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Citation for final published version:

Zhu, Hengrui, Wu, Tao, Li, Chun-Feng, Gleeson, Matthew, Wieser, Penny, Lu, Jianggu, Yang, Ming, Scholpp, Jesse L., Heaton, Daniel, Nelson, Wendy, Shervais, John W., Potter, Katherine E., Buchs, David M., Class, Cornelia, Homrighausen, Stephan, Wang, Xiao-Jun, Kubota, Yusuke, Höfig, Tobias W., Sager, William and Hoernle, Kaj 2026. Unveiling the hotspot-driven magmatic plumbing system of the Center track, Walvis Ridge (South Atlantic Ocean): Insights from mineral geochemistry and isotopic heterogeneity in basalts from IODP Expedition 391. *GSA Bulletin* 10.1130/b38890.1

Publishers page: <https://doi.org/10.1130/b38890.1>

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1 **Unveiling the hotspot-driven magmatic plumbing system of the Center track,**
2 **Walvis Ridge (South Atlantic Ocean): Insights from mineral geochemistry and**
3 **isotopic heterogeneity in basalts from IODP Expedition 391**

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27

28 **Abstract**

29 The Tristan-Gough-Walvis Ridge (TGW) volcanic chain on the African plate is
30 one of Earth's longest linear features on the seafloor, however, the origin of its
31 magmatism remains poorly understood. To better understand its formation, temporal

32 evolution, and geochemical characteristics, six sites were drilled along the TGW chain
33 by International Ocean Discovery Program (IODP) expeditions 391 and 397T: three
34 sites on Valdivia Bank (Site U1575, U1576, and U1577) and three sites in the Guyot
35 Province (Site U1578, Site1584, and U1585). Among these, Site U1578, located on the
36 deep northwestern flank of a “Center track” seamount, represents the deepest basement
37 penetration (>300 m) of both expeditions. According to preliminary geophysical and
38 biostratigraphic constraints, the recovered stratigraphic sections from Site U1578A
39 span >2.0 Myrs during the Paleocene (ca. 62.3-64.8 Ma). Here, we present an integrated
40 study of the mineral major and trace elements, as well as in-situ Sr isotopes of
41 plagioclase (Pl) from Site U1578 basalts. Based on the olivine-spinel (Ol-Sp)
42 thermometry and pyMelt modeling, the calculated mantle potential temperature of the
43 Center track is at least 1400 °C, about 50 °C higher than that of mid-ocean ridge basalts
44 (MORBs), most likely reflecting the influence of a mantle plume. Based on
45 geochemical compositions and crystal textures, we identified three distinct
46 clinopyroxene (Cpx) groups: Group 1 are characterized by high Mg# values (> 83.4),
47 Ni (> 200 ppm) and Cr (> 5000 ppm) contents; Group 2 have intermediate Mg# values
48 (76.6-83.4), Ni (100-200 ppm) and Cr (1000-5000 ppm) contents; Group 3 are
49 characterized by low Mg# values (< 76.6), Ni (< 70 ppm) and Cr (< 30 ppm) contents.
50 Multiple episodes of magma recharge and magma mixing occurred in the plumbing
51 system beneath the Center track, which formed the diversity of mineral textures (i.e.,
52 normal zoned, reverse zoned, sector zoned, and oscillatory zoned). Additionally, Pl
53 from this site are isotopically heterogeneous (0.7038 ± 0.0001 to 0.7046 ± 0.0001) and

54 broadly similar to those of basalts from the northern Tristan track in the Guyot province.
55 A shift to more radiogenic Sr isotopic signatures from the lower to the upper igneous
56 units suggests the increasing incorporation of radiogenic materials into the mantle
57 source, potentially derived from enriched mantle typically associated with basalts on
58 the southern Gough track. These findings indicate that basaltic lavas from the Center
59 track originate from an anomalously hot and geochemically enriched mantle source
60 associated with a mantle plume or hotspot activity. Diverse mineral textures and
61 chemistry record complex fractional crystallization and mixing processes during
62 magma storage prior to eruption.

63

64 Keywords: Walvis Ridge; In-situ Sr isotope; Thermobarometer; Crystal zoning;
65 Magmatic plumbing system.

66

67 **1. Introduction**

68 Most volcanoes on Earth are concentrated at tectonic plate boundaries (i.e.,
69 subduction zones and mid-ocean ridges), and their formation can be effectively
70 explained by plate tectonic processes (Palin and Santosh, 2021). However, “intraplate”
71 volcanism can also occur thousands of kilometers away from tectonic plate boundaries
72 (Turcotte et al., 1997). These regions are usually interpreted as the result of mantle
73 melting processes induced by the rising of hot plumes (Campbell and Griffiths, 1990;
74 Campbell and Griffiths, 1992; Duncan and Richards, 1991; Morgan, 1971) or by
75 lithospheric extension (Anderson and Natland, 2014; Hirano et al., 2019). Oceanic

76 intraplate volcanoes usually have higher crystallization temperatures relative to other
77 tectonic settings (Carter et al., 2024; Jennings et al., 2019; Xu and Liu, 2016; Zhang et
78 al., 2023), and their magma storage depths can vary from very shallow levels (1-5 km)
79 at high-flux volcanoes like Hawaii (DeVitre et al., 2024; Lynn et al., 2024; Wieser et
80 al., 2021), to Moho or uppermost mantle depths in lower-flux systems (Barker et al.,
81 2015; Baxter et al., 2023; Gazel et al., 2025; Gleeson et al., 2025; Gleeson et al., 2020;
82 van Gerve et al., 2024). Additionally, compared to MORBs, intraplate volcanic
83 eruptions are commonly associated with shorter crustal residence times (Bonechi et al.,
84 2024; Petrone et al., 2018) and more rapid magma ascent rates (Klügel et al., 2020; Ma
85 et al., 2024; Peslier et al., 2015), potentially related to their higher volatile contents (Liu
86 et al., 2017). Thus, investigating the magmatic plumbing system of intraplate volcanoes
87 is essential for understanding their ascent and storage processes prior to eruption,
88 ultimately shedding light on the origins of magmatism.

89 Mineral phases within volcanic rocks are important recorders of the complex
90 journey of magma transfer through and storage within Earth's mantle and crust
91 (Cashman et al., 2017; Cooper, 2017; Ganne et al., 2018; Neave et al., 2021; Ubide et
92 al., 2021). For example, detailed evaluation of mineral chemistry, chemical zoning and
93 textural features can identify distinct regions of magma storage at depth, mixing
94 between these regions, the processes occurring during magma ascent, and the
95 timescales associated with reservoir destabilization and eruption triggering (Costa et al.,
96 2020; Ginibre et al., 2007; Mutch et al., 2019; Petrone et al., 2022; Streck, 2008). In
97 particular, Pl and Cpx are common in intraplate basaltic magmas, and show chemical

98 variations correlating with the wide range of pressure-temperature-composition (P-T-
99 X) conditions under which they grow (Bennett et al., 2019; Li et al., 2020; Palummo et
100 al., 2021; Wieser et al., 2024; Zhu et al., 2024). Thus, they are frequently used to
101 decipher complex magma crystallization processes (Chen et al., 2023; Chen et al., 2024;
102 Gleeson et al., 2020; Molendijk et al., 2022; Scholpp et al., 2022; van Gerve et al., 2020;
103 Wei et al., 2022).

104 Tristan-Gough-Walvis Ridge volcanic chain on the African plate is one of Earth's
105 longest linear features on the seafloor (Sager and Hoernle, 2022), which forms a broad
106 and scattered seamount province leading from the Walvis Ridge to the active volcanic
107 islands of Tristan da Cunha and Gough (Fig.1). It was one of the first seamount tracks
108 proposed to have been formed by a hotspot/mantle plume (Wilson, 1963), a narrow,
109 buoyant, columnar upwelling of higher temperature mantle material driving melting
110 (Morgan, 1971). However, recent research indicates that the hotspot model alone cannot
111 fully explain the complex morphology of Walvis Ridge (Sager et al., 2021).
112 Magnetization models reveal quasi-linear polarity zones that cross the Valdivia Bank
113 and follow predicted paleo-Mid-Atlantic Ridge locations (Thoram et al., 2023), and
114 gravity-derived crustal thickness patterns of Guyot Province are consistent with
115 magmatism emplaced in a spreading-ridge environment (Huang and Li, 2024). These
116 observations provide evidence that, in addition to intraplate hotspot volcanism,
117 magmatic activity linked to mid-ocean ridge processes also contributed to the
118 construction of Walvis Ridge at the same time, implying potential temporal and spatial
119 interactions between plume activity and seafloor spreading. Tristan-Gough-Walvis

120 Ridge volcanic chain is one of the most studied underwater volcanic chains in the world,
121 and previous studies have mainly focused on whole-rock major-, trace-element
122 geochemistry and Sr-Nd-Pb-Hf isotopic compositions (Class et al., 2015; Hoernle et al.,
123 2015; Homrighausen et al., 2018; Homrighausen et al., 2020; Rohde et al., 2013a;
124 Salters and Sachi-Kocher, 2010), geochronology (Homrighausen et al., 2019; O'Connor
125 and Jokat, 2015a; O'Connor and Jokat, 2015b; Rohde et al., 2013b), and geophysical
126 investigations (Contreras et al., 2022; Geissler et al., 2017; Sager et al., 2021; Thoram
127 et al., 2023; Thoram et al., 2019). Although some studies have explored the magmatic
128 plumbing system of Walvis Ridge (Le Roex, 1985; Weit et al., 2017), their sample
129 collections have been limited to Gough and Tristan da Cunha islands. The Center track,
130 which is located between the Tristan and Gough tracks (Fig.1), is essential for
131 understanding the origin of magmatism in the TGW volcanic chain. Previously, both a
132 triple-zoned plume model and a mixed center zone model have been proposed (Sager
133 and Hoernle, 2022). The mixed center zone model can be further subdivided into two
134 different scenarios: Rohde et al. (2013a, b) proposed that the Gough and Tristan hotspot
135 tracks resulted from a bifurcated plume, whereas Homrighausen et al. (2023) proposed
136 that the two tracks originated from distinct plumelets rising independently from the
137 deep mantle. In this study, we focus on examining these models through the
138 investigation of the Center track.

139 We examine the in-situ major and trace element compositions of Ol, Sp, Cpx, Pl,
140 and in-situ Sr isotopic compositions of Pl from basalts collected at IODP 391 Site
141 U1578, located in the Center track. We aim to elucidate the complex magmatic

142 plumbing system and to determine the origin of magmatism in the Center track of
143 Walvis Ridge.

144

145 **2. Geological setting and samples**

146 According to the mantle plume model (Morgan, 1971; Richards et al., 1989;
147 Wilson, 1973), Walvis Ridge hotspot volcanism started with the emplacement of the
148 Parana-Etendeka continental flood basalt province at ~132 Ma (Renne et al., 1996).
149 From northeast to southwest, TGW volcanic chain contains a ridge (Frio Ridge), an
150 oceanic plateau (Valdivia Bank), and chains of guyots, and scattered seamounts called
151 the Guyot Province (Sager and Hoernle, 2022). The post-70 Ma Guyot Province is
152 divided into three distinct seamount tracks: the Gough track, the Center track, and the
153 Tristan track (Fig.1). Radiogenic isotope ratios reveal spatial zonation in the chemical
154 composition of these post-70 Ma lavas (Class et al., 2015; Hoernle et al., 2015).
155 Previous studies, based on free-air gravity data and the age-distance relationship of
156 volcanism, suggested that Tristan track was influenced by hotspot interaction with the
157 Mid-Atlantic Ridge (O'Connor and Jokat, 2015a), whereas the Center track and Gough
158 track were generated by hotspots beneath the African plate. However, the mechanism
159 and details of plume-ridge interaction and plume-lithosphere interaction processes are
160 not well understood.

161 To better understand the formation and evolution of Walvis Ridge, IODP
162 Expedition 391 drilled four sites (Sites U1575 to U1578) during December 2021 to
163 February 2022. Site U1578, the focus of this study, is located on the deep northwestern

164 flank of Center track, where Walvis Ridge splits (red circle in Fig.1). Igneous rocks
165 were recovered from Site U1578, representing the deepest basement penetration
166 established during Expedition 391 at 302.14 m. The igneous units comprise pillow and
167 lobate flows, sheet flows, and massive flows with interbedded calcareous sedimentary
168 rocks and volcanoclastic deposits (Fig.2). According to the preliminary geophysical and
169 biostratigraphic studies reported in the expedition report (Sager and Hoernle, 2022), a
170 geologic time span of >2.0 Myrs during the Paleocene (ca. 62.3-64.8Ma) was recovered
171 from Site U1578A. In addition, we observe significant differences in mineral
172 geochemistry between the early and late stages (see Discussion below) and therefore
173 subdivide the Site U1578 samples into upper (Section 20R-30R) and lower units
174 (Section 31R-65R; Fig.2). This subdivision is consistent with the lithostratigraphic
175 division reported in the IODP expedition report (Sager and Hoernle, 2022). The basaltic
176 samples from Site U1578 show a porphyritic texture and contain phenocrysts of Pl, Cpx
177 and Ol. The groundmass often contains Pl, Cpx, skeletal Fe oxide crystals, and Ol.
178 Plagioclase phenocrysts (~10 vol% and 0.5-5 mm) usually exhibit oscillatory zoning
179 (Figs.3a, b). In some samples, Pl phenocrysts show patchy zoning, which consists of
180 irregular corroded cores (Fig.3c). Clinopyroxene phenocrysts (~5 vol% and 0.5-2 mm)
181 also commonly display zoning textures (i.e., core-mantle-rim) and some of them exhibit
182 internal resorption textures (Figs.3d-3f). Note that we use the term “mantle” to refer to
183 the optically distinguishable area between core and rim for zoned Cpx. Most of the Ol
184 phenocrysts (< 5 vol% and 0.5-2 mm) are euhedral to subhedral and usually have Sp
185 and melt inclusions (Figs.3g-3i). Olivine and spinel in this study do not show significant

186 zoning textures. Olivine-hosted Sp crystals range in size from 5 μ m to 20 μ m. Fe oxide
187 crystals in the groundmass appear to be titanomagnetite and ilmenite based on their
188 crystal habit and optical appearance under transmitted and reflected light.

189

190 **3. Analytical methods**

191 3.1 Whole-rock major elements

192 Whole-rock major elements were conducted at the Key Laboratory of Submarine
193 Geoscience, Second Institute of Oceanography (Hangzhou, China). The sample
194 pretreatment of whole rock major element analysis was made by the fusion bead method
195 (Johnson et al., 1999). The flux, which was used for the sample fusion, is a mixture of
196 lithium tetraborate, lithium metaborate, and lithium fluoride (45:10:5), Ammonium
197 nitrate and lithium bromide were used as oxidant and release agents respectively. The
198 melting temperature was 1050 °C and the melting time was 15 minutes. A Zsx Primus
199 II wavelength dispersive X-ray fluorescence spectrometer (XRF) produced by
200 RIGAKU, Japan, was utilized for the analysis of the major elements in the fused pellets.
201 The X-ray tube used in the analysis was equipped with a 4.0-kW end-window Rh target.
202 In this study, the Chinese National standards (GBW07103, GBW07104, GBW07105,
203 GBW07310, GBW07312, GBW07314, and GBW07316) were used to monitor the data
204 quality. The relative standard deviations are within $\pm 5\%$ and all the results are listed in
205 [Table S1A](#).

206

207 3.2 Mineral major- and trace-elements

208 We performed 404 in-situ major and trace element measurements of Ol and Sp
209 crystals extracted from 4 core sections. The composition of Ol and Sp was determined
210 using an Oxford Instruments Ultim Max energy dispersive spectrometry (EDS) detector
211 and an Oxford Instruments Wave wavelength dispersive spectrometry (WDS) detector
212 on the JEOL JSM IT800HL FEG-SEM in the Earth and Planetary Science Department
213 at UC Berkeley. The JEOL FEG-SEM provides a large range of possible beam
214 conditions which can be measured with an in-column Faraday cup, providing access to
215 higher beam currents (~150 nA) suitable for minor-element analysis by WDS and lower
216 beam currents for EDS (~2 nA). For EDS analyses, nepheline (Si), Fe_metal (Fe),
217 diopside (Ca, Mg), MnO (Mn), NiO (Ni), Al₂O₃ (Al), TiO₂ (Ti), and Cr₂O₃ (Cr) were
218 used as calibration standards. Olivine (from Microanalysis Consultants), San Carlos,
219 Mong-Ol, and Chromitite were used to monitor the accuracy of the measurements as
220 well as analytical drift (see [Table S1B](#) for full list of standard sources). During EDS
221 analyses, the beam was rastered over a 1-1.5 μm area using a beam current of 2 nA and
222 an accelerating voltage of 20 kV to determine the FeO-MgO-SiO₂-CaO-MnO-NiO in
223 Ol and FeO-MgO-SiO₂-TiO₂-Al₂O₃-Cr₂O₃ in Sp crystals. Following EDS analyses, the
224 beam current was increased to 150 nA and the aluminum (Al) content in Ol, at the
225 location of each EDS analysis, was determined via WDS. Counts of Al were collected
226 for 90 seconds on peak, and background counts were collected for 45 seconds on either
227 side of the peak. The WDS Al measurements were calibrated on an Al₂O₃ standard.
228 Under these conditions, the analytical uncertainty (1σ) for Al in Ol was ~17 ppm. A
229 profile of points spaced 2-3 μm apart was collected for each Ol-Sp pair, see [Table S1B](#)

230 for more information.

231 In total, we performed 120 in-situ major element measurements of Pl crystals
232 extracted from 11 core sections and 553 measurements of Cpx crystals extracted from
233 15 core sections at Site U1578 (Fig.2). All the results are listed in Tables S1C and D.
234 The chemical compositions of these crystals were measured on polished thin sections
235 by a four-spectrometer Jeol JXA 8100 electron probe microanalyzer (EPMA) in the
236 Key Laboratory of Submarine Geoscience, State Oceanic Administration, Second
237 Institute of Oceanography (Hangzhou, China). Back-scattered electron images were
238 utilized to check the homogeneity of the phenocrysts and the crystal zoning. We used
239 an accelerating potential of 15 kV, a beam current of 20 nA, and a spot size of 5 μm for
240 most analyses. Smaller crystals were analyzed using a 1-3 μm spot. Jadeite (Na),
241 olivine (Mg), hematite (Fe), pyrope garnet (Al), diopside (Si and Ca), orthoclase (K),
242 fluorapatite (P), rutile (Ti), nickel oxide (Ni), and rhodochrosite (Mn) were used for
243 calibration standards (Tables S1C). Element peaks and backgrounds were measured for
244 all elements with counting times of 10 s and 5 s, respectively (except for Fe, and Mn,
245 which were 30 s and 15 s, respectively). Each sample was analyzed 10 times in different
246 locations, and the precision (1σ) is $< 3\%$ for Si, Al, Fe, Mg, Ca, and 8% for Ti, and Na
247 based on repeated standard and sample analyses.

248 A total of 42 Cpx grains from 6 core sections were selected for trace element
249 analysis in this study and all the results, including standard samples, are shown in Table
250 S1E. Trace element analysis of Cpx was conducted at the laser ablation-induced
251 coupled plasma-mass spectrometry (LA-ICP-MS) laboratory in the School of Earth

252 Science, Zhejiang University, China. Analyses were performed using an iCAP-RQ
253 single-collector quadrupole ICP-MS (Thermo Fisher Scientific, Waltham, USA)
254 coupled to an Analyte G2 ArF excimer laser ablation (LA) system equipped with HelEx
255 2 volume sample chamber (Teledyne Cetac Technologies, Omaha, USA). The spot size
256 and frequency of the laser were set to 60 μm and 5 Hz respectively, and all laser analyses
257 were targeted at the cores of individual Cpx grains rather than their rims. The fluency
258 of the laser analyses was 18 J/cm^2 . Each analysis incorporated a background acquisition
259 of approximately 20-30 s followed by 50 s of data acquisition from the sample. Two
260 geochemically distinct reference glasses (BCR-2G, BHVO-2G) were used to cover the
261 possible geochemical spectrum. Every 6 sample analyses were followed by two
262 analyses of NIST SRM 610 to correct the time-dependent drift of sensitivity and mass
263 discrimination. We employed a multi-external-standard without internal-standard
264 approach, processing our data using the ICPMSDataCal software (Liu et al., 2008). The
265 software was used to perform off-line selection and integration of background and
266 analytic signals, as well as time-drift correction and quantitative calibration for trace
267 element analysis. Relative standard deviations (% RSD) of LA-ICP-MS analyses are
268 less than 10% for all trace elements.

269

270 3.3 In-situ Sr isotopes of Pl

271 In-situ Sr isotope analyses of Pl were performed on a Neptune Plus MC-ICP-MS
272 (Thermo Fisher Scientific, USA) coupled with Geolas Pro 193 nm ArF excimer laser
273 ablation system (Coherent, Gottingen, Germany) at State Key Laboratory of

274 Lithospheric Evolution (SKLLE), Institute of the Geology and Geophysics, Chinese
275 Academy of Sciences (IGGCAS). Detailed operating conditions for the laser ablation
276 system and the ICP-MS instrument and data reduction are the same as described by [Xu
277 et al. \(2022\)](#). During the LA-MC-ICP-MS analysis, the laser had an energy density of
278 2.8 J/cm², a frequency of 8 Hz, and a beam spot size of 130 μm. Two natural Pl
279 standards, BDL and JH-56 were used as unknown samples to verify the accuracy of the
280 calibration method for in situ Sr isotope analyses. The measured values of BDL and JH-
281 56 are consistent with the recommended values ([Xu et al., 2022](#); [Yang et al., 2009](#)) and
282 all the results are listed in [Table S1F](#).

283

284 **4. Results**

285 4.1 Whole-rock major elements

286 We present major element data for eight Center track volcanic rock samples from
287 different core sections of Site U1578 ([Fig.2](#); [Table S1A](#)) and all major elements have
288 been recalculated to 100 wt.% on a volatile free basis for further discussion. All
289 volcanic rocks are characterized by relatively high total alkali contents (Na₂O + K₂O =
290 3.7-5.5 wt.%) and plot mainly within the field of alkaline basalt and trachybasalt in the
291 TAS classification diagram ([Fig.4](#)). U1578 volcanic rocks have relatively low Mg#
292 values (39.2-53.0; Mg# = 100×Mg/(Mg + Fe_{tot}), where Mg and Fe_{tot} represent molar
293 proportions) and low MgO contents (4.4-7.8 wt.%; [Table S1A](#)), indicating relatively
294 evolved magmas. Comparisons to the published geochemical data ([Class and le Roex,
295 2008](#); [Cliff et al., 1991](#); [EWART et al., 2004](#); [Gibson et al., 2005](#); [Hoernle et al., 2015](#);

296 [Homrighausen et al., 2019](#); [Homrighausen et al., 2023](#); [Le Roex, 1985](#); [Le Roex et al.,](#)
297 [1990](#); [Rohde et al., 2013a](#); [Salters and Sachi-Kocher, 2010](#); [Willbold and Stracke, 2006](#);
298 [Willbold and Stracke, 2010](#)) indicate that U1578 volcanic rocks plot within the Gough
299 track (including Gough Island) field in the TAS classification diagram ([Fig.4](#)). Most
300 volcanic rocks from the Tristan track (including Tristan Island) have higher total alkali
301 contents and lower SiO₂ contents, belonging to tephrite basanite series.

302

303 4.2 Major and minor-elements of Ol and their Sp inclusions

304 Olivine phenocrysts and their Sp inclusions are common in some Site U1578
305 basalts ([Fig.2](#)). Olivines from upper igneous units (i.e., Section 28R) display higher Fo
306 values (87.8-90.5; $Fo = 100 \times Mg / (Mg + Fe_{tot})$, where Mg and Fe_{tot} represent molar
307 proportions) and MgO contents (46.49-49.10 wt.%), whereas Ols from lower igneous
308 units (i.e., Section 53 and 60R) have lower Fo values (79.9-88.6) and MgO contents
309 (41.86-48.74 wt.%). Olivines show negative correlations between CaO, MnO, and Fo
310 values, and positive correlations between NiO, and Fo values, and no clear correlation
311 between Al₂O₃ and Fo ([Fig.5](#)). Olivines in U1578 volcanic rocks record variations in
312 NiO and MnO concentrations between 0.18-0.44 wt.% and 0.10-0.30 wt.% respectively,
313 and CaO contents range from 0.22-0.45 wt.% ([Table S1B](#)). We suggest that these Ol
314 phenocrysts are not mantle-derived xenocrysts because of their significantly higher
315 CaO and MnO contents relative to those of Ol in mantle xenoliths ([Ren et al., 2004](#);
316 [Thompson and Gibson, 2000](#)). The Ol-hosted Sp inclusions from U1578 are Cr-rich
317 and exhibit Cr# of 38.4-67.7 ($Cr\# = 100 \times Cr / (Cr + Al)$, where Cr and Al represent

318 molar proportions). Al₂O₃ (15.53-33.10 wt.%) and Cr₂O₃ (27.53-48.55 wt.%) contents
319 of the Sp crystals in studied samples plot along the Cr-Al trend (Barnes and Roeder,
320 2001), which Irvine (1967) explained as lines corresponding to Sp equilibrating with
321 Ol of constant composition at constant temperature. Notably, Sp inclusions from upper
322 igneous units usually have higher Cr# values (60.8-67.7, average = 65.99) and lower
323 Al₂O₃ contents (15.53-17.69 wt.%, average = 16.15) than those from lower igneous
324 units (Fig.6b; Table 1). Also, Sp inclusions from upper igneous units plot within the
325 ocean island basalt (OIB) field in the diagram of TiO₂ versus Al₂O₃ (Kamenetsky et al.,
326 2001), which is consistent with the geological setting of Site U1578 (Sager and Hoernle,
327 2022). However, Sp inclusions from lower units plot in the transitional field between
328 OIB and MORB, and even within the MORB field (discussed below). Regarding the
329 OIB field in Fig.6b, it is defined based on multiple OIB localities, including Hawaii
330 (Kilauea, Mauna Loa, Mauna Kea), Reunion, and French Polynesia (Tubuai, Mangaia).

331

332 4.3 Major, trace-elements of Cpx

333 Clinopyroxene crystals from Site U1578 have compositions spanning Wo₃₈₋₄₈En₂₈₋
334 ₄₉Fs₇₋₂₅ (Wo = 100×Ca/(Mg + Fe + Ca); En = 100×Mg/(Mg + Fe + Ca); Fs = 100×
335 Fe/(Mg + Fe + Ca)) and mainly classified as diopside and Ca-rich augite (Fig.7a;
336 (Morimoto et al., 1988). Complex variations in Cpx chemistry are observed between
337 different samples, but there is no obvious correlation between the depths of Site U1578
338 and Cpx compositions (Table S1C). Clinopyroxenes from Site U1578 display positive
339 correlations between CaO and Mg# values, negative correlations between TiO₂, FeO_t,

340 MnO, Na₂O, and Mg#, and no correlation between Al₂O₃ and Mg# (Fig.S1). In this
341 study, we use macrocrysts and microlites to refer to crystals larger than 200 μm and
342 smaller than 200 μm, respectively. Clinopyroxene macrocrysts usually have higher CaO
343 (20.01-21.86 wt.%), Mg# values (74.4-88.0) and lower Na₂O (0.24-0.49 wt.%), FeO_t
344 (4.01-9.18 wt.%), and TiO₂ (0.81-1.89 wt.%) contents than those of Cpx microlites
345 (CaO = 18.06-21.14 wt.%; Mg# = 53.2-79.7; Na₂O = 0.23-0.55 wt.%; FeO_t = 7.01-
346 13.91 wt.%; TiO₂ = 0.91-3.58 wt.%; Table S1C). Clinopyroxene macrocrysts from Site
347 U1578 usually display significant zoning textures (Fig.3) and complex geochemical
348 variations (Fig.S1). The compositional zoning of Cpx macrocrysts could reflect
349 sequential changes in the magmatic environment through time (Ubide et al., 2021) or
350 the recycling of antecrysts in the plumbing system (Molendijk et al., 2022).

351 The abundance of Ni and Cr in all Cpx crystals are positively correlated with Mg#
352 values, whereas V, Co, Zn, and Zr are negatively correlated with Mg# values (Fig.8;
353 Table S1E). No correlation exists between Sc, Sr, and Mg# values. The chondrite-
354 normalized patterns (Sun and McDonough, 1989) of rare earth elements (REE) in Cpx
355 exhibit sub-parallel trends and convex-upward REE patterns (Fig.9a). Clinopyroxene
356 microlites exhibit higher REE contents and more pronounced negative Eu anomalies
357 ($\text{Eu}/\text{Eu}^* = \text{Eu}_N / (\text{Sm}_N \times \text{Gd}_N)^{0.5}$; 0.72-0.89) than Cpx macrocrysts (0.84-1.12; Fig.8f),
358 indicating the abundant crystallization of Pl in the groundmass during late stage melt
359 differentiation (Palummo et al., 2021). In the primitive mantle (PM)-normalized multi-
360 element spectra (McDonough and Sun, 1995), Cpx crystals show negative Ba, Nb, and
361 Sr anomalies (i.e., $\text{Sr}/\text{Nd}_{\text{PM}} = 0.07\text{-}0.56$), positive Th anomaly (i.e., $\text{Th}/\text{Ba}_{\text{PM}} = 4.67\text{-}$

362 147.42) and depletion of Zr and Hf relative to the neighboring elements (i.e., Zr/S_{MPM}
363 = 0.30-0.55, $Hf/S_{\text{MPM}} = 0.43-0.94$; Fig.9b). Using the experimentally measured
364 partition coefficients of REE between coexisting Cpx and melt (Sun and Liang, 2012),
365 we calculate the REE patterns of the parental melts in equilibrium with Cpx (Fig.9a),
366 which closely resemble the average whole-rock OIB composition from Sun and
367 McDonough (1989).

368

369 4.4 Major elements and in-situ Sr isotopic compositions of Pl

370 Plagioclase crystals collected from different units of Site U1578 were analyzed for
371 major elements using EPMA (Fig.2; Table S1D). Plagioclase macrocrysts are mainly
372 bytownites with compositions of $Ab_{13.7-35.0}An_{63.6-85.8}Or_{0.4-1.5}$ ($Ab = 100 \times Na / (K + Na +$
373 $Ca)$; $An = 100 \times Ca / (K + Na + Ca)$; $Or = 100 \times K / (K + Na + Ca)$), whereas Pl microlites
374 are labradorites with compositions of $Ab_{28.8-41.8}An_{56.3-70.1}Or_{1.0-1.9}$ (Fig.7b). Some Pl
375 macrocrysts display normal zoning textures (Fig.3b) with decreasing An values from
376 core to rim (Fig.7c). Negative correlations between An and TiO_2 and FeO are clearly
377 observed (Fig.S2). Ninety in-situ Sr isotopic analyses have been conducted for 16 Pl
378 macrocrysts from Site U1578 (Table S1F), displaying significant variations in $^{87}Sr/^{86}Sr$
379 isotopes (0.7038 ± 0.0001 to 0.7046 ± 0.0001). It is important to note that abrupt
380 changes in the $^{87}Sr/^{86}Sr$ isotope composition between upper igneous units ($0.7044 \pm$
381 0.0001) and lower igneous units (0.7040 ± 0.0001) are observed (Table 1). All reported
382 $^{87}Sr/^{86}Sr$ values are presented with $\pm 2\sigma$ analytical uncertainty

383

384 **5. Discussion**

385 5.1 Evidence for the thermal anomaly of Center track

386 Temperature is a fundamental property of magmas and can be used to identify their
387 origin and tectonic setting (Herzberg, 2011; Herzberg and Asimow, 2008; Putirka, 2005;
388 Zhang et al., 2023). In this study, the Al-in-Ol thermometer from Zhang et al. (2023) is
389 applied to calculate the crystallization temperature of Ol. This thermometer is not
390 affected by crystallization pressure, melt composition, oxygen fugacity, and water
391 content (Coogan et al., 2014; Wan et al., 2008). Details of the Al-in-Ol thermometer
392 method and Monte Carlo error propagation approach used to assess the impact of
393 analytical uncertainty are included in the Supplementary Text S1 and Table S2.
394 Crystallization temperatures calculated from Ol-Sp pairs in U1578 samples show a
395 large range, with an expected increase in temperature with increasing Fo content
396 (1205.6 ± 23.9 °C to 1288.5 ± 23.9 °C; 1SE; Fig.10a). At a given Fo content, Ol in this
397 study exhibits significantly higher crystallization temperatures than those of global
398 MORBs (Coogan et al., 2014; Matthews et al., 2021), but similar to those estimated
399 from Iceland (Matthews et al., 2016; Spice et al., 2016). This observation supports the
400 presence of anomalously hot primary melts and indicates a thermal anomaly in the
401 mantle source beneath the Center track.

402 To further constrain the thermal anomaly in Center track of Walvis Ridge, we
403 calculated mantle potential temperatures (T_p) using pyMelt modeling. PyMelt is an
404 open-source python library for calculating the melting behavior of the Earth's mantle
405 (Matthews et al., 2022). As a forward modeling tool, pyMelt allows users to input

406 mantle potential temperature and mantle source lithology proportions to estimate the
407 primary crystallization temperature. Therefore, to estimate the mantle potential
408 temperature, we need to constrain both the approximate proportion of pyroxenite in the
409 mantle source beneath the Center track and the primary crystallization temperature of
410 the magmas. For the mantle source lithology, two melting lithologies (the KLB-1
411 lherzolite (Katz et al., 2003), and the silica-undersaturated pyroxenite KG1 (Kogiso et
412 al., 1998)) are used. We set four possible source pyroxenite fractions ($px = 10\%$, 15% ,
413 20% , and 25%) based on a previous study from Gough and Tristan da Cunha Island (Li
414 et al., 2025). For the primary crystallization temperature, we extrapolate a liquid line of
415 descent from the Ol-Sp crystallization temperature to $Fo = 91$ (following the method
416 from Matthews et al. (2021)), which would be in equilibrium with a lherzolitic mantle
417 source. This yields a maximum primary magma temperature of 1401°C for the lower
418 unit and 1316°C for the upper unit (Fig.10a). For each lithology condition, we tested
419 13 different mantle potential temperatures in pyMelt and then used linear interpolation
420 to determine the T_p values required to produce the primary magmatic temperatures
421 calculated above (red dashed lines in Fig.10b). The pressure at which melting
422 terminated (representing lithospheric pressure) was set to 1.5 GPa , based on data from
423 Tristan da Cunha Island (Li et al., 2025), and the crystallization pressure was set to 0.4
424 GPa based on constraints from Cpx barometry (Supplementary Text S2). The
425 uncertainties on T_p were estimated by propagating the errors from the Ol thermometer
426 and from extrapolating the liquid line of descent used to calculate the primary
427 crystallization temperature. The results show that the T_p required to recreate the

428 crystallization temperature of the lower unit ranges from 1522 °C (px = 10%) to
429 1580 °C (px = 25%). It should be noted that these calculated T_p values of the lower unit
430 are ~200 °C higher than those of ambient MORB-source mantle (typically ~1350 °C).
431 However, these estimates should be regarded as absolute upper limits of T_p , and the
432 actual values may be significantly lower. Several factors could lead to overestimation,
433 including: (i) a high contribution of pyroxenite-derived melts, which would invalidate
434 the assumption that the primary mantle melts are in equilibrium with Fo₉₁ Ol as the
435 most primitive Ol in such melts would likely have a lower Fo content (Li et al., 2025);
436 and (ii) the abundant crystallization of phases other than Ol (e.g., Cpx and Pl), which
437 would change the crystallization path and reduce the applicability of the Ol-based
438 reverse crystallization method from Matthews et al. (2021). Interestingly, the upper unit
439 yields T_p estimates of 1373°C (px = 15%) to 1424 °C (px = 25%), significantly lower
440 than the lower unit. We consider these upper unit values, derived from an Ol with Fo >
441 90, to be representative of the lower bound of possible mantle potential temperature
442 beneath the Center track (~1400 °C). The difference in T_p between the upper and lower
443 units may reflect variations in the proportion of pyroxenite in the mantle source and/or
444 the spatial variation within the plume itself. Considering relative plate motion, the
445 higher T_p of the lower units may correspond to the center of the plume, whereas the
446 lower T_p of the upper units – which are believed to be 2 Myrs younger (Sager and
447 Hoernle, 2022) – may correspond to the plume edge, an issue that requires further
448 investigation in future work.

449 Taken together, our results suggest that the Center track of Walvis Ridge is derived

450 from a source that is at least 50 °C hotter than the mantle sources beneath mid-ocean
451 ridge settings (Fig.10c). This indicates that a hotspot or mantle plume may have fed this
452 track. The REE patterns of the parental melts in equilibrium with Cpx are elevated in
453 LREE relative to HREE in a manner that is more similar to published OIB than MORB
454 data (Sun and McDonough, 1989), which is also consistent with a hotspot or mantle
455 plume origin (Hofmann and White, 1982). A plume origin is also supported by
456 identification of enriched mantle one (EMI) components in the source region of the
457 Tristan-Gough-Walvis Ridge hotspot track (Hoernle et al., 2015; Homrighausen et al.,
458 2020; Homrighausen et al., 2023) and a rising plumelet just southwest of the Tristan
459 island group based on the local seismic tomography (Schlömer et al., 2017). The
460 elevated $^3\text{He}/^4\text{He}$ isotope signatures in the basalts from both upper and lower units of
461 U1578 ($17.48 \pm 1.45 R_A$; Scholpp et al., 2025) indicate that the presence of a primordial,
462 undegassed mantle component in the mantle source region beneath the Center track -
463 most likely originating from lower-mantle material entrained into an upwelling plume
464 structure (Homrighausen et al., 2019; Homrighausen et al., 2020; Homrighausen et al.,
465 2023). Combined with the preliminary age of Site U1578 (62-64 Ma) (Sager and
466 Hoernle, 2022), we propose that the mantle plume beneath the Center track of Walvis
467 Ridge was active and carried heat upward in rising columns of hot material at 62-64
468 Ma. This helps to explain why the magma production rate of Guyot Province was
469 relatively high at 62-64 Ma (Huang and Li, 2024). In summary, our new data indicates
470 that there was a thermal anomaly in the source region of the Center track, which was
471 most likely related to a mantle plume.

472

473 5.2 Characteristics and origins of the diversity of crystal zonings

474 Mineral textures and associated in-situ compositional variations can record
475 information about crystallization conditions at the scale of an entire magmatic plumbing
476 system and provide crucial information about magma transfer and storage through
477 Earth's mantle and crust (Giacomoni et al., 2016; Li et al., 2020; Nardini et al., 2024;
478 Palummo et al., 2021; Pelullo et al., 2022; Wei et al., 2022; Zhang et al., 2024). In this
479 study, we conduct a K-means cluster analysis on N=42 crystals using eight geochemical
480 parameters of Cpx; Mg#, Na₂O, Sc, Cr, Ni, La, Y and jadeite (Jd) content from LA-
481 ICP-MS data. K-means clustering was performed in Python using the scikit-learn
482 package (Pedregosa et al., 2011). Data were first standardized using StandardScaler,
483 before applying KMeans to classify all crystals into three groups and principal
484 component analysis (PCA) was applied to visualize the clustering results (Fig.S5). Mg#,
485 Sc, Cr, and Ni are sensitive to melt composition and fractionation degree, while La and
486 Y capture the variability in trace element enrichment and depletion of the mantle source.
487 Na₂O and Jd content are included because they are more sensitive to crystallization
488 pressure and temperature (Putirka et al., 2003). Together, these parameters provide a
489 good balance between major and trace element information.

490 The average major and trace element compositions of Cpx from Groups 1-3 are
491 summarized in Table 2, with detailed data available in Tables S1C and S1E. Group 1
492 are characterized by high Mg# values (> 83.4), Ni (> 200 ppm) and Cr (> 5000 ppm);
493 Group 2 have Mg# values of 76.6- 83.4, Ni and Cr contents of 100-200 ppm and 1000-

494 5000 ppm; and Group 3 are characterized by low Mg# values (< 76.6), Ni (< 70 ppm)
495 and Cr (< 30 ppm) contents. The distinct Mg# of these groups was used to classify the
496 N=553 EPMA analyses, given the lack of high-quality trace element data for these spots
497 to perform the full clustering. 21% of these EPMA analyses are classified as Group 1,
498 49% as Group 2 and 30% as Group 3. There is sufficient range in the Mg# of individual
499 crystals that different zones within the same crystal can be linked to different groups
500 (i.e., Group 1-3 in Fig.11a, Group 2-3 in Fig.11b). We calculate the crystallization
501 temperatures and pressures of Cpx using Cpx-melt and Cpx-only thermobarometry (see
502 [Supplementary Text S2](#) for details), and the results are presented in [Table S3](#).

503 When considering zoning across the crystal, pyroxenes show four textural types:
504 (1) euhedral, unzoned crystals; (2) normally zoned crystals with higher Mg#, Cr₂O₃,
505 and CaO, and lower Na₂O, TiO₂, FeO, and MnO contents in their cores/mantles
506 compared to their rims (i.e., [Fig.3d](#)); (3) reverse-zoned crystals with intermediate
507 compositions in their cores and higher Mg# values in their mantles or rims (i.e., [Fig.3e](#));
508 and (4) sector-zoned crystals (i.e., [Fig.3f](#)). In this study, sector zoning in Cpx consists
509 of Si-Mg-rich, Al-Ti-poor hourglass forms $\{-1\ 1\ 1\}$ and Si-Mg-poor, Al-Ti-rich prism
510 forms $\{h\ k\ 0\}$ ([Table S1](#)). In addition, sector zoning in Cpx at Site U1578 follows typical
511 sector partitioning in alkaline magma systems ([Downes, 1974](#); [Kouchi et al., 1983](#);
512 [Leung, 1974](#); [Mollo et al., 2020](#); [Ubide et al., 2019](#)), suggesting low degrees of
513 undercooling in a dynamic crystallization regime ([Di Stefano et al., 2020](#); [Klügel et al.,](#)
514 [2020](#); [Masotta et al., 2020](#); [Tapu et al., 2022](#)). However, the absence of sector zoning
515 in most Cpx phenocrysts suggests crystallization under even lower degrees of

516 undercooling during slower magma ascent, where growth rates were reduced to
517 approach near-equilibrium conditions (Ubide et al., 2019).

518 Within each of these zoned crystals, individual points making up transects can be
519 classified into Group 1-3 using the clustering algorithm. Some normally zoned crystals
520 show a core with high Mg# (> 83.4; Group 1), a mantle with intermediate composition
521 (Mg# = 76.6-83.4; Group 2) and a rim with low Mg# (< 76.6; Group 3; Fig.11a). Within
522 the three zones, the composition is relatively constant, with sharp transitions between
523 zones. Other crystals have an intermediate core (Mg# = 76.6-83.4; Group 2) and a rim
524 characterized by low Mg# (< 76.6; Group 3), and the width of this rim can differ
525 (Figs.11b, 11c). Some reversely zoned crystals have intermediate Group 2 cores, with
526 a mantle of more primitive Group 1 compositions, before dropping back to more
527 evolved Group 2 compositions with a continuous drop in Mg# towards the rim (Fig.
528 11d). Others have very evolved Group 3 cores, passing into more primitive Group 2
529 mantles, and finally dropping in Mg# towards the rim (Fig.11e). In summary, Cpx
530 crystals at Site U1578 show various zoning patterns and sharp geochemical
531 composition changes from their cores, and mantles to their rims, suggesting these Cpx
532 crystals have undergone complex crystallization processes within different magmatic
533 environments marked by the plateaus of relatively constant composition (Ginibre et al.,
534 2007; Jerram and Martin, 2008; Streck, 2008; Ubide et al., 2021).

535 According to the similar trace element patterns (Fig.9) and continuous variation
536 trends on the plots of trace element (i.e., Cr, Ni, and V) against Mg# (Fig.8), we suggest
537 that the three different Cpx groups (Group 1, Group 2 and Group 3) in the Center track

538 of Walvis Ridge crystallized from co-genetic magmas (Jankovics et al., 2016; Li et al.,
539 2020; Liu et al., 2022; Wei et al., 2022; Zhang et al., 2024). Based on the distinct core
540 compositions of Cpx from Group 1, Group 2, and Group 3 (Fig.11), we propose that
541 three different magmatic environments (MEs) existed beneath the Center track, named
542 ME1, ME2, and ME3, respectively. The high Mg# Group 1 cores of normally zoned
543 Cpx (i.e., Fig.11a) are interpreted to crystallize in a more primitive magmatic
544 environment (ME1), which is consistent with their higher Cr (> 5000 ppm) and Ni (>
545 200 ppm) contents (Fig.8). Subsequently, they were incorporated into more evolved
546 magmas where they grew low Mg# mantles and rims (i.e., ME2). Normally zoned Cpx
547 with intermediate Group 2 cores (i.e., Figs.11b, 11c) are interpreted to crystallize in a
548 more evolved magmatic environment (ME2) than crystals with Group 1 cores. The
549 relatively sharp compositional changes between different zones indicates rapid
550 transition to a new set of conditions stabilizing a different Cpx composition, here
551 inferred to represent a new magmatic environment. The rims of the normally zoned Cpx
552 shown in Fig.11c are relatively thin (<20 μm) compared to the thicker (~75 μm) rims
553 of the Cpx shown in Fig.11b. Based on the Fe-Mg interdiffusion rates in Cpx (Müller
554 et al., 2013), the thin rims of Cpx in Fig.11c and the lack of a compositional plateau
555 indicate that they likely formed during ascent, cooling and crystallization of the
556 groundmass, rather than prolonged storage in ME3 (Dimanov and Wiedenbeck, 2006;
557 Mollo et al., 2010; Zhang et al., 2024). On the contrary, the plateau in Mg#, CaO and
558 TiO₂ for the Group 3 rim in Fig.11b is more consistent with crystallization from a cooler,
559 evolved melt in a relatively stable environment (ME3) (Cao et al., 2022; Ma et al.,

560 2024).

561 Reverse-zoned Cpx is generally related to magma mixing, suggesting an open
562 system behavior of the magmatic system (Ginibre et al., 2007; Liu et al., 2022;
563 Molendijk et al., 2022; Streck, 2008). In BSE photomicrographs of reverse-zoned Cpx
564 shown in Fig.11d, a subhedral light grey crystal core (Group 2) is surrounded by a dark
565 grey mantle with higher contents of Mg# and Cr₂O₃ and lower contents of TiO₂ and
566 FeO (Group 1). We suggest that the internal patchy shape reflects thermal erosion as a
567 consequence of hot, mafic magma input from ME1 into the resident evolved crustal
568 magma reservoir (Ma et al., 2024; Palummo et al., 2021), which was followed by
569 crystallization forming the more euhedral darker rim. Reverse-zoned Cpx with highly
570 evolved Group 3 cores (Fig.11e) crystallized in the most evolved magmatic
571 environment (ME3), before the injection of more mafic magma from ME2 that partially
572 dissolved the pre-existing crystal core during the evolved stage. This further suggests
573 that magma mixing was relatively common within ME3. It should be noted that,
574 although the cores of reverse-zoned Cpx shown in Fig.11e and the rims of some
575 normally zoned Cpx (i.e., Fig.11b) are all classified as Group 3 based on Mg# alone,
576 they are chemically distinct. The composition of cores showing Group 3 characteristics
577 have a very small range of Mg# values (72.5-76.6). In contrast, the rim analyses
578 classified as Group 3 have a much wider range of Mg# values (62.5-76.6), and extend
579 up to far higher TiO₂ and Al₂O₃ contents (Fig.S9). We interpret the cores to have formed
580 in the relatively stable ME3 with a relatively narrow range of compositions, and the
581 rims to have formed during additional crystallization during ascent and syn-eruptive

582 quenching at the seafloor (Giacomoni et al., 2021; Giacomoni et al., 2016; Streck, 2008).

583 The complex magma recharge and mixing processes revealed by Cpx are also
584 supported by the chemical zoning in Pl phenocrysts (Fig.S10). Plagioclase phenocrysts
585 from Site U1578 are usually oscillatory zoned and can be divided into fine banding and
586 coarse banding based on the width of the zoning (Cao et al., 2022; Downes, 1974;
587 Streck, 2008). The widths of oscillatory coarse banding in Site1578 Pl lie typically
588 above 50 μm and can exceed 200 μm in some cases (i.e., Figs.S10a, S10d), which
589 usually results from magma mixing (Streck, 2008; Yang et al., 2019) or polybaric-
590 isothermal convection in a large magma chamber (Ustunisik et al., 2014). While the
591 latter mechanism is possible given that experimental data indicates a change of ~ 3 mol%
592 An per kbar (Ustunisik et al., 2014), this would not explain the observed Cpx Mg#
593 zoning. It seems more likely that the multiple abrupt increases in An% content (up to
594 10%) from core to rim of oscillatory coarse banding are likely to be produced by the
595 same process as abrupt zoning in Cpx; magma mixing between more evolved and
596 primitive melts from the different magmatic environments.

597 Some oscillatory zoning in Pl phenocrysts record at least three episodes of
598 compositional variations (i.e., Fig.S10a), indicating that they have undergone at least
599 three regularly paced small injections of magmas that were slightly hotter and more
600 compositionally primitive than the Pl growth environment (Ginibre et al., 2002; van
601 Gerve et al., 2020). Compared with the coarse oscillatory banding, the fine oscillatory
602 zoning with low-amplitude anorthite content (0-3 An%) and short width (< 20 μm) (i.e.,
603 Figs.S11b, S11c) may be largely attributed to crystallization kinetics (Ginibre et al.,

604 2002; Loomis, 1982; Tsune and Toramaru, 2007) and crystallization in a stable magma
605 reservoir with stable oxidation conditions according to the stable Fe contents in fine
606 oscillatory zoning (Tapu et al., 2022). However, we cannot rule out the possibility that
607 such fine-scale zoning also records frequent, small-volume injections of new magma
608 into the reservoir, which could produce subtle chemical variations without forming
609 coarse oscillatory bands. Repeated smaller injections could perhaps account for some
610 of the ‘spikeyness’ in the profiles for Cr and Ti in Cpx (Fig.11), that is smeared away
611 by faster diffusion of Fe-Mg so not seen in Mg# transects.

612

613 5.3 In-situ Sr isotopes in Pl and their implications on the petrogenesis of Site 1578
614 basalts

615 Whole-rock Sr-Nd-Pb-Hf-La-Ce-O isotopic data of rocks from Tristan-Gough
616 hotspot track have revealed that the underlying plume structure is likely sourced from
617 a heterogeneous lower mantle region, containing contributions from EM1-type mantle
618 (Boyet et al., 2019; Gibson et al., 2005; Harris et al., 2000; Hoernle et al., 2015;
619 Homrighausen et al., 2019; Homrighausen et al., 2020; Salters and Sachi-Kocher, 2010;
620 Willbold and Stracke, 2010). As we discussed above, however, magma bodies in the
621 crust result from mixing of multiple batches of mantle-derived magmas (Batiza, 1984;
622 Cipriani et al., 2004), and the degree to which we can reconstruct the heterogeneity in
623 mantle source regions from basalts is uncertain (Lambart et al., 2019). Moreover,
624 conducting systematic geochemical analyses from the lower to upper igneous units at
625 one drill site can shed new light on mantle heterogeneity on the centimeter to kilometer

626 scale (Lambart et al., 2019; Lissenberg et al., 2024) and help us determine the temporal
627 geochemical evolution of the mantle source (Höfig et al., 2014; Sano et al., 2011).

628 In this study, Site U1578 Pl are isotopically heterogeneous, with a clear contrast
629 between upper and lower units (Fig.12; upper = $\sim 0.7044 \pm 0.0001$; lower = $\sim 0.7040 \pm$
630 0.0001). These $^{87}\text{Sr}/^{86}\text{Sr}$ values are significantly higher than those from basalts from
631 South Mid-Atlantic Ridge (SMAR; $^{87}\text{Sr}/^{86}\text{Sr} = 0.7030\text{-}0.7035$; Castillo and Batiza,
632 1989; Fontignie and Schilling, 1996; Graham et al., 1996; Ito et al., 1987; Regelous et
633 al., 2009; Zhang et al., 2021). Within a single crystal, variations are mostly unresolvable
634 within analytical uncertainty (Fig.S11). To assess whether the $^{87}\text{Sr}/^{86}\text{Sr}$ variations in Pl
635 between different depth intervals are statistically significant, we performed two-sample
636 t-tests for all possible pairs of Pl grains from nine core sections (see Table S5). The
637 results are visualized in a heatmap (Fig.S12), where pairwise comparisons with p-
638 values < 0.05 indicate statistically significant differences. The five deepest samples (37-
639 60 R) are all significantly different from the four upper samples (22R-29R). When
640 considering all measurements within the upper vs. lower units, these are also
641 statistically different with p-values < 0.05 .

642 We wish to investigate the origin of these differences between lower and upper
643 units. Sr isotopes can be modified when Pl are metasomatically altered by seawater.
644 However, the combination of low loss on ignition (LOI) contents (Table S1A) and
645 petrographic observations indicating fresh mineral phenocrysts (Figs.3, S13) suggests
646 minimal secondary overprint, indicating that $^{87}\text{Sr}/^{86}\text{Sr}$ ratio in the Pl is likely to retain
647 primary magmatic signatures. Moreover, the lack of differences in LOI contents

648 between the lower and upper units further indicates that the significant Sr isotopic
649 variations are not attributable to alteration. Considering the slow diffusion rate of Sr in
650 Pl and the relatively short residence time in the magma reservoirs of intraplate
651 volcanoes (Costa et al., 2020), we infer that their $^{87}\text{Sr}/^{86}\text{Sr}$ ratios are unlikely to be
652 significantly modified through reequilibration with the host melt. Thus, it seems likely
653 that the shift towards more radiogenic Sr isotopes in the upper units do reflect the
654 infiltration of higher $^{87}\text{Sr}/^{86}\text{Sr}$ isotopic components in the crust or plume itself (Edwards
655 et al., 2019; Waight and Tørnqvist, 2018).

656 It is still unclear whether the Center track represents a distinct geochemical end
657 member or simply represents mixing between the Tristan- and Gough-type (Sager and
658 Hoernle, 2022). To further explore the contribution of Tristan and Gough tracks to the
659 Center track, we compare our data with the published Sr isotope ratios of volcanic rocks
660 along the Tristan and Gough tracks (Hoernle et al., 2015; Rohde et al., 2013a; Salters
661 and Sachi-Kocher, 2010). Plagioclases from lower igneous units of Site U1578 have
662 significantly lower Sr isotope ratios than Gough track samples, but plot within the range
663 of the Tristan track samples from DSDP Sites 527 and 528 (Fig.13a), indicating they
664 may share similar sources with Tristan track samples. We propose that the shift to more
665 radiogenic Sr from lower to upper igneous units is likely attributed to incorporation of
666 more radiogenic Gough-type components into the plume source, rather than a process
667 occurring in the crust beneath Walvis Ridge. However, more work is required to test
668 this idea further (i.e., Nd-Hf-Pb isotope). It should be noted that the presence of
669 sedimentary interbeds (Figs.2, 12), and magnetic reversals recorded within Site U1578

670 (Sager and Hoernle, 2022) imply episodic volcanism over an extended period (>2 Myrs)
671 (Knight et al., 2004; Potter et al., 2019). Although the whole-rock compositions
672 between upper and lower igneous units show no obvious differences (Fig.4), Ol from
673 upper igneous units are characterized by higher Fo, NiO contents and crystallization
674 temperature (Figs.5, 10), indicating that waning activity on the center track is
675 characterized by eruption of magma carrying primitive crystal cargoes.

676 In conclusion, our in-situ Sr isotope data in Pl favor a hotspot/mantle plume origin
677 for the Center track, given their similarity to the Tristan track and deviation from
678 MORB. Additionally, Pl from Center track is broadly similar to those of Tristan-type
679 end-member, but the shift to more radiogenic Sr isotopic signatures from the lower to
680 the upper igneous units suggests an increasing contribution of radiogenic melts into the
681 volcanic plumbing system, potentially derived from an increased supply of material
682 from the Gough-type end-member and/or changes in the supply of melts from distinct
683 mantle sources due to variations in melting conditions (i.e., lower mantle T_p).

684

685 5.4 The complex magmatic plumbing system beneath the Center track

686 Based on the lines of evidence discussed above, we proposed that three different
687 magmatic environments were present beneath the Center track (Fig.14). The diverse
688 zoning patterns of Cpx reveal a complex magmatic plumbing system beneath the Center
689 track, involving processes such as fractional crystallization, magma mixing, recharge,
690 and ascent. Multi-stage magma recharge is consistent with the observed chemical
691 zoning patterns and Sr isotope compositions in Pl phenocrysts. To better constrain the

692 relative mixing of different mantle-derived melts contributing to different magmatic
693 systems (upper vs lower), we apply a quantitative binary mixing model using Sr
694 isotopic composition versus whole-rock Ba/Sr ratios for samples with LOI < 3 wt.%
695 (Table S6). We selected the most isotopically enriched and depleted whole-rock
696 compositions from 58-72 Ma Tristan-Gough tracks as endmembers (Homrighausen et
697 al., 2019; Rohde et al., 2013a). Sr and Ba exhibit similar geochemical behavior in
698 basalts and have comparable mineral-melt partition coefficients (Philpotts and
699 Schnetzler, 1970), therefore, their ratio is only minimally affected by fractional
700 crystallization. The whole-rock Sr isotope-Ba/Sr mixing diagram (Fig.13b) shows that
701 Tristan track samples are characterized by isotopically depleted values, whereas Gough
702 track samples are marked by enriched isotope ratios. For the Center track samples,
703 results indicate that the upper units (0.7042-0.7046) require 20-30% contribution from
704 the enriched Gough melts, whereas the lower units (0.7038-0.7042) require only 10%.
705 This implies that the upper units incorporated 10-20% more of melts from the enriched
706 Gough component relative to the lower units and variations in the contribution of
707 isotopically enriched melts played a key role in shaping the temporal and compositional
708 evolution of the Center track lavas.

709 Furthermore, the potential influence of the South Mid-Atlantic Ridge on the
710 Center track should not be overlooked. Notably, the $^{87}\text{Sr}/^{86}\text{Sr}$ values and An contents of
711 the Pl crystal in Sample 82 deviate markedly from the binary mixing trends (Fig. S2C)
712 towards lower $^{87}\text{Sr}/^{86}\text{Sr}$ ratios. In addition, in binary compositional plots (Figs.5, 6), 19
713 Ol and 6 Sp analyses fall within the typical MORB fields. We suggest that these crystals

714 are most likely inherited as scavenged xenocrysts from the Atlantic lower crust, rather
715 than crystallized from the Center track parental magma. Petrographic observations
716 reveal that the Pl crystal was larger (~0.8 cm) than other Pl phenocrysts (~0.2 cm) and
717 display corroded margins (Fig.S13C), indicating reaction with the host basaltic liquid
718 (Sial, 1977). Moreover, the occurrence of crystal clots within the same thin section (<2
719 cm apart; Fig.S13C) further supports the interpretation that numerous captured crystals
720 were entrained during magma ascent (Lissenberg et al., 2019; Sano et al., 2011). Taken
721 together, we infer that the occurrence of these anomalous mineral crystals in Sample 82
722 does not indicate direct involvement of South Mid-Atlantic Ridge-derived melts, but
723 rather represents the participation of lower oceanic crustal material into the magmatic
724 system beneath the Center track of Walvis Ridge.

725

726 **6. Conclusions**

727 We present the major, trace element, and Sr isotope data for Ol, Cpx, and Pl at Site
728 U1578 to elucidate the origin of magmatism in the Center track of Walvis Ridge and its
729 complex magmatic plumbing system. Our results lead to the following conclusions:

730 (1) Based on the Ol-Sp thermometry and pyMelt modeling, the calculated mantle
731 potential temperature of the Center track is at least 1400 °C, about 50 °C higher than
732 that of MORB, most likely reflecting the influence of a mantle plume.

733 (2) Based on the Cpx Mg# values and trace element compositions, three distinct
734 Cpx groups are identified. Diverse zoning in Cpx and Pl indicates that multi-stage
735 magma recharges and magma mixing are common magmatic processes within the

736 plumbing system beneath the Center track of Walvis Ridge.

737 (3) Site U1578 Pl are isotopically heterogeneous (0.7038 ± 0.0001 to $0.7046 \pm$
738 0.0001) and broadly similar to those of Tristan-type end-member. The shift to more
739 radiogenic Sr isotopic signatures from the lower to the upper igneous units suggests the
740 increasing incorporation of radiogenic materials into the mantle source, potentially
741 derived from the Gough-type end-member. This study helps us determine the temporal
742 geochemical evolution of mantle source heterogeneity beneath the Center track of
743 Walvis Ridge.

744

745 **Declaration of Competing Interest**

746 We have no known competing financial interests or personal relationships that
747 could have appeared to influence the work reported in this paper.

748

749 **Acknowledgments**

750 The authors are grateful to all scientists on board Expedition 391 of the
751 International Ocean Discovery Program, as well as the crew of the D/V *JOIDES*
752 *Resolution* and the IODP technical staff for their role in obtaining samples. We thank
753 Simon Matthews, Wei-Ran Li and Yishen Zhang for their valuable guidance on the
754 pyMELT modeling and Ol-Sp thermometer. We thank Mihai Ducea and Shan Li for the
755 editorial handling, Gang Zeng, and Tong Hou for their constructive comments that
756 significantly improved this work. The authors are also grateful to the support of the
757 Hainan Observation and Research Station of Ecological Environment and Fishery

758 Resource in Yazhou Bay. This study was supported by the National Key Research and
759 Development Program of China (2023YFF0803404), the National Natural Science
760 Foundation of China (Nos. 42072069 and 42176055), the Zhejiang University Seek
761 Truth Eagle Scholarship (No. 12234074). EDS analyses were supported by an NSF IF
762 grant 2218698 to MG and PW.

763

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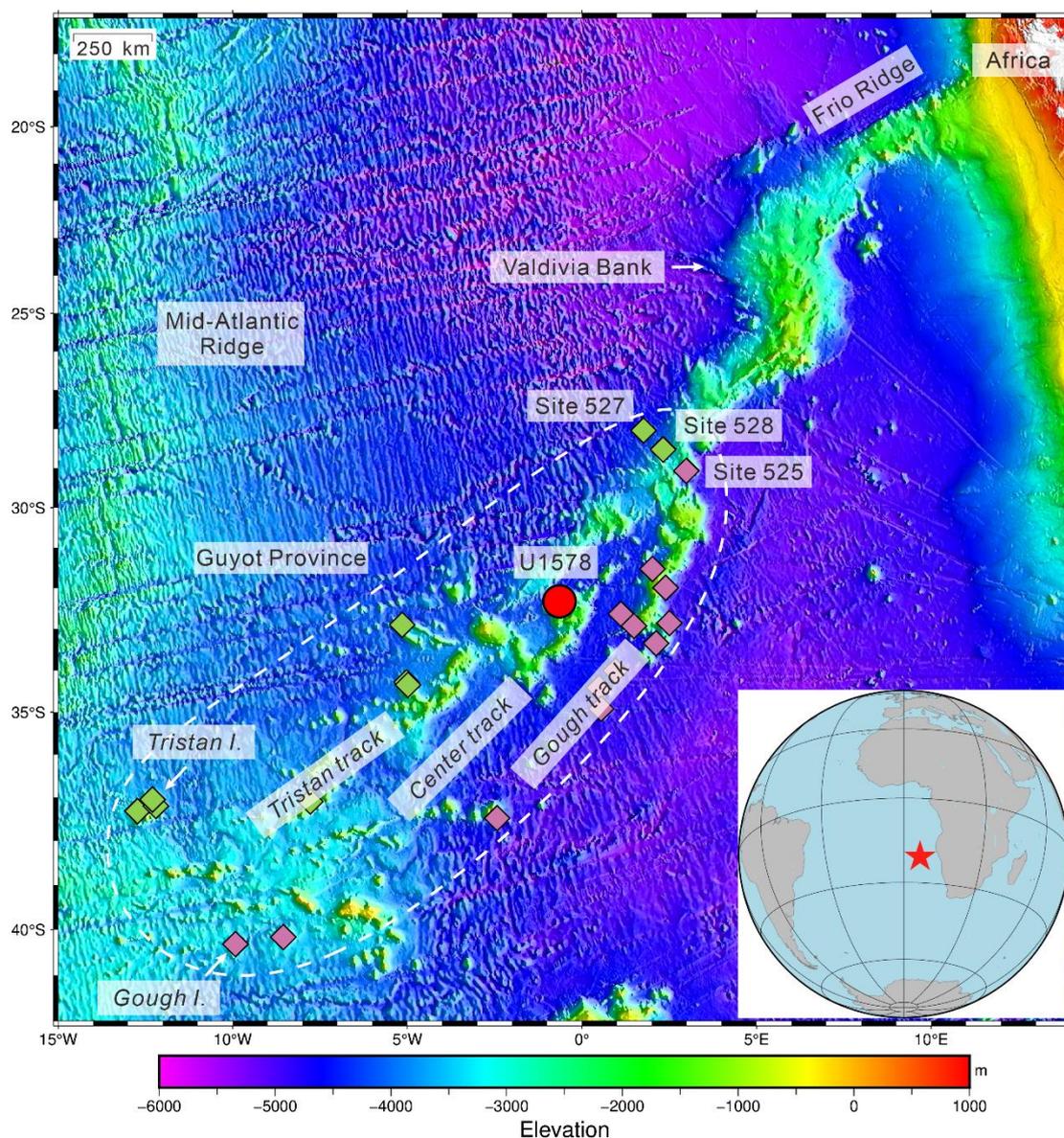
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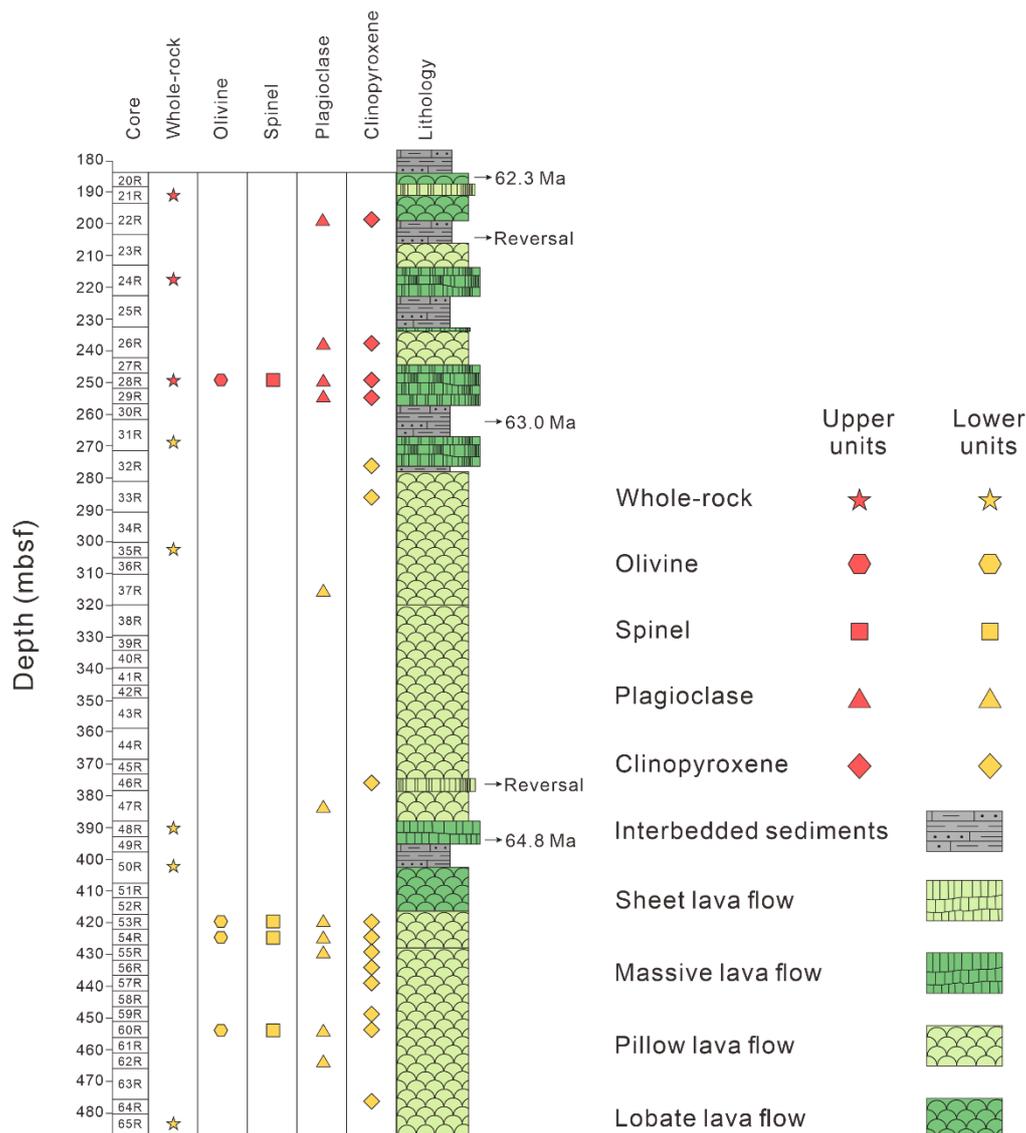
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1131 **Figure and Table Captions**



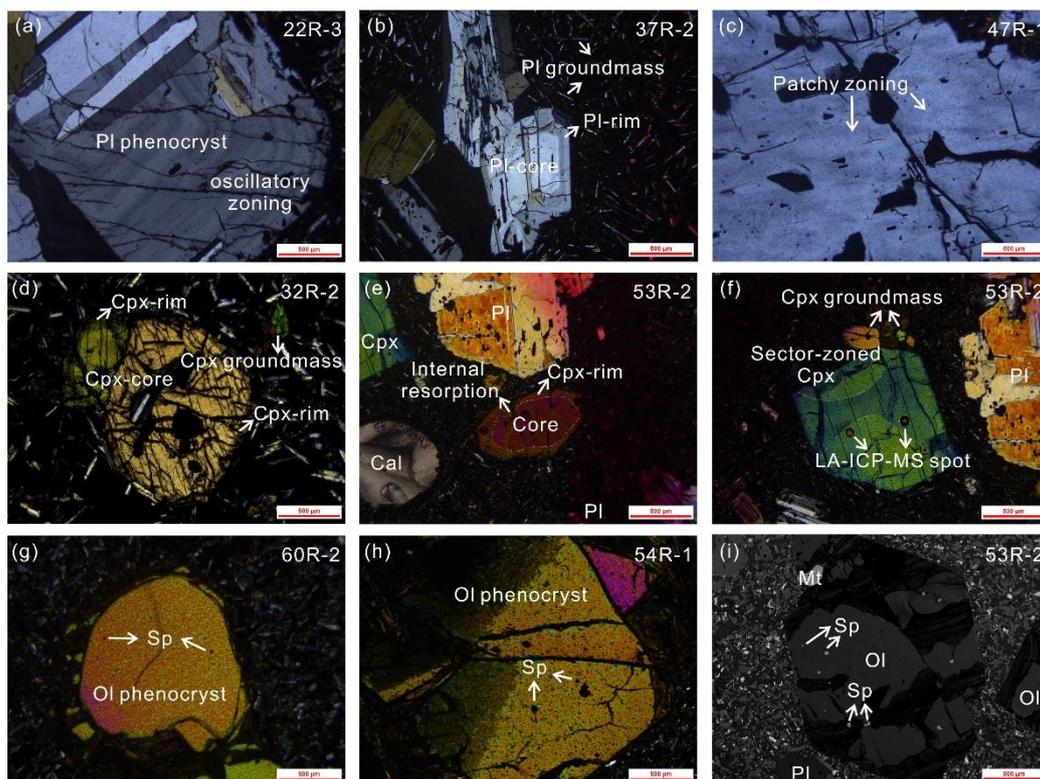
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1133 Fig. 1 Walvis Ridge bathymetry and feature names with the location of IODP Site
1134 U1578 (Sager and Hoernle, 2022). Inset shows the location of Walvis Ridge relative to
1135 the South Atlantic modified after Thoram, S., et al. (2013). The green and purple
1136 diamonds represent the locations of previous studies from Tristan track and Gough track
1137 (Hoernle et al., 2015; Homrighausen et al., 2019; Rohde et al., 2013a; Rohde et al.,
1138 2013b; Salters and Sachi-Kocher, 2010).



1139

1140 Fig. 2 Lithology of Site U1578 at Walvis Ridge, showing both interbedded sediments
 1141 and igneous units (Sager and Hoernle, 2022). Site U1578 penetrated 302.14 m of the
 1142 igneous basement, and it represents the deepest basement penetration established
 1143 during Expedition 391 and 397T. The igneous units comprise pillow and lobate flows,
 1144 sheet flows, and massive flows with interbedded sediments. The age estimates and
 1145 position of magnetic reversals are based on the Expedition 391 preliminary report
 1146 (Sager and Hoernle, 2022). The star, hexagon, square, triangle, and diamond represent
 1147 the whole-rock, Ol, Sp, Pl, and Cpx in this study. Mbsf, meters below the sea floor.



1148

1149 Fig. 3 Cross-polarized light and BSE images of Walvis Ridge volcanic rocks from Site

1150 1578. Zoned Pl phenocrysts under cross-polarized light: (a) oscillatory texture; (b)

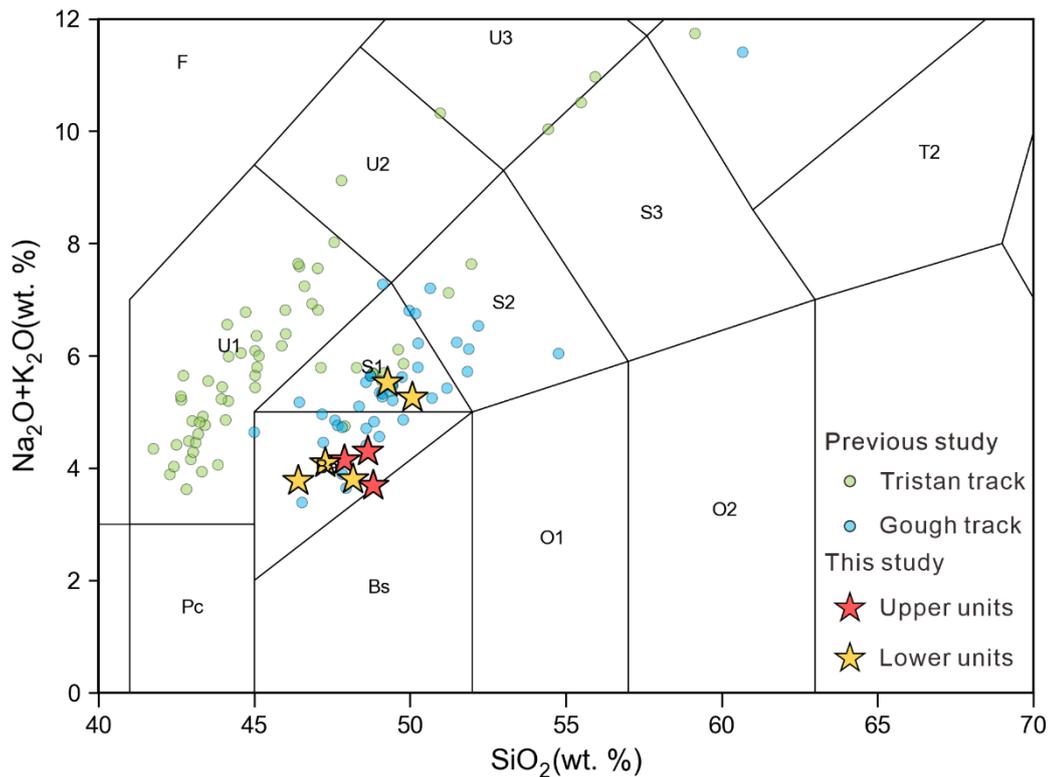
1151 normal zoning texture; (c) patchy texture. Zoned Cpx phenocrysts under cross-

1152 polarized light: (d) normal zoning texture; (e) reverse zoning texture; (f) sector zoning

1153 texture. Petrography of large Ol phenocrysts containing abundant euhedral Sp

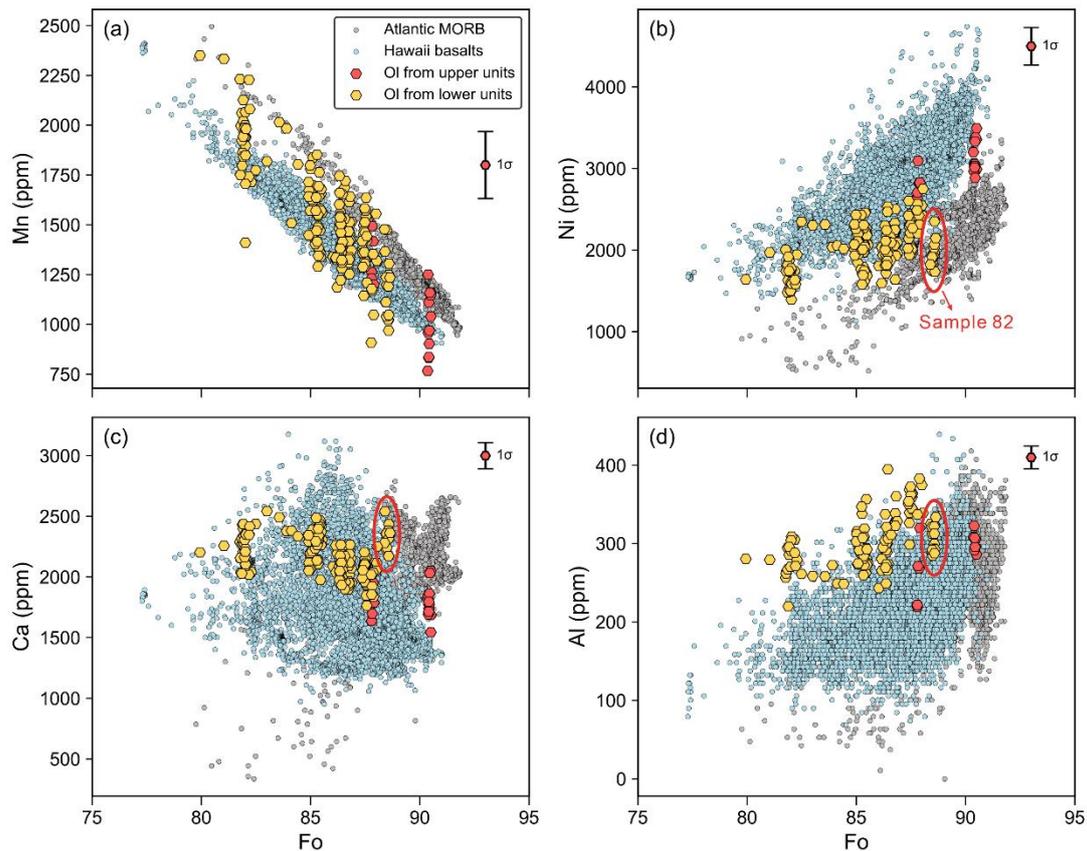
1154 inclusions (g-h: microscopic photo; i: BSE image). Clinopyroxene (Cpx), plagioclase

1155 (Pl), olivine (Ol), spinel (Sp), calcite (Cal), magnetite (Mt).



1156

1157 Fig. 4 Total alkalis vs silica (TAS) diagram for volcanic rocks at Site 1578 compared
1158 with other volcanic rocks from Tristan track and Gough track. The red and yellow stars
1159 represent the samples from upper and lower units, respectively.



1160

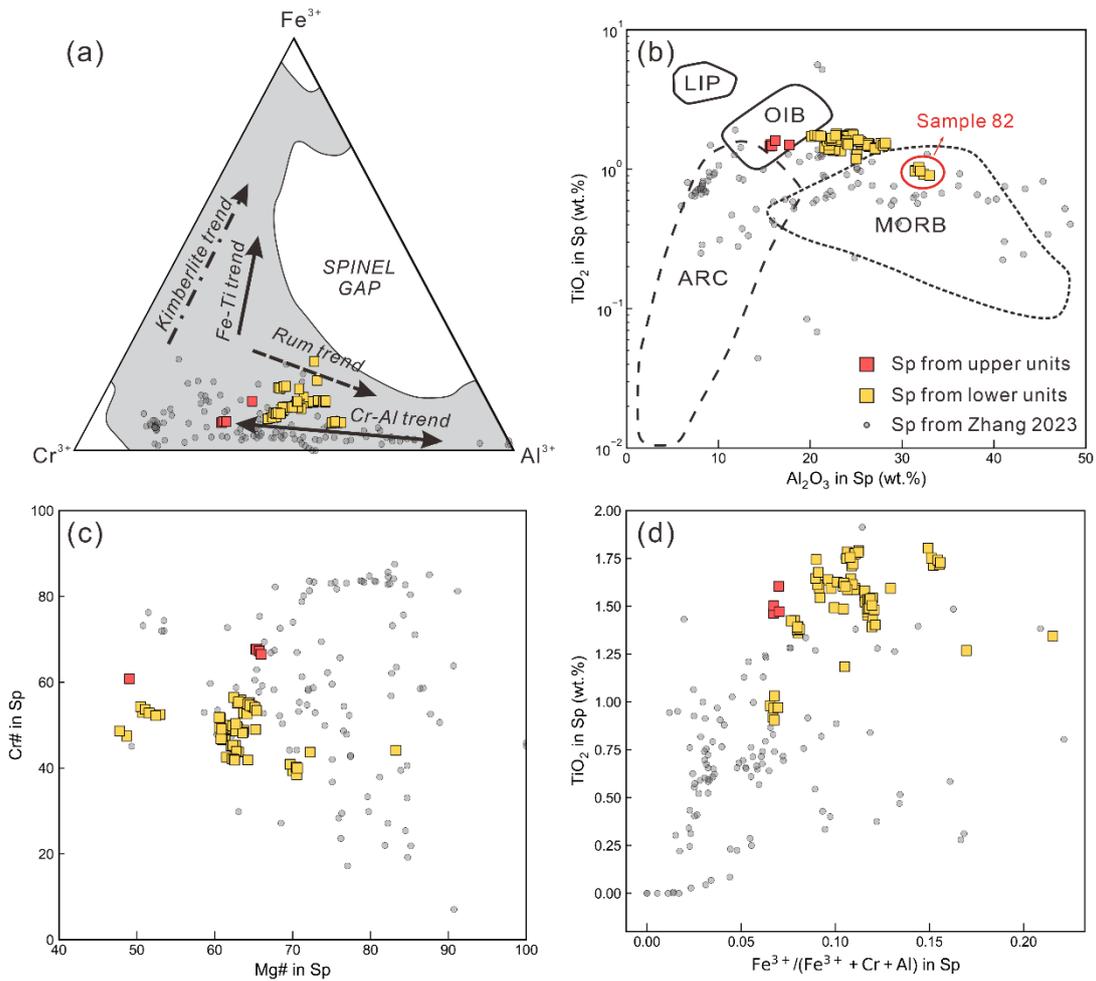
1161 Fig. 5 Olivine Fo content vs Mn (a), Ni (b), Ca (c), and Al (d) concentration in Ol (ppm).

1162 The red and yellow hexagons represent the Ol from upper and lower igneous units,

1163 respectively (see Fig.2). The grey and blue points represent the Ol from Atlantic MORB

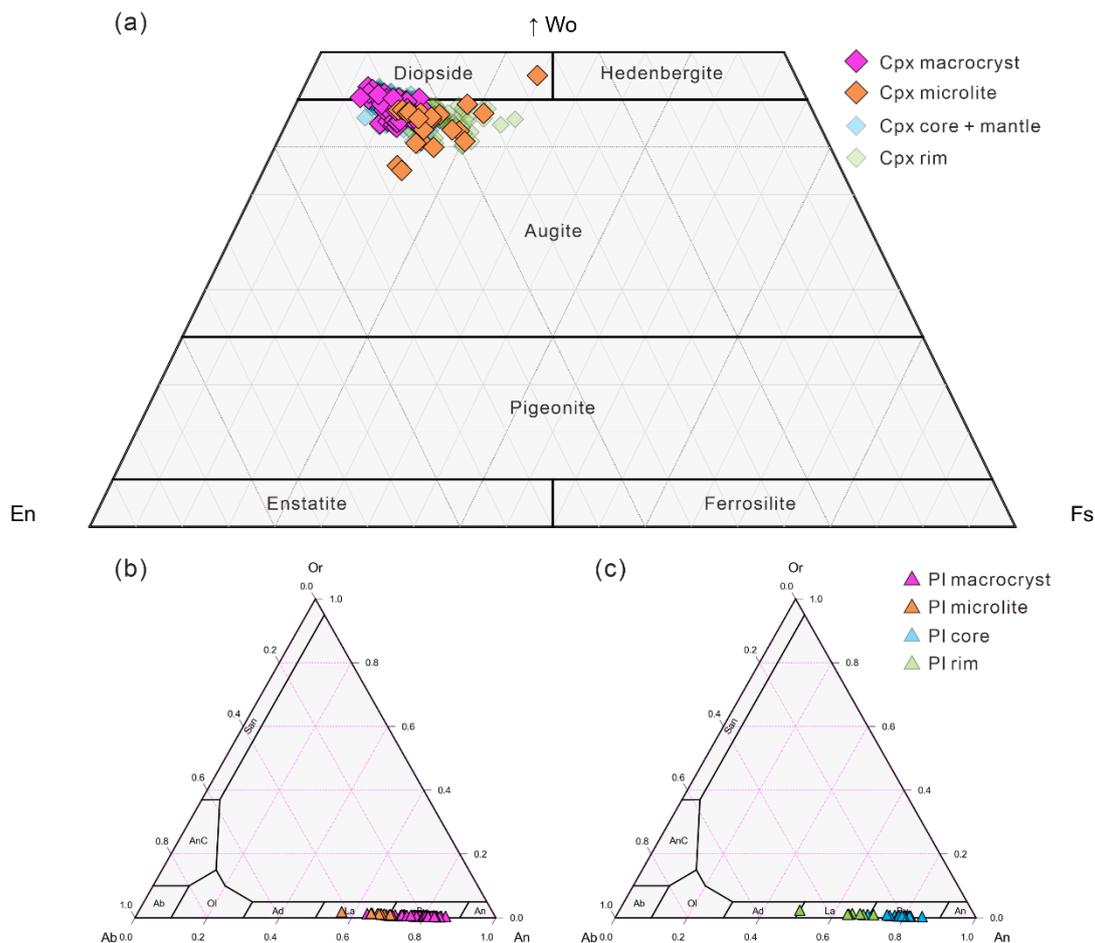
1164 and Hawaii basalts, respectively. Representative one standard deviation analytical

1165 uncertainties are shown as black bars.



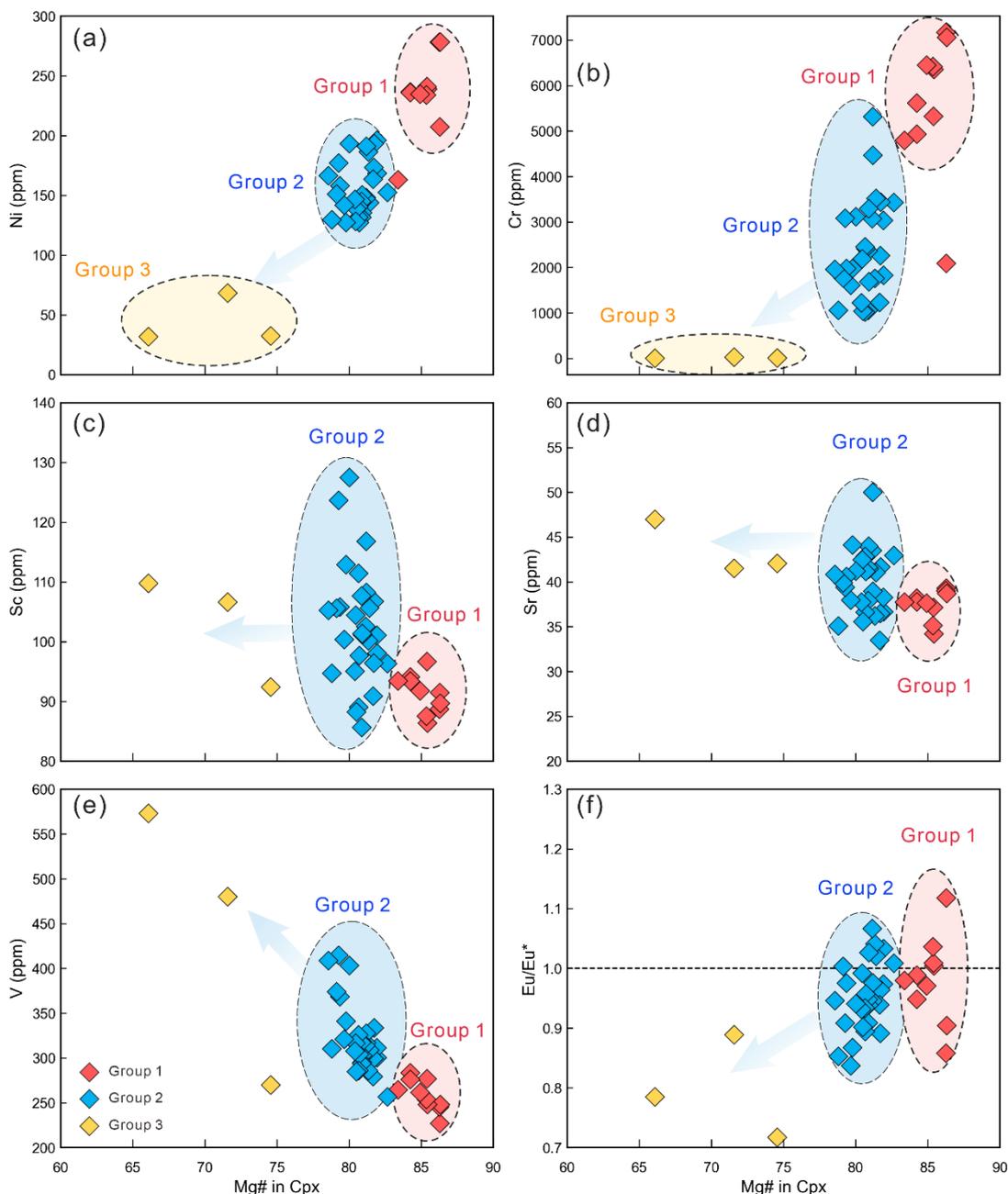
1166

1167 Fig. 6 The compositional variation diagram for Sp. (a) Cr^{3+} - Al^{3+} - Fe^{3+} triangular
1168 diagram for distinguishing the evolution trend of Sp (after Barnes and Roeder, 2001);
1169 (b) Diagram of TiO_2 versus Al_2O_3 for the affinity of Sp to the tectonic settings (after
1170 Kamenetsky et al., 2001); (c) Diagram of Cr# in Sp versus Mg# in Sp; (d) Diagram of
1171 TiO_2 in Sp versus $\text{Fe}^{3+}/(\text{Fe}^{3+} + \text{Cr} + \text{Al})$ in Sp. The red and yellow squares represent the
1172 Sp from upper and lower igneous units, respectively. The grey points represent the Sp
1173 from Zhang et al. (2023) used to calibrate the Al-in-Ol thermometer.



1174

1175 Fig. 7 Compositions of Cpx and Pl from the basaltic lavas cored at Site U1578, Walvis
1176 Ridge. The data points are individual measurements obtained for each crystal, plotted
1177 in the enstatite-ferrosilite-wollastonite (En-Fs-Wo) ternary diagrams ((a); [Morimoto et](#)
1178 [al., 1988](#)), and the anorthite-albite-orthoclase (An-Ab-Or) ternary diagrams (b and c),
1179 which is based on the ternary plot python package ([Harper et al., 2015](#); [Wieser et al.,](#)
1180 [2022](#)). Note that “mantle” in (a) refers to the optically distinguishable area between
1181 core and rim for zoned Cpx.



1182

1183 Fig. 8 Plots of Ni (a), Cr (b), Sc (c), Sr (d), V (e), and Eu/Eu* (f) versus Mg# in Cpx.

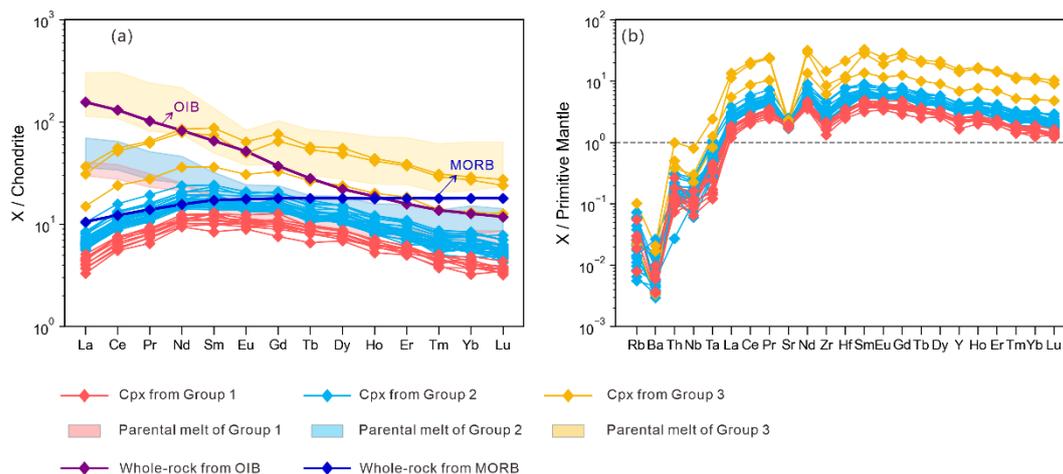
1184 The horizontal dashed line in (f) represents whether the Cpx shows Eu anomalies

1185 ($Eu/Eu^* = Eu_N / (Sm_N \times Gd_N)^{0.5}$). The blue trend lines represent the variations of trace

1186 elements during the fractionation history. The red, blue, and yellow diamonds represent

1187 the Cpx from Group 1, Group 2 and Group 3, respectively.

1188



1189

1190 Fig. 9 Chondrites-normalized REE diagrams (a) and primitive mantle-normalized trace

1191 element abundances diagrams (b) for Cpx in this study. All the data used here were

1192 acquired by LA-ICP-MS. The red, blue, and yellow areas in (a) show REE patterns for

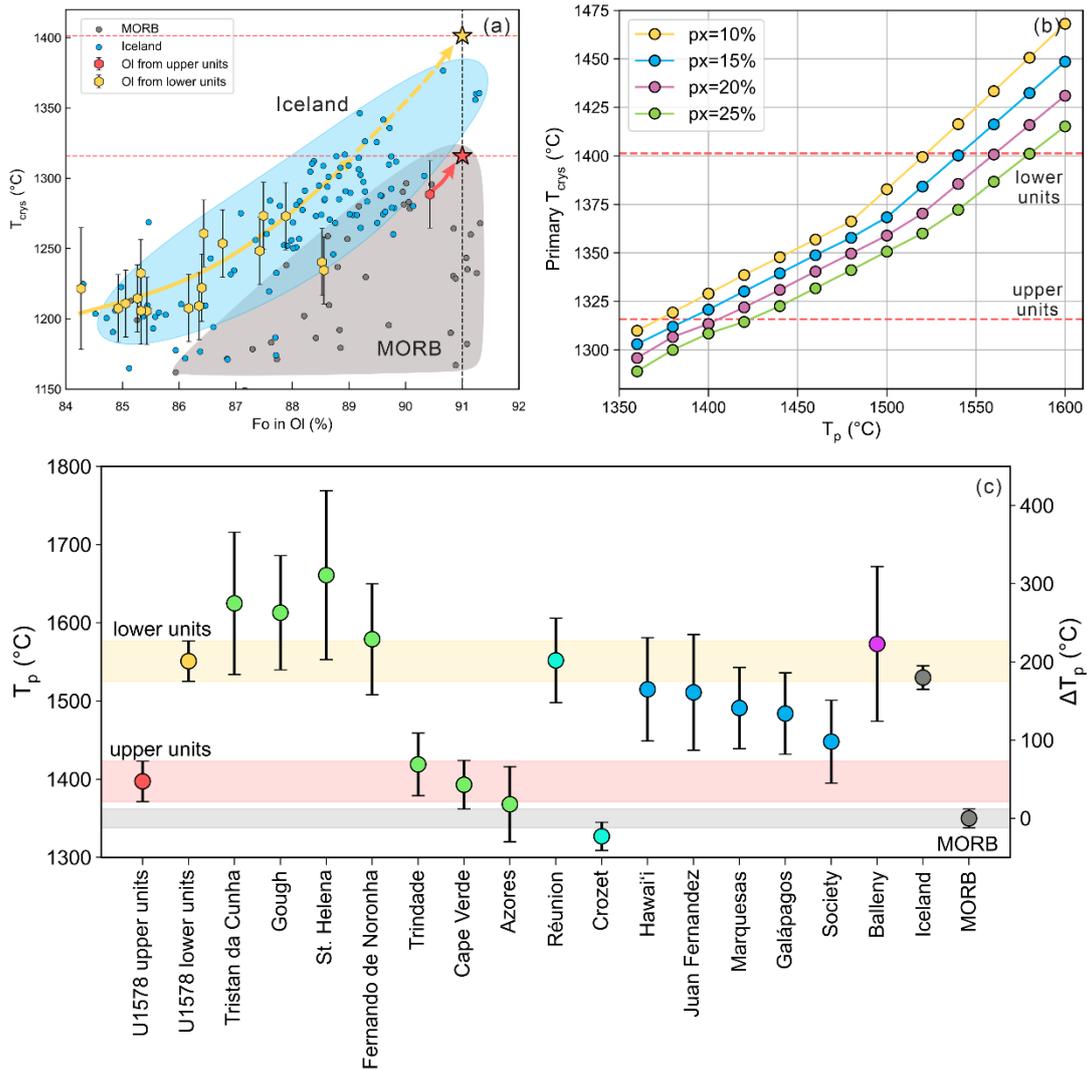
1193 hypothetical melts in equilibrium with Cpx from Group 1, Group 2, and Group 3

1194 respectively, calculated using partition coefficients ($K_D(\text{Fe-Mg})^{\text{Cpx-melt}}$) from Sun and

1195 Liang (2012). OIB, MORB data and chondrite values are from Sun and McDonough

1196 (1989). Primitive mantle normalization values are from McDonough and Sun (1995).

1197 This figure is based on the REE plot python package (Williams et al., 2020).



1198

1199 Fig. 10 (a) Al-in-Ol crystallization temperatures vs. Ol Fo content for Site U1578,

1200 Walvis Ridge in comparison with those for Iceland (Matthews et al., 2016; Spice et al.,

1201 2016) and MORB (Coogan et al., 2014; Matthews et al., 2021). The yellow and red

1202 curves show the result of extrapolating a liquid line of descent from the crystallization

1203 temperature and Ol composition to Fo₉₁ Ol from lower and upper units. (b) pyMELT

1204 simulation results showing primary crystallization temperatures calculated by inputting

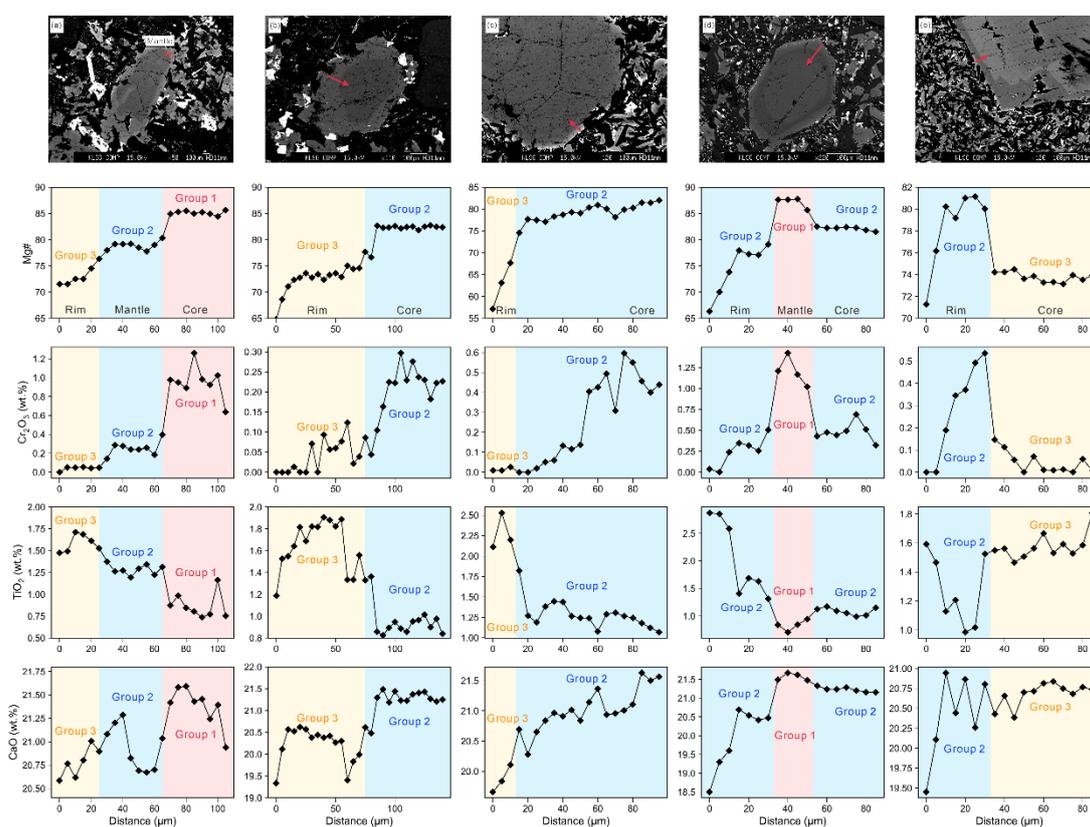
1205 different mantle potential temperatures at four possible source pyroxenite fractions (px

1206 = 10%, 15%, 20%, and 25%). (c) Comparison of mantle potential temperatures from

1207 the Center track with those of mid-ocean ridge basalts, Icelandic basalts, and global

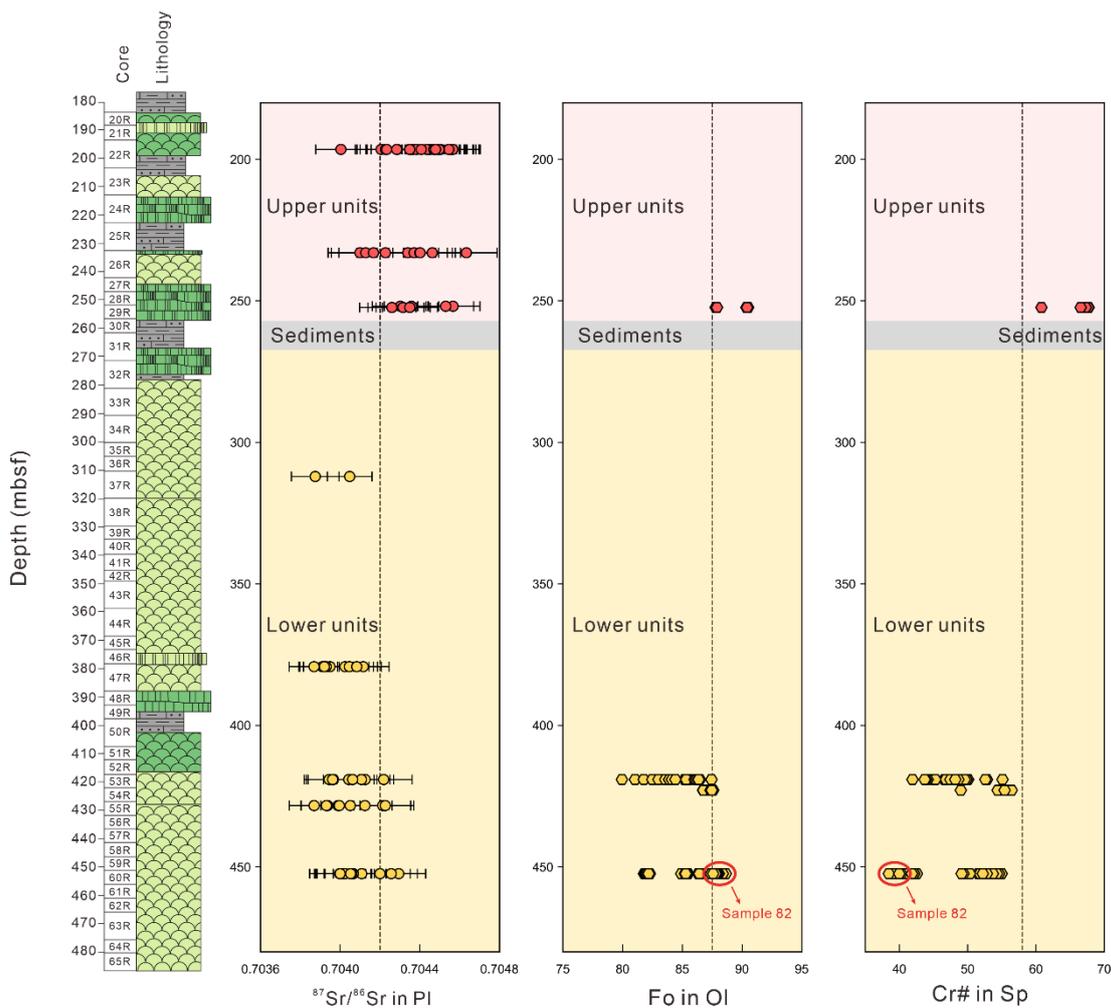
1208 ocean island basalts (Li et al., 2025). The error bars on T_p were estimated by
1209 propagating the errors from the Ol thermometer and from extrapolating the liquid line
1210 of descent used to calculate the primary crystallization temperature.

1211



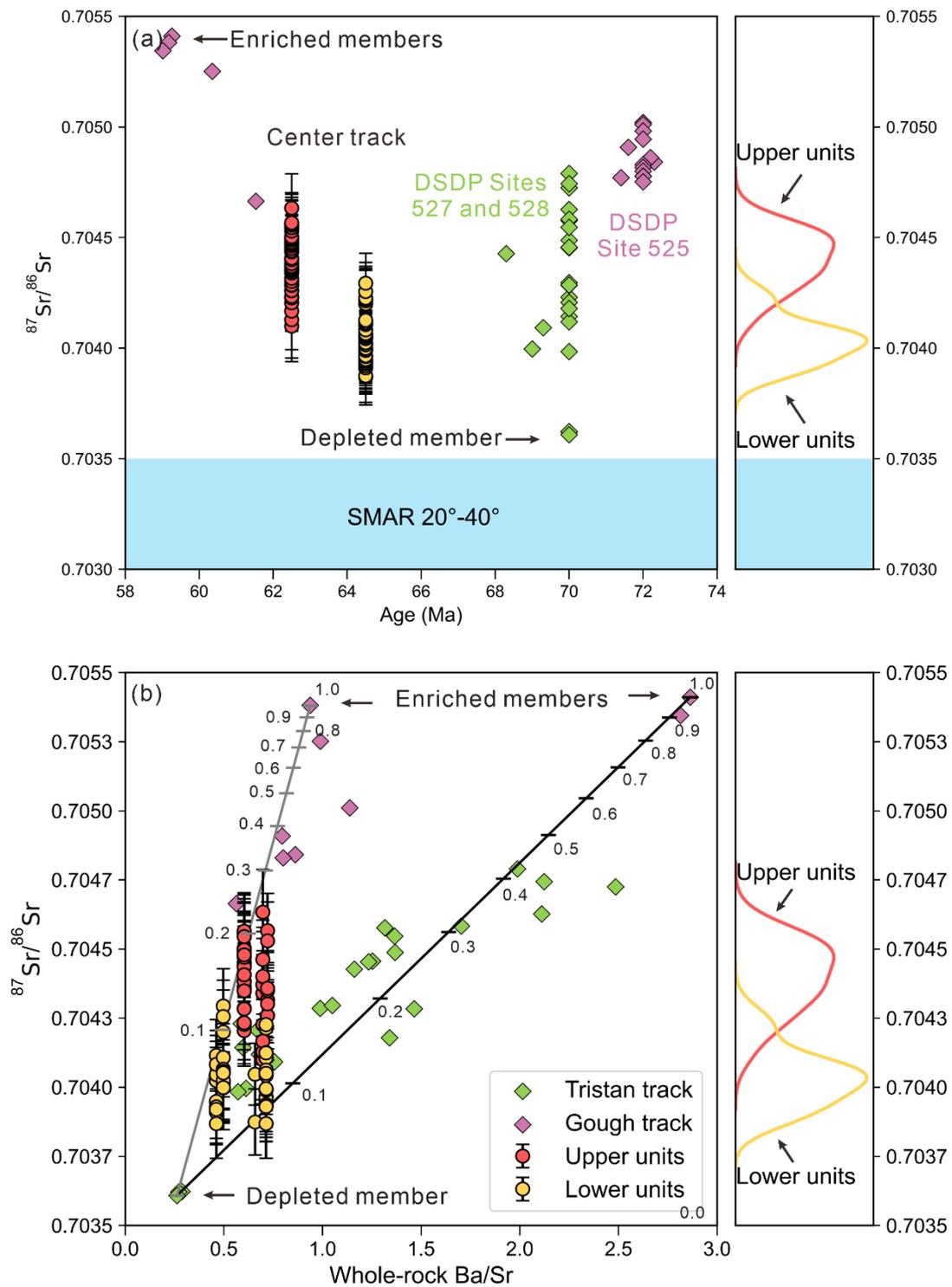
1212

1213 Fig. 11 Representative BSE images and analytical points with major element
1214 compositions for zoned Cpx from volcanic rocks at Site 1578, Walvis Ridge. (a-c) are
1215 normal zoned Cpx and (d-e) are reverse zoned Cpx. The red arrows represent the
1216 location of the electron microprobe traverses across the zoning. The red, blue, and
1217 yellow areas represent the Cpx from Group 1, Group 2 and Group 3, respectively.



1218

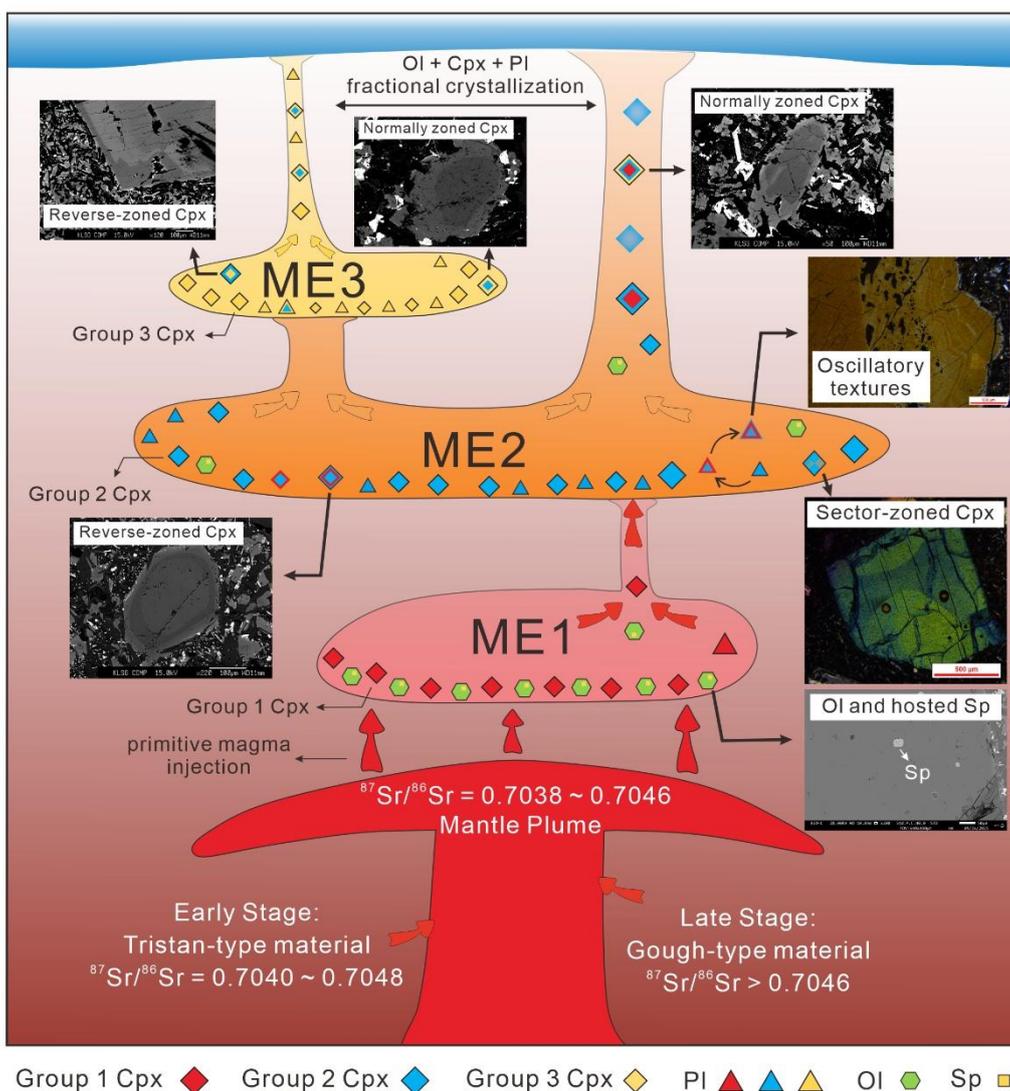
1219 Fig. 12 Downhole chemical variations in Sr isotopes in Pl, Fo in Ol, and Cr# in Sp. The
 1220 red and yellow circles/hexagons represent the Pl/Ol from the upper and lower igneous
 1221 units of Site U1578, respectively. The red, grey, and yellow areas on the stratigraphic
 1222 column represent the upper igneous units, sediments, and lower igneous units,
 1223 respectively.



1224

1225 Fig. 13 (a) Ages for each sample versus $^{87}\text{Sr}/^{86}\text{Sr}$ ratios from the Center track, Tristan
 1226 track, and Gough track. The red and yellow circles represent the PI from the upper and
 1227 lower igneous units of Site U1578, respectively. The green and purple diamonds

1228 represent the literature whole-rock data from the Tristan and Gough tracks (Hoernle et
 1229 al., 2015; Homrighausen et al., 2019; Rohde et al., 2013a; Rohde et al., 2013b; Salters
 1230 and Sachi-Kocher, 2010). Sample locations for these analyses are shown on Fig.1. (b)
 1231 Binary mixing model of whole-rock Sr isotope ratios versus Ba/Sr concentration, with
 1232 the KDE curves on the right showing the distribution of Center track Sr isotope ratios
 1233 of Pl.



1234
 1235 Fig. 14 A petrologic model for the magmatic plumbing system of the Center track in
 1236 Walvis Ridge according to in-situ major, trace elements and Sr isotopes of Ol, Cpx, and

1237 Pl from IODP Site U1578 in this study. Depths and connections between reservoirs are
1238 largely schematic, and reservoirs may be separated vertically, horizontally, or both.

1239

1240 Table 1 The average compositions of Ol, Sp and Pl from Site U1578 upper units, lower
1241 units and Sample 82. – available upon request

1242

1243 Table 2 The average compositions of Group 1, 2 and 3 Cpx from Site U1578. – available
1244 upon request

1245

1246 **Supplementary Text S1. Ol-Sp thermometry and Monte Carlo error propagation**

1247 Prior to temperature calculations, we ensure equilibrium conditions between Ol and Sp by
1248 conducting high-resolution compositional transects (~2 μm spacing) across each Ol-Sp pair.
1249 Equilibrium is assessed based on established $\text{Fe}^{3+}/\text{Fe}^{2+}$ and Cr–Al–Mg partitioning behavior (Coogan
1250 et al., 2014; Zhang et al., 2023). Olivine and spinel are generally euhedral, also suggesting they
1251 formed at equilibrium. Data points used for thermometry, indicated by grey bars in Fig.S3, are
1252 averaged and subsequently used for temperature calculations. This profiling approach avoids the
1253 influence of secondary fluorescence (Matthews et al., 2016), thereby yielding more robust and
1254 representative results.

1255 Temperatures are calculated using the Ol-Sp equilibrium thermometry equations provided by
1256 Zhang et al. (2023), and the results are reported in Table S2A. The thermometric calibration spans
1257 a wide range of P–T–composition conditions (1174–1606 $^{\circ}\text{C}$, 0.1–1350 MPa, up to 7.4 wt.% H_2O)
1258 and encompasses the compositional range of Sp from most OIB samples, including our dataset
1259 (Fig.6). We apply equations 5 and 7 of Zhang et al. (2023) and use a statistical Z-test, following the
1260 method outlined in their study, to determine the most appropriate crystallization temperature for
1261 each Ol-Sp pair. Using the same method, we also calculate crystallization temperatures for Ol-Sp
1262 pairs from other OIBs (Li et al., 2025; Matthews et al., 2021), MORBs (Coogan et al., 2014;
1263 Matthews et al., 2021), and Iceland samples (Matthews et al., 2016; Spice et al., 2016) to compare,
1264 and all results are reported in Tables S2B–D.

1265 To quantify uncertainty, we apply Monte Carlo error propagation for four representative Ol-Sp
1266 pairs from Center track. Each Monte Carlo simulation is repeated 1,000 times, generating a
1267 distribution of temperature estimates from which the mean and standard deviation were calculated.

1268 By employing a high beam current (150 nA) and extended counting times, the analytical uncertainty
1269 (1σ) for Al in Ol was reduced to ~17 ppm. Using the Python script of Zhang et al. (2023), we
1270 propagated the analytical uncertainties of all elements measured in Ol and Sp (Table S1B), and the
1271 results show that the propagated uncertainty in temperature estimates ranges from 5.6 to 10.8 °C
1272 (Table S2E). According to Zhang et al. (2023), the uncertainty on the Ol-Sp Al-exchange
1273 thermometer from experimental calibration is ± 23.9 °C. Therefore, the combined uncertainty —
1274 calculated as $\sigma_{\text{overall}} = \sqrt{(\sigma_{\text{equation}}^2 + \sigma_{\text{analysis}}^2)}$ — is only slightly (0–2 °C) higher than the
1275 calibration uncertainty alone. These results demonstrate that analytical uncertainties associated with
1276 EDS-WDS measurements have a limited impact on the calculated temperature values.
1277

1278 **Supplementary Text S2. Cpx-melt and Cpx-only thermobarometry**

1279 In order to understand the chemical and structural evolution of magmatic plumbing system
1280 further, pressures and temperatures are estimated by clinopyroxene-only (Cpx-only) and
1281 clinopyroxene-melt (Cpx-melt) thermobarometry by using the open-source Python3 package
1282 Thermobar v. 1.0.41 (Wieser et al., 2022). Cpx-only thermobarometry is from Jorgenson et al.
1283 (2022), and Cpx-melt barometry and thermometry are from Neave and Putirka (2017) and Putirka
1284 (2008, eq. 33). All calculated results are shown in Tables S3A and B. Note that crystals with strong
1285 sector zoning are excluded from our calculation (MacDonald et al., 2023; Ubide et al., 2019).

1286 For the Cpx-melt thermobarometry, we use the major element data for the Center track volcanic
1287 rocks from different core sections of Site U1578 (Fig.2; Table S3A) as the liquid compositions and
1288 liquid compositions are matched with every Cpx composition in Thermobar. We then use thresholds
1289 of 0.06, 0.05, and 0.03 for the DiHd, EnFs, and CaTs (DiHd: diopside + hedenbergite, CaTs: Ca-
1290 Tschermak, and EnFs: enstatite + ferrosilite) components respectively (Mollo et al., 2013; Putirka,
1291 1999) to filter the disequilibrium data (Fig.S6). This study performs the calculation using the total
1292 amount of Fe (Fe_{tot}) in the Cpx and liquid. Considering $N=553$ Cpx and $N=8$ Liquid, there are a
1293 total of $N=4424$ Liq-Cpx pairs. Finally, the Cpx-Liquid melt matching algorithm finds a total of
1294 $N=618$ Cpx-Liq matches by using the specified filter and $N=397$ Cpx match to one or more liquids
1295 (Table S3A). Then we plot $N=397$ Cpx-Liq pairs on a Rhodes diagram and each liquid composition
1296 is the average of all equilibrium liquids for one Cpx. Nearly all values fall within the fields using a
1297 $K_D^{Cpx-Liq}$ of 0.28 ± 0.08 (Putirka, 2008). For P-T estimates, we use only Cpx-liquid pairs that passed
1298 the applied equilibrium tests and apply the Cpx-melt thermometer presented in Putirka (2008) (eq.
1299 33, $SEE = \pm 45$ °C) and Cpx-melt barometer presented in Neave and Putirka (2017) ($SEE = \pm 140$

1300 MPa). All calculated results are shown in [Table S3A](#) and [Fig.S7a](#).

1301 Clinopyroxene-only thermobarometry is also one of the most practical tools to reconstruct
1302 crystallization pressures and temperatures of Cpx, and it does not require any information of
1303 coexisting silicate melt. Thus, P-T conditions of N=553 Cpx are all calculated by the machine
1304 learning-based Cpx thermobarometry ([Jorgenson et al., 2022](#)) (SEE = ± 320 MPa and ± 51 °C) using
1305 the open-source Python3 package Thermobar v. 1.0.41 ([Wieser et al., 2022](#)). All calculated results
1306 are shown in [Table S3B](#) and [Fig.S7b](#). We also try the calculations for other different existing
1307 thermobarometry models: the Cpx-melt barometer (eq. 30, SEE = ± 160 MPa) of [Putirka \(2008\)](#) and
1308 the thermometer (eq. 33, SEE = ± 45 °C) of [Putirka \(2008\)](#) ([Table S3C](#)); the Cpx-melt barometer
1309 (SEE = ± 260 MPa) and thermometer (SEE = ± 40.0 °C) of [Petrelli et al. \(2020\)](#) ([Table S3D](#)); the
1310 Cpx-melt barometer (SEE = ± 270 MPa) and thermometer (SEE = ± 44.9 °C) of [Jorgenson, C. et al.](#)
1311 ([2020](#)) ([Table S3E](#)); and the Cpx-only barometer (eq. 1, SEE = ± 166 MPa) and thermometer (eq. 2,
1312 SEE = ± 36.6 °C) of [Wang, X. et al. \(2021\)](#) ([Table S3F](#)). These models are chosen because they are
1313 the most appropriate for alkali magma compositions from Site U1578.

1314 For Group 1 Cpx crystals, the Mg# values of their equilibrium melt are much higher than their
1315 host rocks ([Fig.S6d](#)), thus we could not find an appropriate melt composition to calculate their P–T
1316 conditions when using Cpx-melt thermobarometry. P–T crystallization conditions obtained from
1317 Group 2 and 3 Cpx-melt pairs show that they have similar pressures and temperatures with different
1318 Mg# contents, suggesting they crystallized in separate magma reservoirs at the same depth ([Fig.S7a](#)).
1319 Based on Cpx-only thermobarometry ([Jorgenson et al., 2022](#)), Group 1 Cpx may have originated
1320 from relatively deeper and hotter magmatic environments, which could be interpreted as antecrysts
1321 crystallized in primitive magma and subsequently captured by evolved magma ([Chen et al., 2018](#);

1322 [Li et al., 2020](#)). However, when considering the uncertainties associated with thermobarometry, we
1323 argue that it is not possible to clearly resolve two independent magma reservoirs at different depths
1324 beneath the Center track. Taken together, the entire plumbing system depth can be constrained to 0
1325 -20 km (assuming an average crustal density of 2850 kg/m³ following [Huang and Li \(2024\)](#)) and
1326 EPMA analytical uncertainties have a minimal effect on P-T estimation based on the Monte Carlo
1327 error propagation results ([Supplementary Text S3; Table S4](#)), which do not significantly impact our
1328 interpretation. Combined with seismological evidence for a Moho depth of about 12 km ([Geissler
1329 et al., 2017; Huang and Li, 2024](#)), our results suggest that the entire plumbing system is located
1330 within the crust and uppermost mantle near the Moho ([Fig.S7](#)), which is consistent with
1331 observations from Tristan da Cunha Island ([Weit et al., 2017](#)).

1332

1333

1334 **Supplementary Text S3. Monte Carlo error propagation for Cpx thermobarometry**

1335 Recent studies have highlighted that Cpx barometers can be sensitive to analytical uncertainties
1336 inherent in electron probe microanalysis (EPMA) of mineral compositions (Wieser et al., 2023;
1337 Wieser et al., 2024). To assess the impact of such uncertainties on calculated temperatures and
1338 pressures in this study, we conduct Monte Carlo error propagation using the open-source Python3
1339 package Thermobar v.1.0.41 (Wieser et al., 2022).

1340 We employ the *add_noise_sample_lphase* function in Thermobar to generate synthetic
1341 datasets for both mineral and melt compositions. These datasets simulate analytical uncertainty by
1342 creating normally distributed compositions around each measured value. The 1σ percentage errors
1343 are estimated from repeated EPMA analyses of Cpx and repeated XRF analyses of associated melt
1344 phases. For each valid Cpx-Liq pair, 1000 synthetic compositions were generated. For example, N
1345 = 397 equilibrium Cpx-Liq pairs passed our filtering criteria, resulting in $N = 397,000$ synthetic
1346 pairs used for thermobarometric calculations. The results of Monte Carlo error simulation (Fig.S8)
1347 are summarized below.

1348 Cpx-melt thermobarometer: ~95% of the propagated temperature uncertainties ($1\sigma_{\text{analysis}}$) fall
1349 within $\pm 10^\circ\text{C}$ (Fig. S8a and Table. S4A), and pressure uncertainties within ± 100 MPa (Fig. S8b and
1350 Table. S4B). The uncertainty of the Cpx-melt thermometer ($1\sigma_{\text{equation}}$) presented in Putirka (2008)
1351 (eq. 33) and barometer presented in Neave and Putirka (2017) are $\pm 45^\circ\text{C}$ and ± 140 MPa. The overall
1352 errors calculate from combining errors on the thermobarometer equations and the analyses, using
1353 $\sigma_{\text{overall}} = \sqrt{(\sigma_{\text{equation}}^2 + \sigma_{\text{analysis}}^2)}$, is only slightly ($0\text{--}1.2^\circ\text{C}$, $0\text{--}32$ MPa) higher than errors induced
1354 by the thermobarometer equations alone.

1355 Cpx-only thermobarometer: ~95% of propagated temperature uncertainties are within $\pm 30^\circ\text{C}$

1356 (Fig. S8c and Table. S4C), and pressure uncertainties within ± 70 MPa (Fig. S8d and Table. S4D).

1357 The uncertainty of the Cpx-melt thermometer and barometer presented in Jorgenson et al., (2022)

1358 are ± 51 °C and ± 320 MPa. The overall errors calculate from combining errors on the

1359 thermobarometer equations and the analyses, using $\sigma_{\text{overall}} = \sqrt{(\sigma_{\text{equation}}^2 + \sigma_{\text{analysis}}^2)}$, is only

1360 slightly (0–8.2 °C, 0–8 MPa) higher than errors induced by the thermobarometer equations alone.

1361 These results demonstrate that the analytical uncertainties associated with EPMA have a

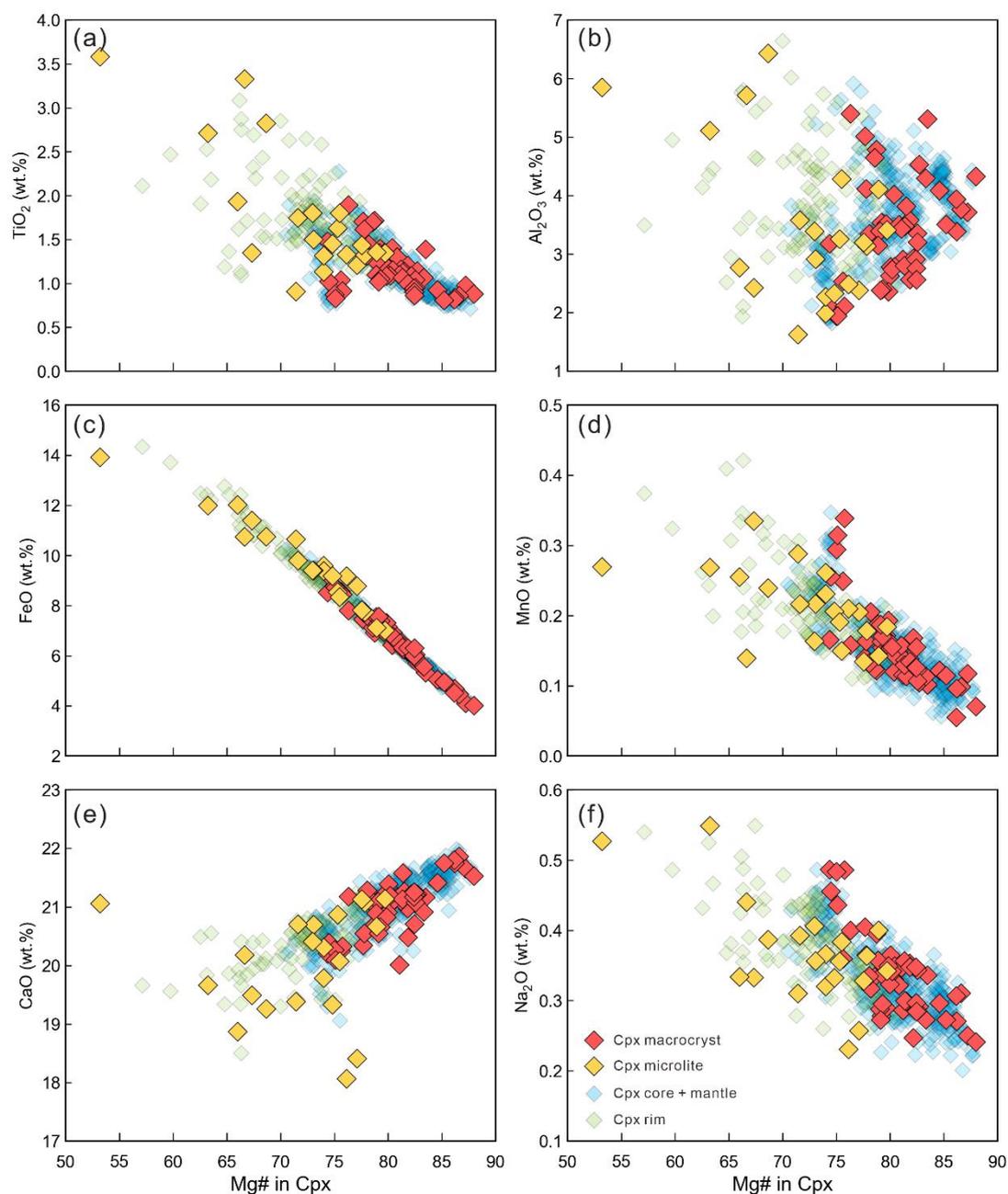
1362 limited impact on the calculated P–T values. Therefore, they do not significantly influence our

1363 thermobarometric interpretations.

1364

1365

1366 **Supplementary Figure S1**



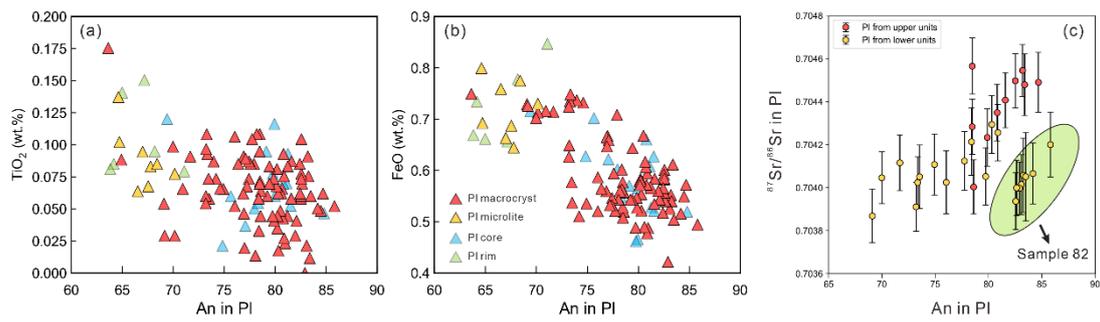
1367

1368 Fig. S1 Plots of TiO_2 (a), Al_2O_3 (b), FeO (c), MnO (d), CaO (e), and Na_2O (f) versus Mg# in Cpx.

1369 All the data used here were acquired by EPMA.

1370

1371 **Supplementary Figure S2**



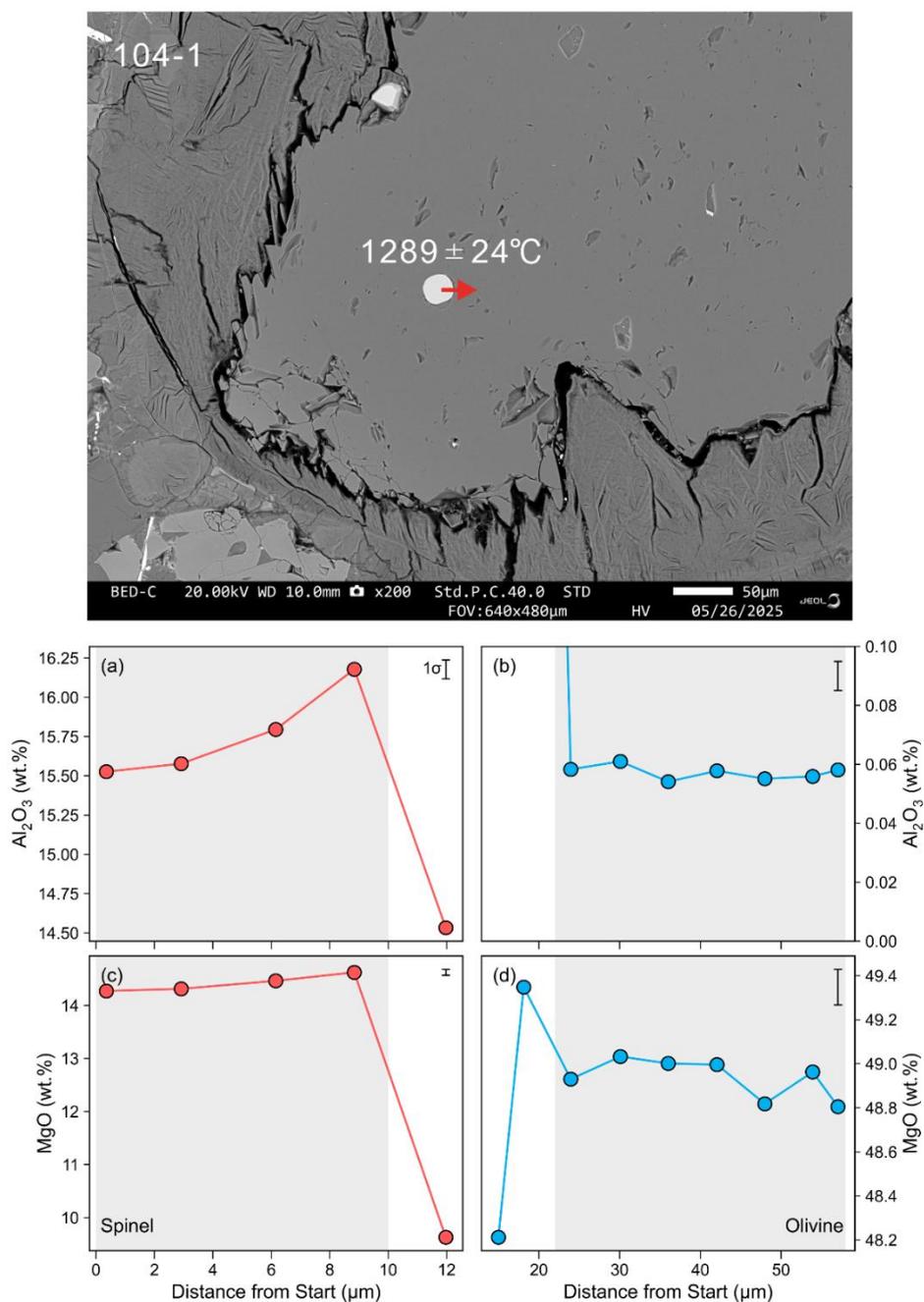
1372

1373 Fig. S2 Compositions of Pl macrocrysts and microlites from the basaltic lavas cored at Site U1578,

1374 Walvis Ridge. Plots of TiO₂ (a), FeO (b) and Sr isotope (c) versus An (%) in Pl.

1375

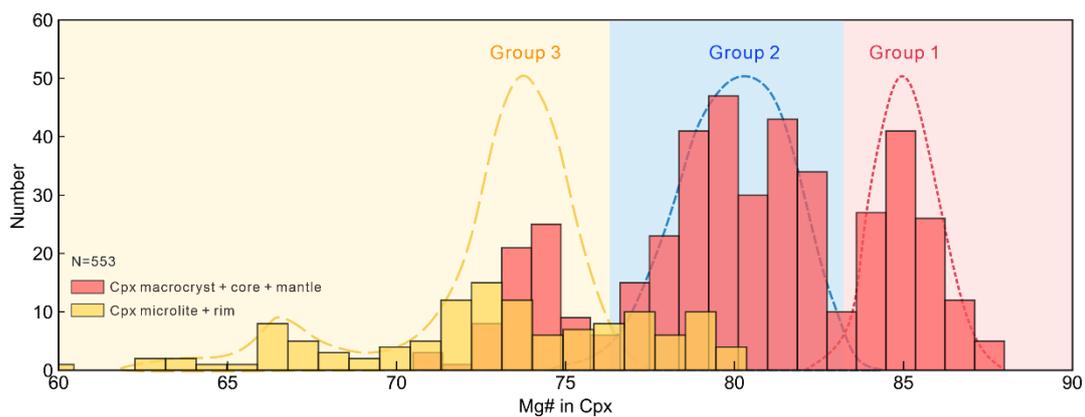
1376 **Supplementary Figure S3**



1377

1378 Fig. S3 Backscatter Electron Image for Center track Ol crystal 104-1. A profile was measured by
1379 SEM-EDS-WDS through both Sp and Ol as shown the red arrow. The grey bars indicate points
1380 included in the thermometry calculations. Representative one standard deviation analytical
1381 uncertainties are shown as black bars. The temperature calculated for this profile and its uncertainty
1382 are shown on the BSE image.

1383 **Supplementary Figure S4**



1384

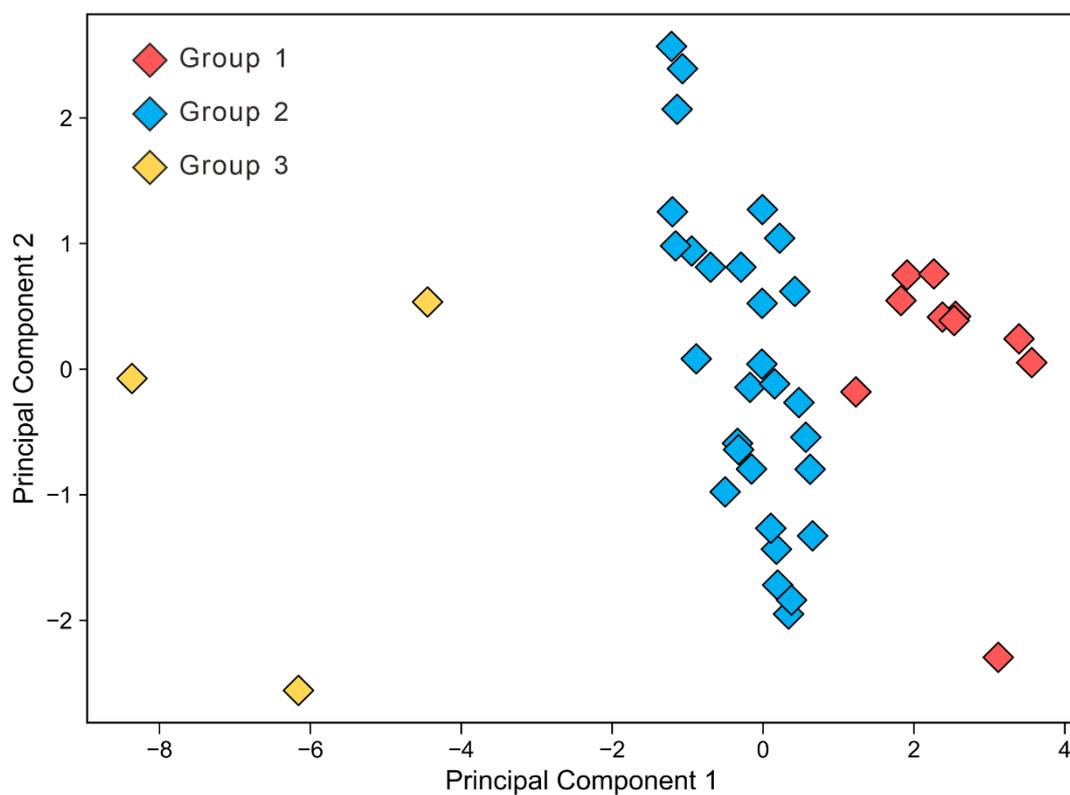
1385 Fig. S4 Distribution of Mg# contents of Cpx macrocrysts and microlites from the basaltic lavas

1386 cored at Site U1578, Walvis Ridge. The red, blue, and yellow areas represent the range of Mg# in

1387 Cpx from Group 1, Group 2 and Group 3, respectively.

1388

1389 **Supplementary Figure S5**

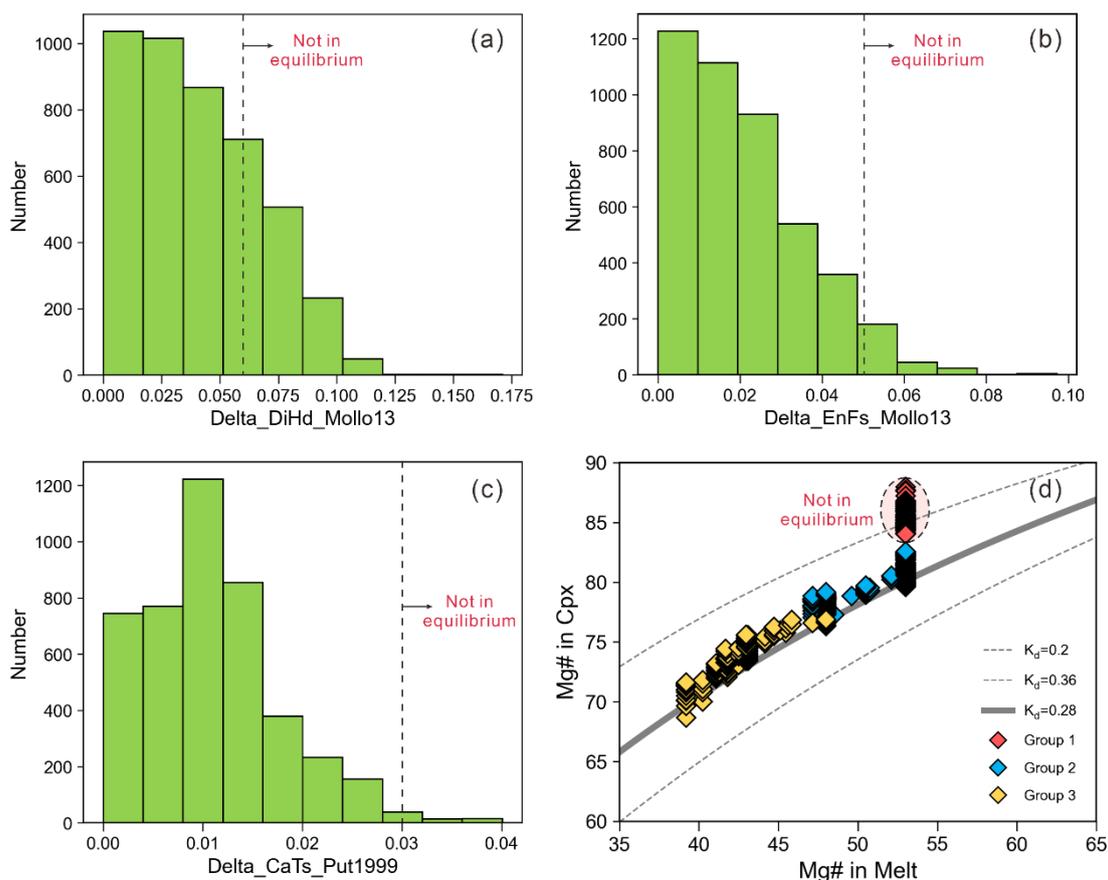


1390

1391 Fig. S5 K-means cluster analysis of Cpx based on eight geochemical parameters: Mg#, Na₂O, Sc,
1392 Cr, Ni, La, Y and jadeite content of Cpx. The red, blue, and yellow diamonds represent the Cpx from
1393 Group 1, Group 2 and Group 3, respectively.

1394

1395 **Supplementary Figure S6**

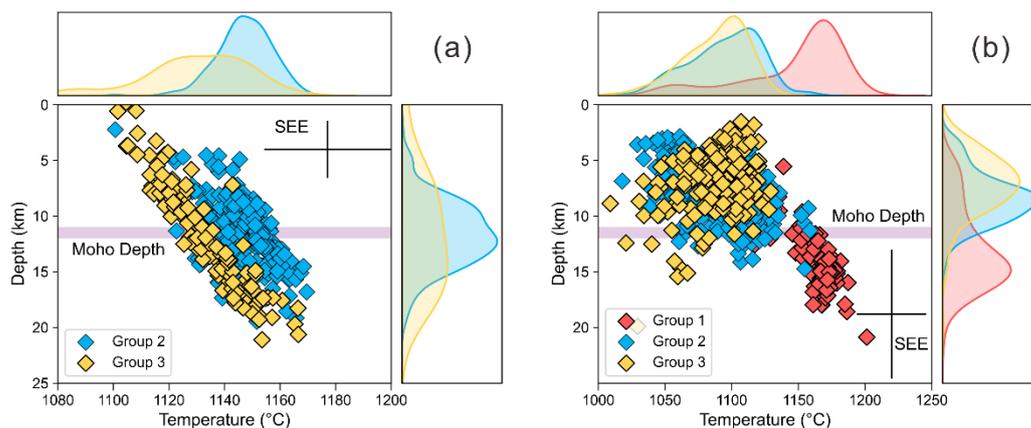


1396

1397 Fig. S6 The Cpx-melt equilibrium test. The tests were carried out by calculating the measured DiHd
1398 (a), EnFs (b), and CaTs (c) values. We use thresholds of 0.06, 0.05, and 0.03 for the DiHd, EnFs,
1399 and CaTs components respectively to filter the disequilibrium data (Mollo et al., 2013; Putirka, 1999;
1400 Putirka, 2008). The Cpx and whole-rock values are plotted in the Rhodes diagram to illustrate this
1401 equilibrium test (Rhodes et al., 1979). The grey and dashed lines define the equilibrium curves
1402 between Cpx and melt (d), which were calculated using $K_D(\text{Fe-Mg})^{\text{Cpx-melt}} = 0.28 \pm 0.08$ (Putirka,
1403 2008). The red, blue, and yellow diamonds represent the Cpx from Group 1, Group 2 and Group 3,
1404 respectively.

1405

1406 **Supplementary Figure S7**

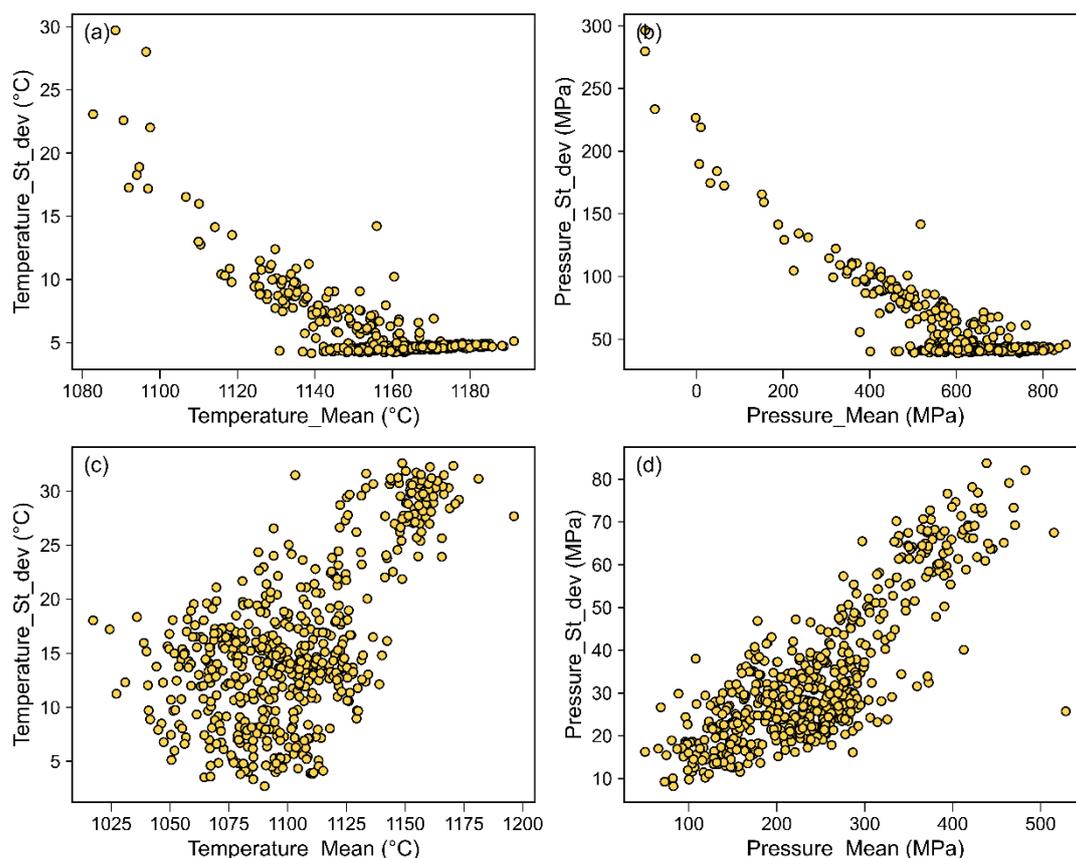


1407

1408 Fig. S7 Plots of crystallization depth versus temperature of Group 1, 2, 3 Cpx. The red, blue, and
1409 yellow diamonds represent the Cpx from Group 1, Group 2 and Group 3, respectively. (a) Cpx-melt
1410 thermobarometry (Neave and Putirka, 2017; Putirka, 2008); (b) Cpx-only thermobarometry
1411 (Jorgenson et al., 2022). The depth of Moho is about 12 km according to the seismological evidence (Geissler
1412 et al., 2017; Huang and Li, 2024).

1413

1414 **Supplementary Figure S8**

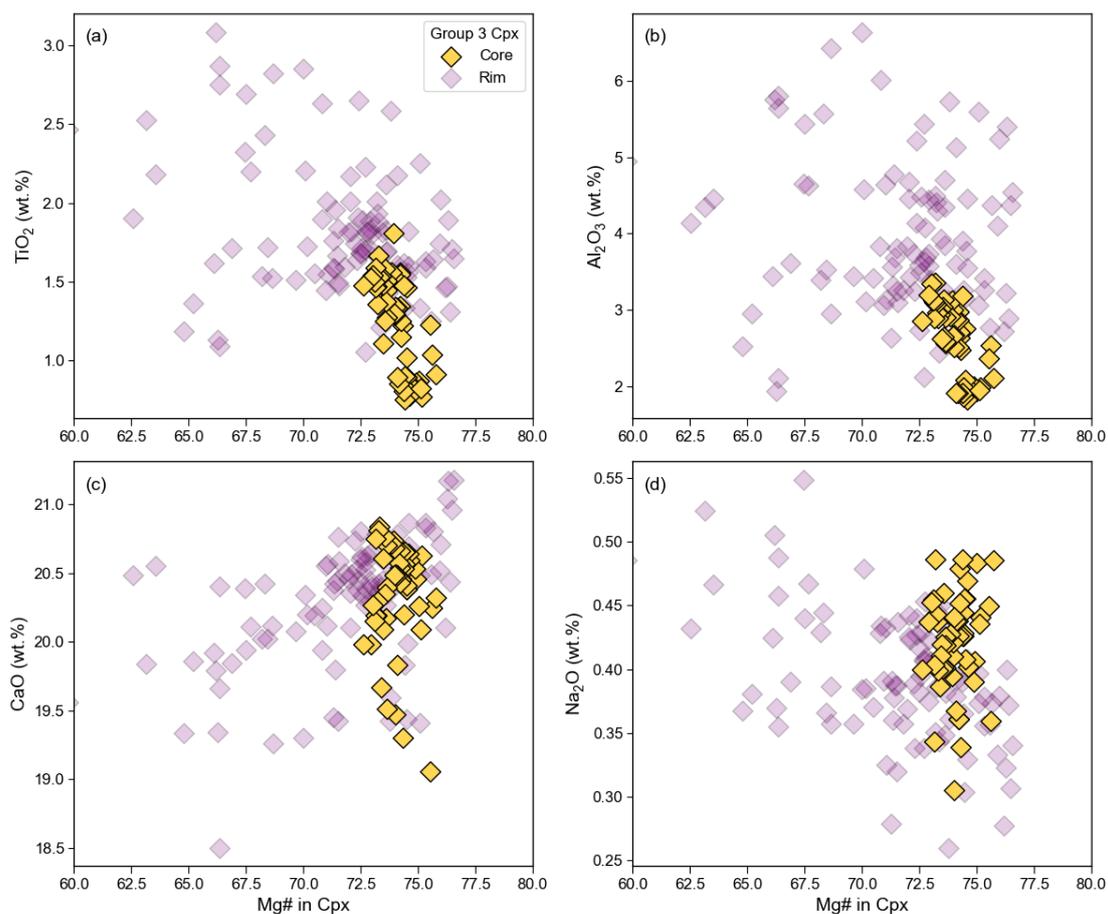


1415

1416 Fig. S8 Monte Carlo error propagation results for Cpx thermobarometry. Panels (a) and (b) show
1417 temperature and pressure uncertainties for the Cpx–melt thermobarometer, with ~95% of
1418 propagated temperature errors within ± 10 °C and pressure errors within ± 100 MPa. Panels (c) and
1419 (d) display uncertainties for the Cpx-only thermobarometer, with ~95% of temperature errors within
1420 ± 30 °C and pressure errors within ± 70 MPa.

1421

1422 **Supplementary Figure S9**



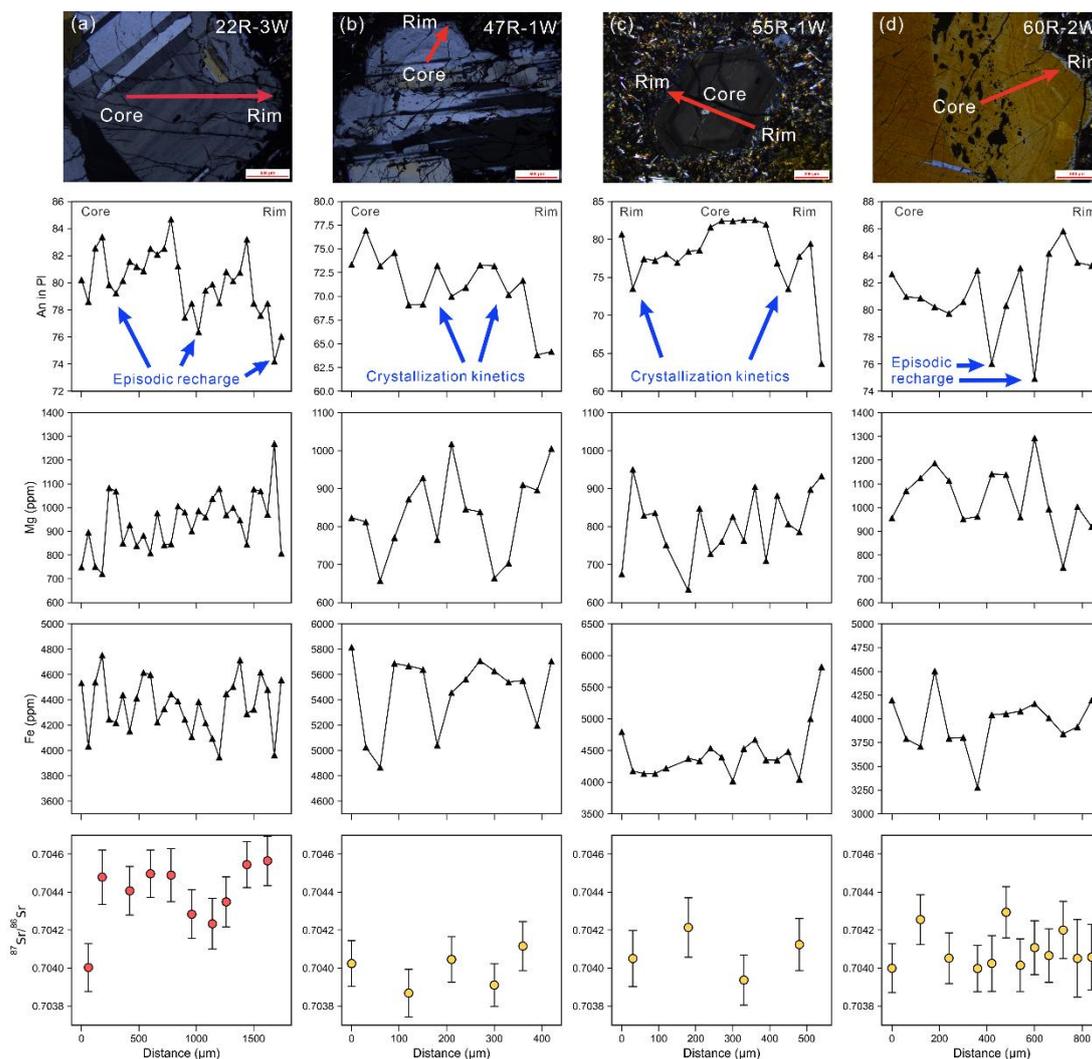
1423

1424 Fig. S9 Plots of TiO₂ (a), Al₂O₃ (b), CaO (c), and Na₂O (d) versus Mg# in Group 3 Cpx. All the

1425 data used here were acquired by EPMA.

1426

1427 **Supplementary Figure S10**



1428

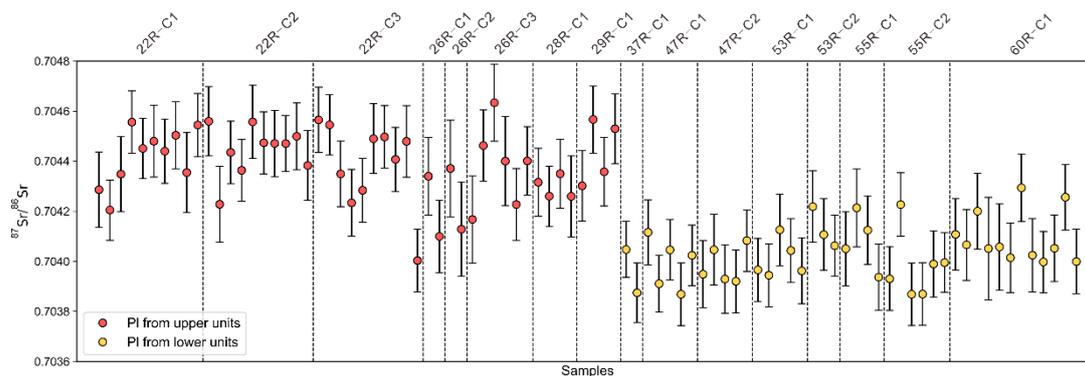
1429 Fig. S10 Representative BSE images and analytical points with major element compositions and Sr

1430 isotopes for zoned PI from volcanic rocks at Site 1578, Walvis Ridge. The red arrows represent the

1431 location of the electron microprobe traverses across the zoning.

1432

1433 **Supplementary Figure S11**

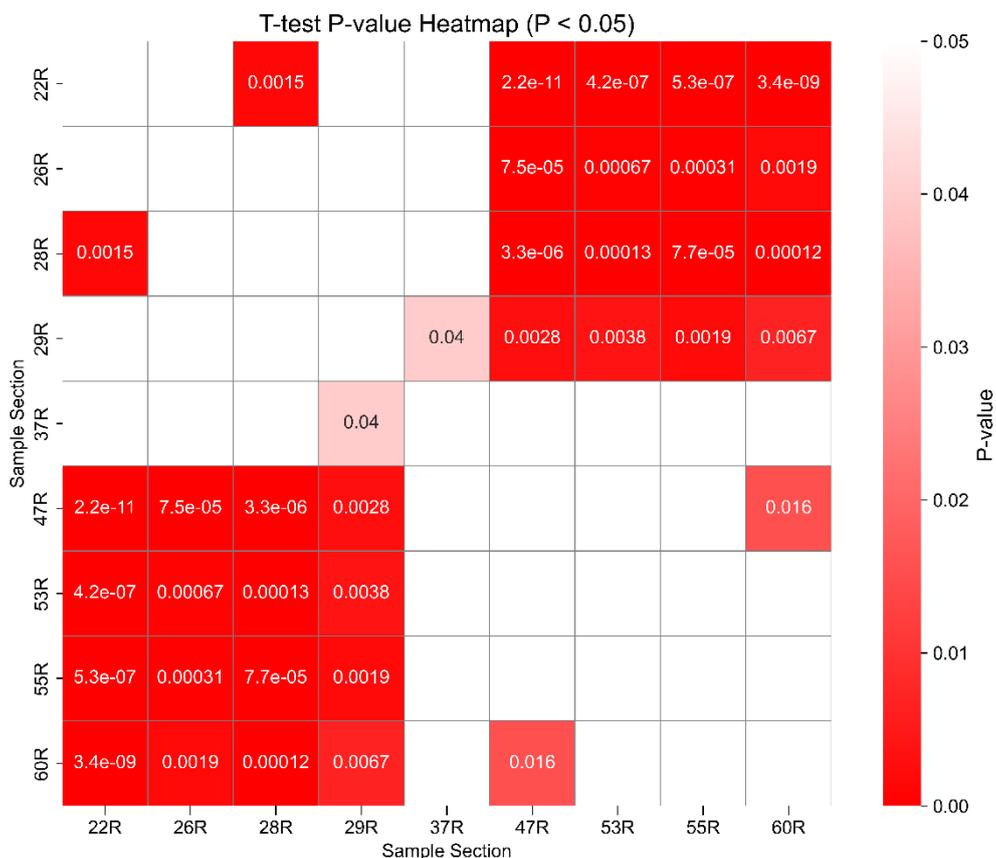


1434

1435 Fig. S11 Sr isotopic compositions and associated uncertainties for Pl phenocrysts from volcanic
1436 rocks at Site U1578, Walvis Ridge. Within a single crystal, variations are mostly unresolvable within
1437 analytical uncertainty. Sample labels such as “22R-C1” denote the core section (22R) and the first
1438 analyzed Pl crystal from that section (C1).

1439

1440 **Supplementary Figure S12**

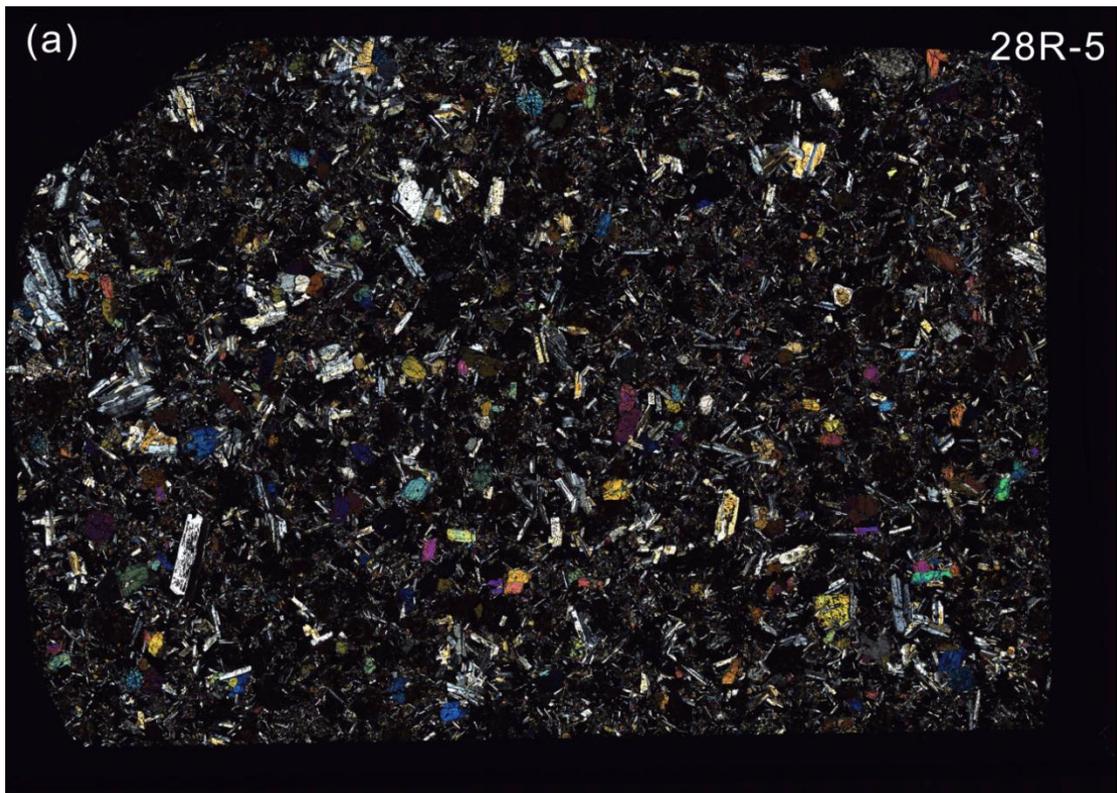


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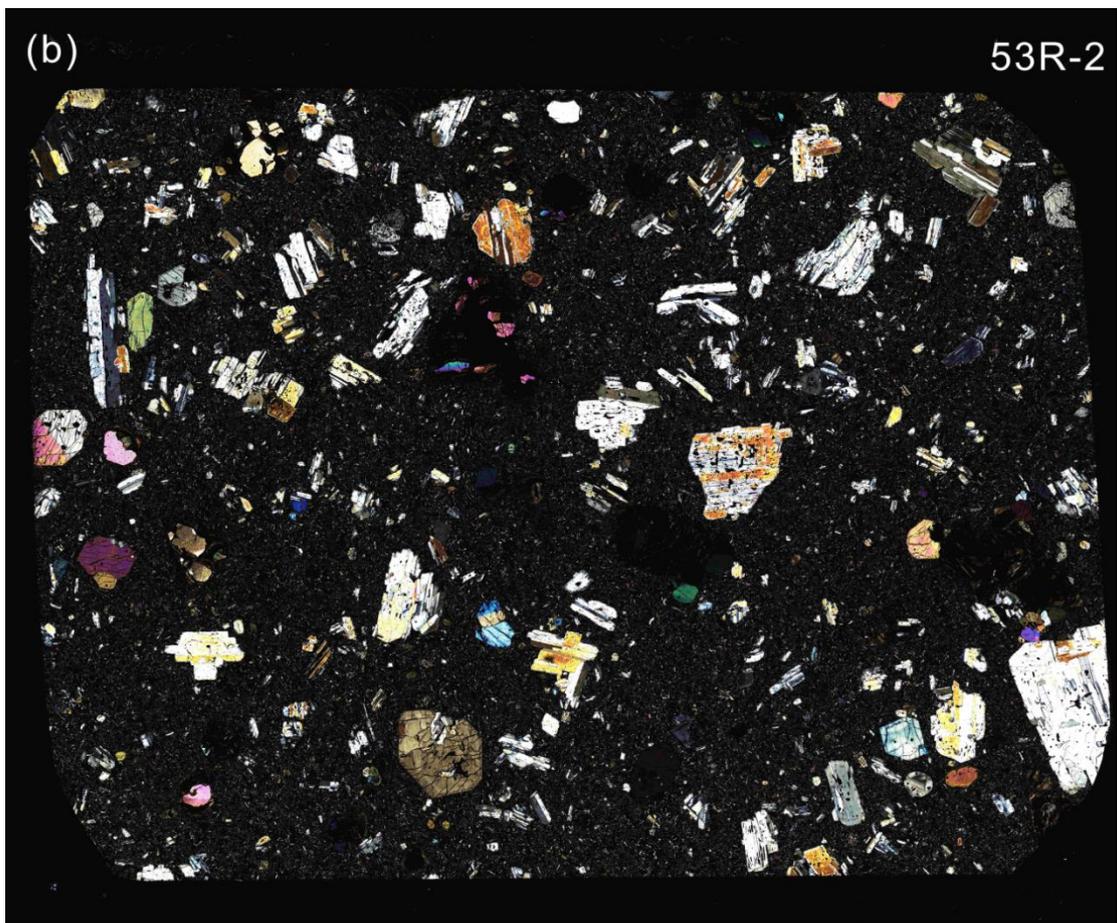
1442 Fig. S12 Heatmap showing statistical differences in $^{87}\text{Sr}/^{86}\text{Sr}$ values between Pl grains from nine
 1443 different core sections. The heatmap summarizes the results of two-sample t-tests performed on all
 1444 36 pairwise combinations, with numbers indicating p-values less than 0.05 that highlight statistically
 1445 significant differences. Detailed t-test results, including t-statistics and p-values, are provided in
 1446 [Table S5](#).

1447

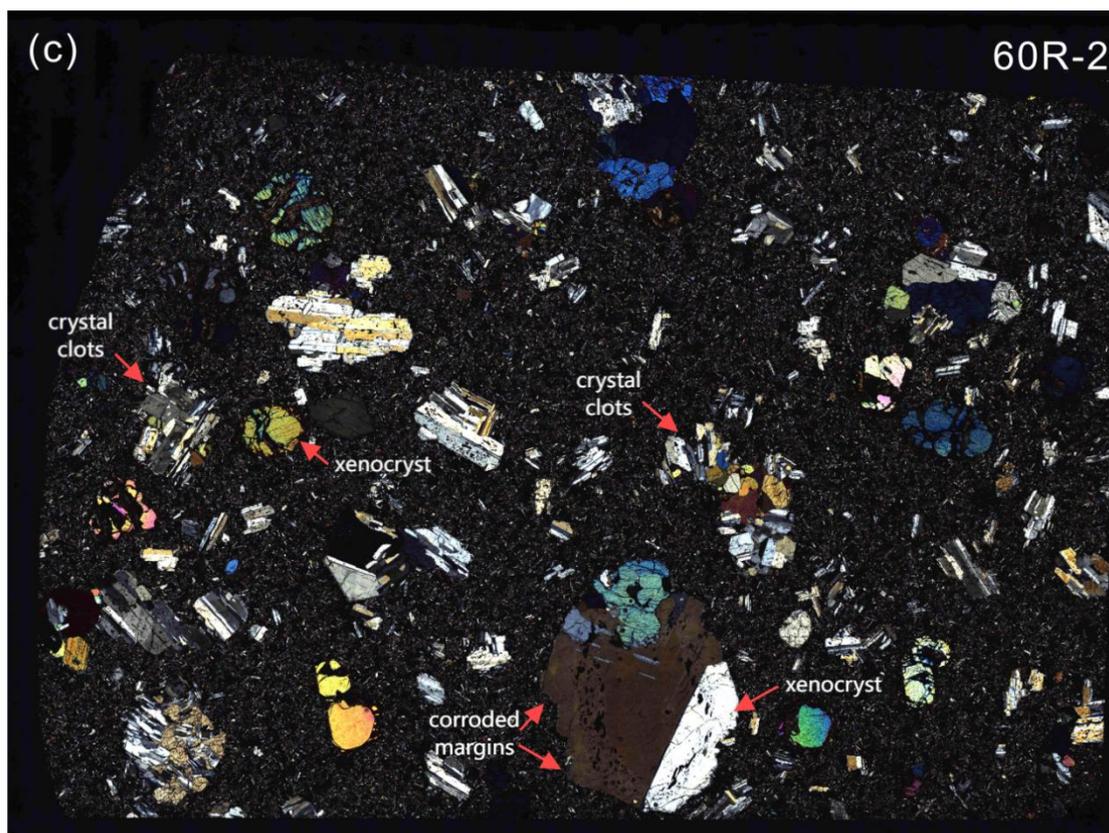
1448 **Supplementary Figure S13**



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1450



1451

1452 Fig. S13 Cross-polarized light photomicrographs of three thin sections (28R-5 from the upper unit,

1453 53R-2 and 60R-2 from the lower unit) demonstrating fresh Pl and Cpx phenocrysts with sharp

1454 crystal boundaries, absence of alteration rims, and no secondary overgrowths or pseudomorphs.

1455 Whole thin section scans made using a WITec alpha300R Raman microscope with a motorized stage.

1456

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