched in rhodium, resulting in low Pt/Rh values (Ryan et al., 1990; Hoashi et al.,

597 1993; Pateav and Jakobsen, 2004) and that samples from the T-J boundary present Pt/Rh values
598 that are higher than known iron meteorites, further arguing against the HSEs at Čanj being derived
599 from such a source.

600 Lastly, there appears to be a relation between the paleogeographic location of the T-J
601 sections and HSE enrichment, with the sections that were closer to CAMP (i.e., Partridge Island
602 and NW Tethys sections; Fig 7, 8) showing more absolute Ir enrichment than that of Panthalassic
603 sections. Consequently, this might indicate that the magnitude of HSE deposition is related to
604 CAMP proximity. Whether this indicates the means of deposition to be HSE-rich aerosols as a
605 consequence of volcanic outgassing (Tanner et al., 2008; Tegner et al., 2020; Whiteside et al.,
606 2021), proximal ash fall, and/or CAMP rock weathering (e.g., Tanner et al., 2008), remains to be
607 determined, though the earlier observation that HSE enrichment is related to the clay fraction,
608 implicates weathering of juvenile CAMP basalts as the main cause behind HSEs entering the
609 system. Nonetheless, this apparent correlation of Čanj with both other T-J sections and CAMP,
610 and the incongruity with both chondritic and iron meteorites (in contrast to the distinct correlation
611 of K-Pg sections with chondritic values rather than Deccan Trap basalts) clearly favors a CAMP
612 origin of the HSE compositions observed at Čanj, and thus that the HSE influx at the T-J boundary
613 is terrestrially, rather than extra-terrestrially, derived.

614

615 *6.4. A further examination of the terrestrial signal*

616 With the HSE signal strongly implying a terrigenous origin of the element enrichment at the Čanj
617 section, it becomes imperative to determine whether this is consistent with other geochemical
618 markers. Despite being lithologically diverse, both the Halobia limestone and Passée Jaspeuse Fm.
619 show trace element patterns typical of sedimentary lithologies, with LREE being relative enriched

620 when compared to the HREEs, and an overall flat lying MREE to HREE pattern (Fig. 4). Such a
621 trace element pattern is the common result of the upper continental crustal input of REEs
622 dominating the signal (Piper, 1974; McLennan, 2001). The observed difference between the two
623 formations, with the Passée Jaspeuse being overall more enriched in trace element content, is likely
624 caused by dilution by carbonate in the Halobia limestones compared to the more argillaceous
625 Passée Jaspeuse. Notably, however, the Halobia limestones show distinct Ce depletion in most
626 samples relative to neighboring trace elements (Fig. 4). Such a shift in Ce content may suggest that
627 the Halobia limestones were deposited in well-oxygenated seawater, as cerium can undergo
628 oxidation in seawater from soluble Ce(III) to the highly insoluble quadrivalent state (Piper, 1974;
629 Bellanca et al., 1997). Its subsequent fixation in particulate matter, including organics, is thought
630 to be responsible for distinctive depletion of Ce (Bellanca et al., 1997). In contrast, under reducing
631 conditions, Ce remains in its soluble Ce(III) state; therefore, marine sediments precipitated in
632 reducing environments will not show a relative Ce depletion (Piper, 1974; Bellanca et al., 1997).
633 Following this assertion, as both the boundary clay layer and the Passée Jaspeuse do not indicate
634 a Ce depletion, their Ce content might suggest that they were deposited under more oxygen-
635 depleted conditions.

636 The main clay layer, when compared to the other lithologies, shows a REE pattern
637 markedly different to that expected for a typical sedimentary rock (Fig. 4; e.g., McLennan et al.,
638 2001), presenting a distinct V-shaped pattern, with relative LREE and HREE enrichment
639 compared to the middle REEs. HREE enrichment is not typically associated with crustal material
640 (e.g. Rudnick and Fountain, 1995, Rudnick and Gao, 2003), and therefore implies the involvement
641 of lithologies with a relative HREE enrichment, such as mantle derived material that exhibits
642 higher Lu/La ratios linked to the decreasing incompatibility from La to Lu (e.g. Salter and Stracke,

643 2004). Interestingly, a similar V-shaped trace element pattern has been observed in a limestone
644 layer just below the ICIE level in the Kendlbachgraben section (Austria) (Pálffy and Zajzon, 2012).
645 At that site, the paired LREE and HREE enrichment was interpreted to reflect a sedimentary
646 component for the former, and a magmatic component, likely derived from CAMP, for the latter
647 (Pálffy and Zajzon, 2012). To further investigate this assertion, we have plotted different CAMP
648 section compositions in Figure 4. CAMP volcanics from N America and SW Europe,
649 paleogeographically the most proximal part of the LIP to Čanj, have trace element patterns
650 documenting relative HREE enrichment (Marzoli et al., 2011; Callegaro et al., 2014), which is
651 consistent with the clay layer signature and that of the Kendlbachgraben section. CAMP volcanics
652 from NW Africa, which were paleogeographically further from Čanj, exhibit much more variation
653 in their trace element patterns and are overall less enriched in HREE content (Callegaro et al.,
654 2017). This observation is corroborated by the HSE normalized trace element patterns of the Čanj
655 clay interval being largely consistent with Moroccan CAMP lavas (Fig. 9C). In contrast, marine
656 K-Pg sites show trace element patterns that lack a relative HREE enrichment (see e.g., Ebihara
657 and Miura, 1996; Shrivastava et al., 2013; Lorocho et al., 2016; Sial et al., 2018). This shows a
658 strong contrast between the trace element enrichments of the extinction horizon at Čanj to that of
659 K-Pg sections. As such, these observations support a genetic link of HREE enrichment at the ETE
660 interval in relation to CAMP proximity and mark geochemical changes related to sedimentary
661 influx and not just a change in facies.

662 To further emphasize this point, we note that the Hg enrichment observed at Čanj is
663 stratigraphically correlative with the ICIE, consistent with peaks observed in both terrestrial and
664 marine records of the ETE globally (Thibodeau et al., 2016; Percival et al., 2017; Lindström et al.,
665 2019; Kovács et al., 2020). Mercury is typically deposited in sediments bound to organic material

666 but can also be associated with sulfides or clays (Shen et al., 2020). Given the paucity of organic
667 matter (typically <0.1 wt%) and sulfides (<0.1 wt%) at Čanj, it is likely that Hg is primarily bound
668 to clays in this record. However, some argillaceous intervals do feature slightly higher organic
669 contents (although still very low: typically, <0.2 wt%). In this context, the record of the Hg peak
670 in the main clay layer might suggest that it results from the lithological change to more argillaceous
671 and marginally more organic-rich sediments. However, only two of the four samples in the clay
672 layer record the distinct Hg peak, despite all four featuring a similar Al₂O₃ content, and nor are
673 those two Hg-rich levels marked by notably elevated TOC (indeed, the highest TOC sample has a
674 relatively low mercury concentration; Appendix B). Moreover, there are no comparable Hg
675 enrichments in any other argillaceous levels studied here, suggesting an exceptional input of
676 mercury to the Budva Basin during the extinction event. Given that this mercury peak is
677 stratigraphically correlative with those from other sites, which have been linked to CAMP activity,
678 and that volcanism is a major natural source of mercury to the surface environment (Grasby et al.,
679 2019; Percival et al., 2021), it also suggests that the Čanj input was part of the same LIP-related
680 global Hg-cycle perturbation documented as occurring during the extinction. Interestingly, several
681 other T-J records show further mercury enrichment above the ICIE (e.g., Percival et al., 2017;
682 Kovács et al., 2020; Yager et al., 2021), potentially reflecting continued CAMP activity. Whilst
683 Hg contents in Passée Jaspeuse sediments are slightly higher than in the underlying Halobia
684 limestones, there is no clear peak above the extinction interval at Čanj. In contrast, two low
685 magnitude peaks (albeit each consisting of only one data point) below the main clay layer may
686 support previous hypotheses of pre-extinction perturbations to the global Hg cycle (Lindström et
687 al., 2019).

688

689 6.5. Osmium isotope record of the Čanj section

690 The geochemical evidence presented above supports the involvement of CAMP over an
691 other (extra-)terrestrial influx. To further test this hypothesis, the $^{187}\text{Os}/^{188}\text{Os}$ values of the Čanj
692 section were examined. Due to the relatively short residence time of Os in seawater (~ 10 ky;
693 Peucker-Ehrenbrink and Ravizza, 2000), trends in sedimentary $^{187}\text{Os}/^{188}\text{Os}_i$ record rapid changes
694 in osmium fluxes to the global ocean where sudden shifts to unradiogenic values might indicate
695 juvenile flood basalt weathering, submarine volcanism, and/or impact events ($^{187}\text{Os}/^{188}\text{Os} = \sim 0.13$;
696 Allégre et al., 1999) and shifts to radiogenic values might indicate enhanced terrestrial weathering
697 of continental crust (average modern-day riverine runoff $^{187}\text{Os}/^{188}\text{Os} = \sim 1.4$; Peucker-Ehrenbrink
698 and Jahn, 2001; see also Sato et al., 2013; Goderis et al., 2021).

699 The $^{187}\text{Os}/^{188}\text{Os}_i$ values at Čanj display an overall gradual decrease upsection, suggesting a
700 rising input of unradiogenic osmium to the global ocean during the T-J interval, with a transient
701 shift to an even more unradiogenic composition at the extinction interval (Fig. 6). The majority of
702 this shift appears to initiate just below the clay layer, with the Halobia limestone featuring
703 relatively stable $^{187}\text{Os}/^{188}\text{Os}_i$ values between 0.76 and 0.61 up to -0.33 m, above which it shifts to
704 more unradiogenic values with a distinct sharp shift to the most unradiogenic composition in the
705 main clay layer (Fig. 6). Notably, above the transient shift in $^{187}\text{Os}/^{188}\text{Os}_i$ values associated with
706 the clay layer, the gradual decrease resumes, with no return towards more radiogenic compositions
707 observed in the remainder of our sampling interval. This trend highlights that a gradual increase
708 in the flux of more unradiogenic osmium to seawater began prior to the ICIE, but with a sharp
709 transient increase in this influx during the extinction interval. This pattern is largely consistent
710 with that of St. Audrie's Bay, which records a very similar overall decreasing $^{187}\text{Os}/^{188}\text{Os}_i$ trend
711 (Cohen and Coe, 2002; Fig. 10). There, the $^{187}\text{Os}/^{188}\text{Os}_i$ trend was interpreted to reflect the

712 continued influx of unradiogenic Os from CAMP volcanism (Cohen and Coe, 2002). However,
713 Cohen and Coe (2002) did not sample the stratigraphic interval marked by the ICIE (as determined
714 by Hesselbo et al., 2002), precluding direct comparison of the seawater $^{187}\text{Os}/^{188}\text{Os}_i$ records at the
715 extinction horizon between St. Audrie's Bay and Čanj. Nonetheless, the data suggest comparable
716 behavior of the $^{187}\text{Os}/^{188}\text{Os}_i$ values in sections of the NW Tethys and Northern Europe (Fig. 10).

717 Interestingly, whilst the Panthalassic Kurusu section records a similar decline in
718 $^{187}\text{Os}/^{188}\text{Os}_i$ values, this shift is shown as taking place during the Late Norian to middle Rhaetian
719 interval, with a sharp return to more radiogenic compositions during the latest Rhaetian, and no
720 significant unradiogenic shift at the radiolarian defined Rhaetian – Hettangian boundary (Kuroda
721 et al., 2010; Fig. 10). Although this pattern is similarly interpreted as the initial gradual input of
722 unradiogenic Os, likely through weathering of CAMP basalts, and followed by the increased
723 radiogenic Os input from enhanced continental weathering outweighing the unradiogenic Os input
724 during the late Rhaetian (Kuroda et al., 2010), the timing contrasts with data from Čanj and St.
725 Audrie's Bay. It should be noted that there is currently no evidence for CAMP magmatism as early
726 in the Rhaetian as implied by the Kurusu section, although the geological record of the LIP is
727 limited by relatively poor preservation of the igneous units. If all three sections do indeed record
728 CAMP basalt weathering, the differences between Tethyan and Panthalassic sections might indicate
729 an apparent difference in Os isotopic behavior between the Čanj and St. Audrie's Bay on the one
730 hand, and the Kurusu section on the other hand, potentially highlighting a different response to the
731 mechanism behind the extinction in the Panthalassic Ocean when compared to the Tethys Ocean
732 and Northwestern Europe. However, this conclusion would demand that the global ocean was
733 heterogeneous with respect to osmium during the Triassic-Jurassic interval, in contrast to today.
734 This assessment highlights the complexity of the events surrounding the T-J boundary, with the

735 general trends appearing to be comparable globally, with HSE, REE, Hg and unradiogenic Os
736 enrichment, however, with a magnitude of enrichment that is neither uniform nor consistent (Fig.
737 8 and 10).

738 The spike of more unradiogenic Os composition in the Čanj clay layer correlates with the
739 extinction horizon and ICIE, which was coeval with the onset of CAMP eruptions (Deenen et al.,
740 2010; Ruhl et al., 2011; Blackburn et al., 2013). As such this unradiogenic shift likely marks the
741 main onset of CAMP. Whilst this shift was not previously detected at St. Audrie's Bay or Kurusu,
742 the relatively low-resolution dataset at the former site and uncertain stratigraphic position of the
743 ICIE at the latter means that this sharp decline in $^{187}\text{Os}/^{188}\text{Os}_i$ values could nonetheless be present
744 at both. Further studies of sites where the ICIE and extinction horizon are stratigraphically well
745 constrained are needed in order to confirm whether the unradiogenic Os isotope shift recorded in
746 the Čanj clay layer is preserved globally. As it might represent a convincing stratigraphic
747 correlation to the onset of CAMP.

748 This caveat notwithstanding, the overall Os curve at Čanj is consistent with the timing of
749 known CAMP magmatism, with the initial onset predating the ETE by at least 100 kyr (Deenen et
750 al., 2010; Davies et al., 2017; Heimdal et al., 2020), expressed as a small gradual shift to
751 unradiogenic values, while the first known major eruptions coincided with the ETE (Deenen et al.,
752 2010; Ruhl et al., 2011; Blackburn et al., 2013). The continuation of CAMP basalt outpourings for
753 at least 600 kyr after the extinction event (Marzoli et al., 1999; Blackburn et al., 2013; Marzoli et
754 al., 2018) is also consistent with the gradual unradiogenic Os shift observed. Hypothetically, the
755 gradual decline in $^{187}\text{Os}/^{188}\text{Os}_i$ values could have resulted from a fall in continental weathering
756 rates and associated runoff of radiogenic Os. However, such a decrease in continental weathering
757 is unlikely at a time of elevated atmospheric CO_2 and associated climate warming (McElwain et

758 al., 1999; Steinhorsdottir et al., 2011) leaving increased input of mantle osmium from weathering
759 of CAMP basalts as the most likely cause. Consequently, we interpret that the sharp shift in
760 $^{187}\text{Os}/^{188}\text{Os}_i$ values at the ETE horizon, as well as the more prolonged shift to a more unradiogenic
761 osmium-isotope composition of seawater observed in the global T-J record, resulted from CAMP
762 weathering. These observations agree with the HSE, REE and Hg data presented in this study and
763 not only argues for the CAMP as the most likely culprit behind the extinction event, but also
764 establishes the Čanj section as one of the best continuous deep marine T-J sections on the planet.

765

766 **7. CONCLUSION**

767 At Čanj, Montenegro, a well preserved and continuous T-J section was studied where deep-marine
768 Triassic carbonates of the Halobia limestone Formation are abruptly interrupted by a ~6 cm finely
769 laminated clay layer, before transitioning to the argillaceous Jurassic red beds of the Passée
770 Jaspeuse. Within this clay layer a negative carbon isotope excursion, relative HREE enrichment,
771 Hg increase, HSE enrichment and sharp shift to unradiogenic Os ratios are recorded. It is
772 interpreted that this sudden change in lithology marks the end-Triassic extinction event. The
773 observed geochemical variations along the Čanj section are consistent with the Central Atlantic
774 Magmatic Province as the main driver behind the end-Triassic extinction. This marks the Čanj
775 section as a rare T-J section that exquisitely encapsulates the geochemical record of the end
776 Triassic extinction interval in the Tethyan deep sea, especially when compared to the overall
777 inconsistent global record.

778

779 **Declaration of competing interest**

780 The authors declare that they have no known competing financial interests or personal
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782

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799

800 **Data Availability**

801 All data used are listed in the references and appendices.

802

803 **Supplementary data**

804 Supplementary data to this article can be found online at: XXXXXXXXX

805

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1226

1227 **Figure Captions**

1228 Figure 1 (A) The Čanj section, Montenegro. (B) Zoomed in view of the Čanj clay layers.

1229

1230 Figure 2. Geological map of the area surrounding the Čanj section. Based on the Institute for
1231 Geological and Geophysical Research 1 : 100000 Geological Map of former Yugoslavia (Osnovna
1232 Geološka Karta SFRJ), Budva, K34-62, 1969 and Bar, K34-64, 1976 from the Federal Geological
1233 Institute, Belgrade. ALB = Albania, BIH = Bosnia and Herzegovina, GRC = Greece, MKD =
1234 Republic of North Macedonia, MNE = Montenegro, SRB = Serbia, XKX = Kosovo.

1235

1236 Figure 3. Lithological log (this study), radiolarian biostratigraphy showing the last occurrence of
1237 Triassic radiolarian and first occurrence of Jurassic radiolarian (Črne et al., 2011), and $\delta^{13}\text{C}_{\text{carb}}$
1238 trends (Črne et al., 2011; this study) for the Čanj section (Budva Basin, Montenegro). Pale red
1239 shading indicates the stratigraphic extent of the negative $\delta^{13}\text{C}_{\text{carb}}$ excursion inferred to be
1240 equivalent to the Initial CIE. Bold dark blue line indicates the moving five-point average for the
1241 new $\delta^{13}\text{C}_{\text{carb}}$ dataset. Question mark indicates that the exact position of the Hettangian –
1242 Sinemurian boundary is unknown.

1243
1244 Figure 4. CI-chondrite-normalized trace-element concentrations, with normalization values from
1245 Sun and McDonough (1989). CAMP compositions for NW Europe and NW Africa from Callegaro
1246 et al. (2014, 2017) respectively, N America data is from Marzoli et al. (2011). CAMP = Central
1247 Atlantic Magmatic Province.

1248
1249 Figure 5. $\delta^{13}\text{C}_{\text{carb}}$, Hg, Al_2O_3 , and $\text{Hg}/\text{Al}_2\text{O}_3$ trends from Čanj. Pale brown shading indicates clay-
1250 rich layers analyzed for Hg concentrations.

1251
1252 Figure 6. Selected highly siderophile element variations for the Čanj section. Shown in black is
1253 absolute element concentrations. In red is the selected element normalized to $\text{Th} \cdot 10^{-3}$ as
1254 normalizing to Th in pbb would effectively remove all peaks. $^{187}\text{Os}/^{188}\text{Os}_i$ is age corrected to 201.3
1255 Ma.

1256
1257 Figure 7. Stratigraphic correlation of Čanj with other T-J records from the NW Tethys and
1258 Panthalassia, based on ammonite, conodont, and radiolarian biostratigraphy, and $\delta^{13}\text{C}$

1259 chemostratigraphy. The palaeogeographic location of each site is indicated; the map is adapted
1260 from Greene et al. (2012). Čanj biostratigraphy is from Črne et al. (2011); log and $\delta^{13}\text{C}$ data are
1261 from this study. All Kuhjoch data are from Ruhl et al. (2009). All Val Adrara data are from Bachan
1262 et al. (2012). All Csóvár data are from Pálffy et al. (2007). Note that constraints on the stratigraphic
1263 extent of radiolarian biozones (and, therefore, the radiolarian T-J turnover) are limited at Csóvár.
1264 Kennecott Point biostratigraphy is from Ward et al. (2001, 2004); all other information from
1265 Williford et al. (2007). Katsuyama biostratigraphy from Carter and Hori (2005); all other
1266 information from Fujisaki et al. (2018). Kurusu biostratigraphy is from Hori et al. (2007);
1267 lithological information from Du et al. (2020); $\delta^{13}\text{C}$ data are from Kuroda et al. (2010).

1268

1269 Figure 8. Comparison of highly siderophile element distribution at different localities. Partridge
1270 Island data from Percival et al. (2017), Tanner and Kyte (2005), and Tanner et al. (2008). Kuhjoch
1271 data from Ruhl et al. (2009) and Tanner et al. (2016). Čanj data from this study. Kurusu data from
1272 Hori et al. (2007). Katsuyama data from Fujisaki et al. (2018).

1273

1274 Figure 9. Pd/Ir vs Pt/Ir diagram adapted from Tegner et al. (2020) for the Čanj section. (A) Čanj
1275 section subdivided per formation, including the clay layer. Shown for comparison are CI (Fischer-
1276 Gódde et al., 2010) and ordinary chondrites (McDonald et al., 2001), CAMP volcanism from
1277 Morocco and Iron meteorites (Tegner et al., 2020). (B) Čanj section (shown as dark-grey field)
1278 compared to sediments from sections containing the end-Triassic extinction event in the Fundy
1279 Basin, Canada (Tanner and Kyte, 2005), Kurusu, Japan (Hori et al., 2007), Sidi Rahal, Morocco
1280 (Whiteside et al., 2021) and Exeter, USA (Whiteside et al., 2021). Fields for the K-Pg impact layer,
1281 Deccan Trap volcanism and Iron meteorites are from Tegner et al. (2020) and references therein.

1282 (C) CI chondrite normalized HSE pattern for Čanj and selected other compositions. CI chondrite
1283 normalization value from Palme et al. (2014).

1284

1285 Figure 10. $^{187}\text{Os}/^{188}\text{Os}_i$ site comparison. St Audrie's Bay data from Cohen and Coe (2002) and
1286 Hesselbo et al. (2002), Kuhjoch data from Ruhl et al. (2009) and Tanner et al. (2016). Čanj data
1287 from this study. Kurusu data from Kuroda et al. (2010). Katsuyama data from Fujisaki et al. (2018).
1288 Error bars (2se) are smaller than line thickness (see Appendix C)

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